

In search of the (ground) truth & kinetic plasma simulators

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Diogo Ferreira*, Paulo Alves**

* IST, Lisbon, Portugal, ** UCLA

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Motivation & questions

Some (of my) questions:

Can ML help us speed up standard plasma simulators?

Our early attempts: ML replaces Monte Carlo modules in PIC - Badiali et al., JPP 2022; Amaro et al., arXiv:2406.02491

Can we build faster ML based simulators?

Rethinking architecture of simulators to match ML uniqueness: 1D collisional plasma model - Carvalho et al.; MLST 2023

What can we learn from data-driven approaches + ML?

Learning physics (following Alves & Fiuza) e.g. collision operators: Carvalho et al., in preparation for submission to JPP

Can standard plasma simulators provide “high quality data” for data-driven discovery?

Capturing collisions in PIC codes: D. Carvalho et al., in preparation

Can we understand qualitative modifications of plasma behavior from “Learning what we already know”

e.g. Waterbag vs Maxwellian; nonlinear waves vs unstable (and then turbulent) scenarios; nonrelativistic to relativistic, Casimir invariants evolution

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“There are unknown unknowns” (and “know unknowns”), Jon Arons citing D. Rumsfeld

MC models in PIC simulations

New simulator models - 1D GNN collisional plasma model

Learning advection and diffusion coefficients

The (ground) truth? - collisions in PIC codes

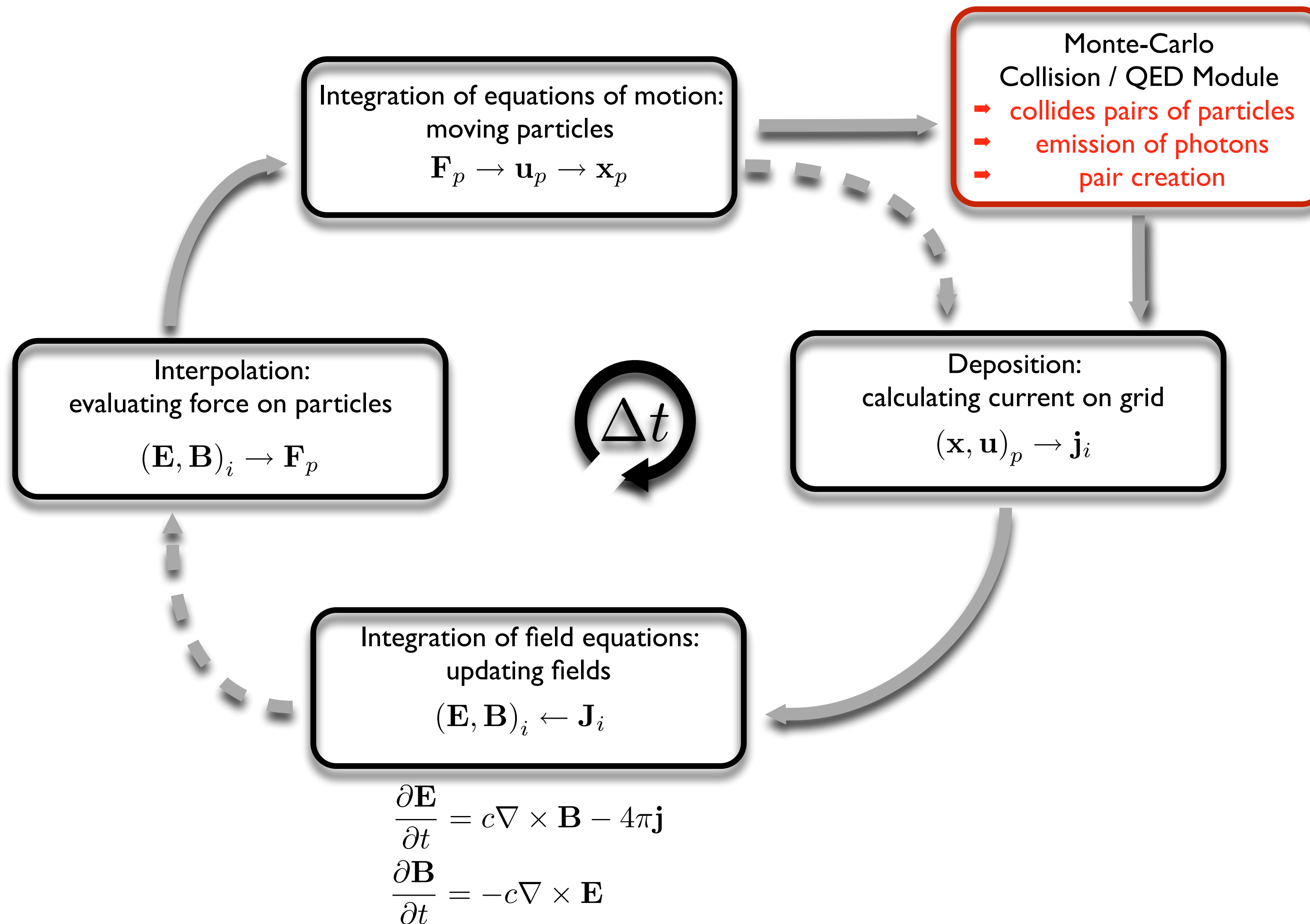
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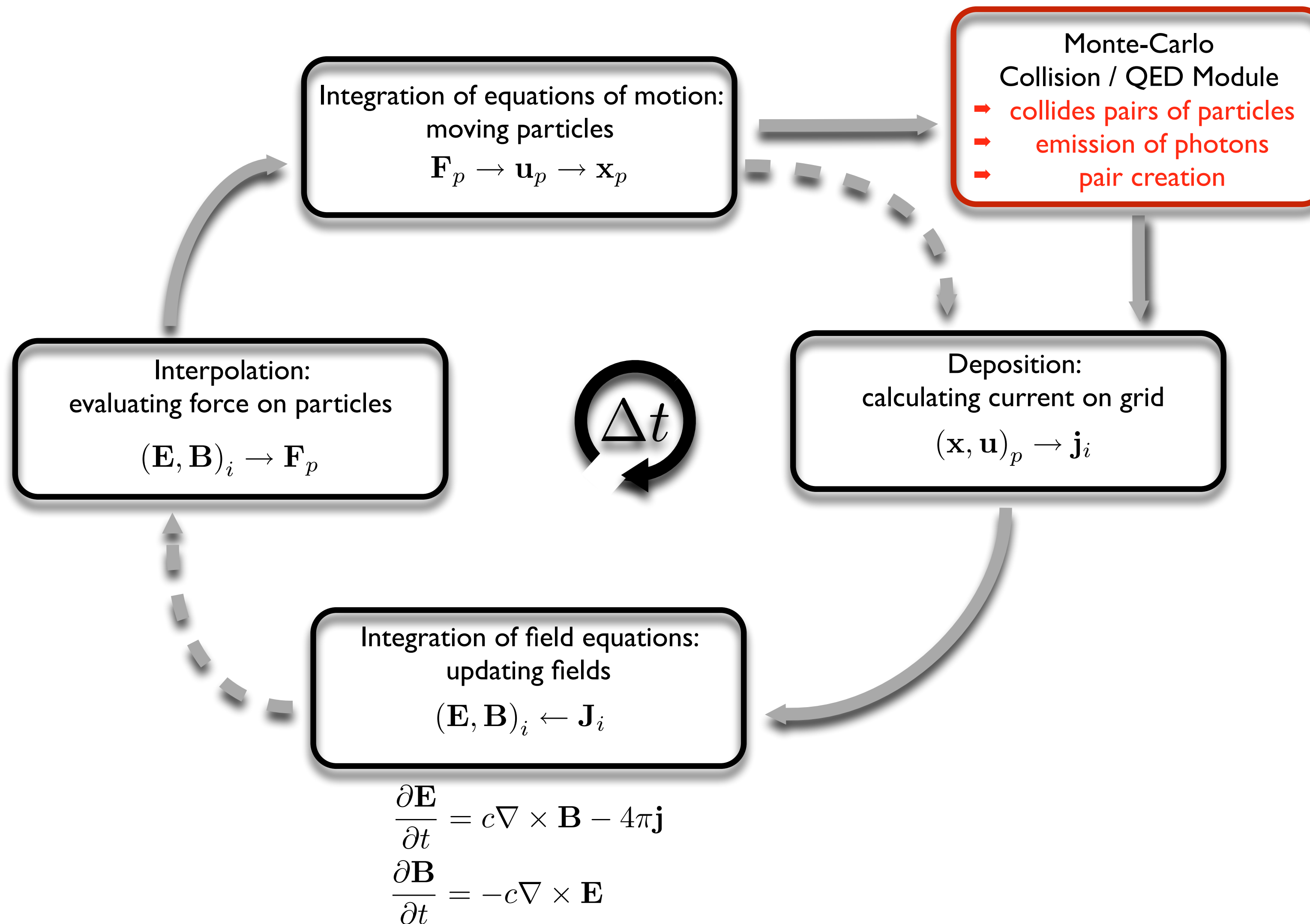
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Collisions/QED Physics are modelled using Monte-Carlo routines



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Problems

Computationally intensive

Memory & run-time

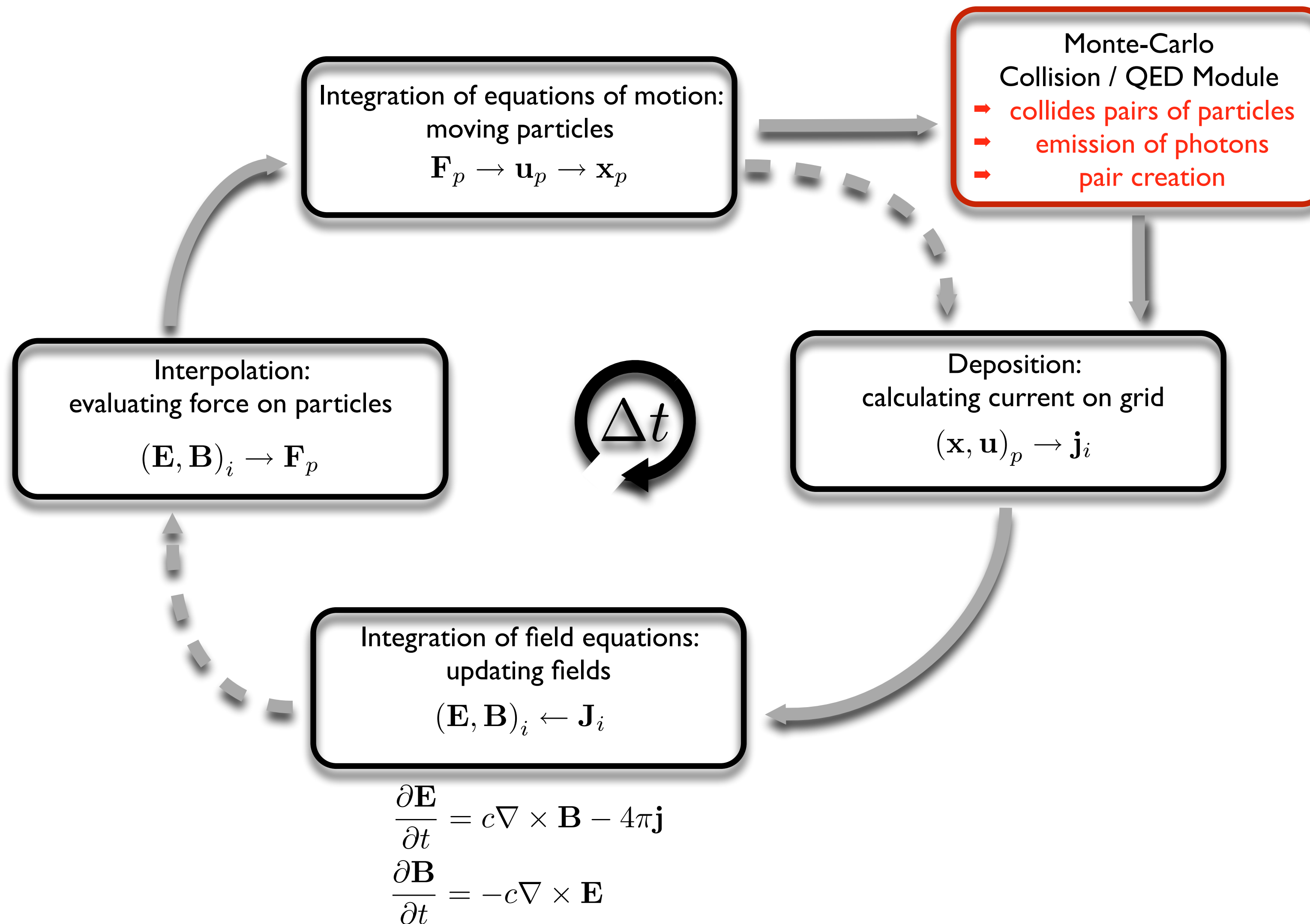
Can lead to numerical issues

e.g. increased numerical heating

Theory valid in limited scenarios

e.g. small angle-scattering

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Computationally intensive

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Can lead to numerical issues

e.g. increased numerical heating

Theory valid in limited scenarios

e.g. small angle-scattering

Can ML tackle these issues?

Reduce computational cost

Design new (stable) numerical algorithms

Learn corrections to existing theory

How are cross-sections calculated?

Theory

Usually impracticable at run-time

Interpolation Tables

Fast to query

Limited to few input parameter values

Chebyshev Polynomials

Exponentially convergent

Impractical for ≥ 3 D input parameter space

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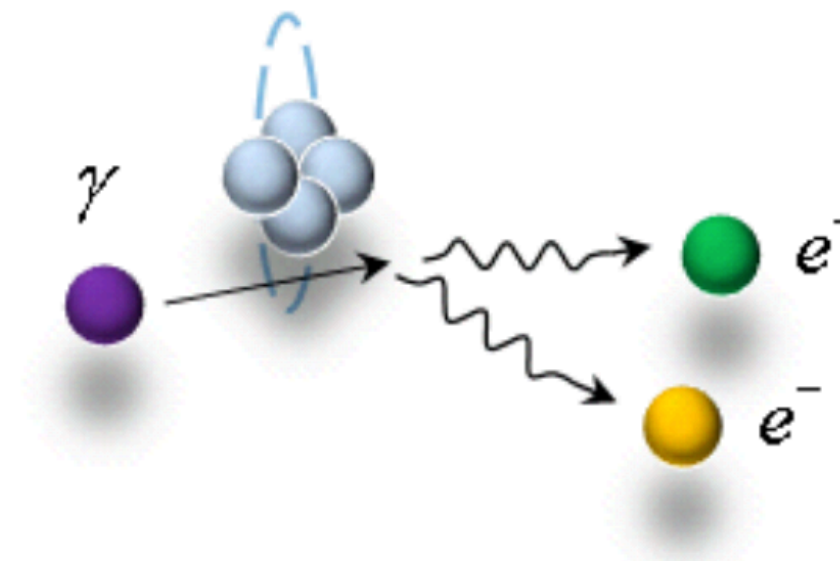
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Impractical for $\geq 3D$ input parameter space

Input parameter spaces are often $> 3D$

Bethe-Heitler



Usual MC Inputs

Photon Energy

Ion Atomic Number

Inputs that should also influence cross-section

Angle of incoming photon

Plasma temperature

Plasma density

Ionisation degree

Local electromagnetic fields

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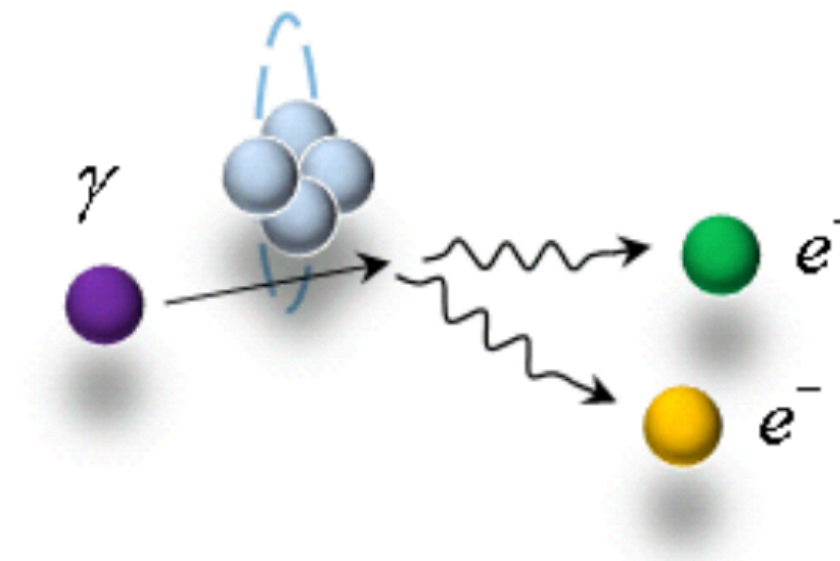
Neural Networks*

Memory efficient for any input parameter space

Run-time dependent on model size

Input parameter spaces are often $> 3D$

Bethe-Heitler



Usual MC Inputs

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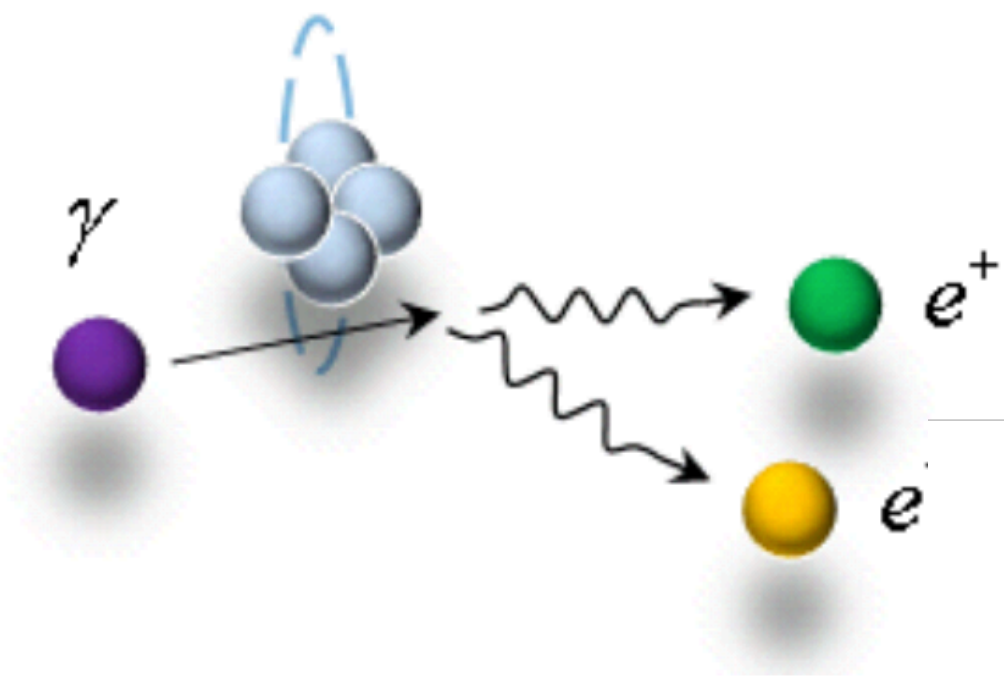
Plasma temperature

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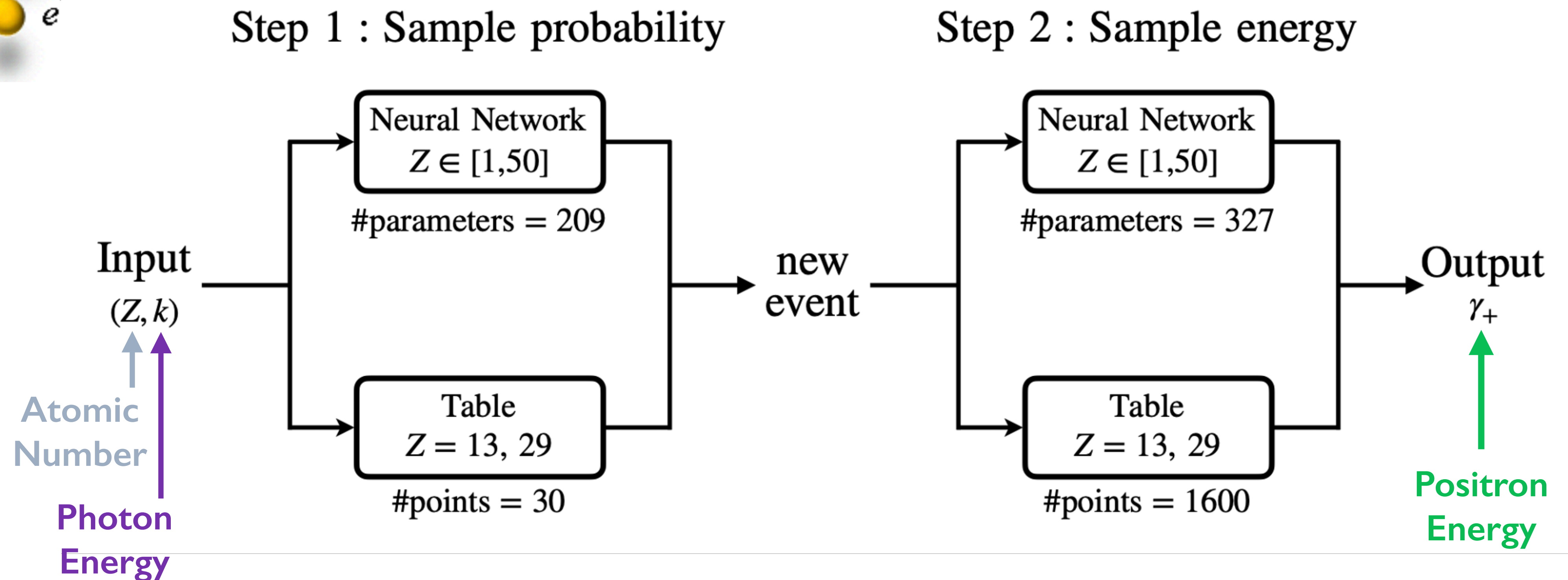
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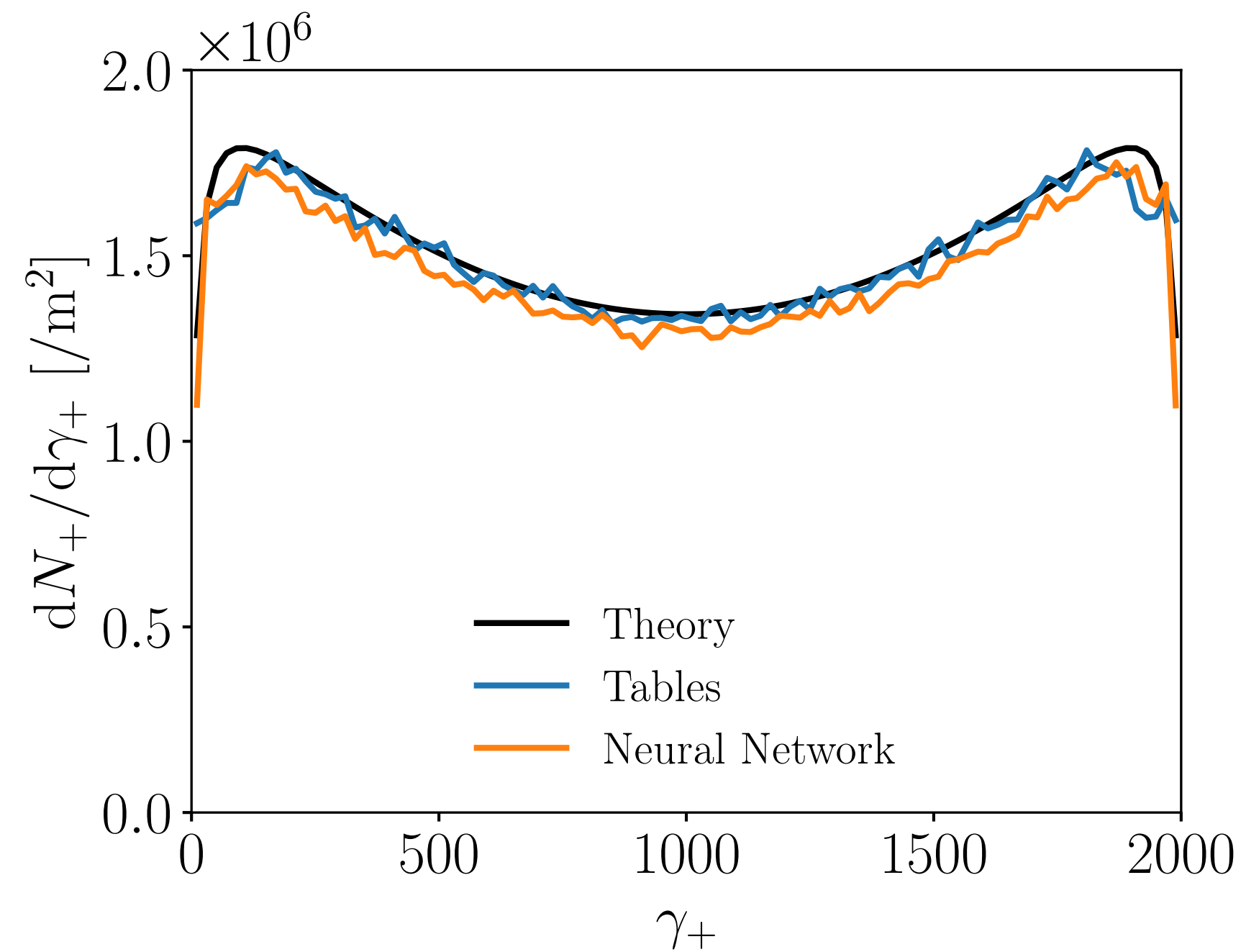
* C. Badiali et al., J. Plasma Phys. 88(6) (2022) and O. Amaro et al., arXiv:2406.02491 (2024)



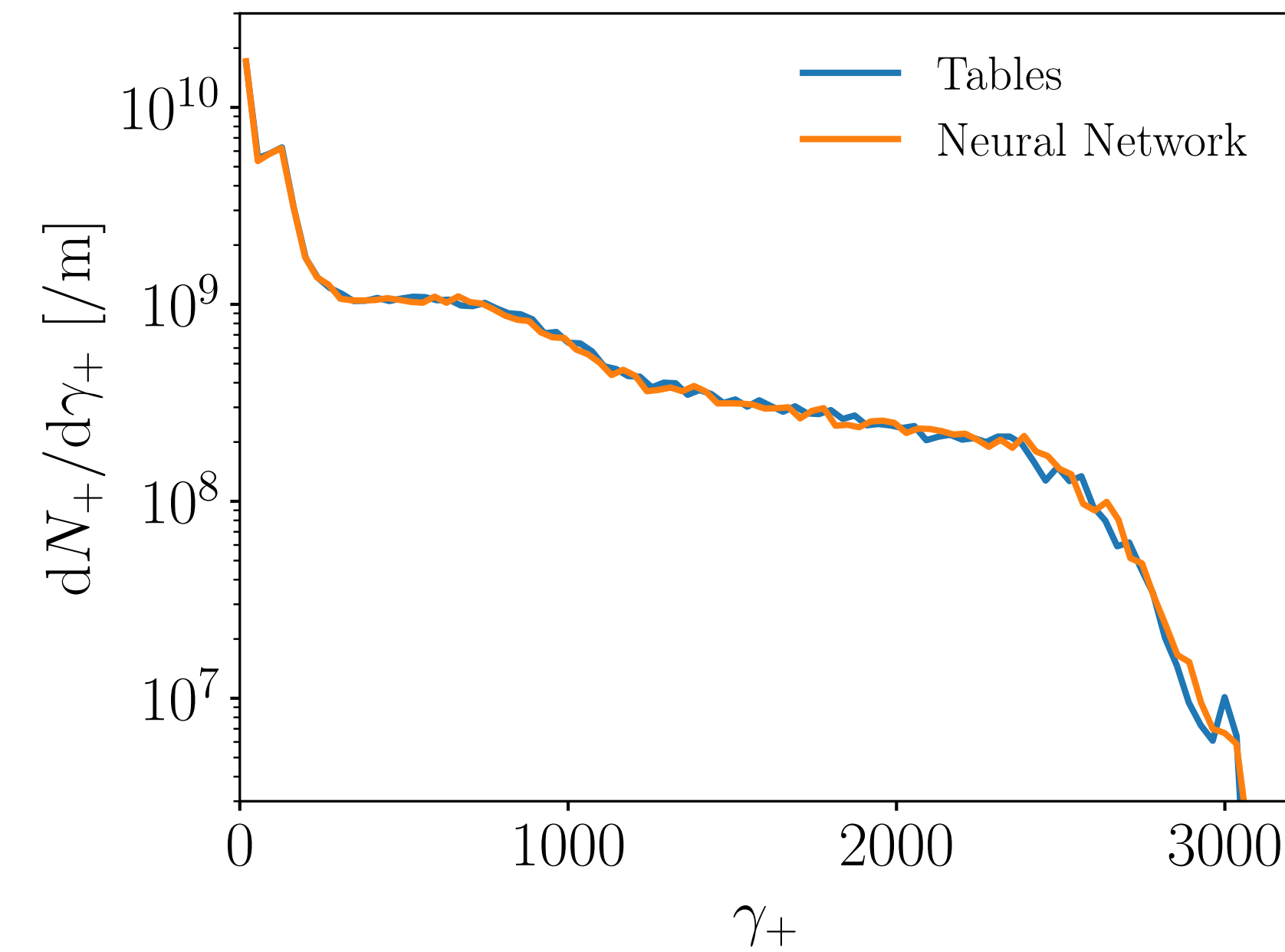
OSIRIS NN Implementation



Benchmark 1D (Early time evolution)



Production 2D (Laser-solid target long-time evolution)



Light-weight and fast OSIRIS-SFQED

Neural Networks are **as accurate** as pre-calculated tables

Require **x100 less memory** to store and are of **comparable runtime**

“With four parameters I can fit an elephant, and with five I can make him wiggle his trunk.” von Neumann

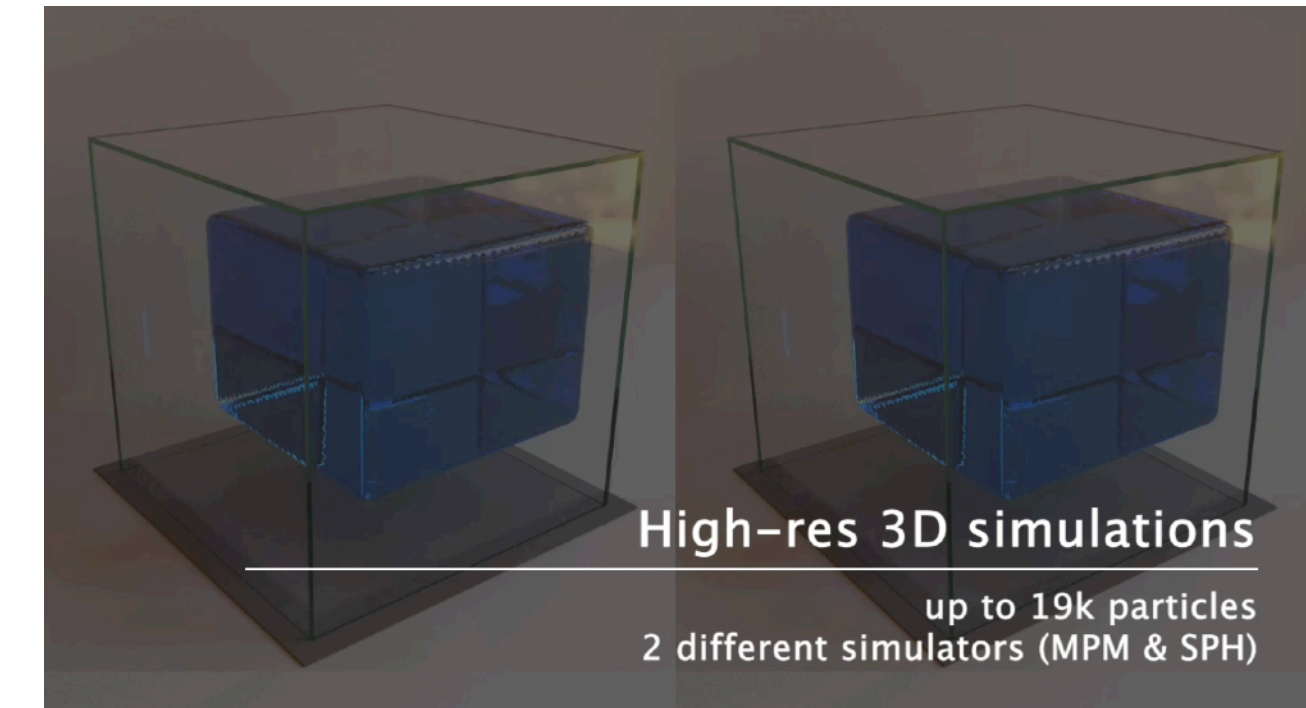
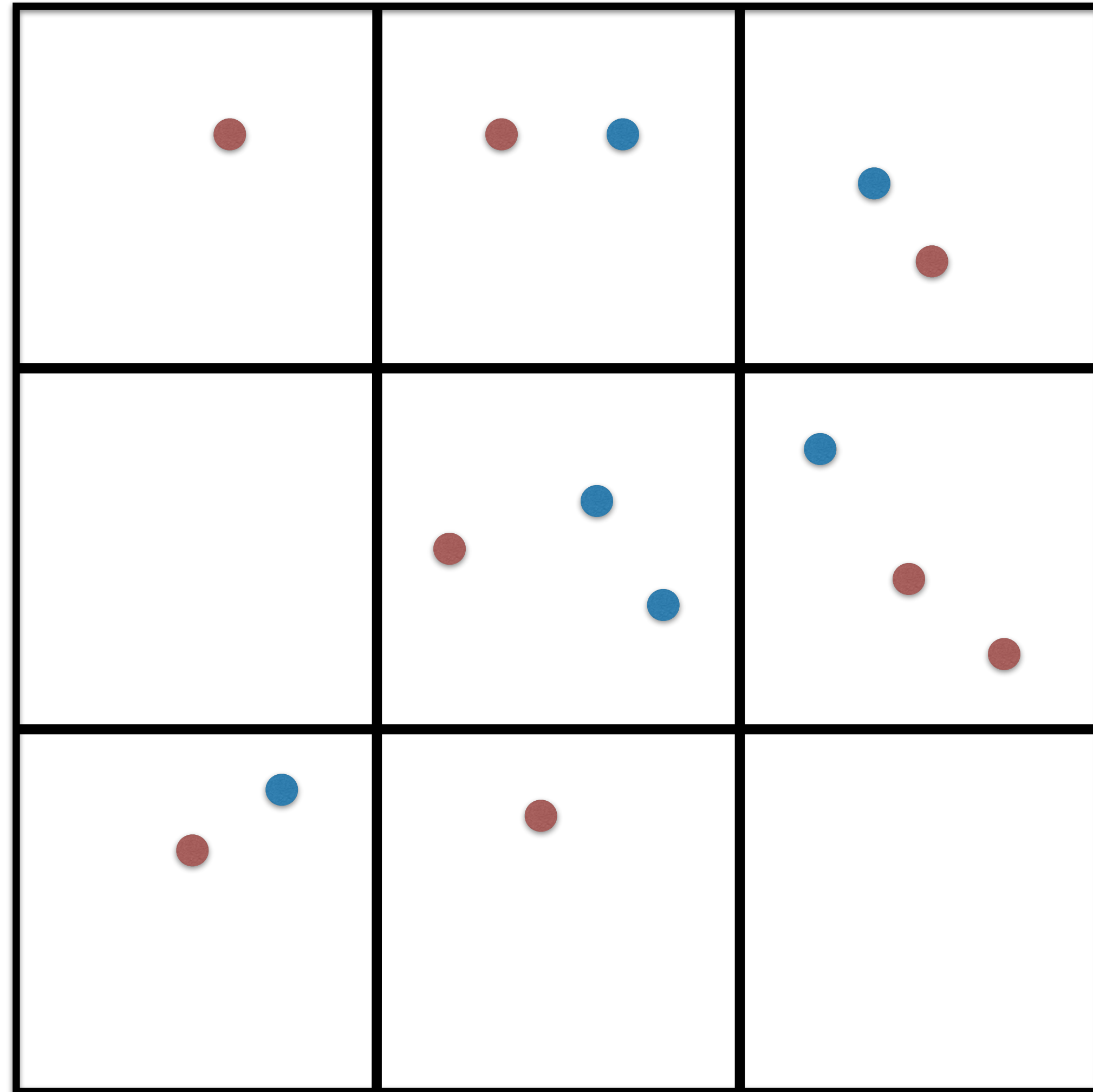
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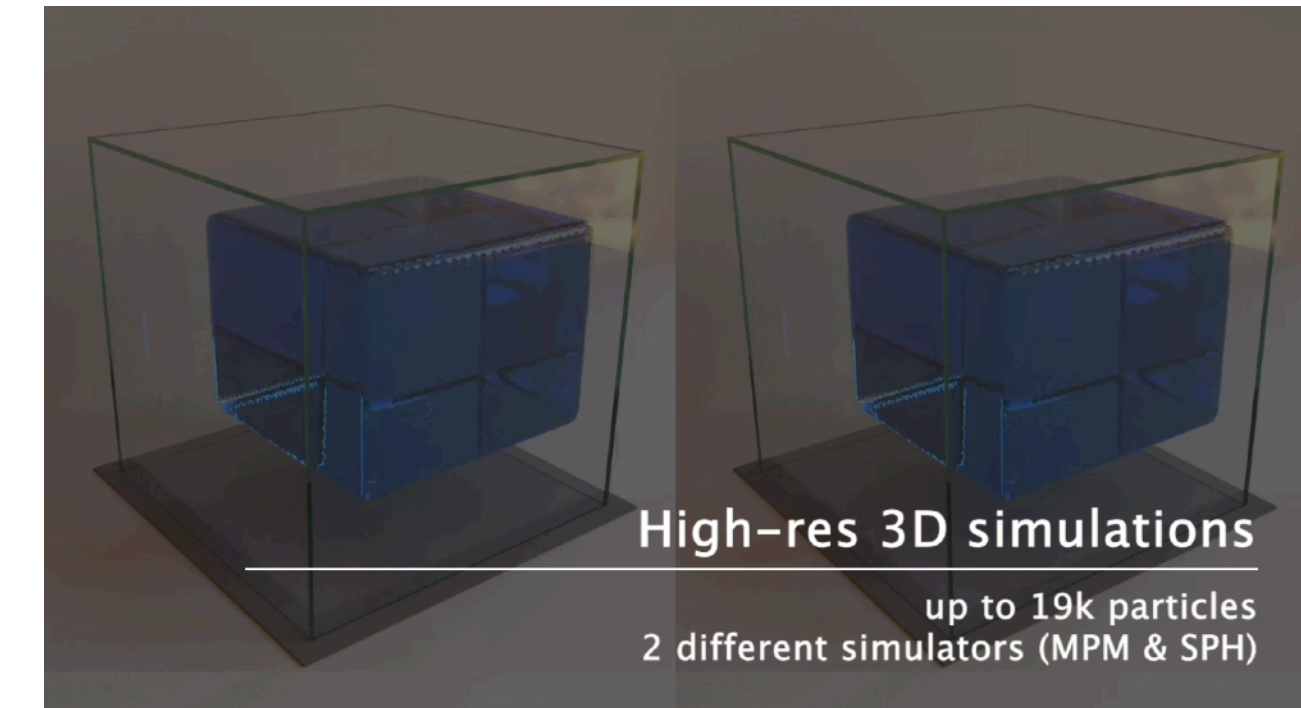
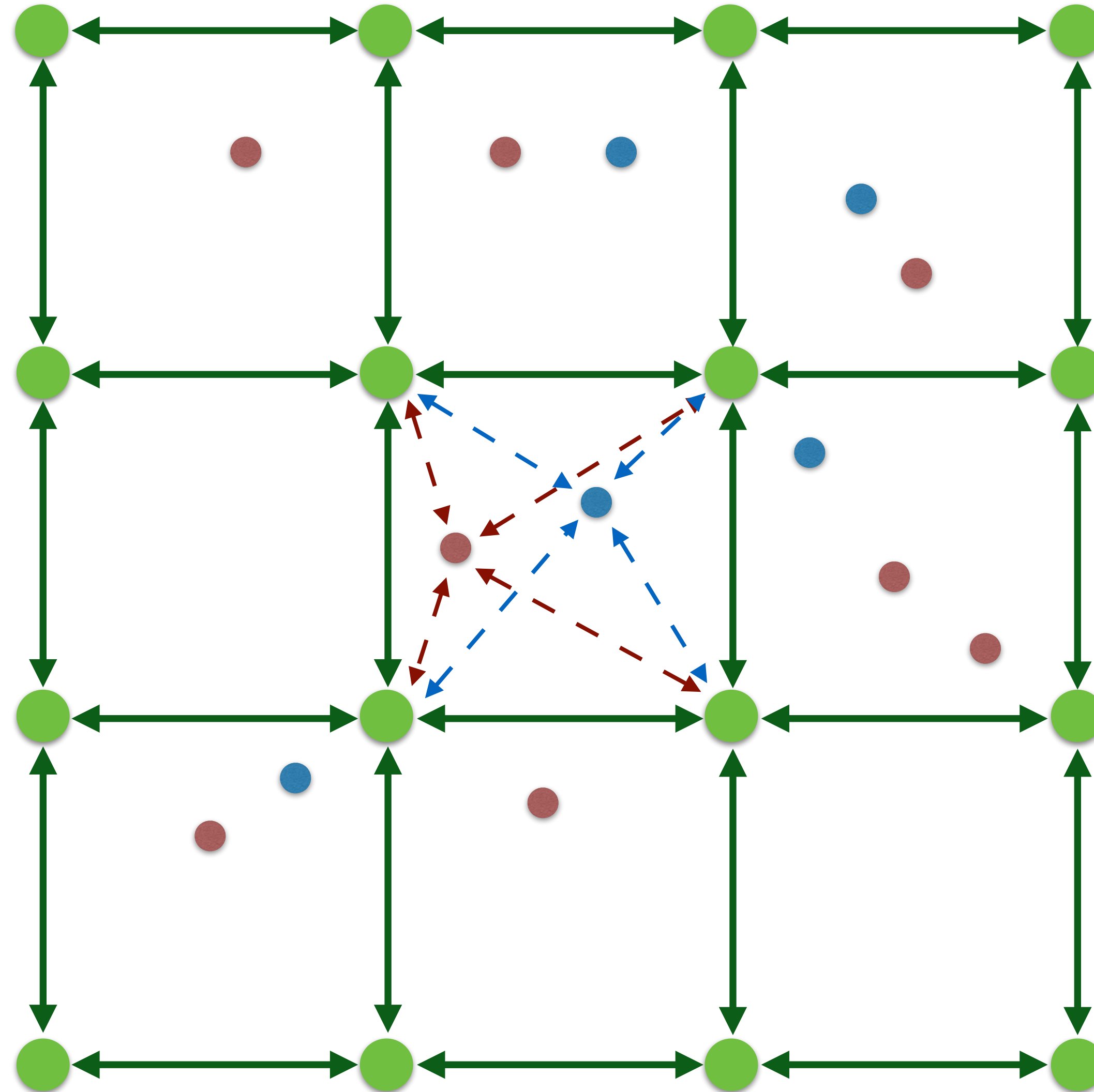
The (ground) truth? - collisions in PIC codes

PIC codes (and others) can be seen from a graph perspective



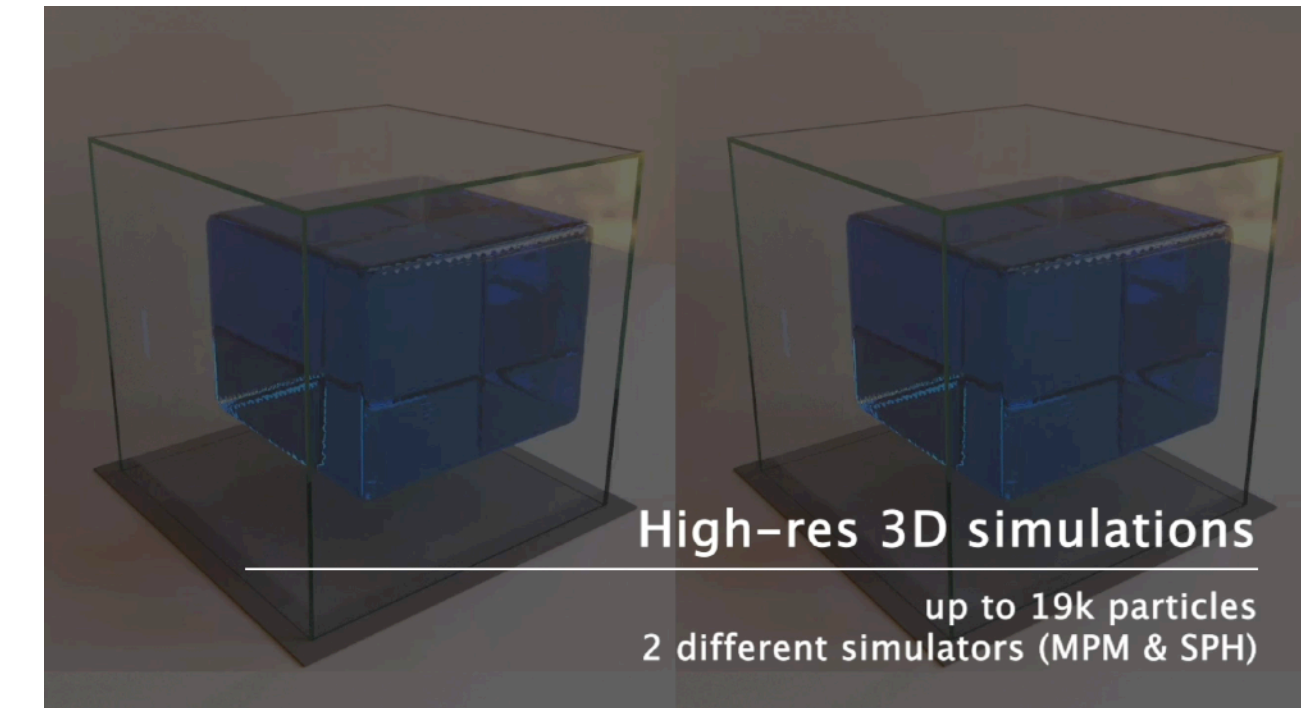
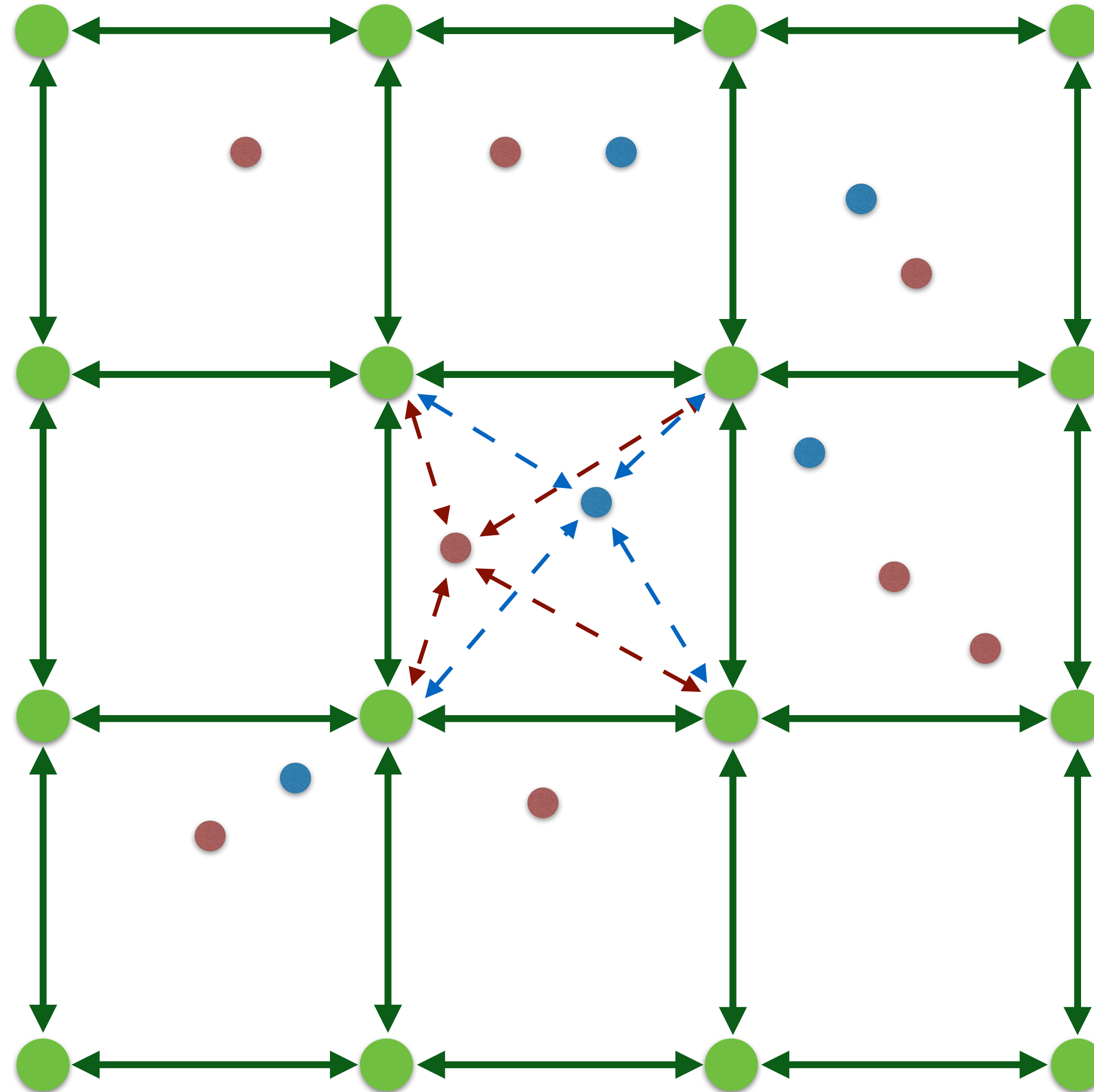
A. Sanchez-Gonzalez et al., ICML PMLR 8459–8468 (2020)
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The Breaking of Finite
Amplitude Plasma
Oscillations

by

John Dawson

June 1959

MATT - 4



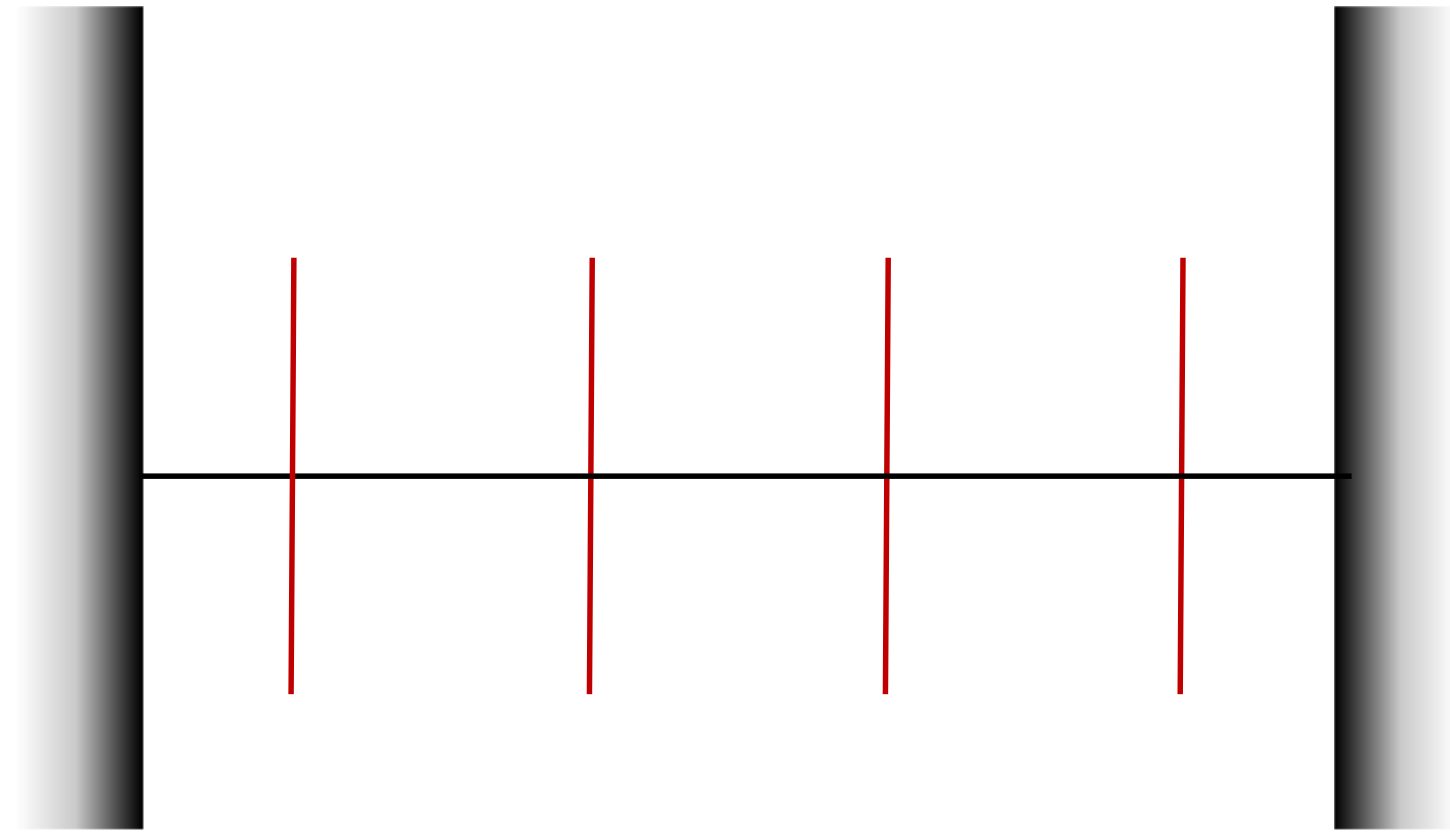
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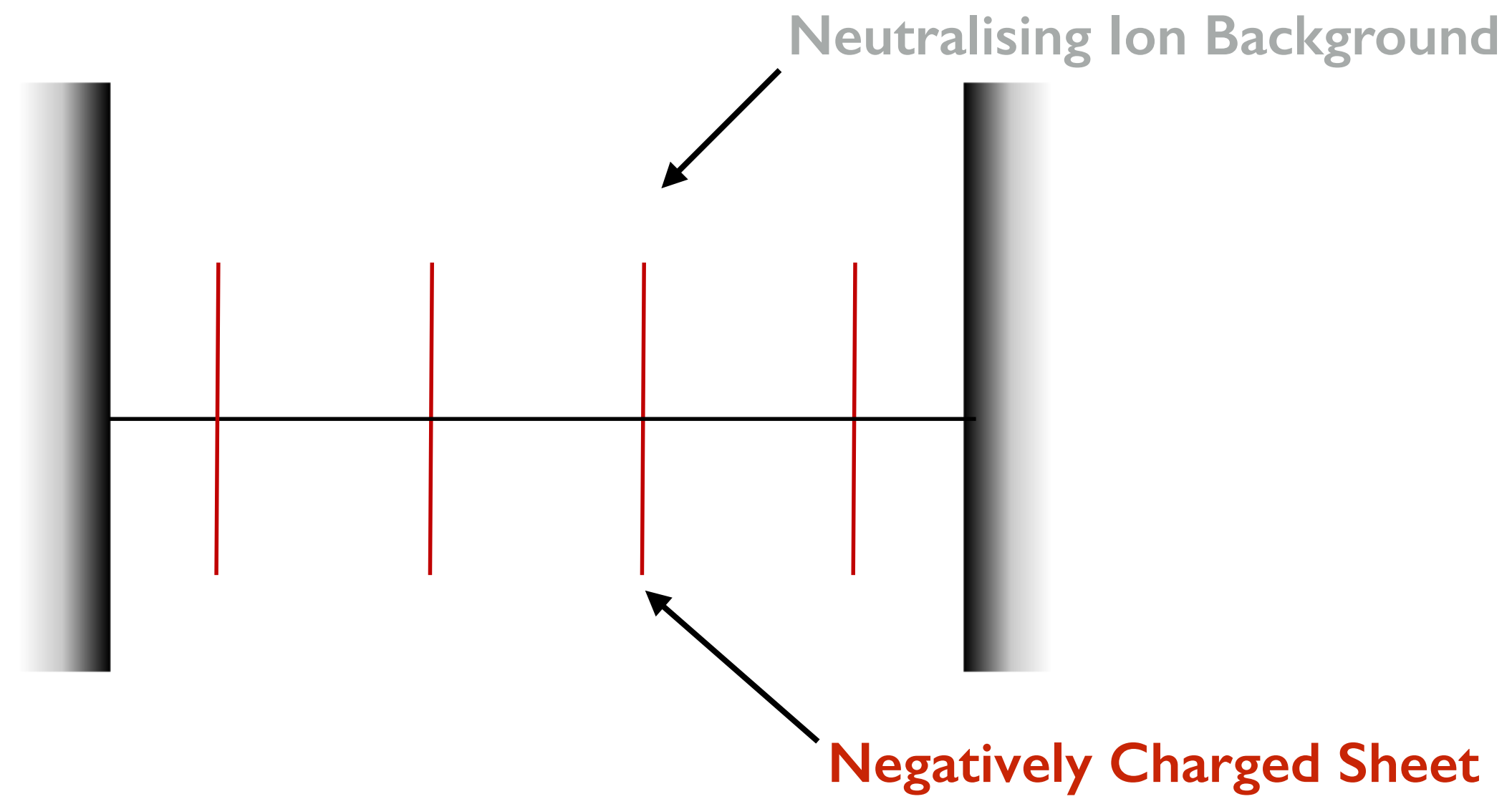
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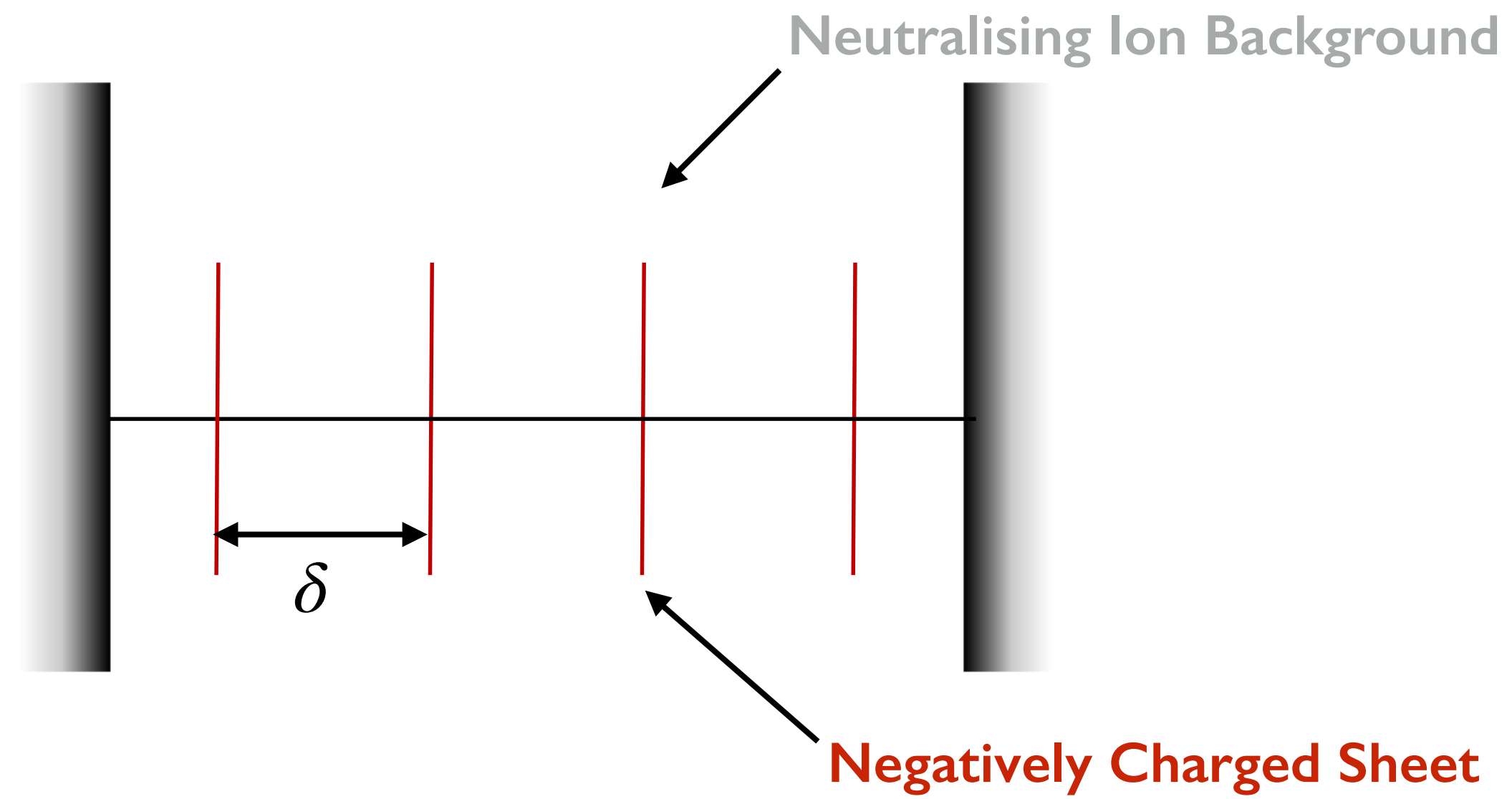
1D Plasma Electrostatic Sheet Model



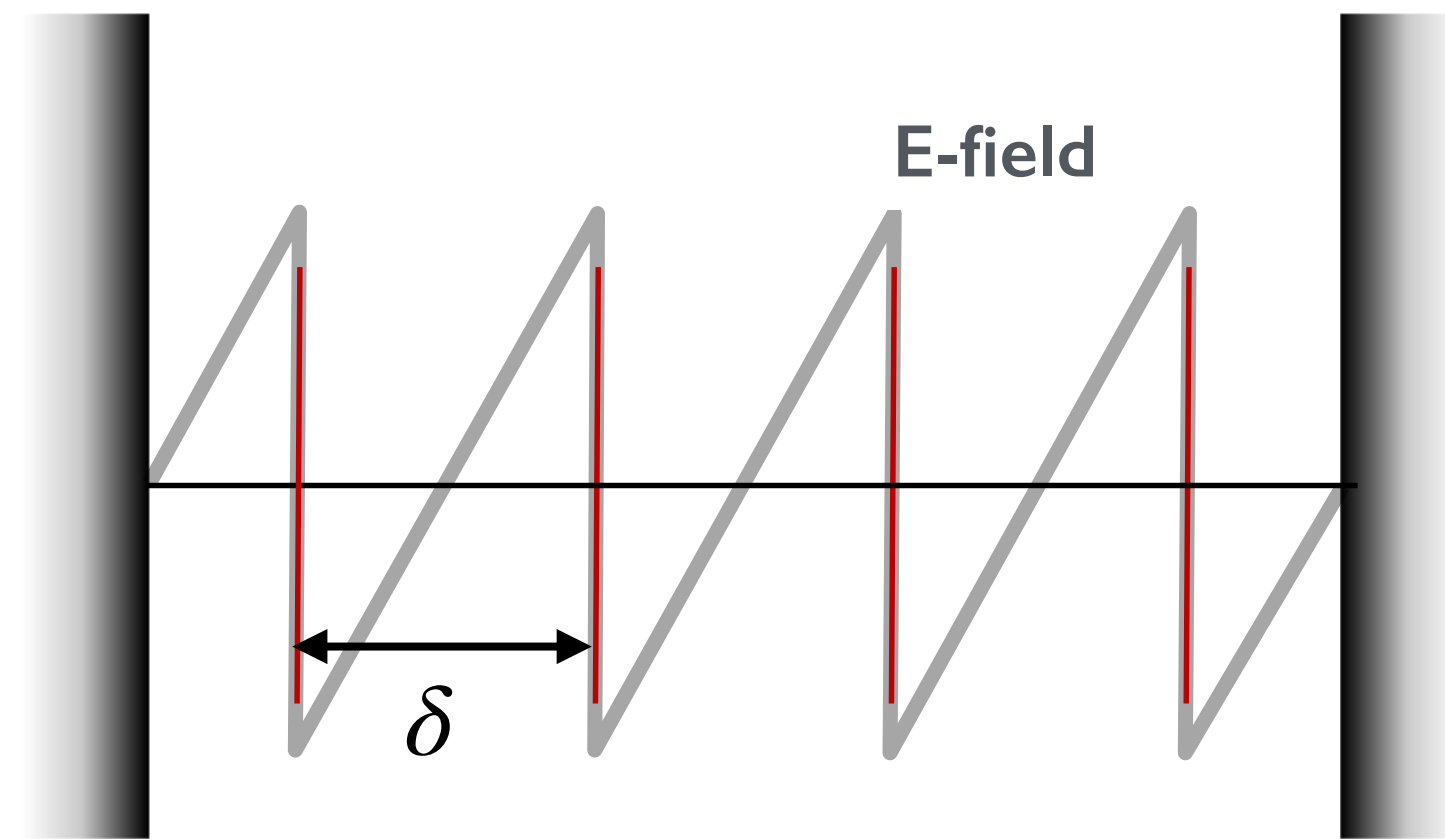
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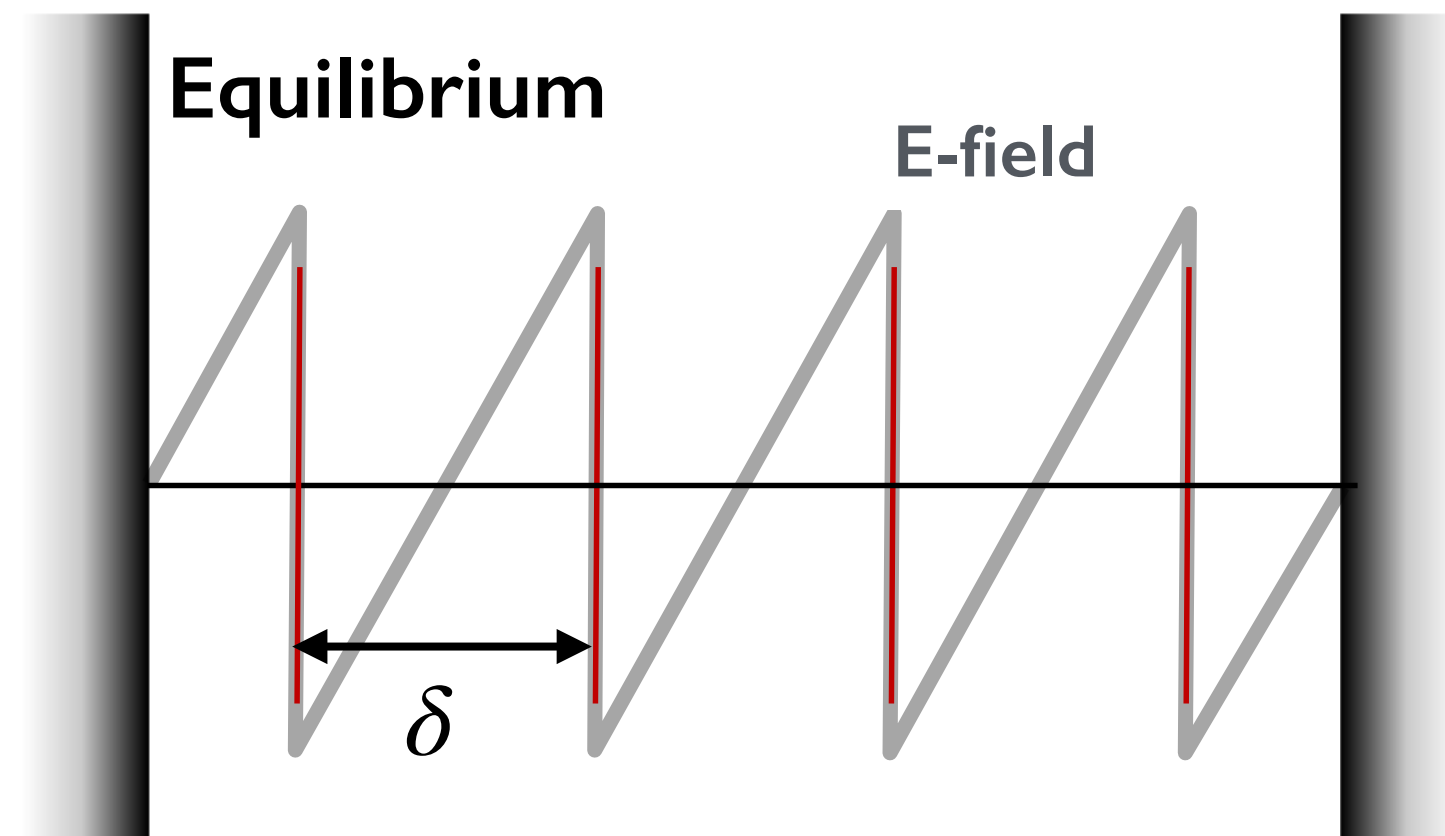
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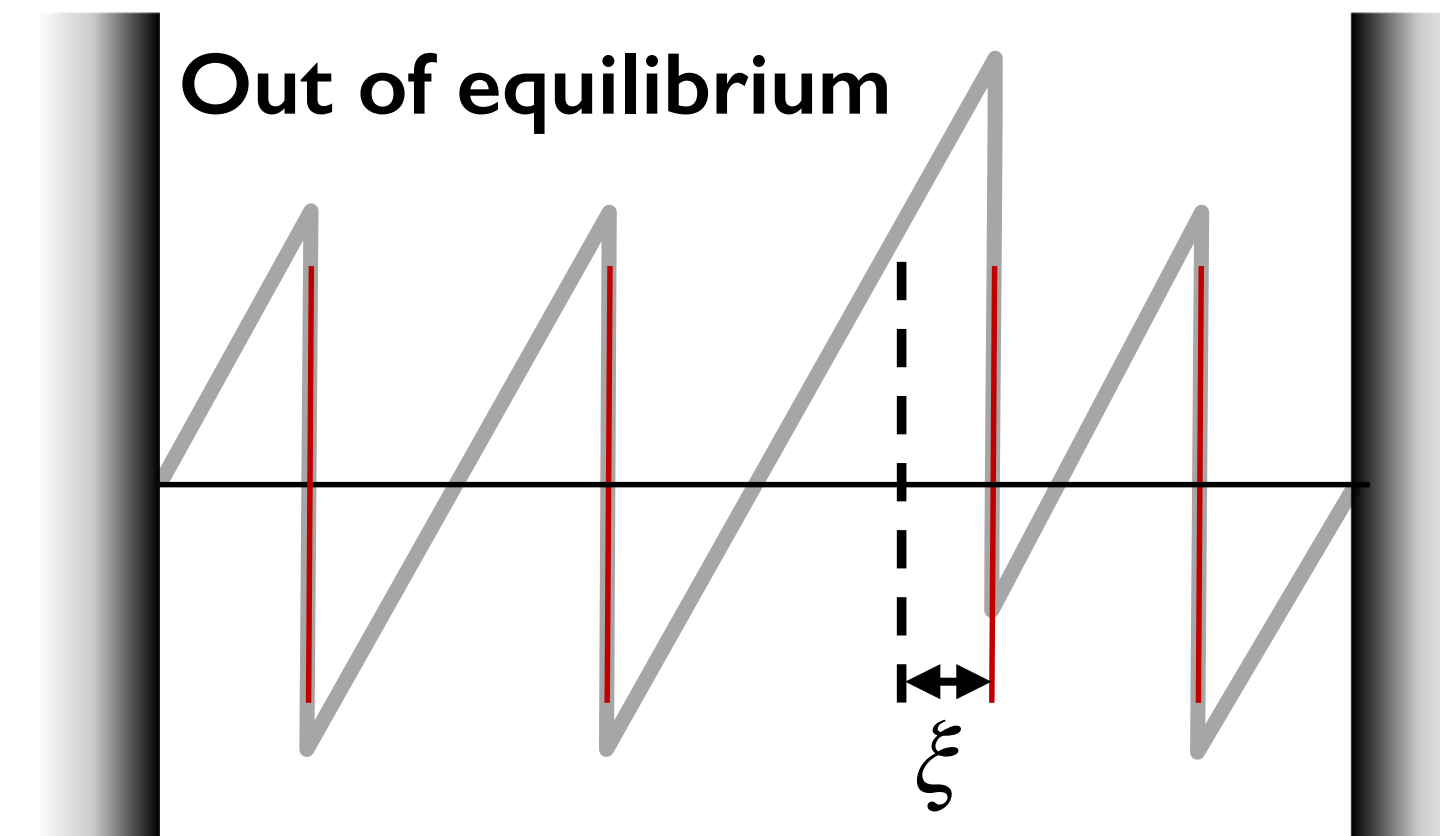
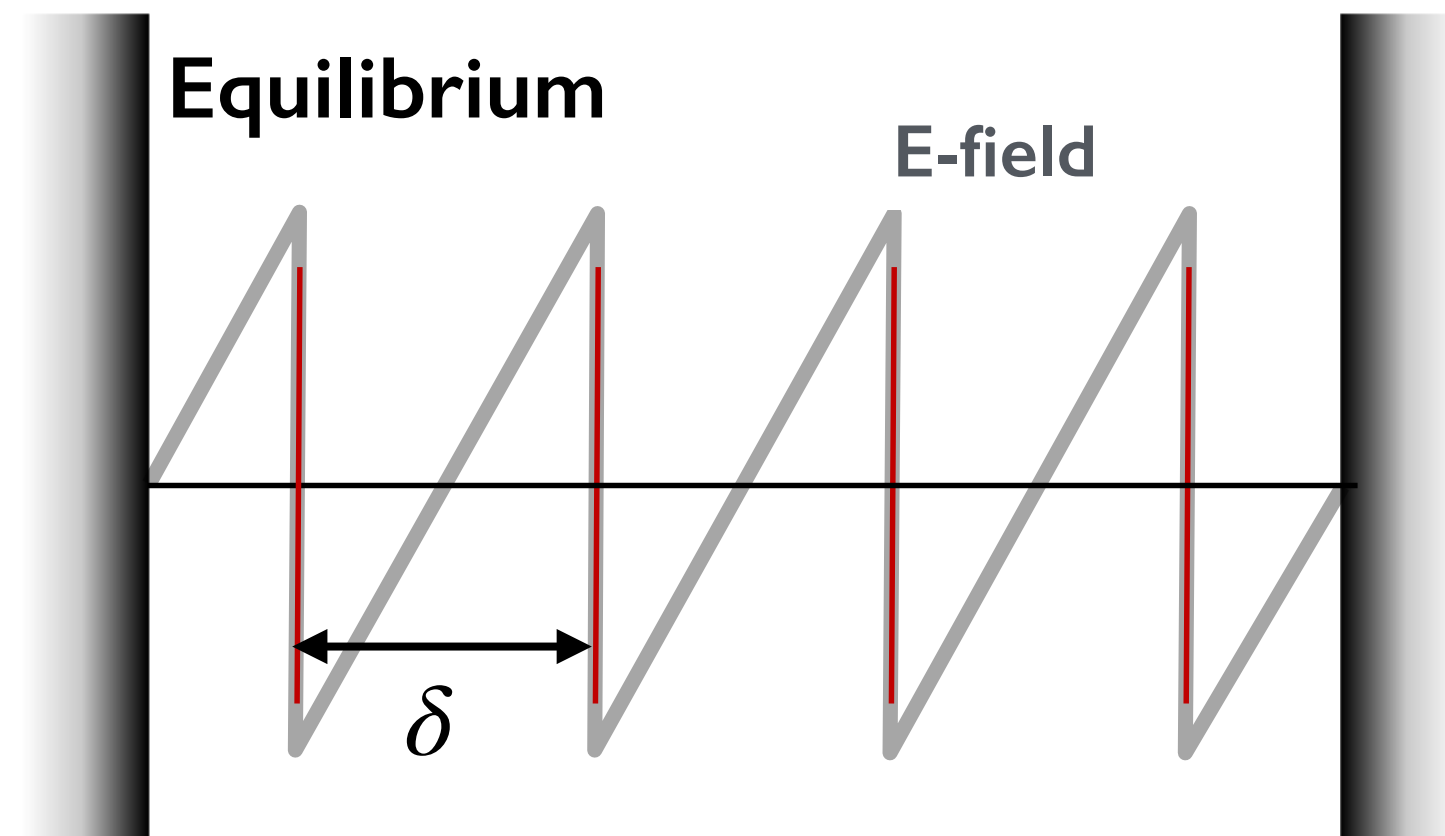
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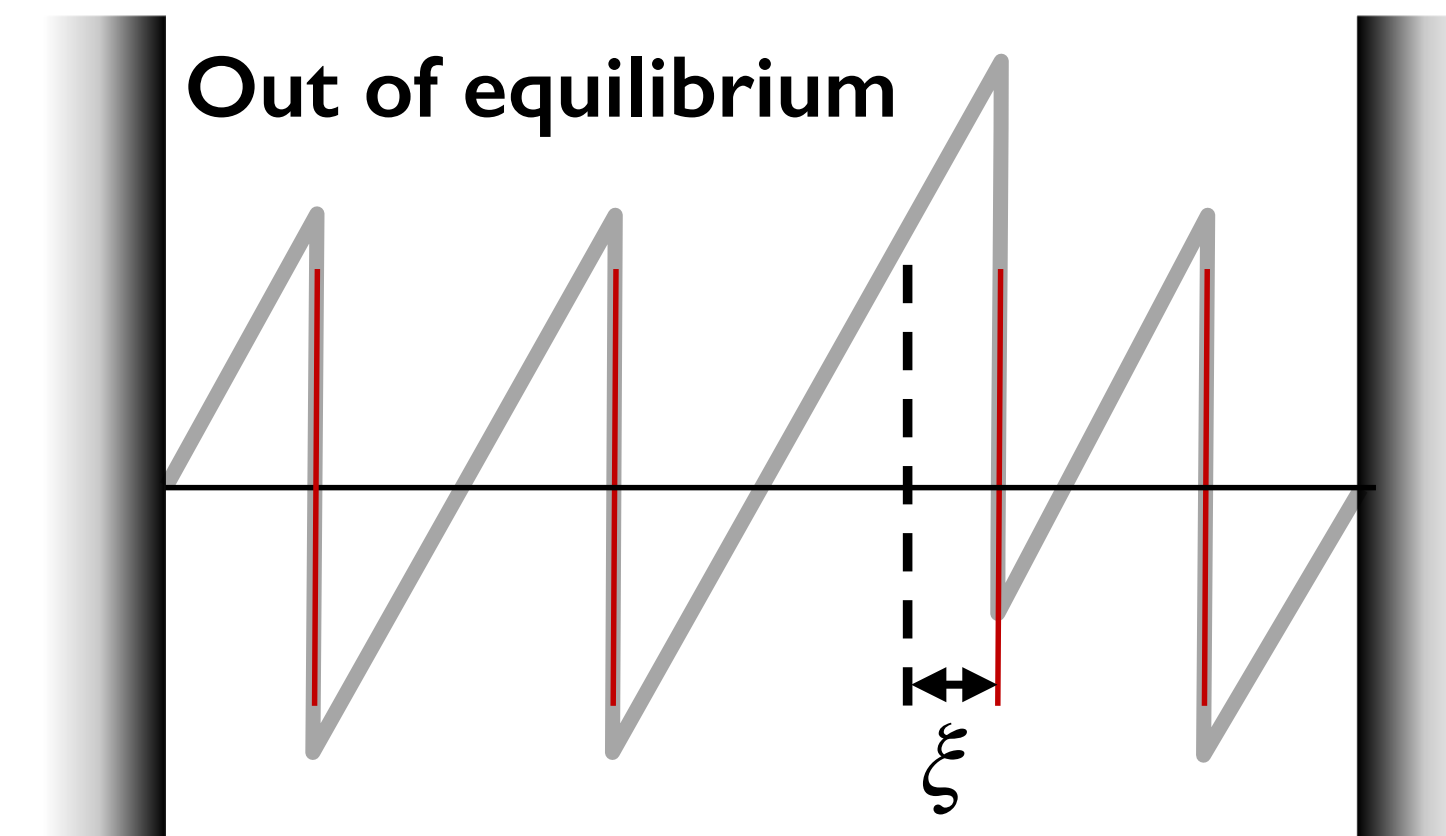
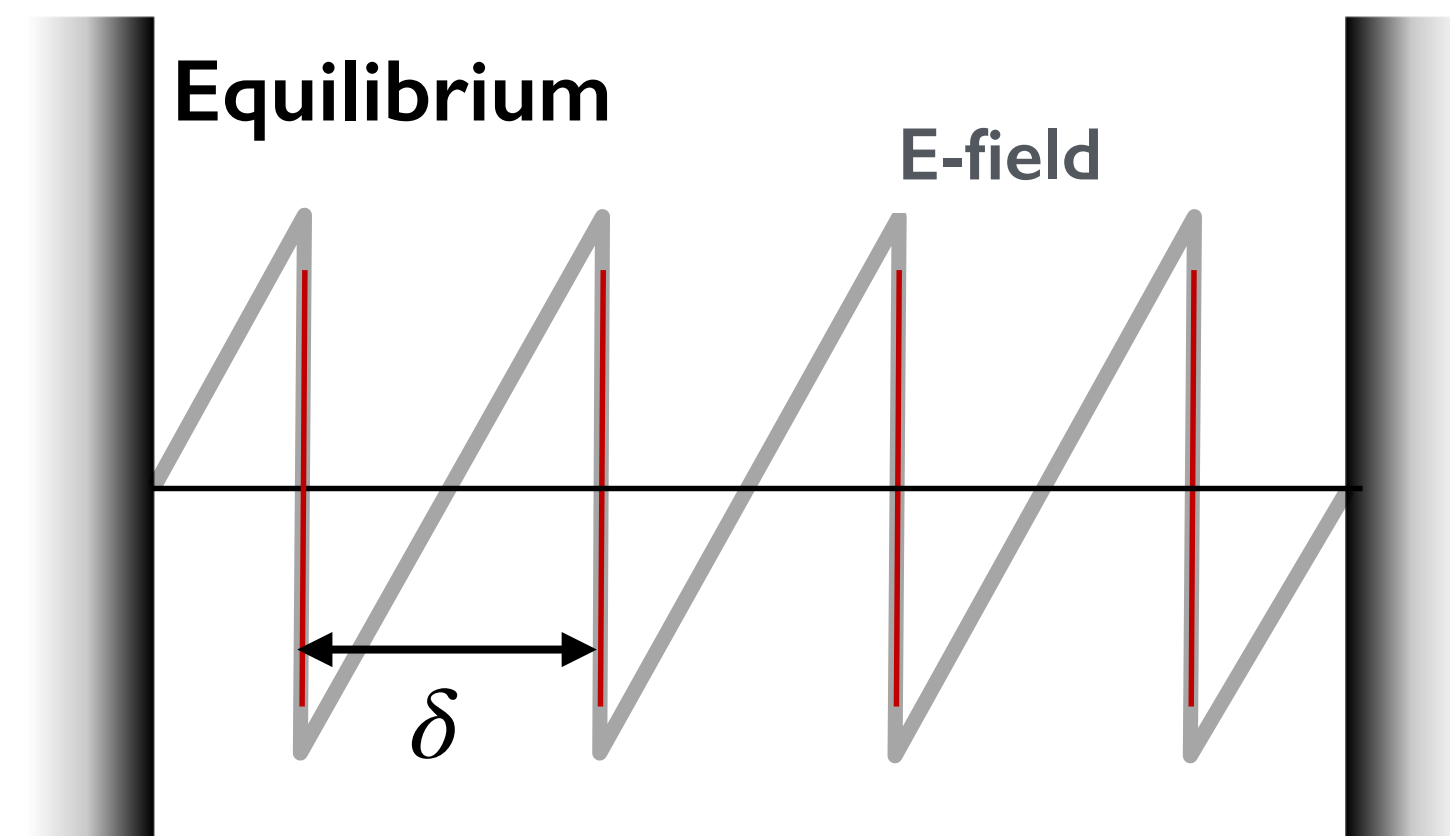


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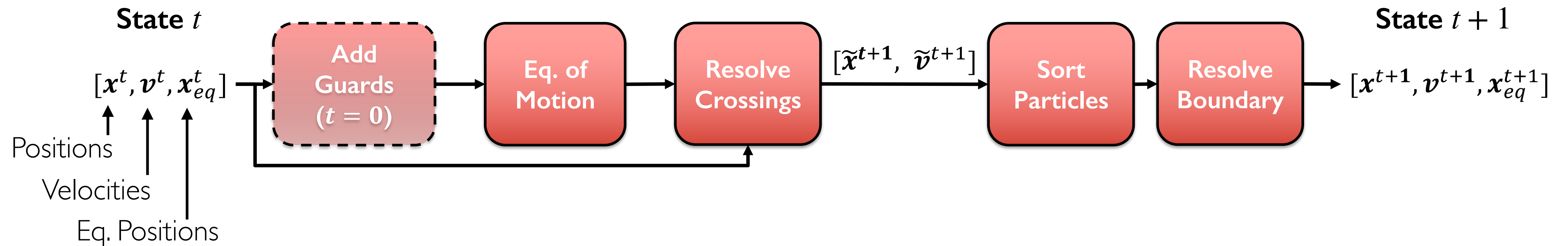


$$\ddot{\xi} = -\frac{4\pi e^2 n_0}{m_e} \xi = -\omega_p^2 \xi$$

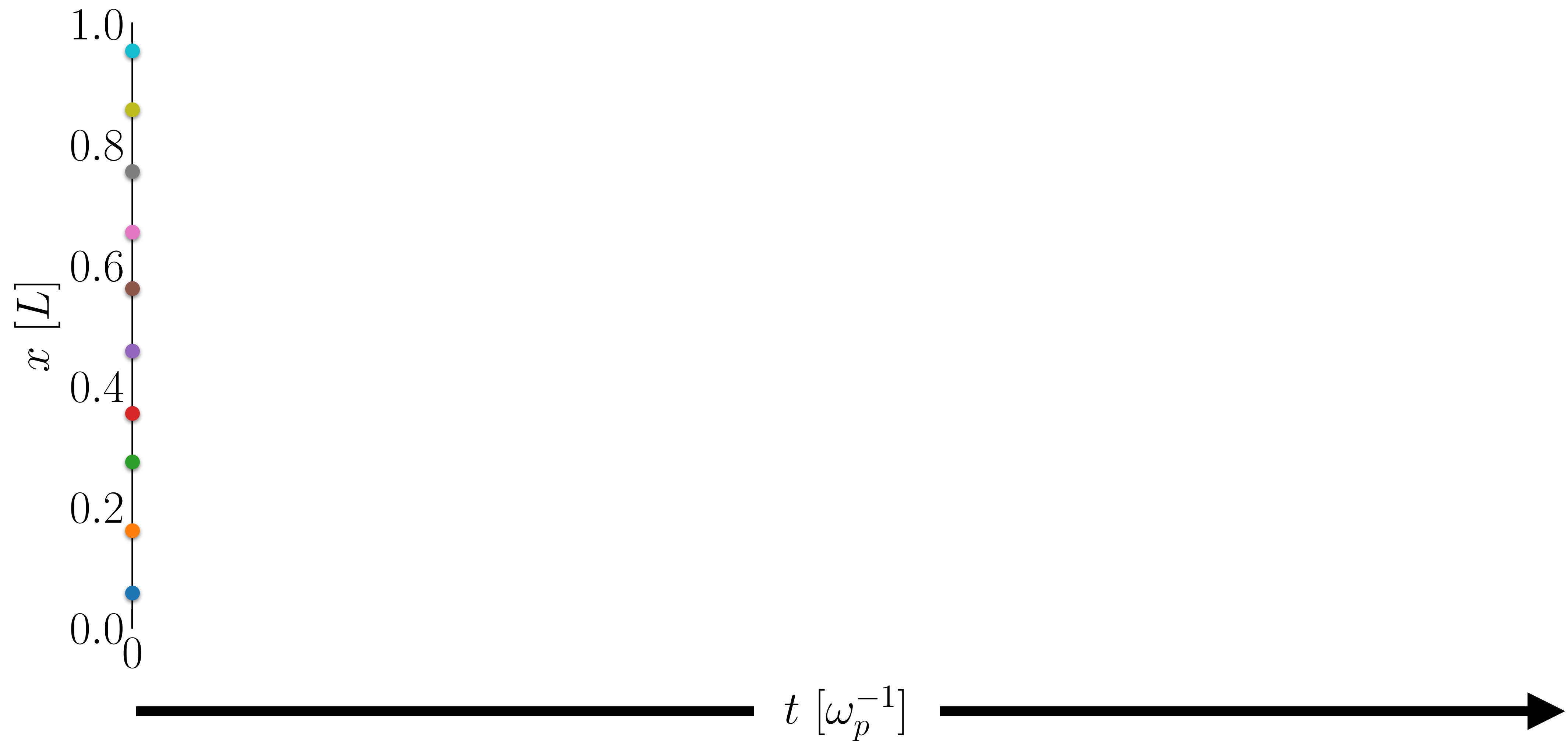
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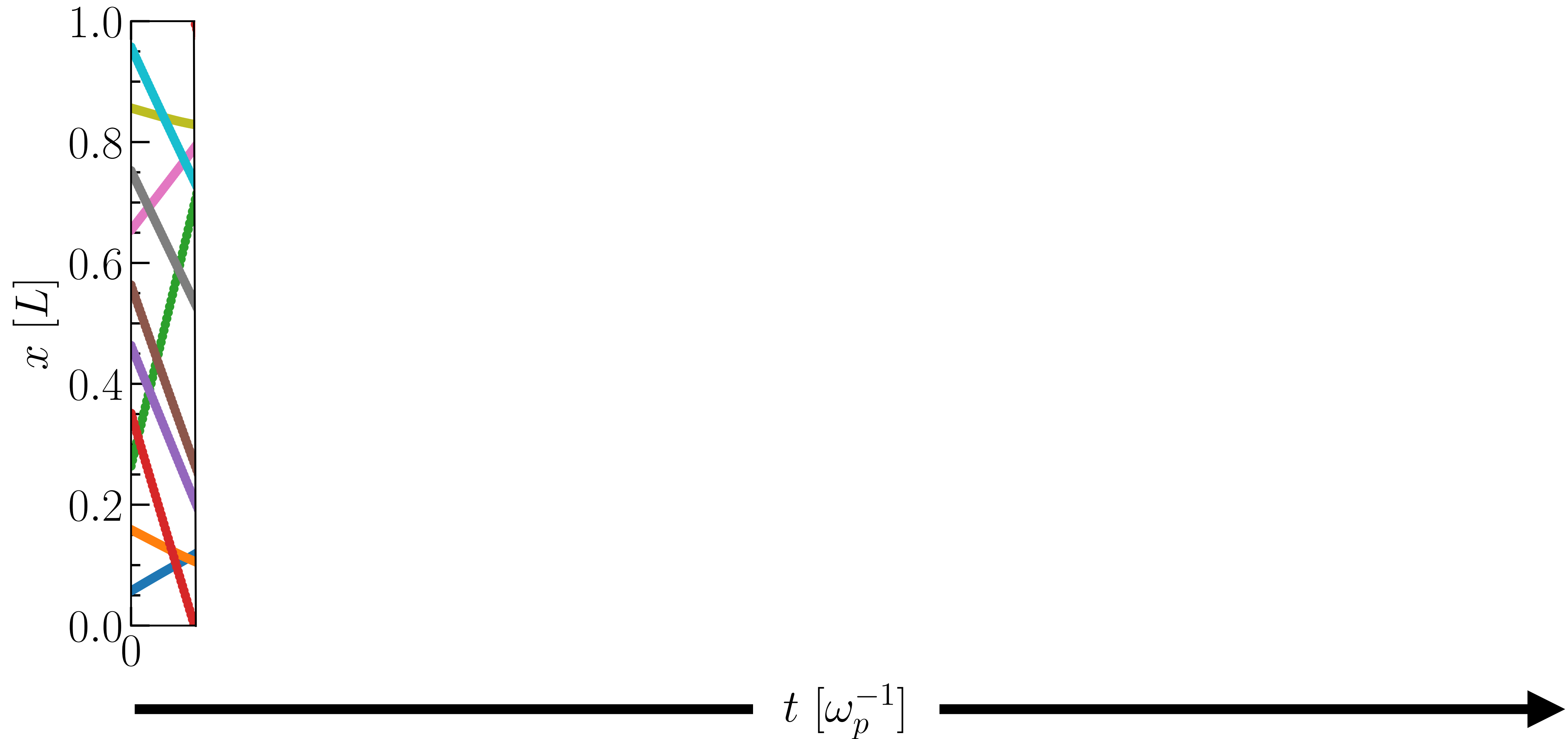
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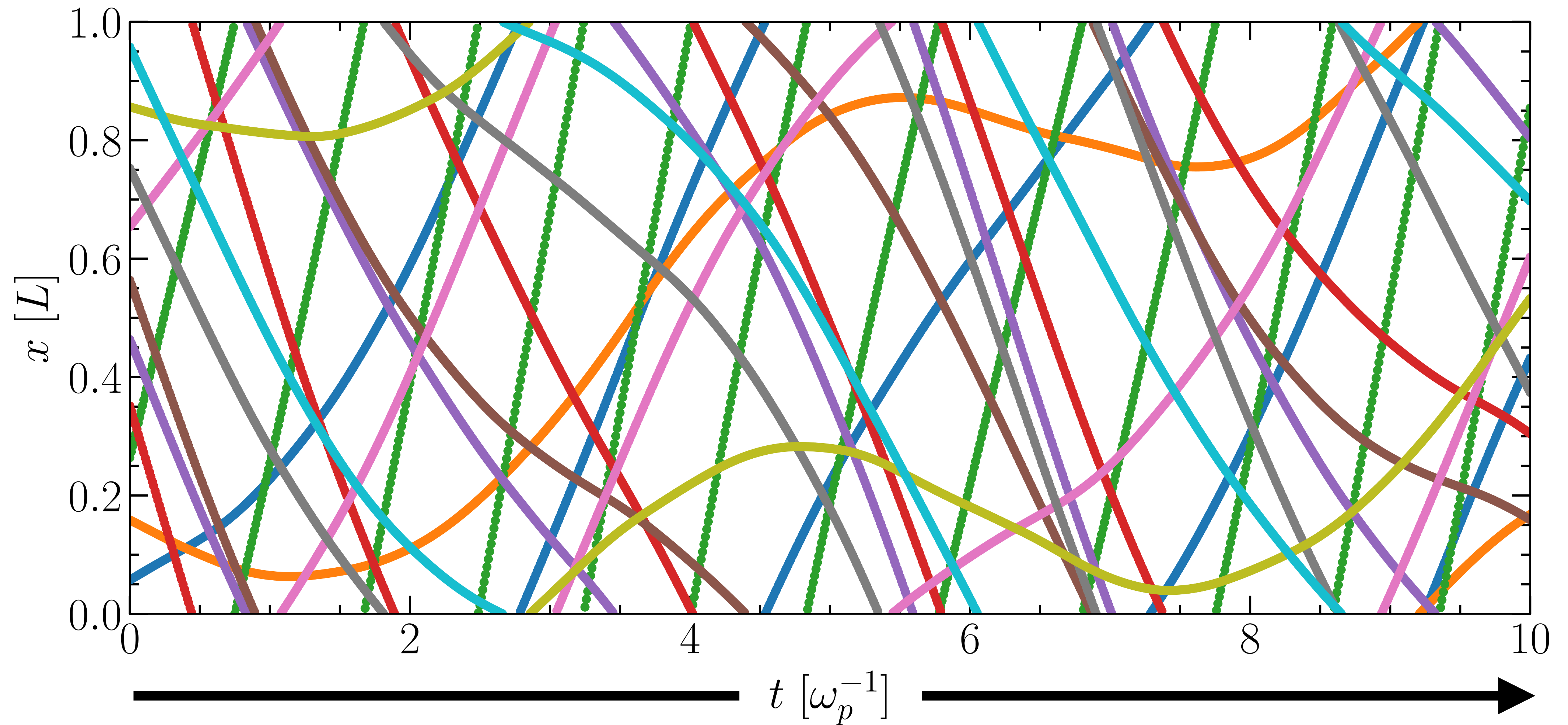
Example of Simulation



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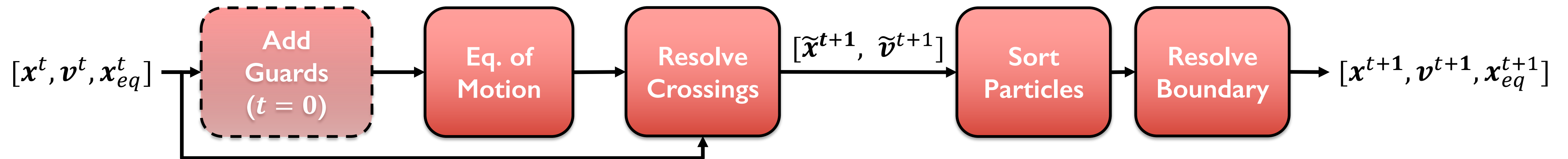


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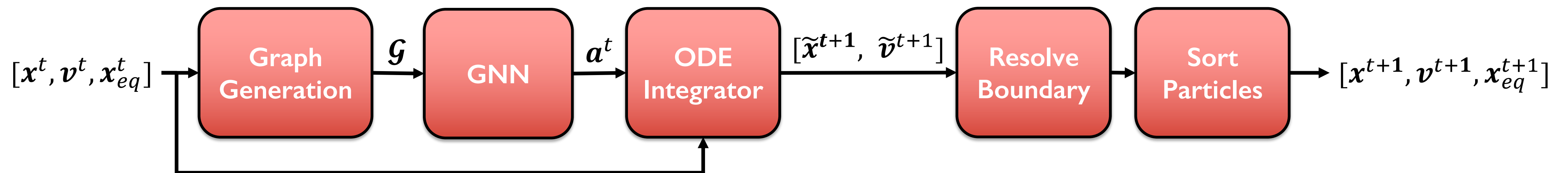


1D Plasma ESM Graph Network Simulator

Sheet Model

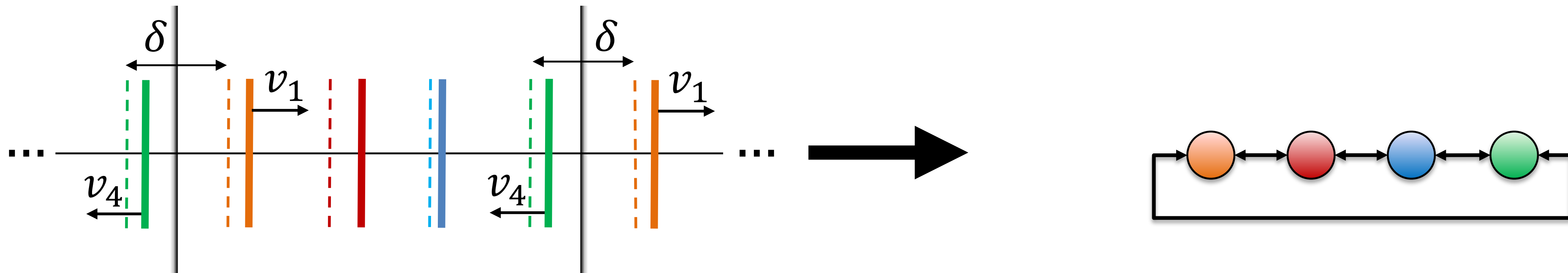


Graph Network Simulator



How do we represent the plasma as a graph?

Periodic Boundaries



i Node

$$\mathbf{n}_i^t = [\xi_i^t, v_i^t]$$

$i \rightarrow j$ Edge

$$\mathbf{r}_{ji}^t = [x_i^t - x_j^t]$$

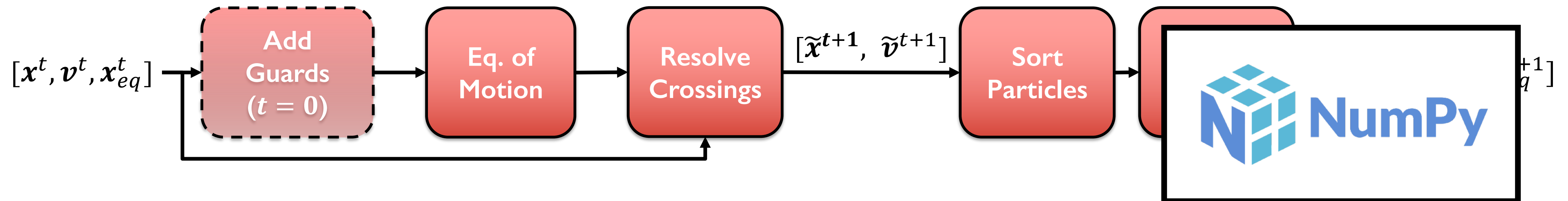
Target

$$a_i^t = \frac{v_i^{t+1} - v_i^t}{\Delta t}$$

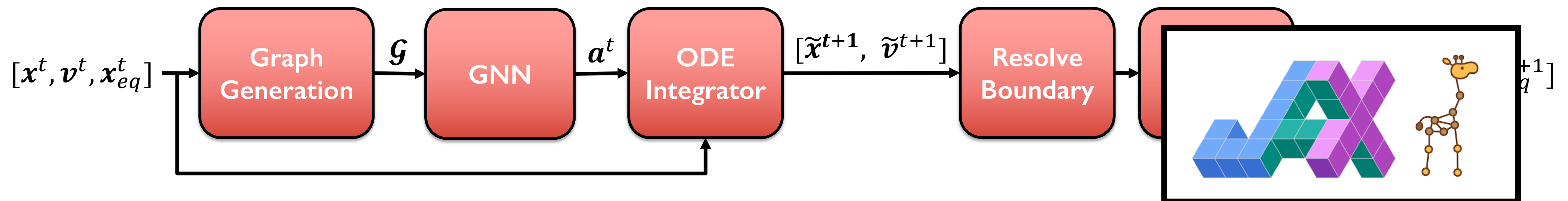
All values are normalised to the intersheet spacing δ

1D Plasma ESM Graph Network Simulator

Sheet Model



Graph Network Simulator



Code: <https://github.com/diogodcarvalho/gns-sheet-model>

<https://github.com/google/jax>

<https://github.com/deepmind/jraph>

GNS generalizes to different number of sheets and boundary conditions

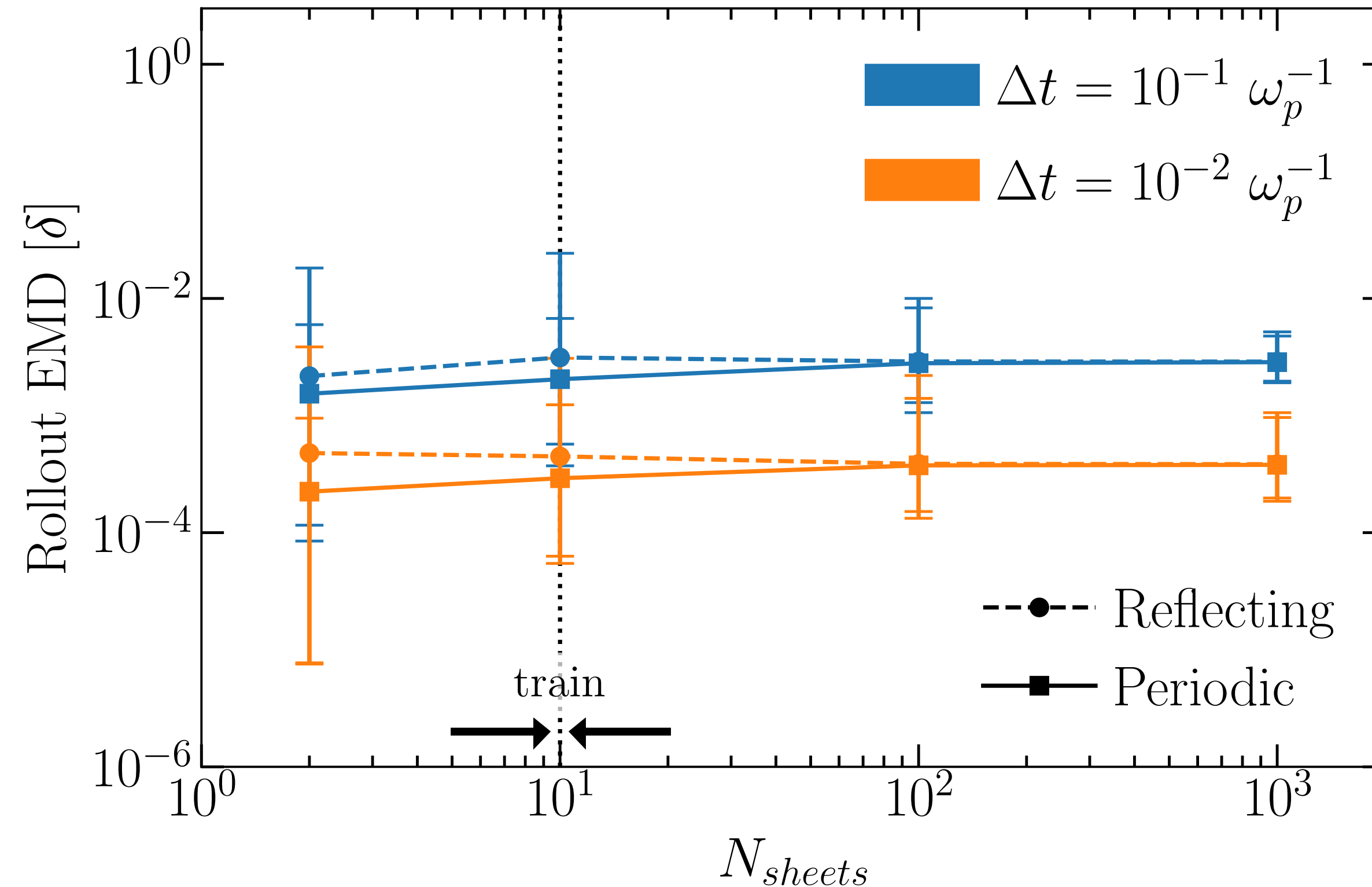
Trained on **subsampled high temporal resolution data** $\left(\Delta t_{orig} = 10^{-4} \omega_p^{-1}\right)$ of **10 sheets**
moving inside a **periodic box** $\left(t_{sim} = 10 \omega_p^{-1}\right)$

Initial positions and velocities are randomly sampled from a uniform distribution

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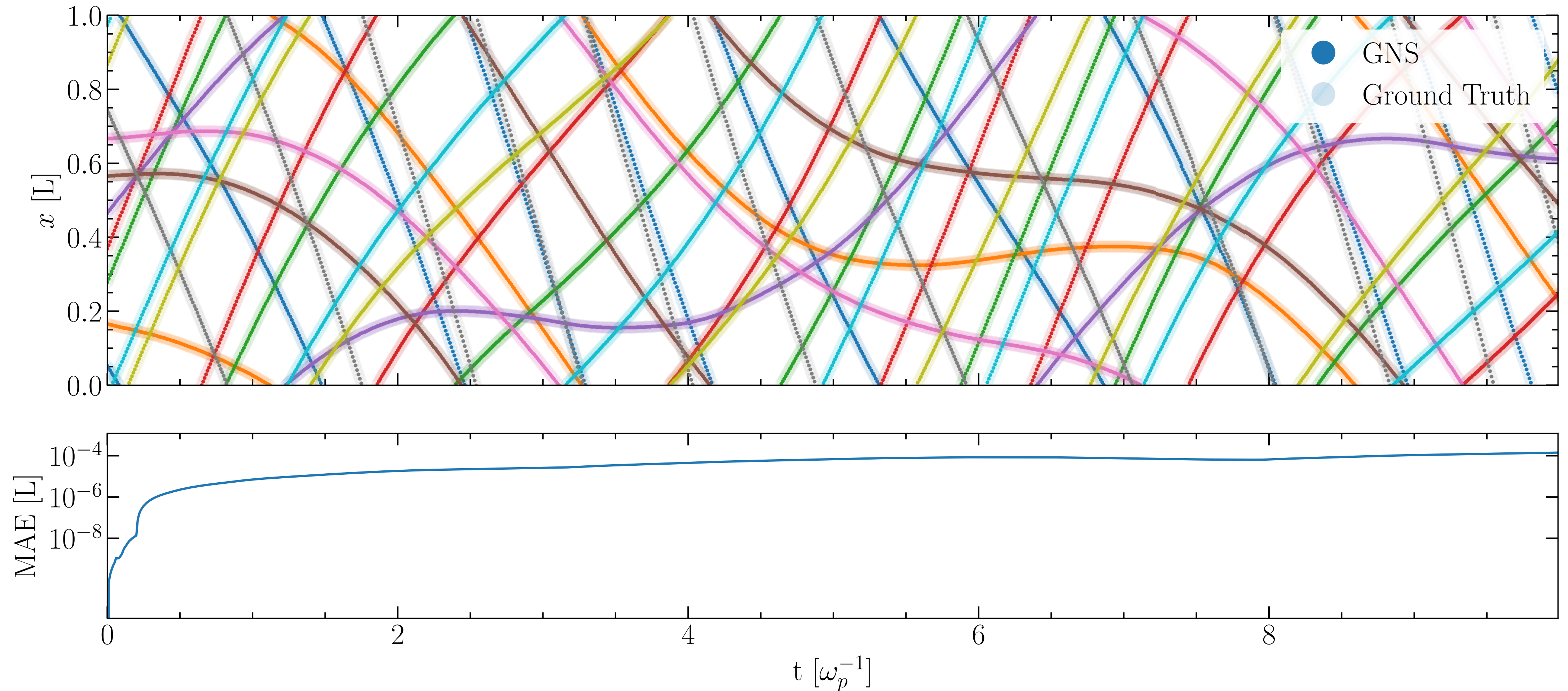
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GNS rollout errors are very small

$$\Delta t = 10^{-2} \omega_p^{-1}$$

$$\text{Rollout MAE} = 5.6 \times 10^{-4} \delta$$

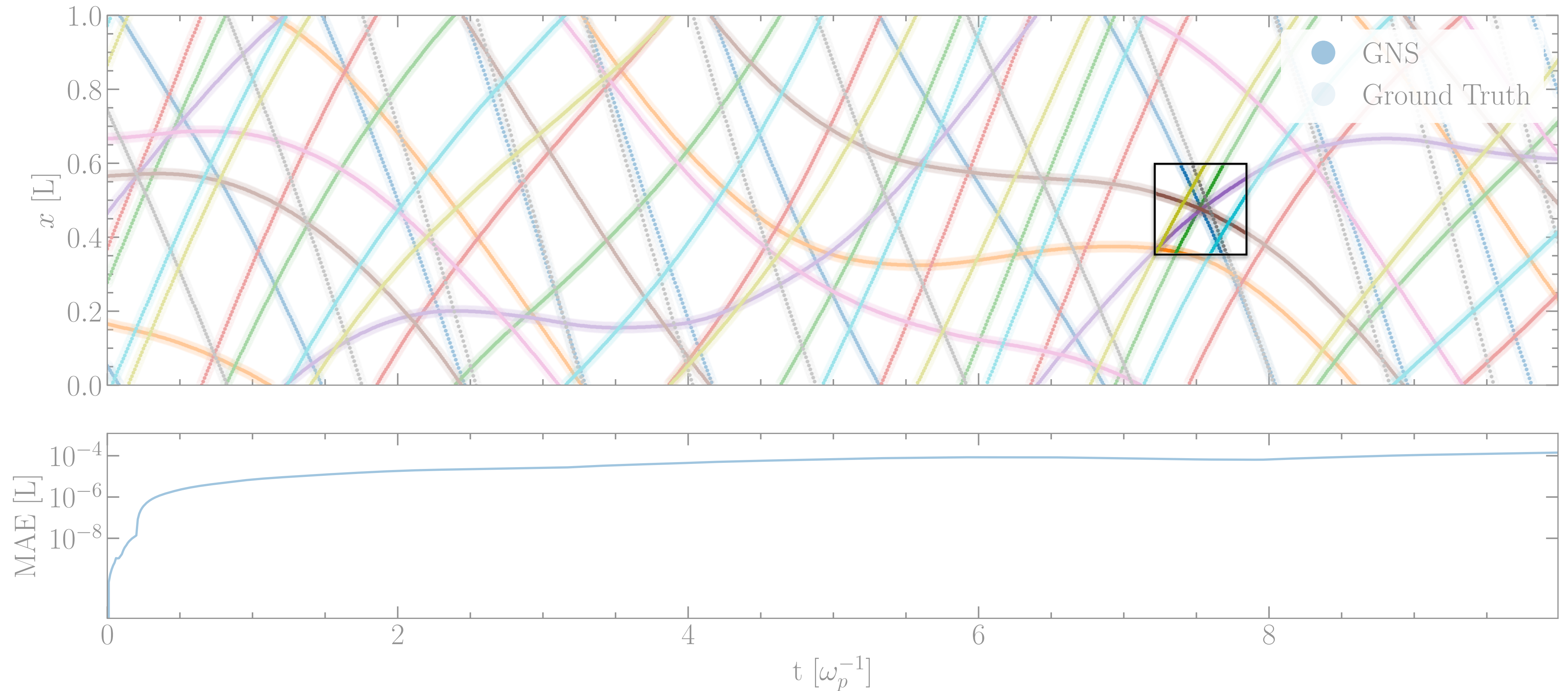


Example shown corresponds to the worst rollout error observed in the test set

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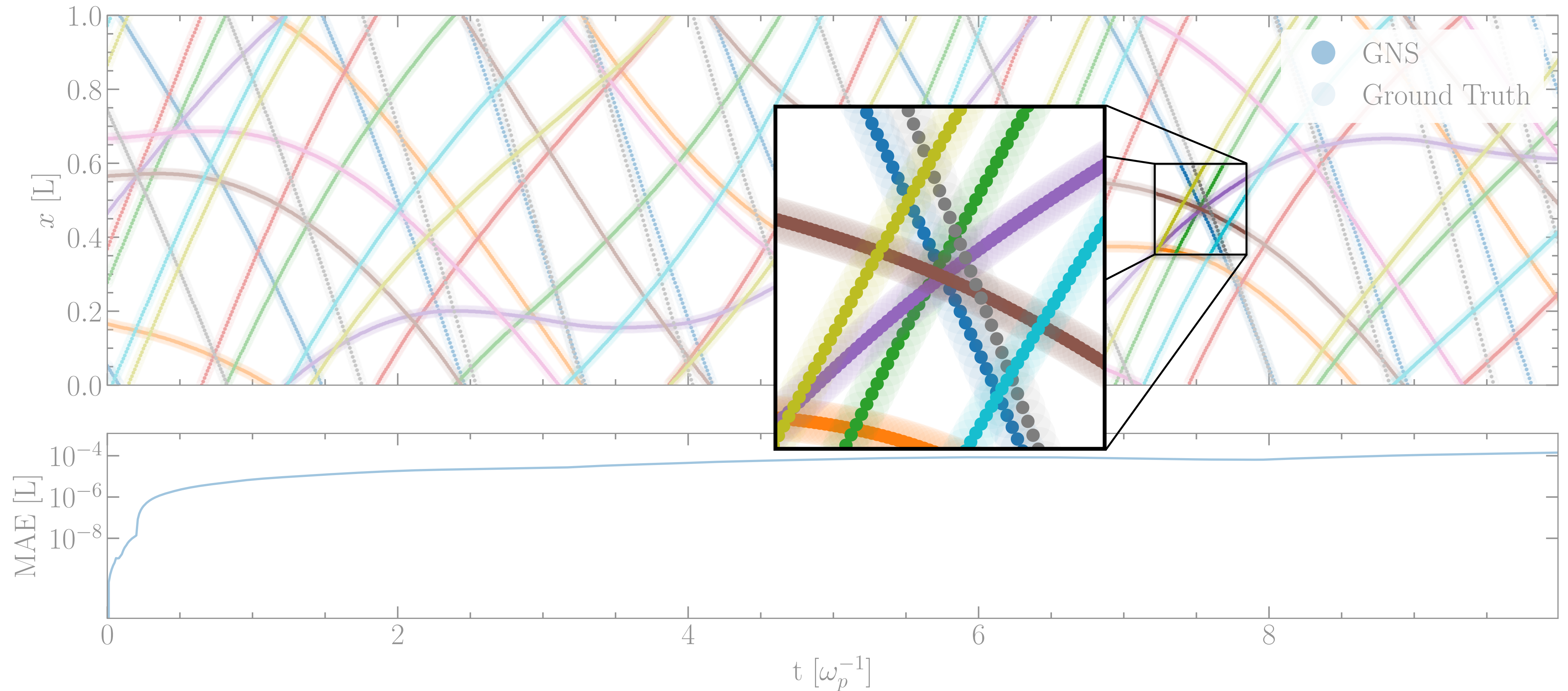


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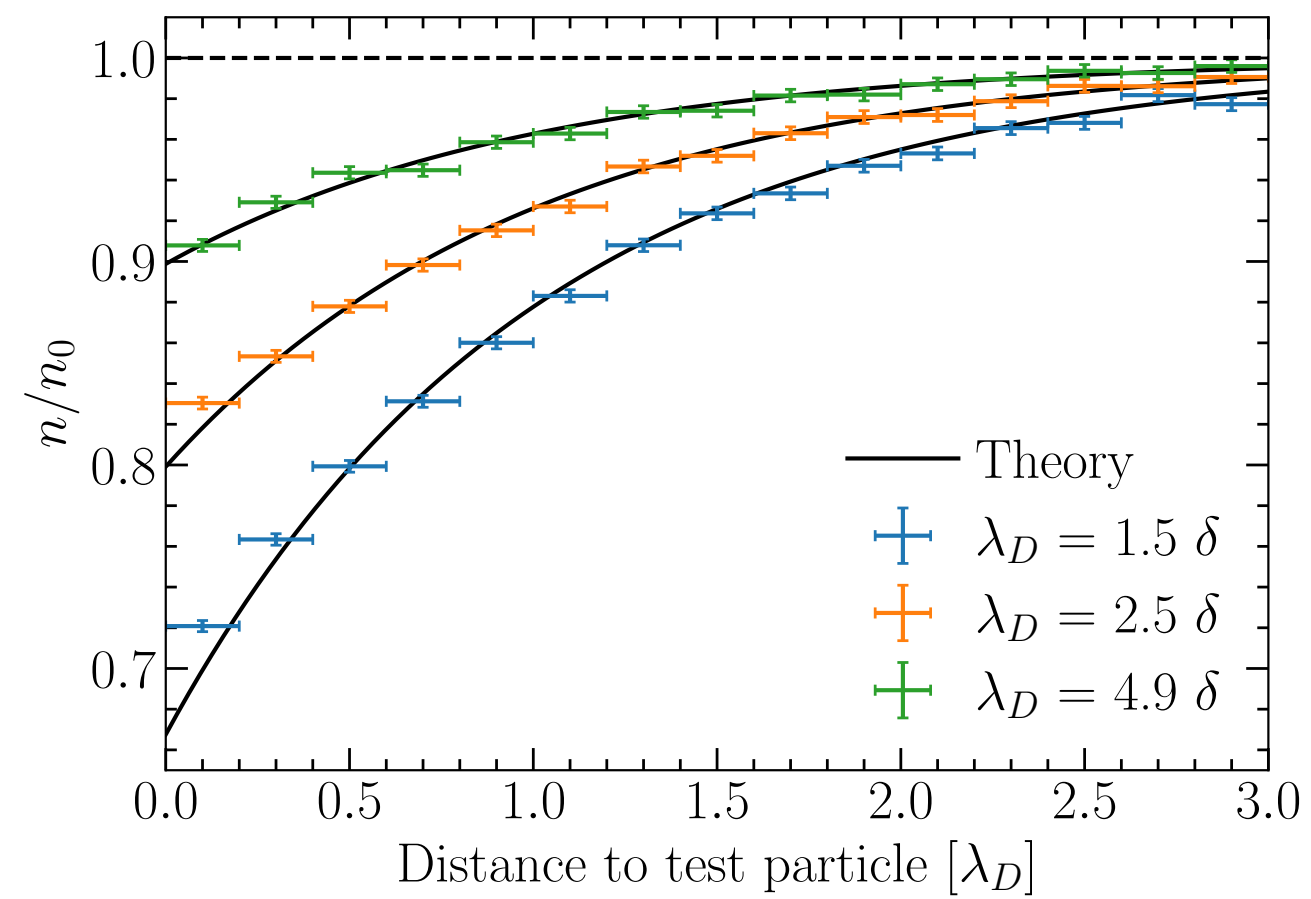
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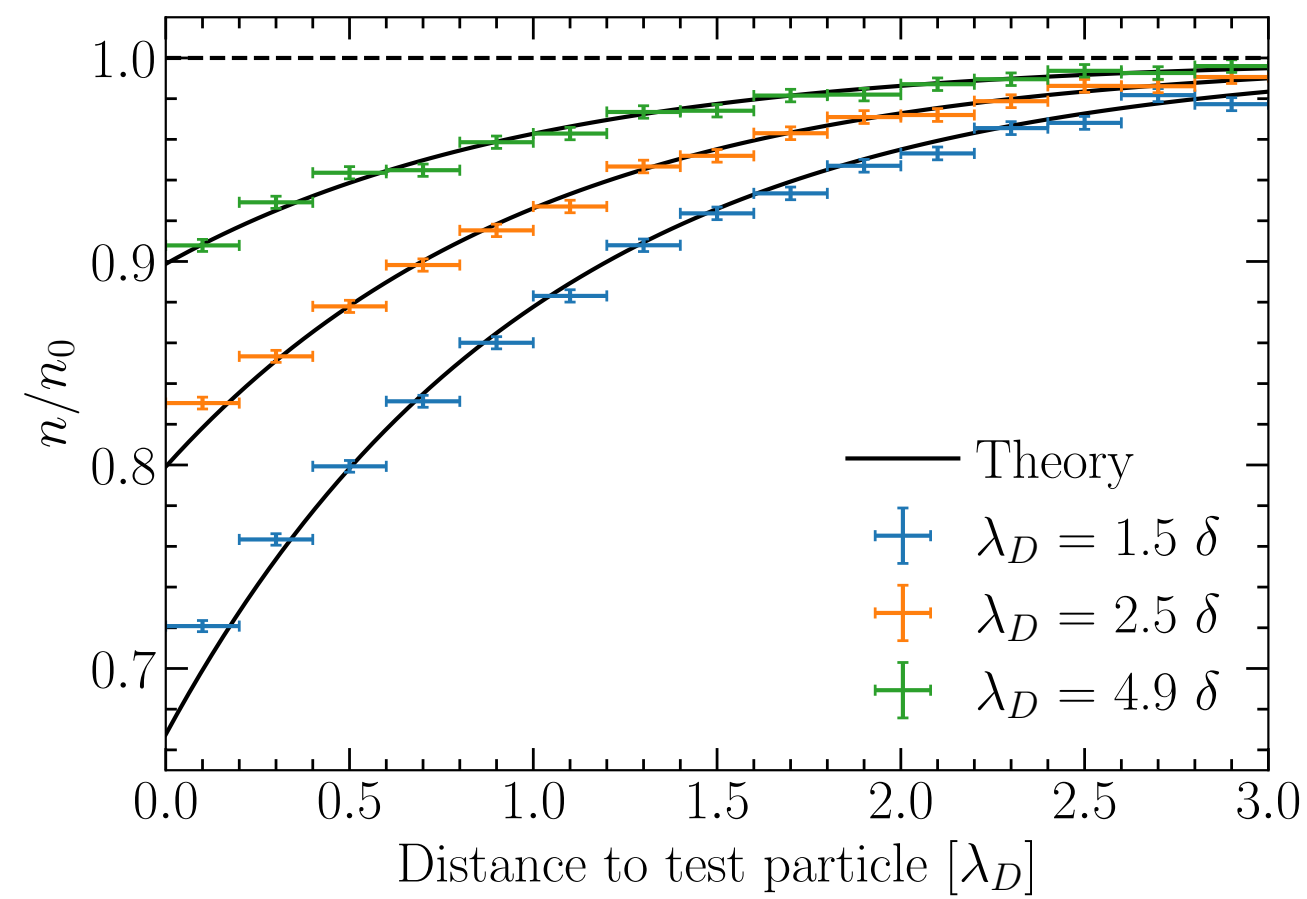
GNS recovers a broad range of kinetic plasma processes

Debye Shielding

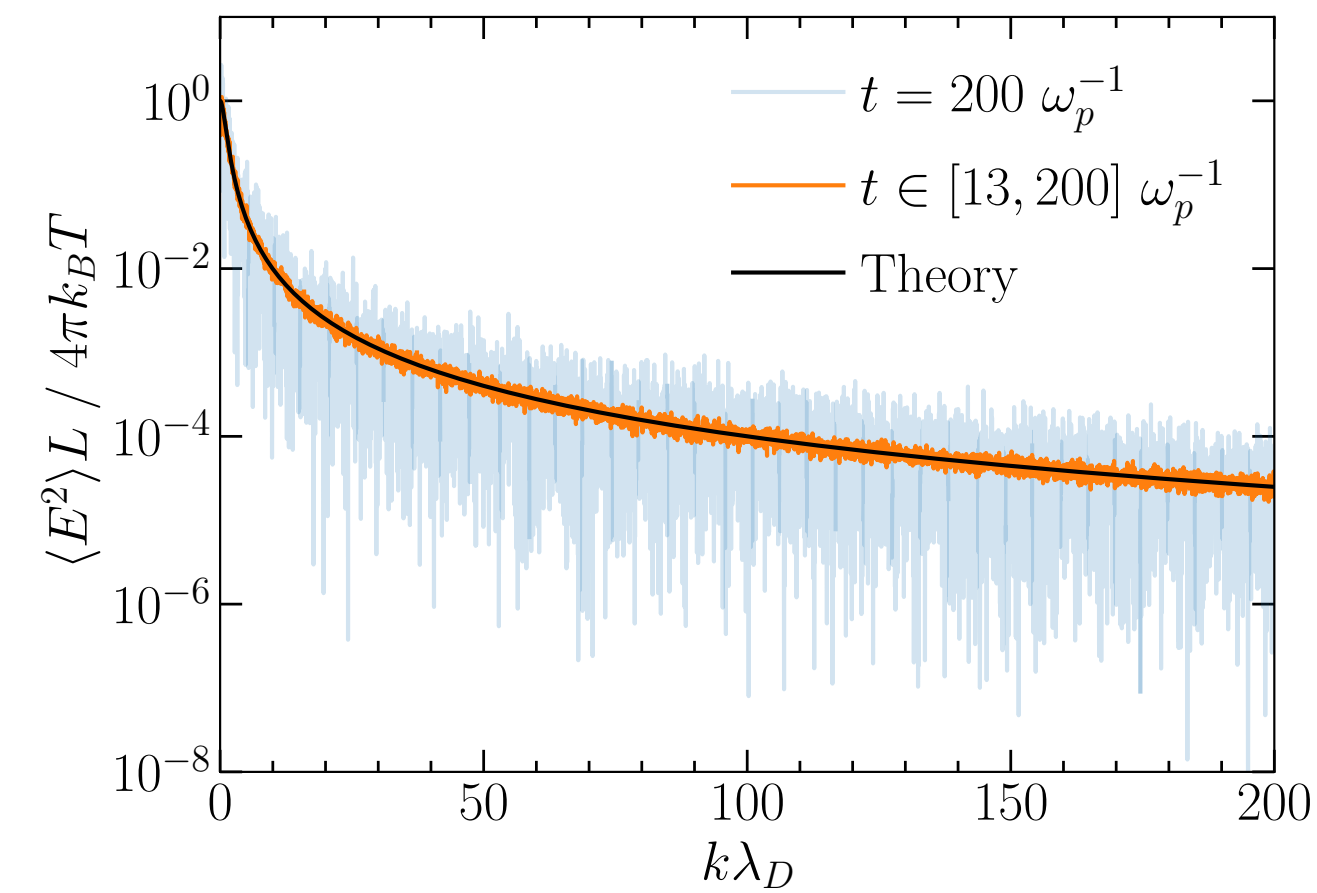


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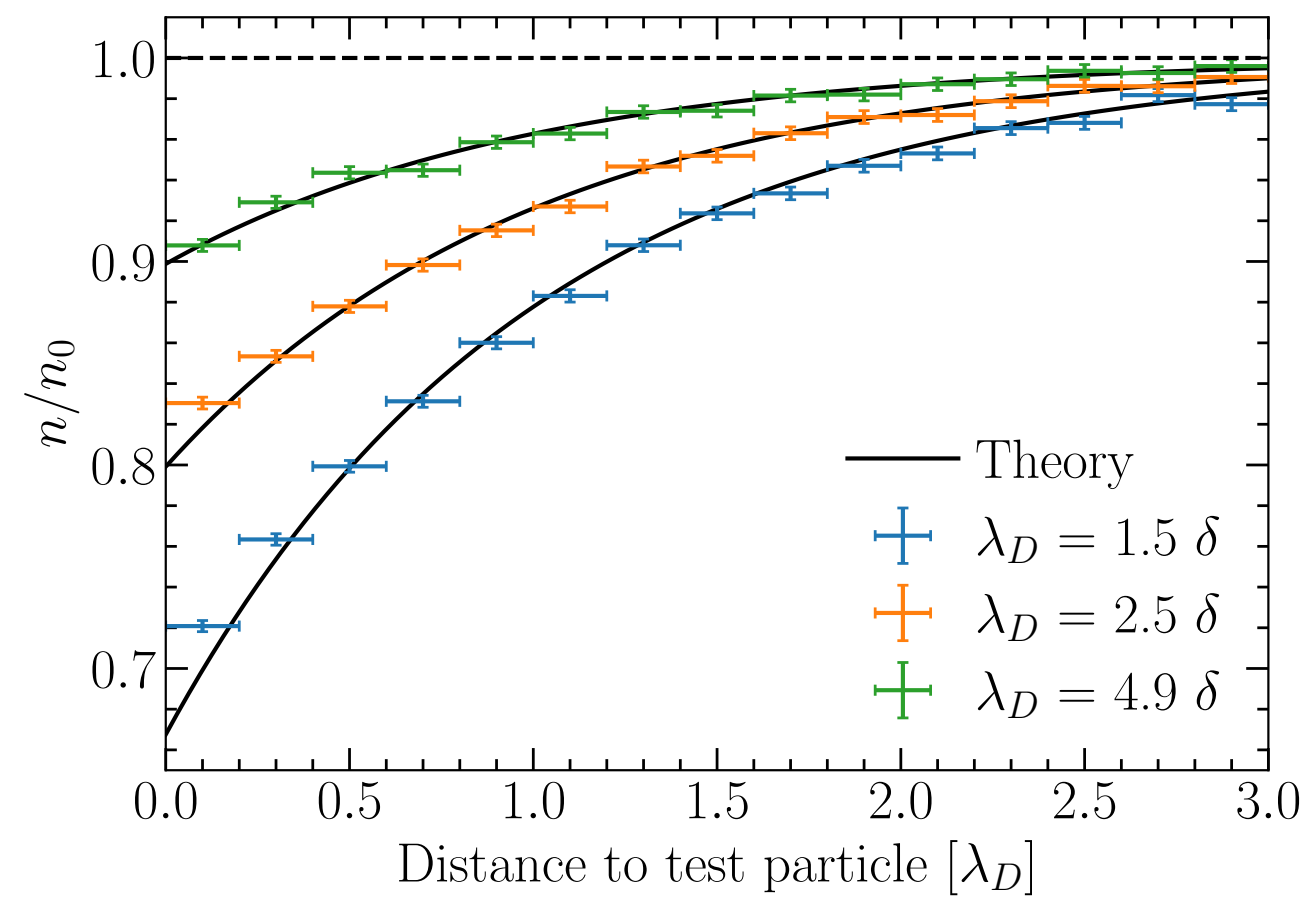


Electrostatic Fluctuations

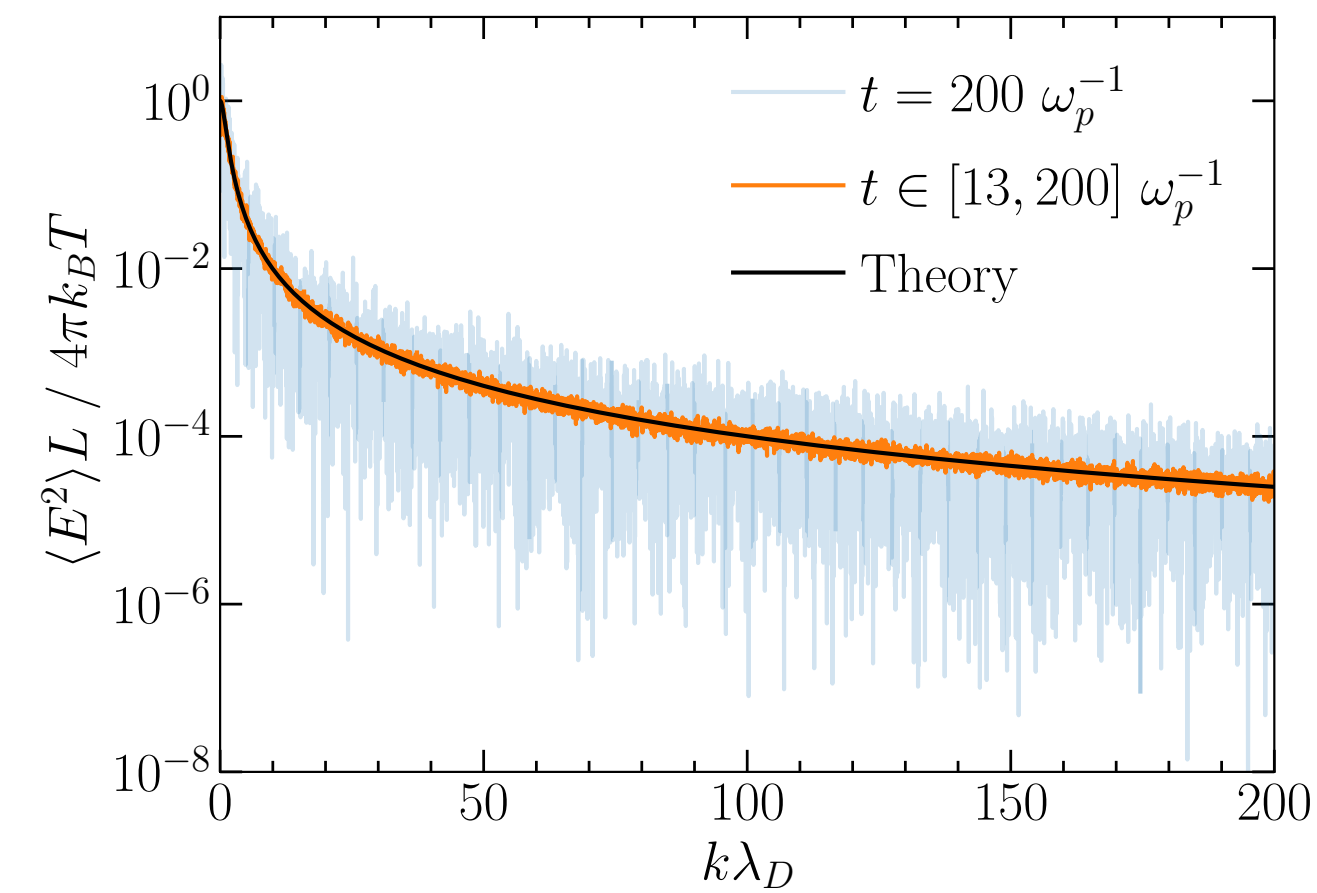


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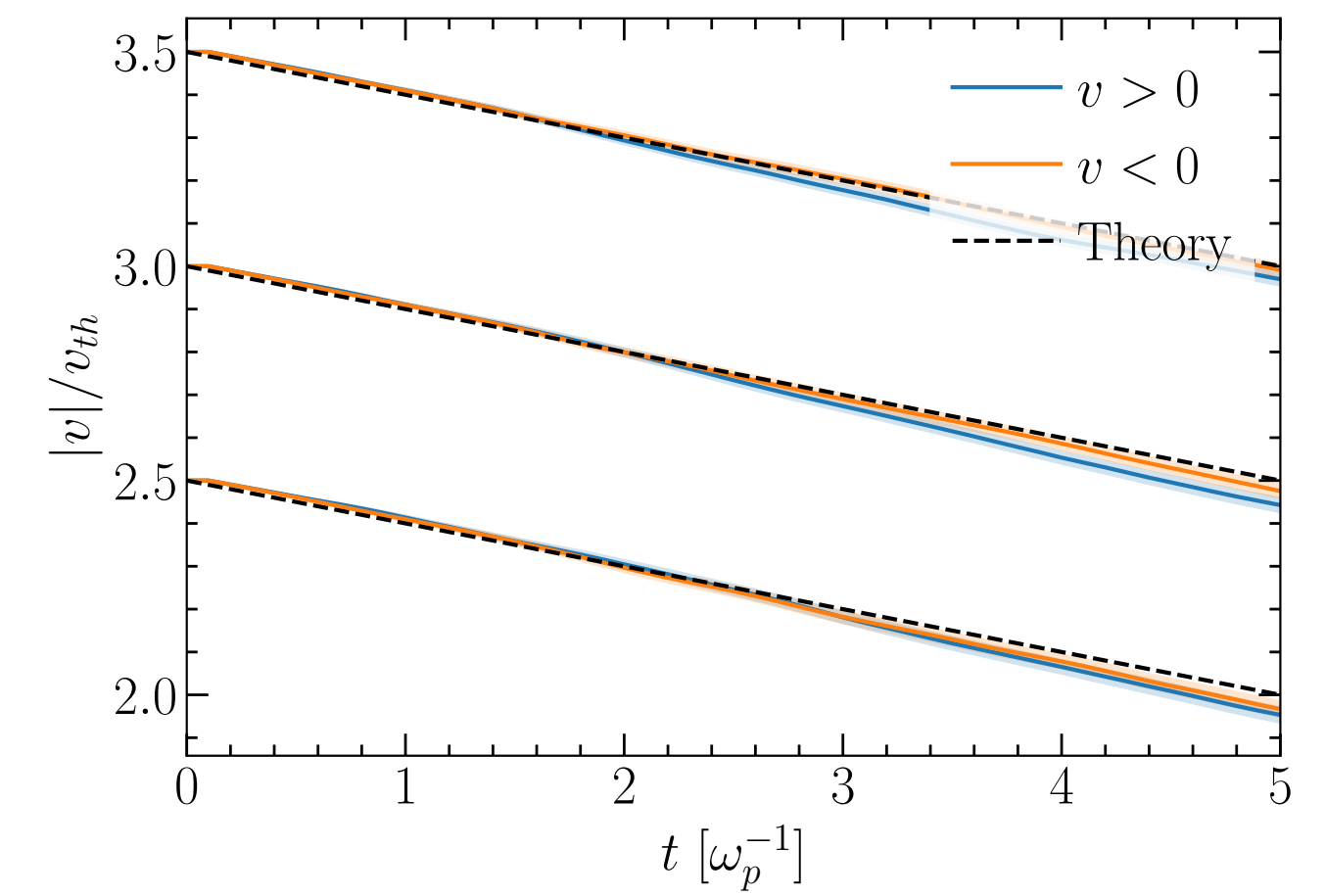
Debye Shielding



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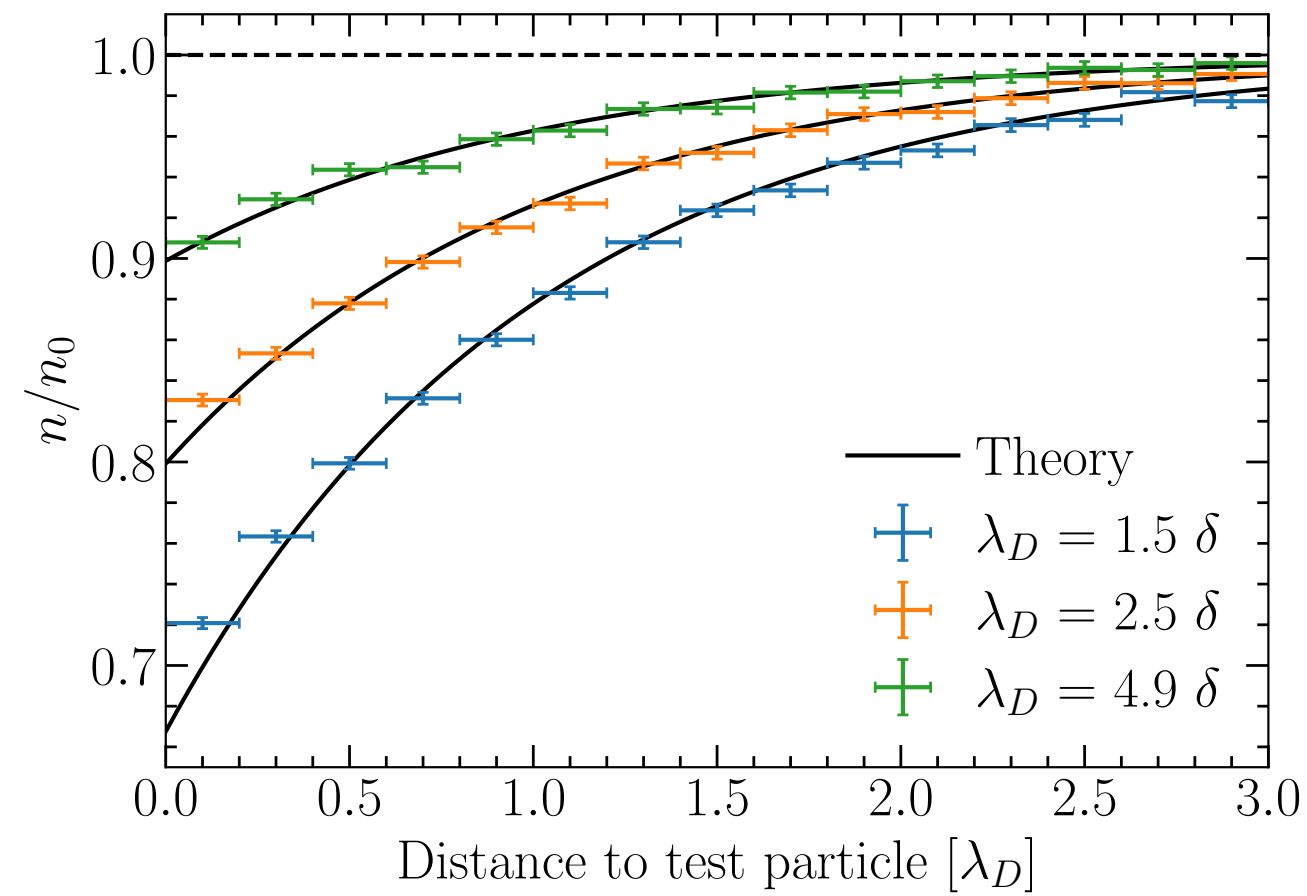


Drag on a Fast Sheet

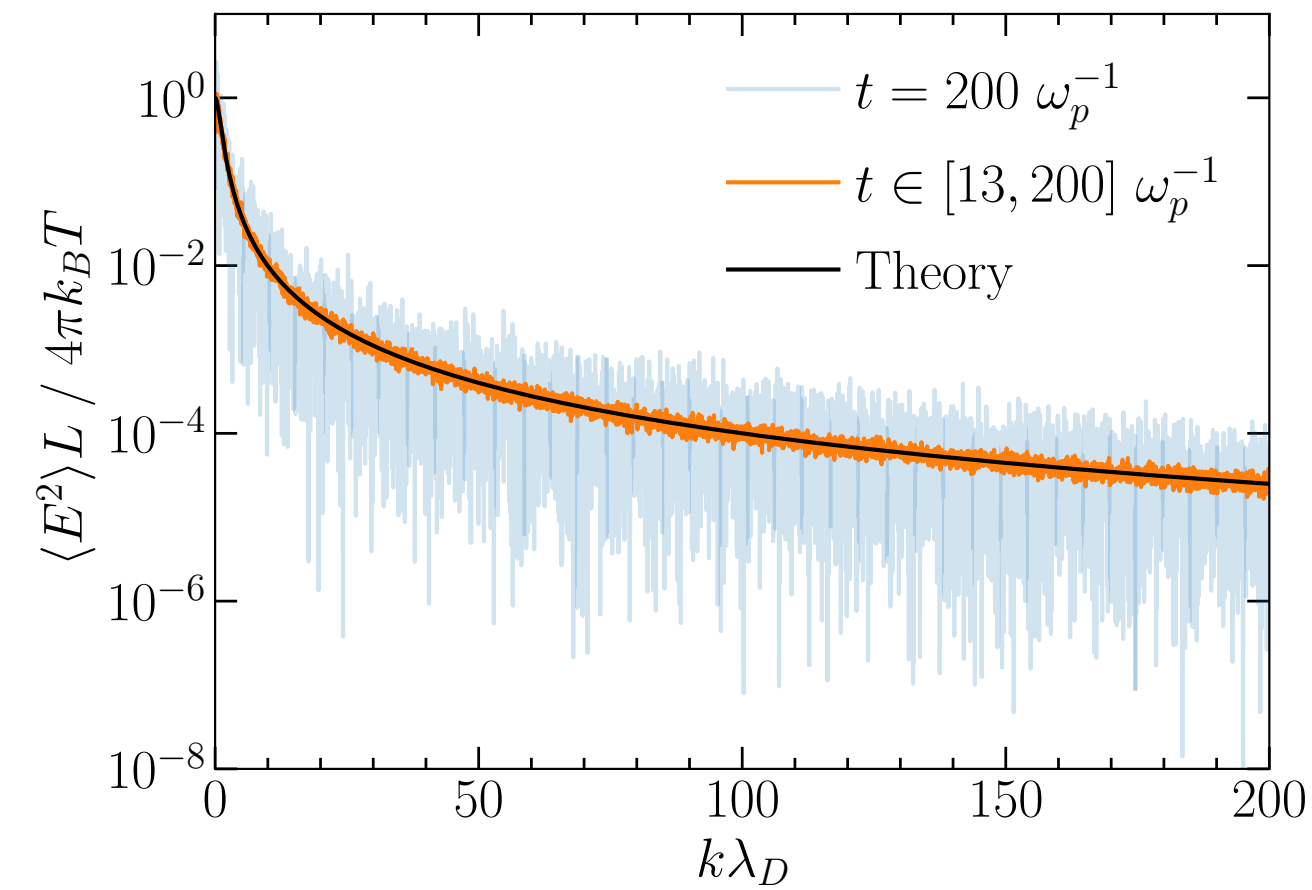


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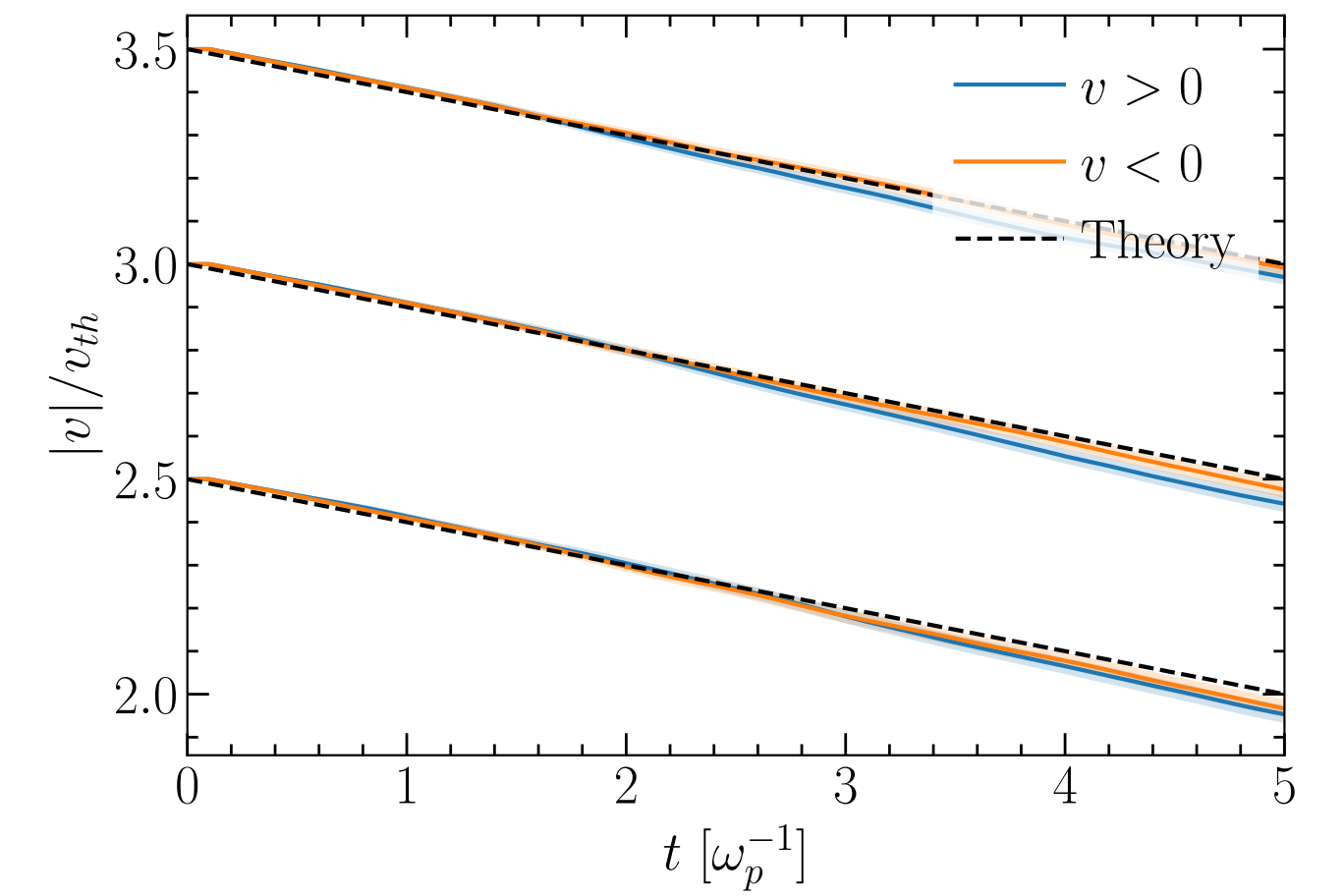
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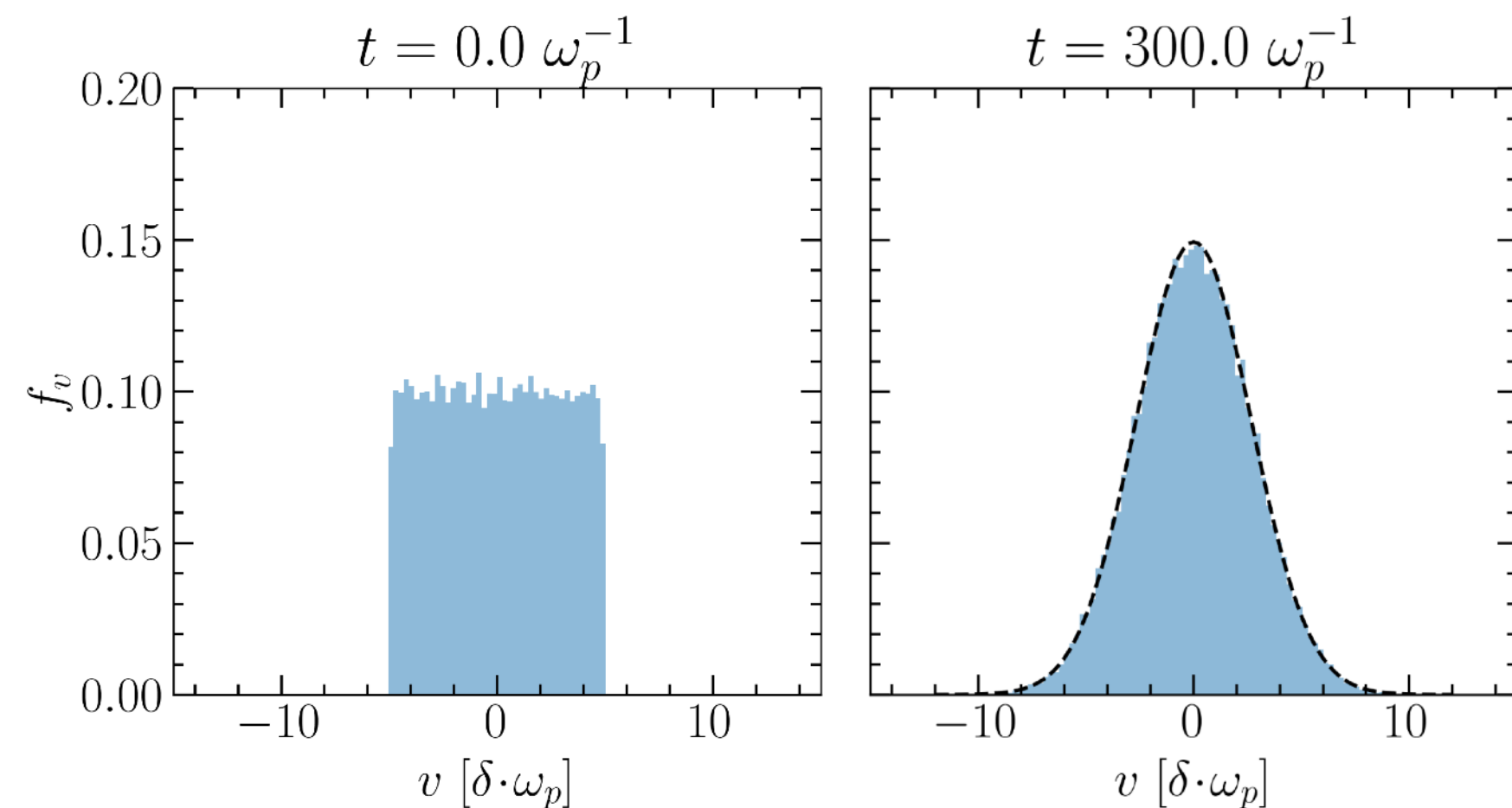
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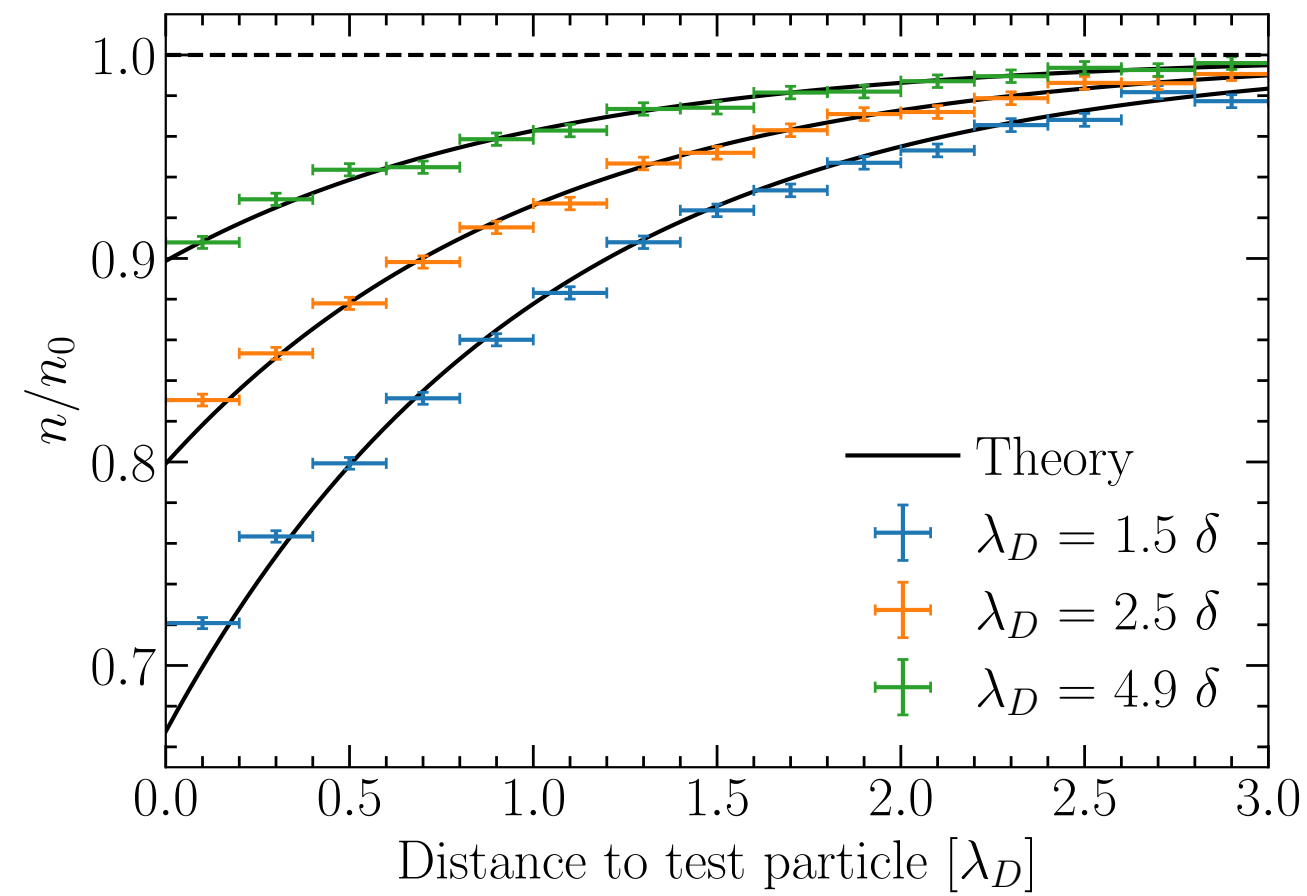


Plasma Thermalization

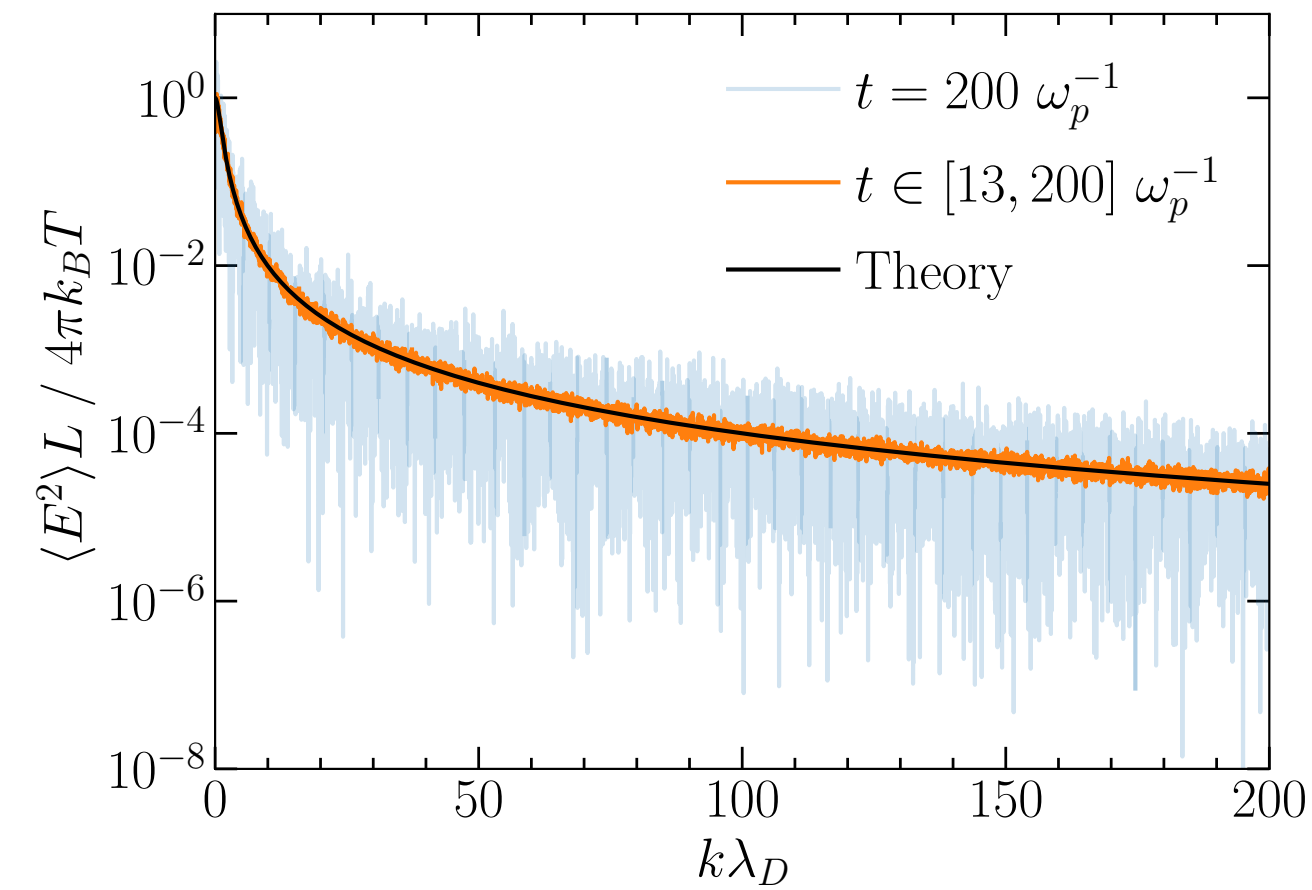


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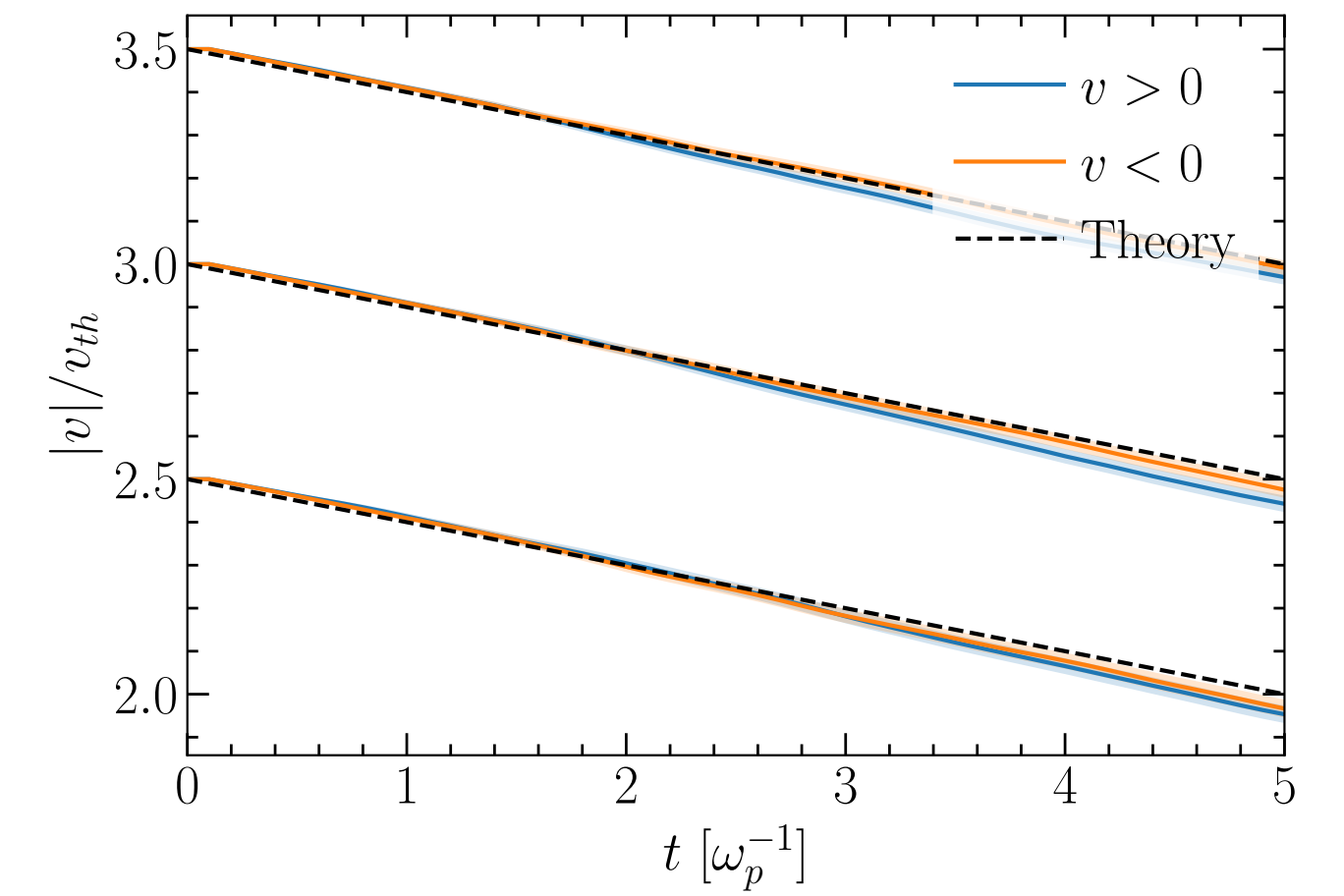
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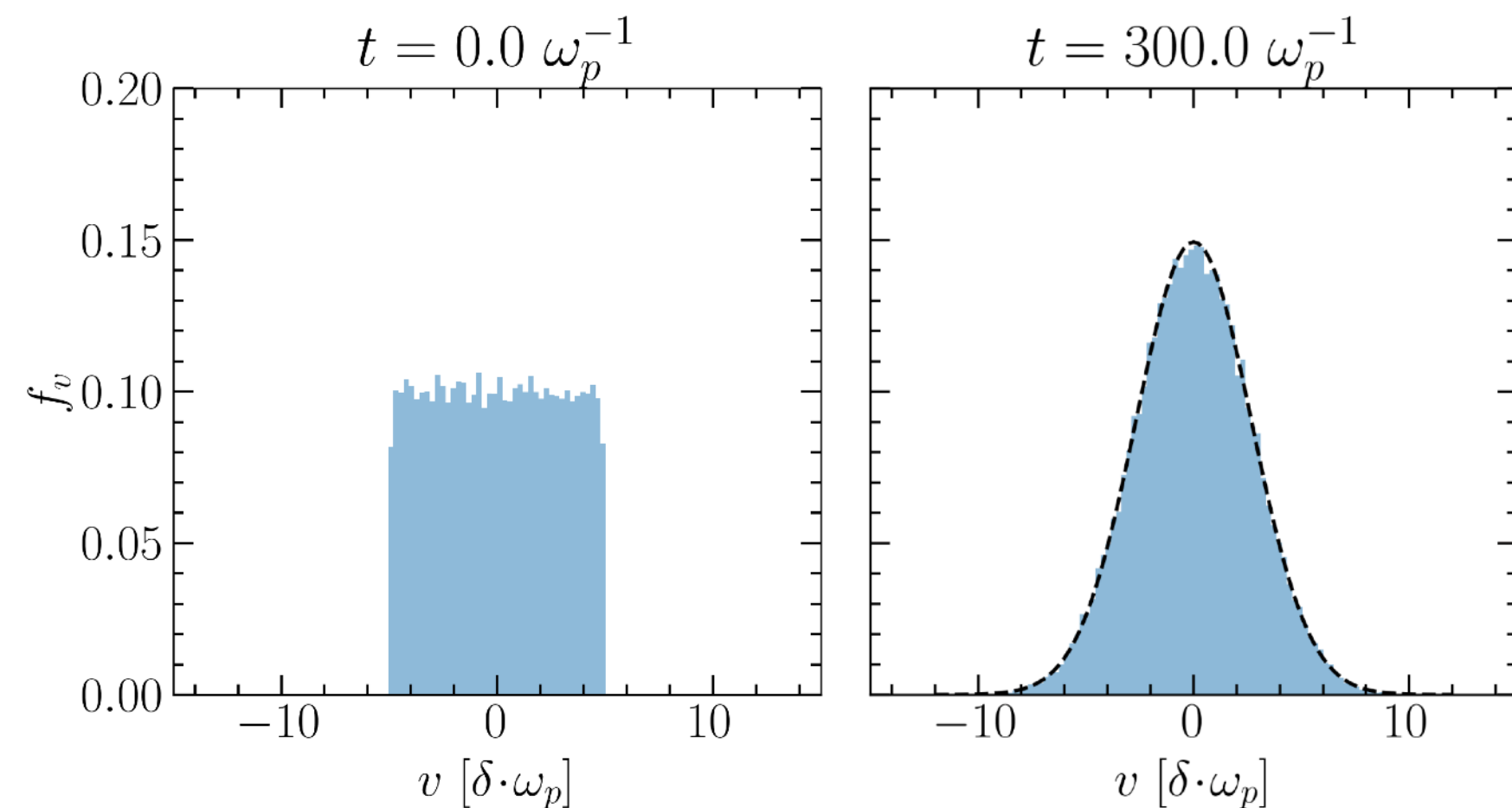
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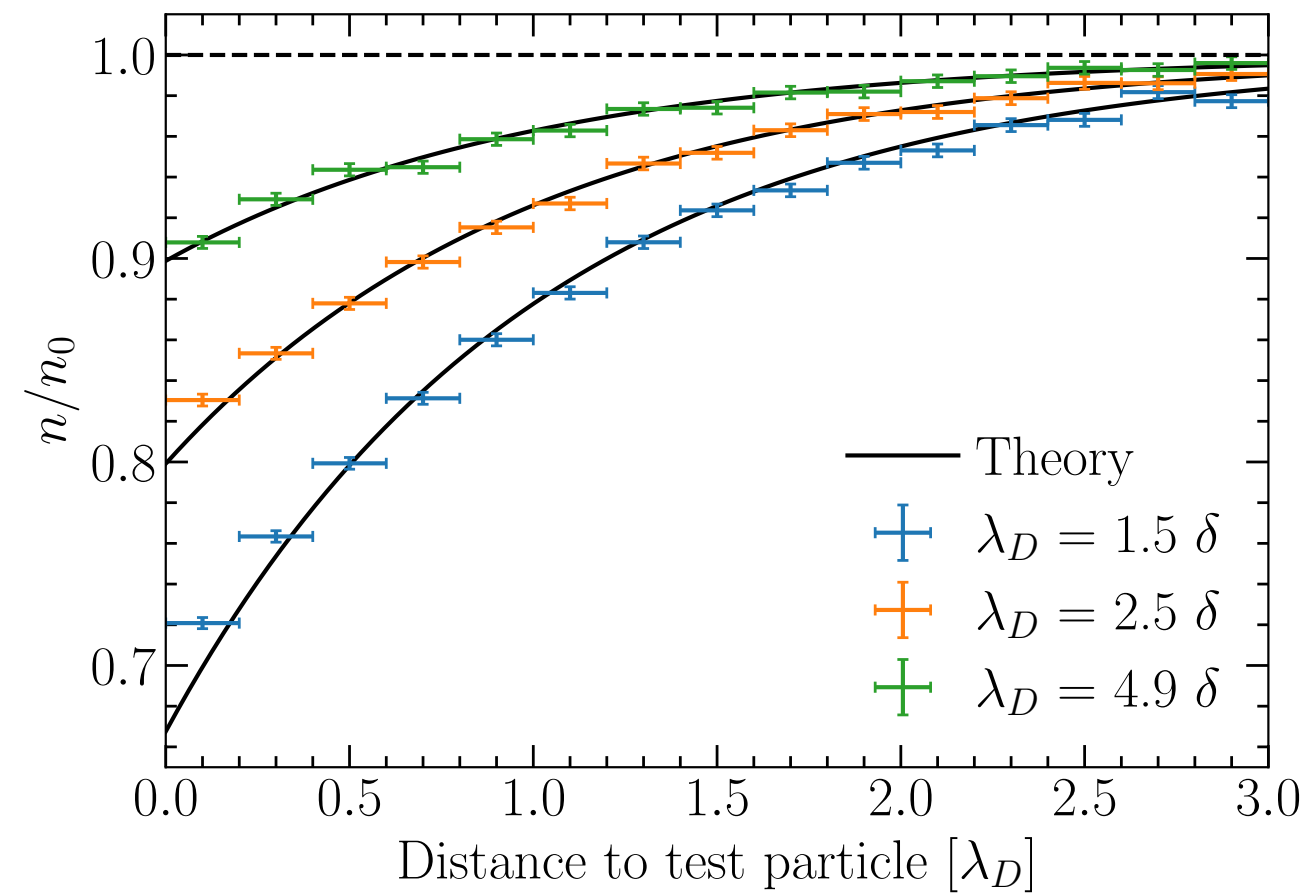


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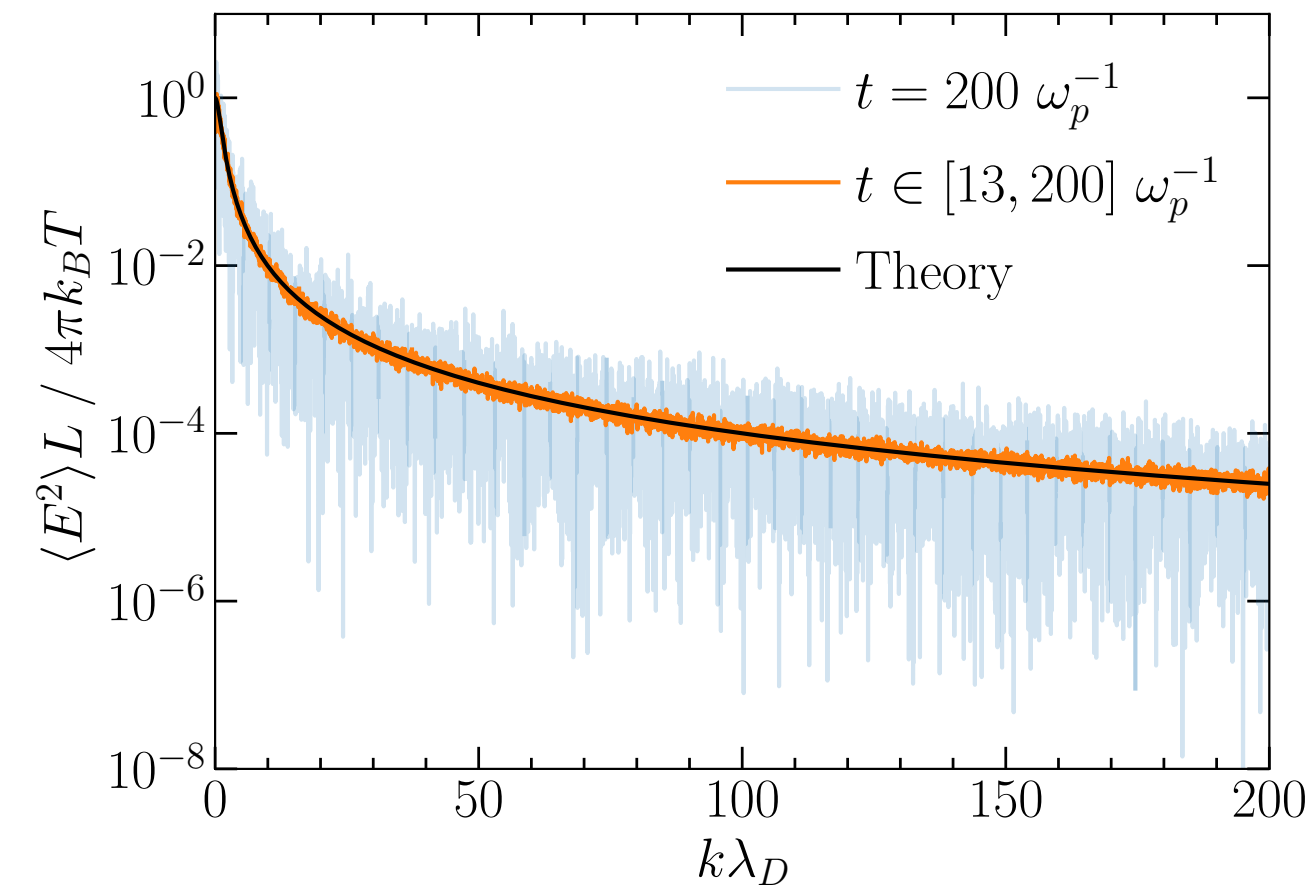


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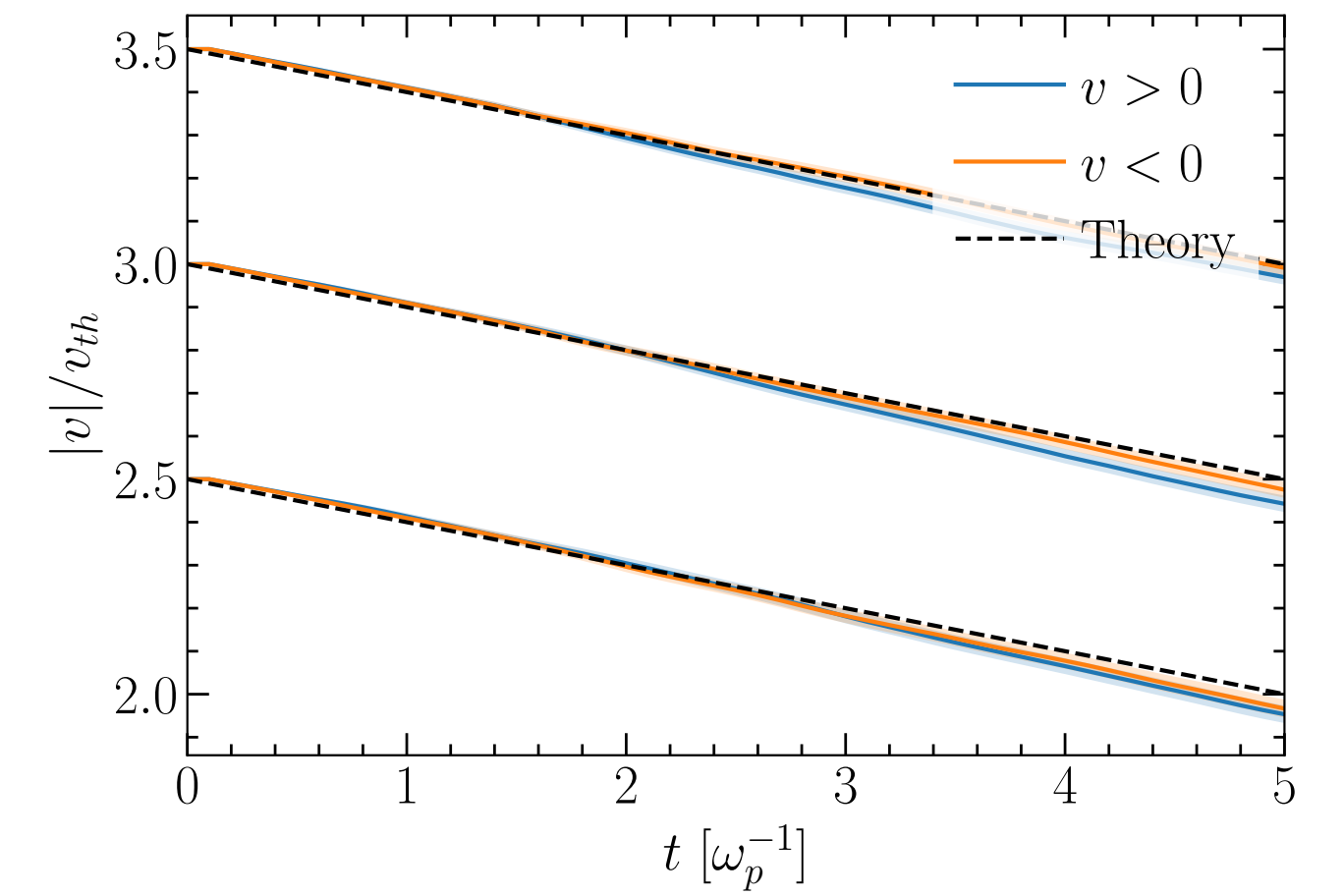
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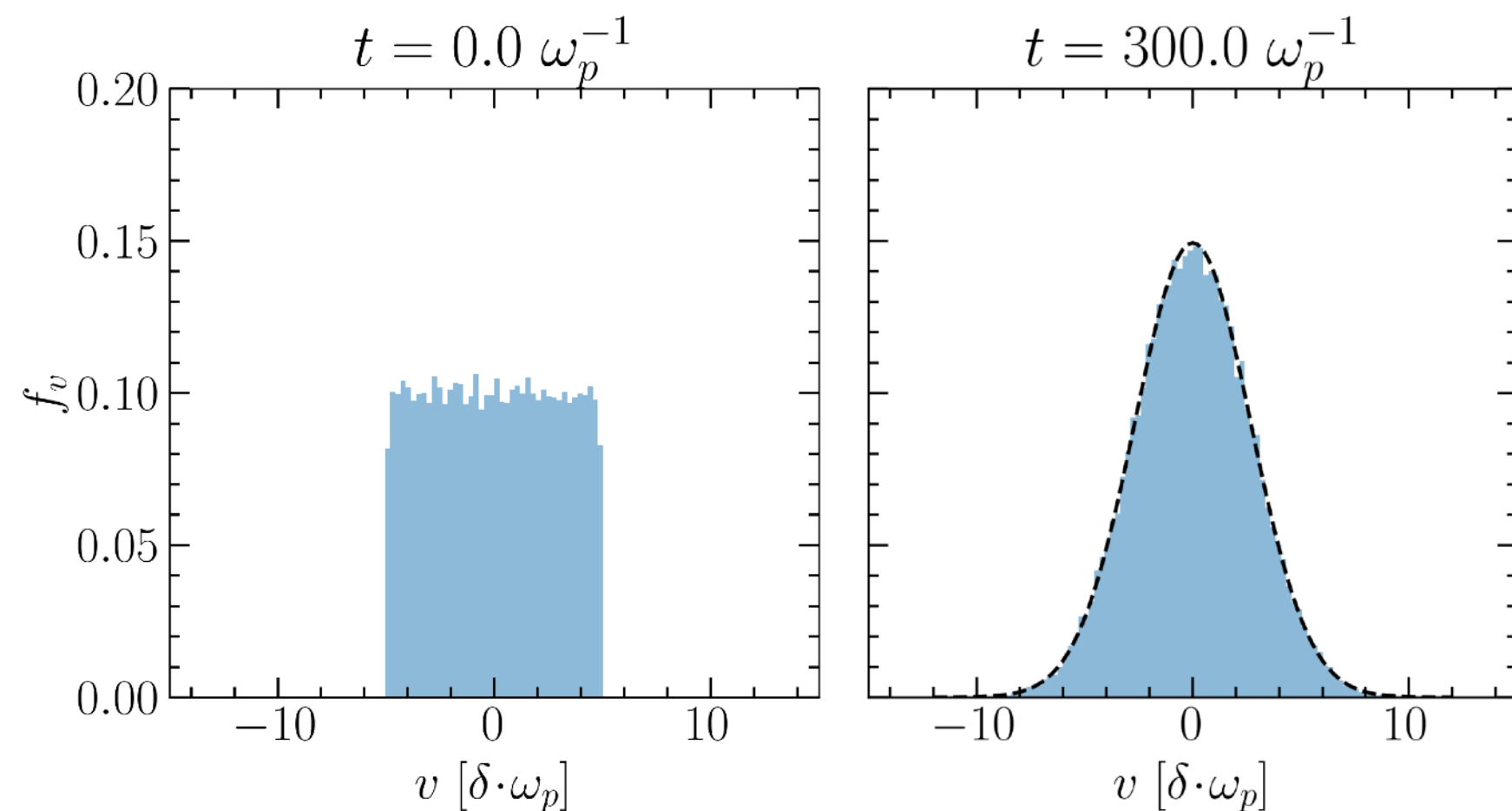
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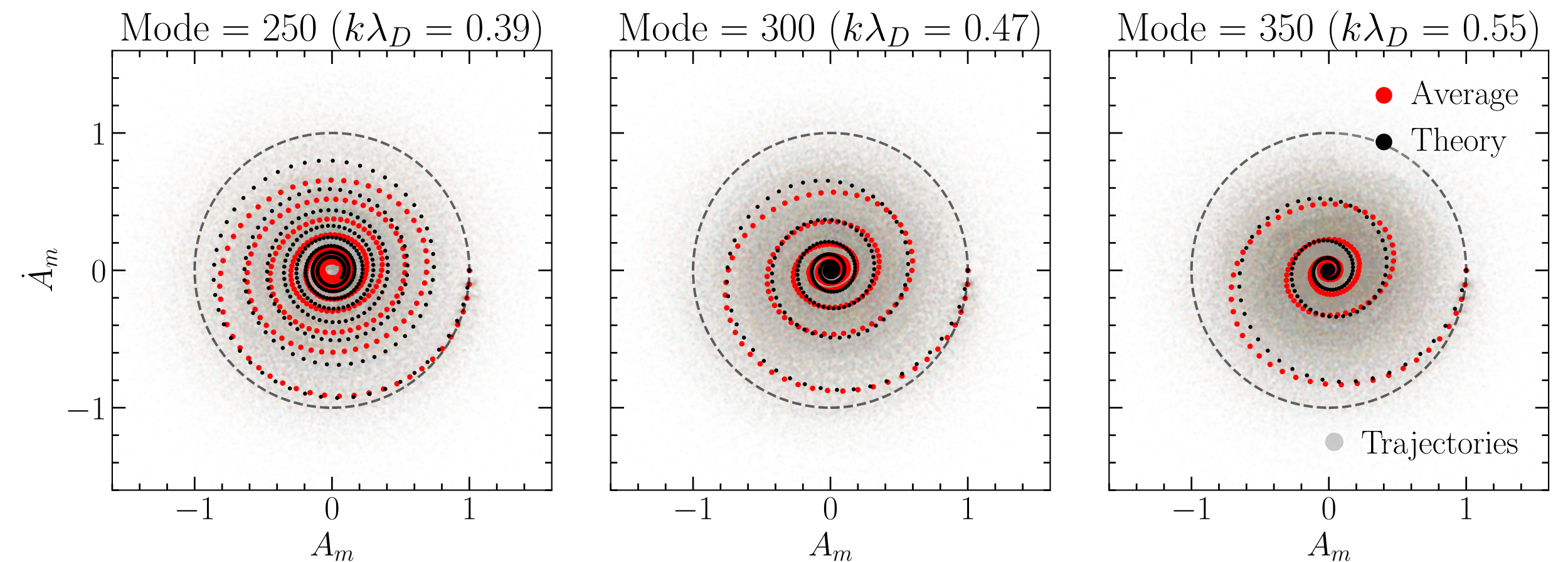
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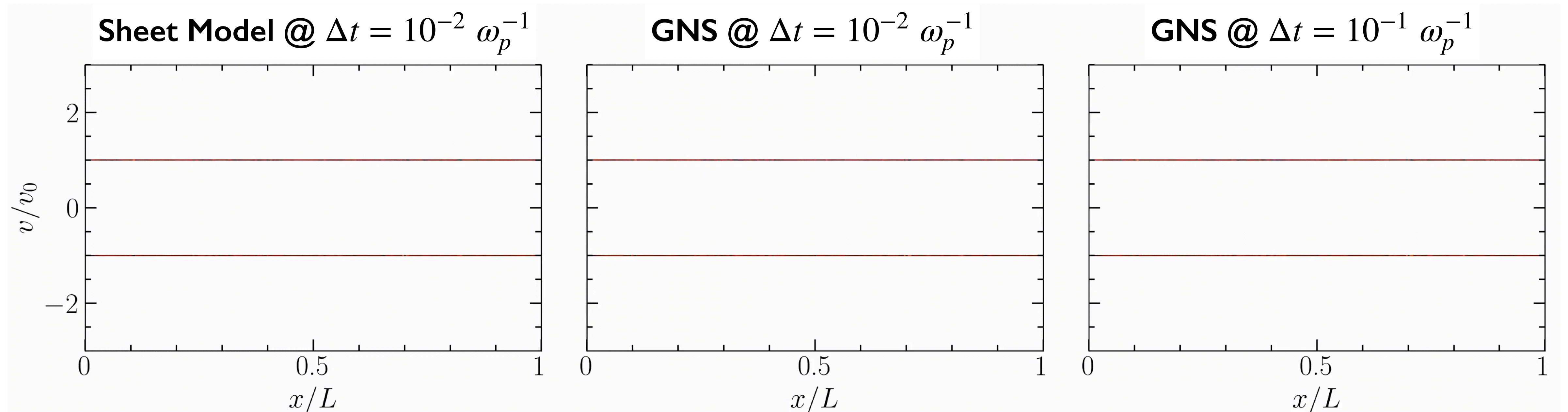


Landau Damping



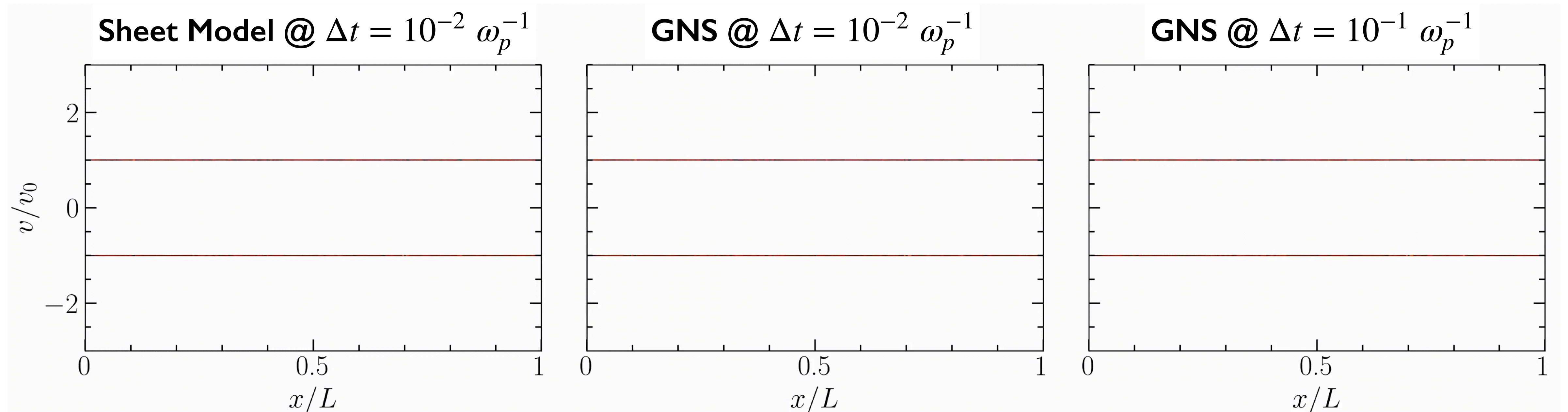
GNS recovers the Two Stream Instability

Parameters: $N_{sheets} = 10,000$ (vs $N_{sheets}^{train} = 10$) $v_0 \approx 500 \delta \cdot \omega_p$ (vs $v_{max}^{train} = 20 \delta \cdot \omega_p$)



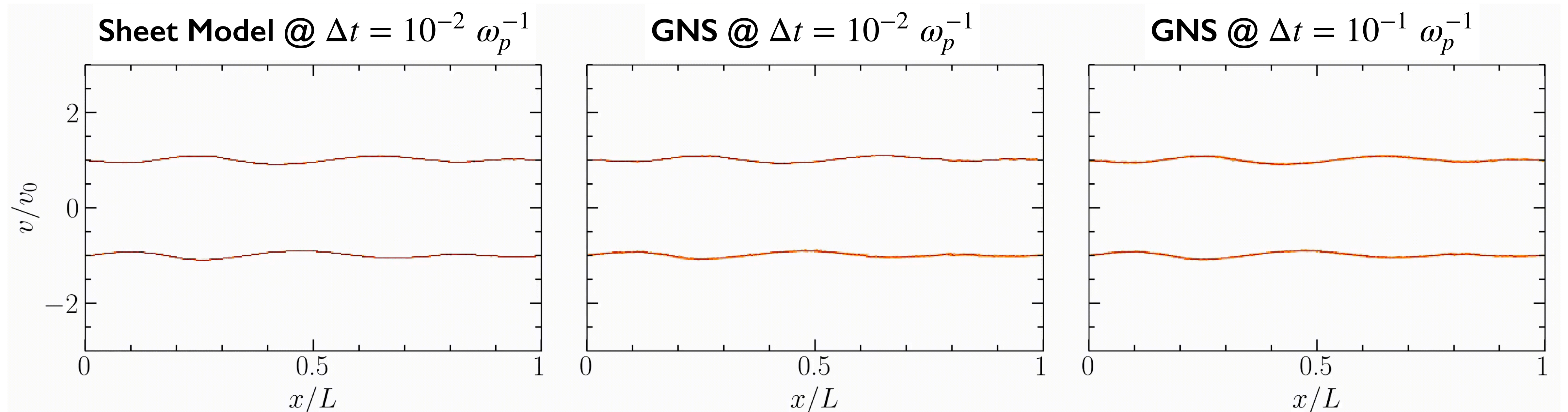
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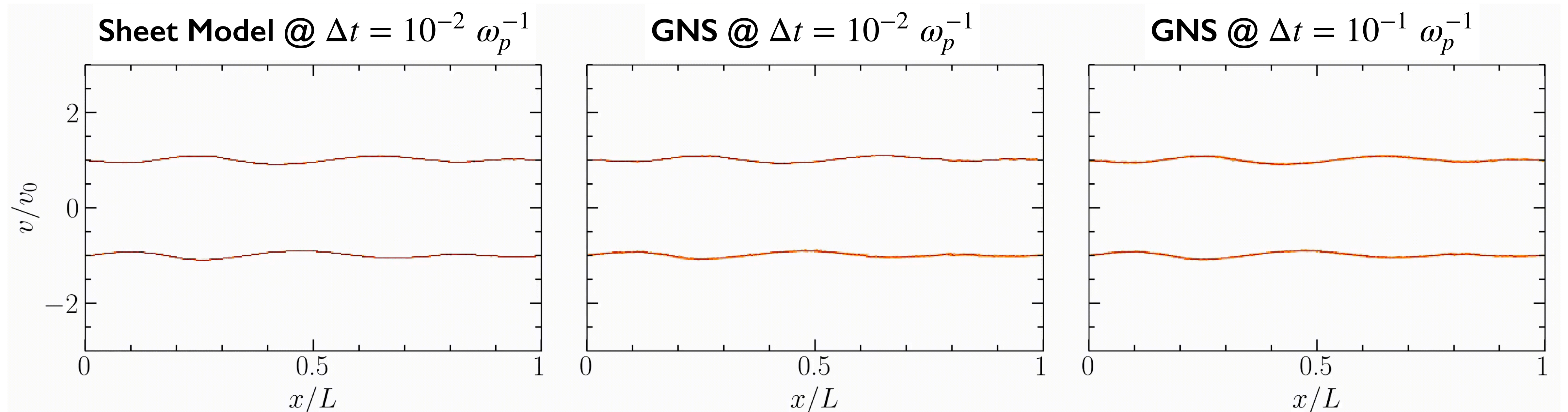
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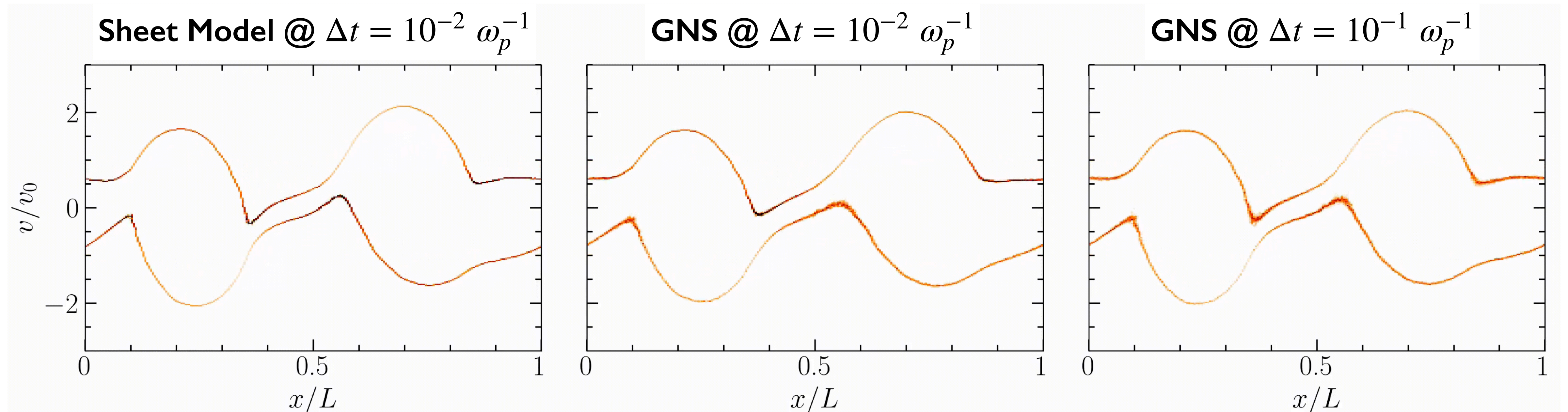
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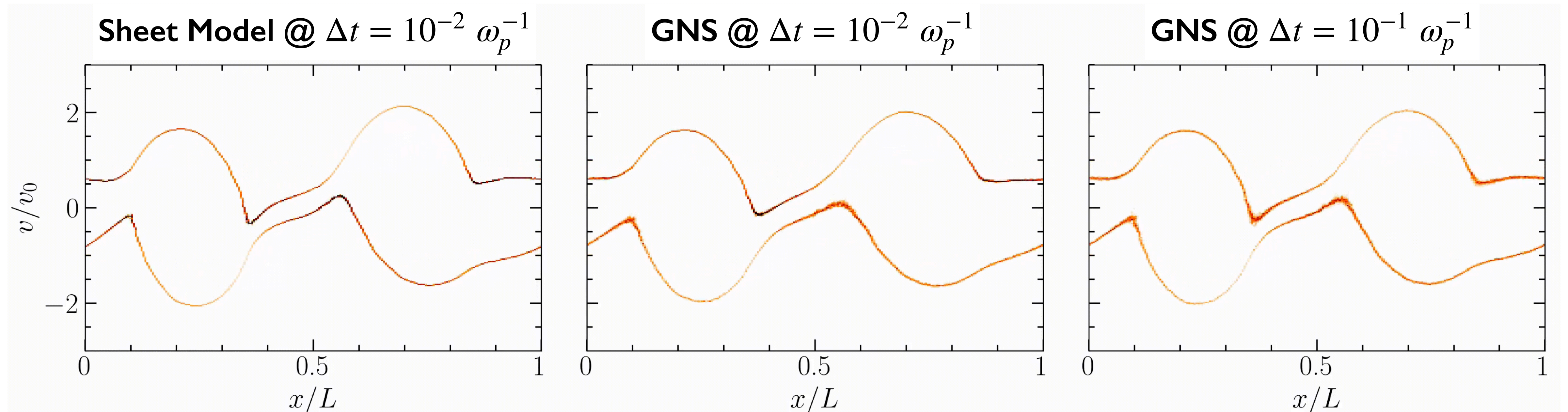
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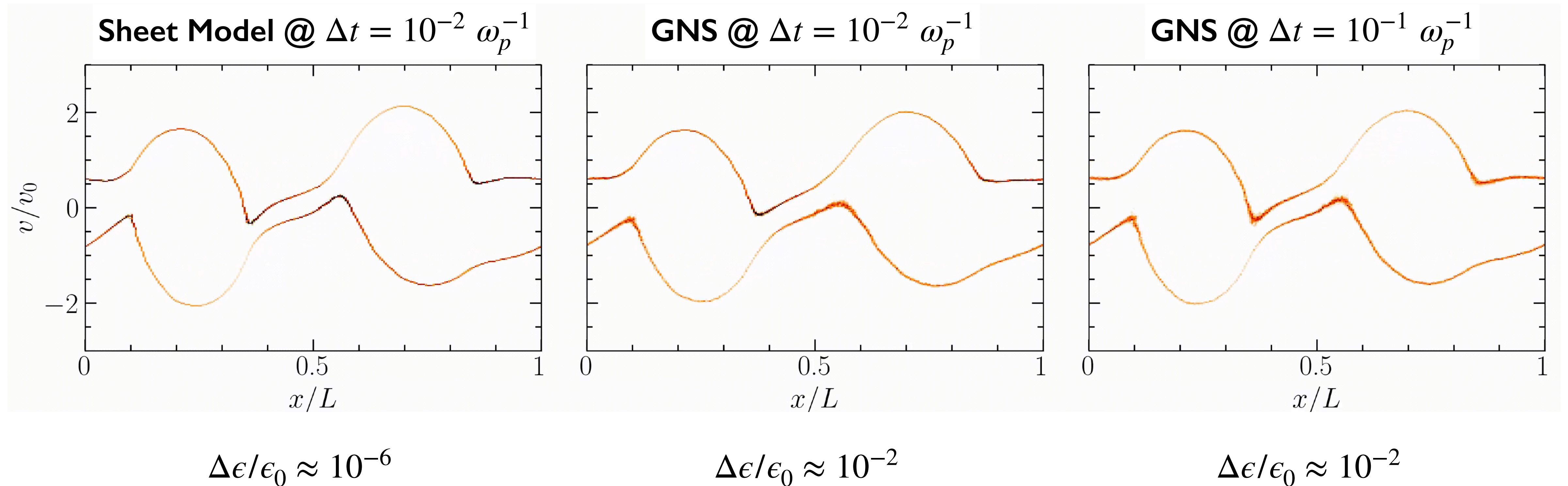
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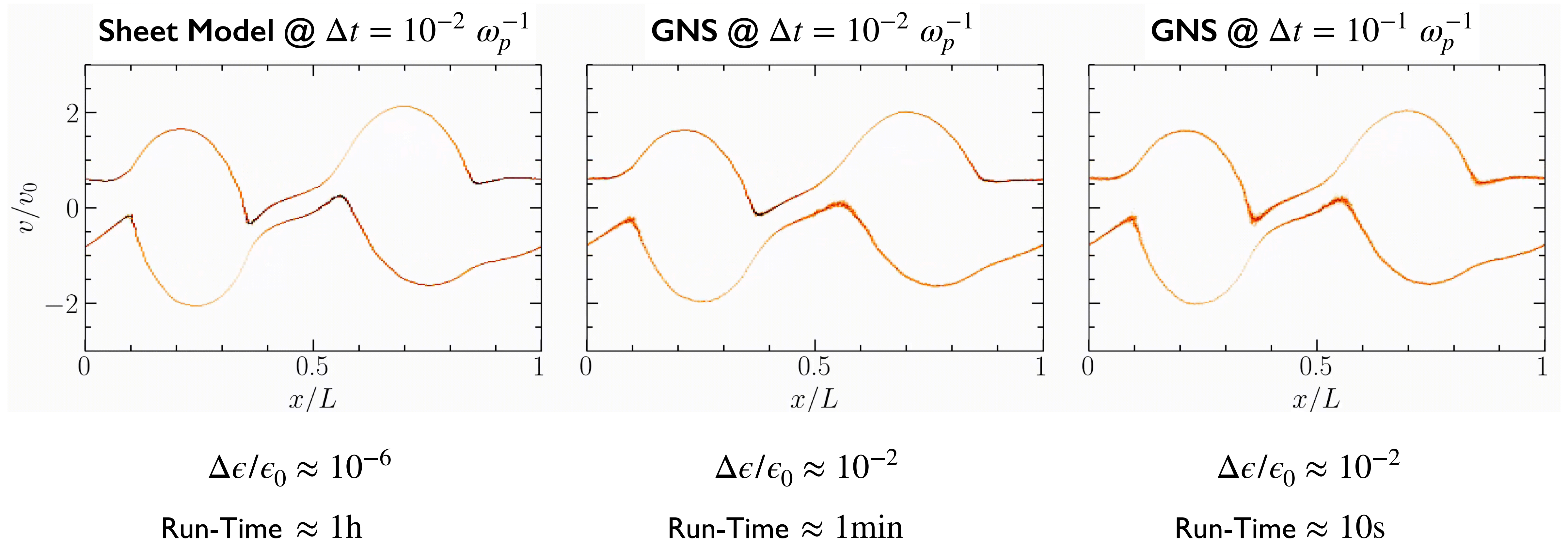
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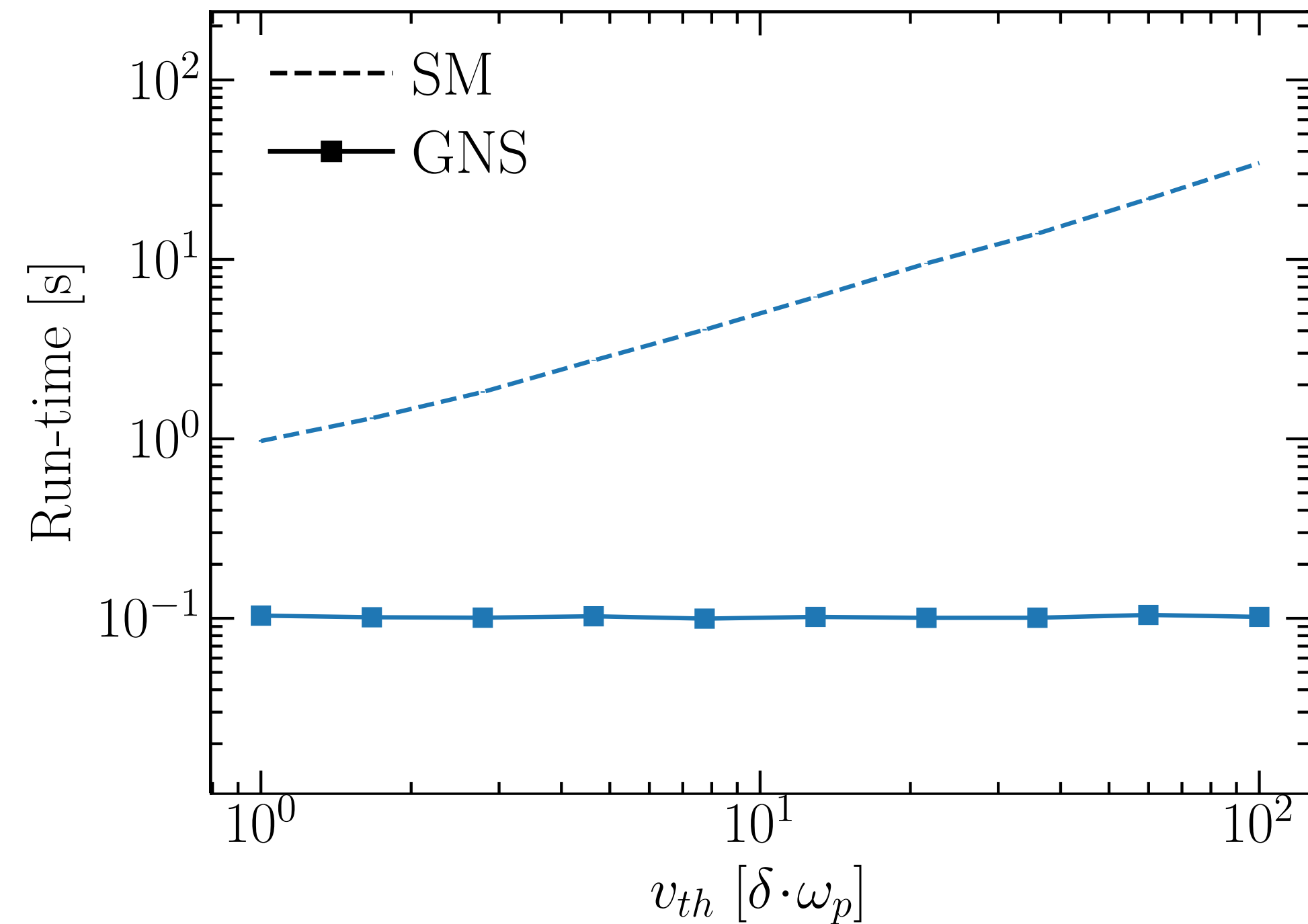
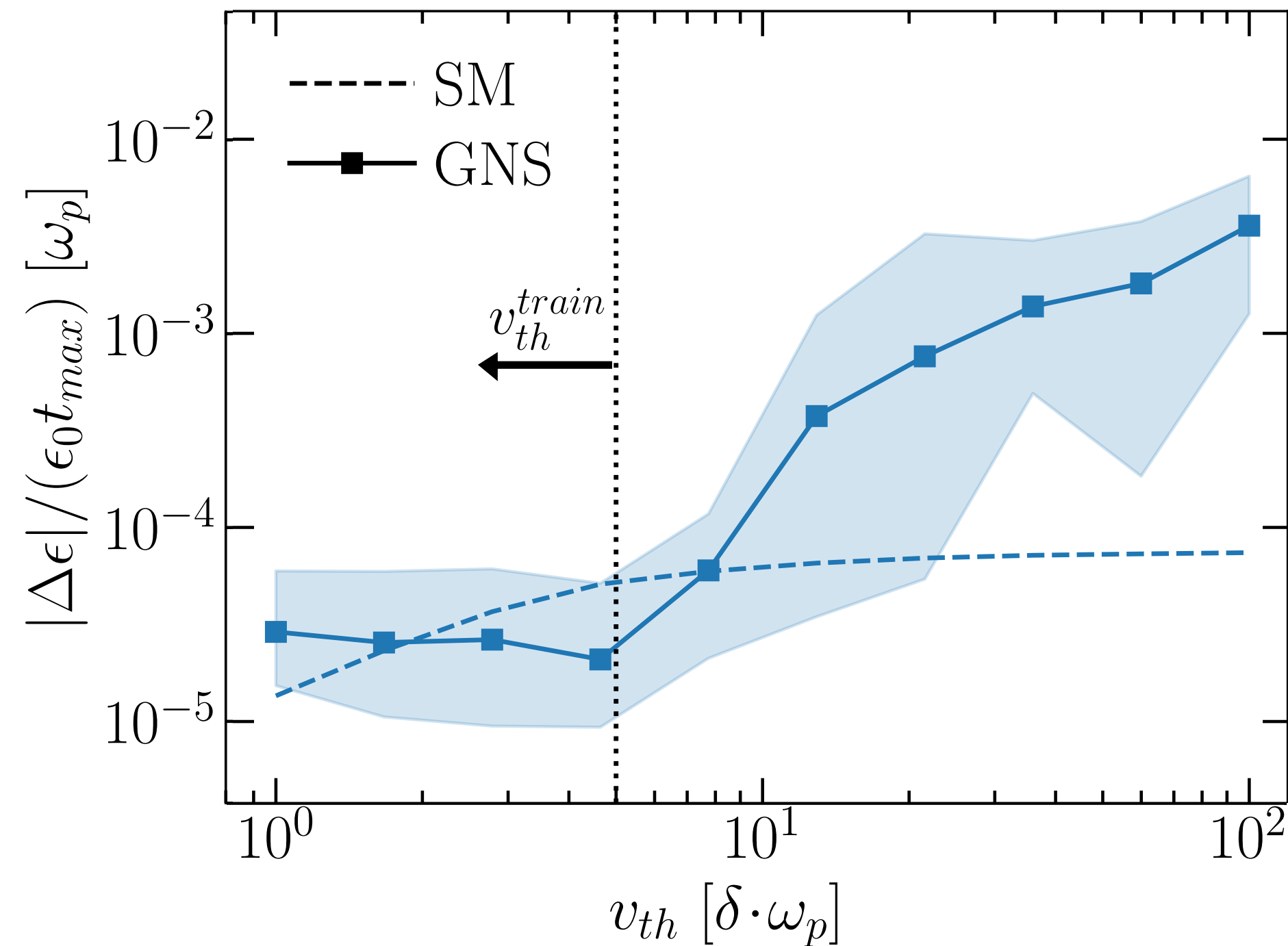
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GNS conserves energy similarly to Sheet Model while being significantly faster*

Parameters: $N_{sheets} = 1000$, velocities sampled from thermal distribution, $\Delta t_{GNS} = 10^{-1} \omega_p^{-1}$



***Note: GNS is implemented in JAX** (GPU), Sheet Model is implemented in NumPy (CPU)

MC models in PIC simulations

New simulator models - GNN collisional plasma model

Learning advection and diffusion coefficients

The (ground) truth? - collisions in PIC codes

What is the ground truth in PIC simulations?

Klimontovich + Maxwell's equations

$$\frac{\partial N}{\partial t} + \mathbf{v} \cdot \nabla_{\mathbf{x}} N - \frac{q}{m} (\mathbf{E}^m + \mathbf{v} \times \mathbf{B}^m) \cdot \nabla_{\mathbf{v}} N = 0$$

This is the particle-in-cell algorithm (with finite-size particles):

statistical mechanics is well-known (e.g. H. Okuda and C. Birdsall, (1970), R. Hockney (1971), M. Touati et al. (2022), S. Jubin et al., (2024))

Born-Infeld electrodynamics

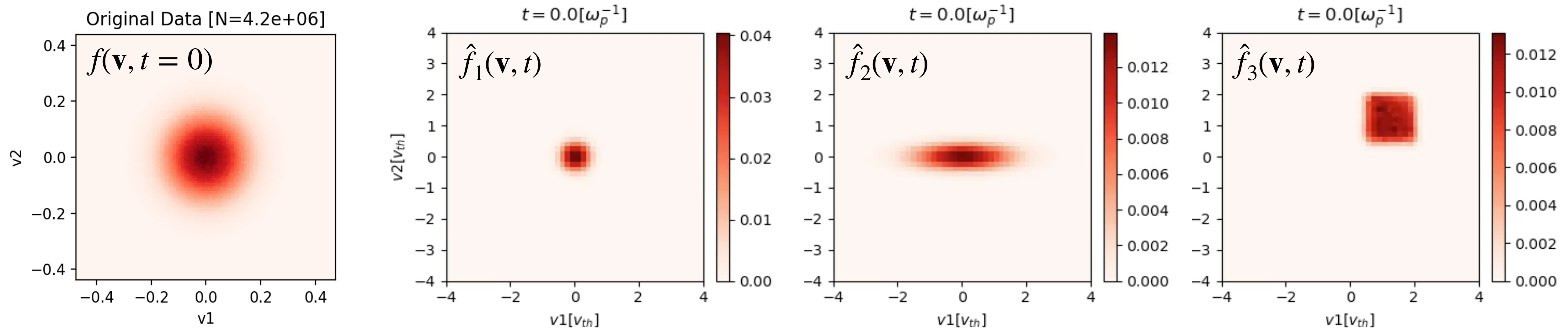
Numerical collision operator has been derived in previous works: Can this be learned from the simulation data in the weakly collisional regime?

Can we describe phase-space dynamics using a Fokker-Planck operator?

PIC (Klimontovich)

$$\frac{\partial N}{\partial t} + \mathbf{v} \cdot \nabla_{\mathbf{x}} N - \frac{q}{m} (\mathbf{E}^m + \mathbf{v} \times \mathbf{B}^m) \cdot \nabla_{\mathbf{v}} N = 0 \quad \xrightarrow{\quad ? \quad} \quad \frac{\partial \hat{f}(\mathbf{v}, t)}{\partial t} = - \nabla_{\mathbf{v}} \cdot \left(\mathbf{A} \hat{f} \right) + \frac{1}{2} \nabla_{\mathbf{v}} \nabla_{\mathbf{v}} \cdot \left(\mathbf{\hat{D}} \hat{f} \right)$$

What if we want (Fokker-Planck)?



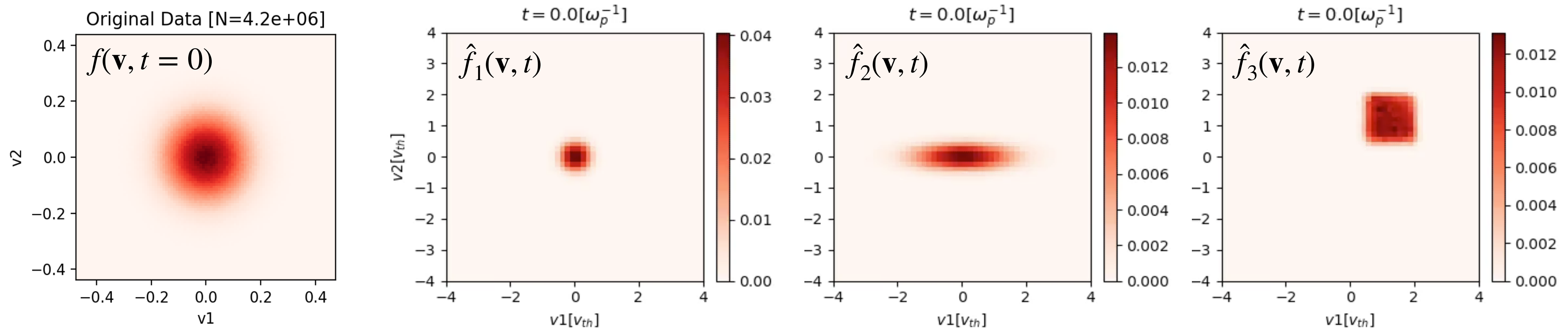
Thermal Plasma

Can we describe phase-space dynamics using a Fokker-Planck operator?

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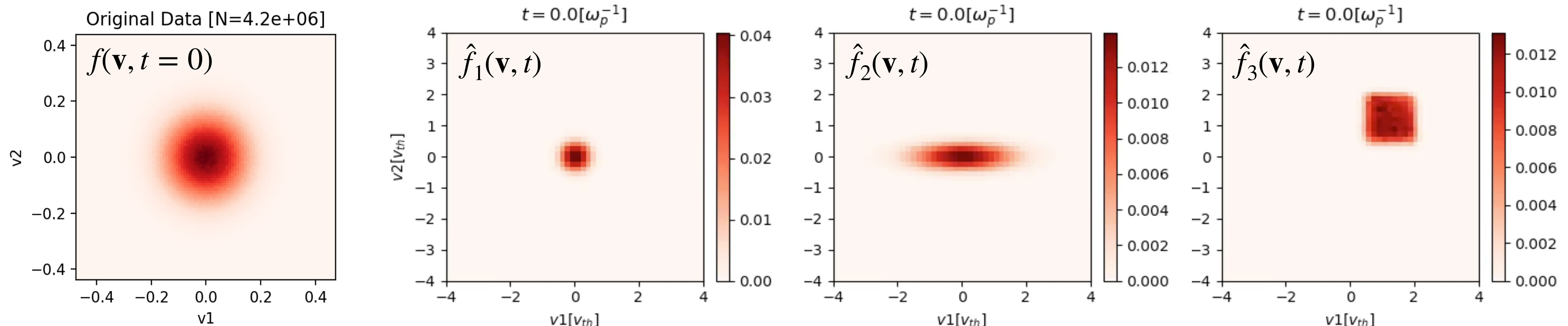
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What if we want (Fokker-Planck)?



Thermal Plasma

How do we estimate \mathbf{A} (advection) and $\mathbf{\overleftrightarrow{D}}$ (diffusion)?

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Option 1: From raw particle data

The “correct” approach if possible

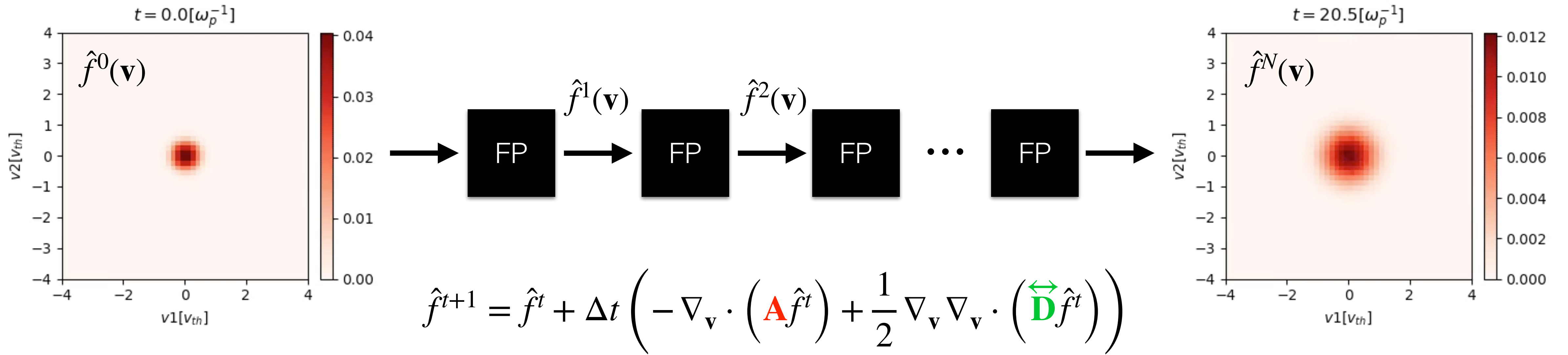
Not feasible for larger systems (memory-wise) unless it is done at run-time

Option 2: From the phase-space evolution of sub-populations

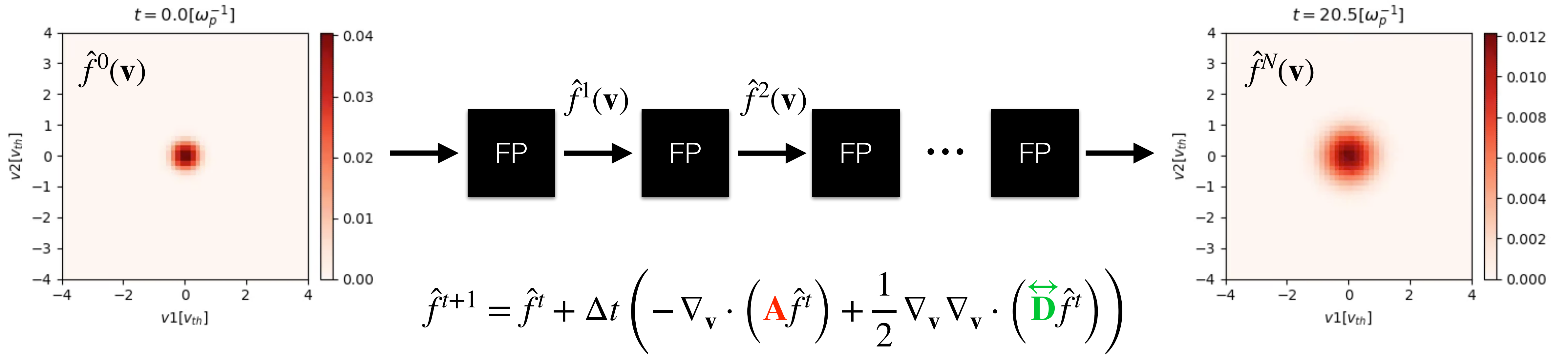
Can be done in post-processing with a differentiable solver

Ill-posed problem: non-unique solution for coefficients

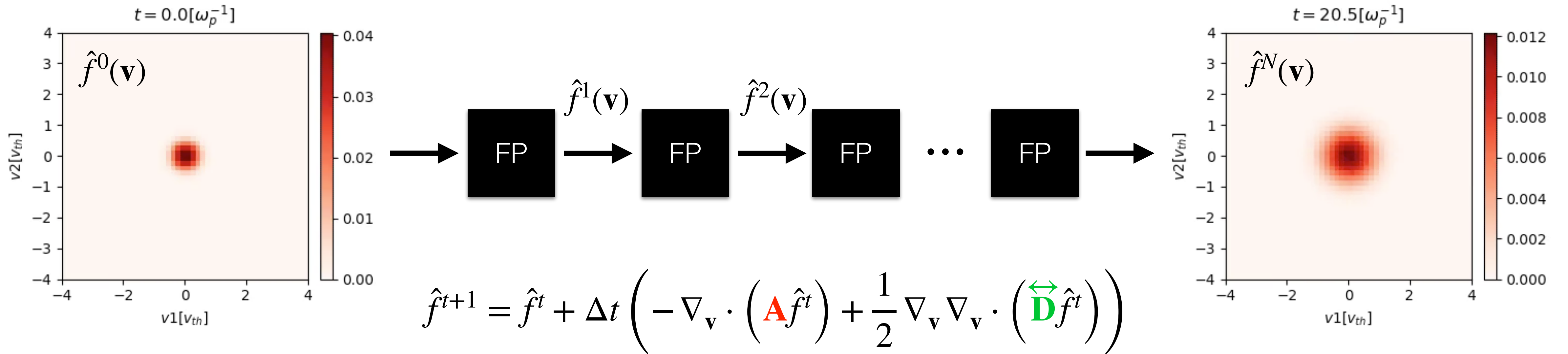
Learning advection / diffusion from evolution of sub-populations



Learning advection / diffusion from evolution of sub-populations



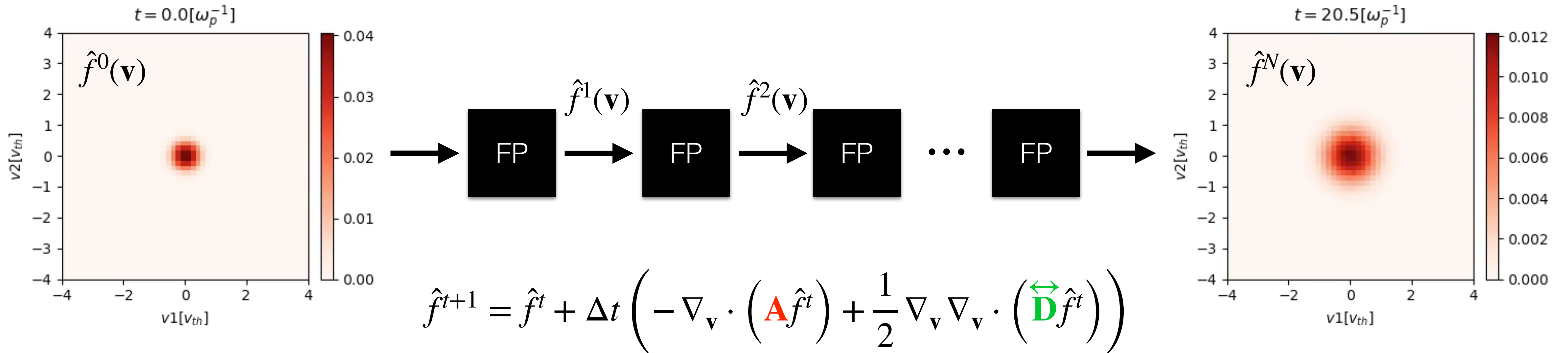
Learning advection / diffusion from evolution of sub-populations



We can make the **Fokker-Planck solver differentiable** and frame this as an **optimisation task**

$$\min_{\mathbf{A}, \mathbf{\tilde{D}}} \left\| \hat{f}_{predicted}^N - \hat{f}_{true}^N \right\|$$

Learning advection / diffusion from evolution of sub-populations



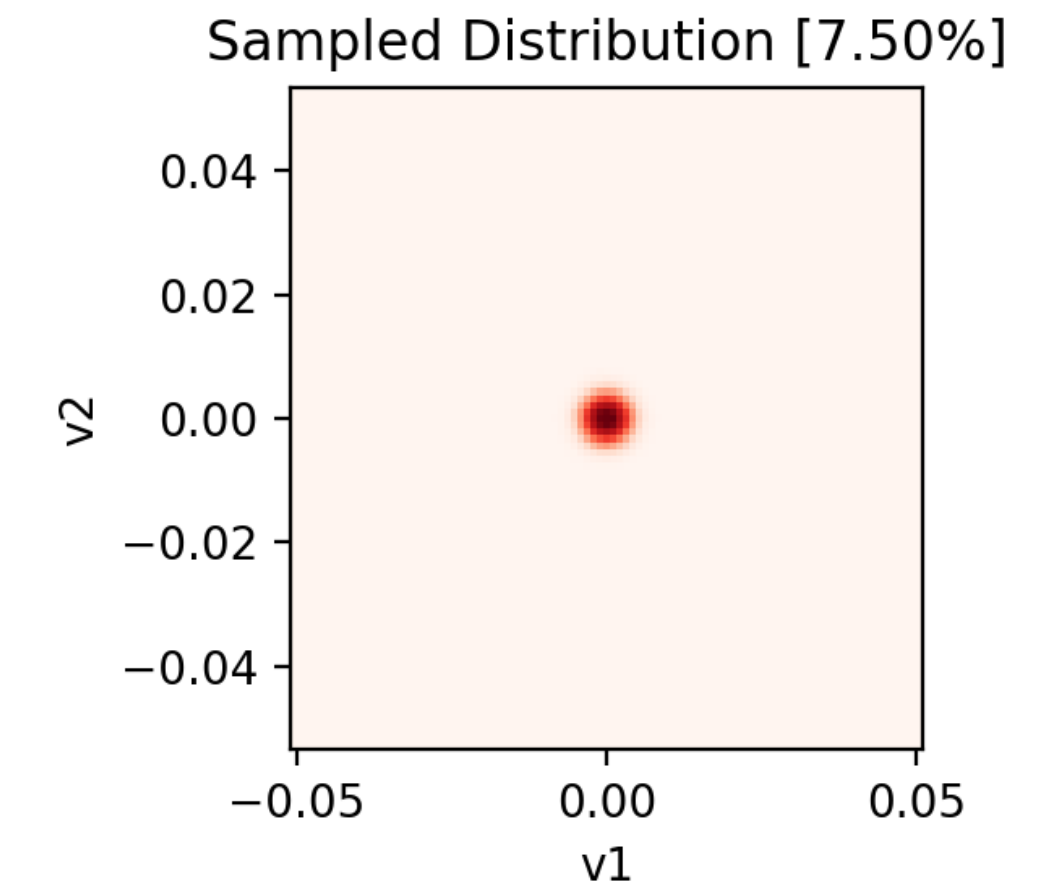
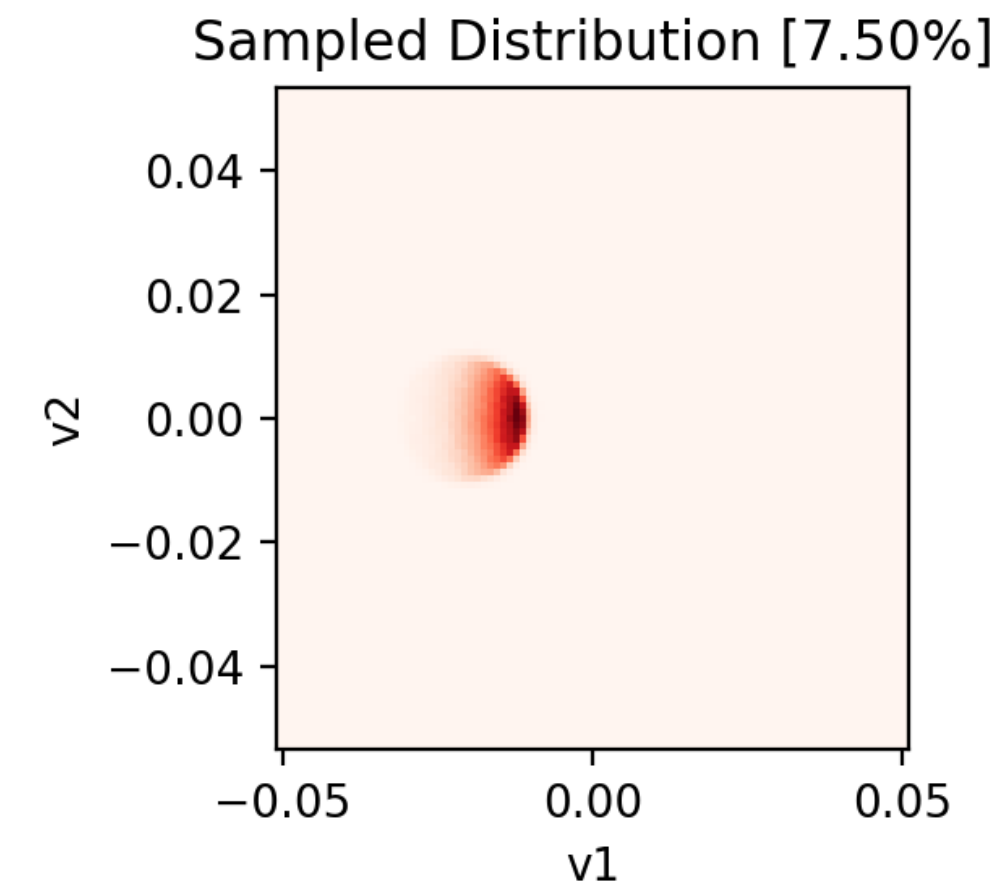
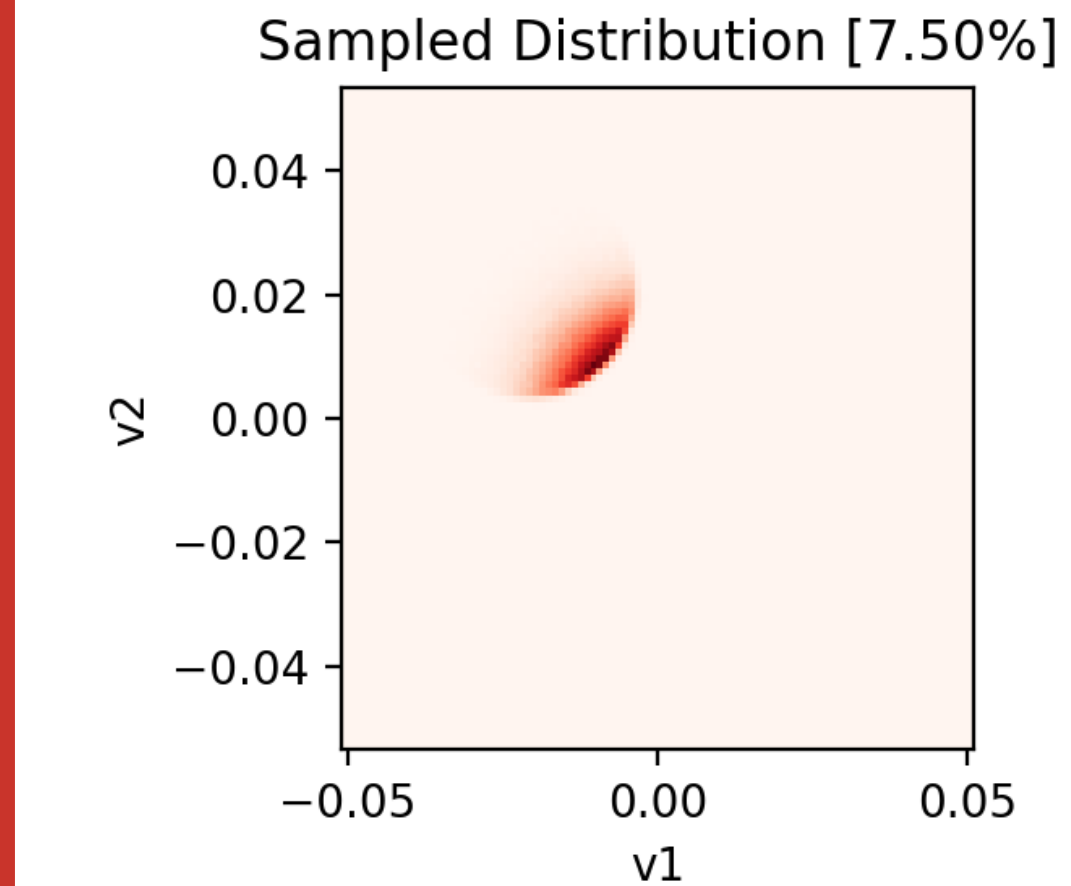
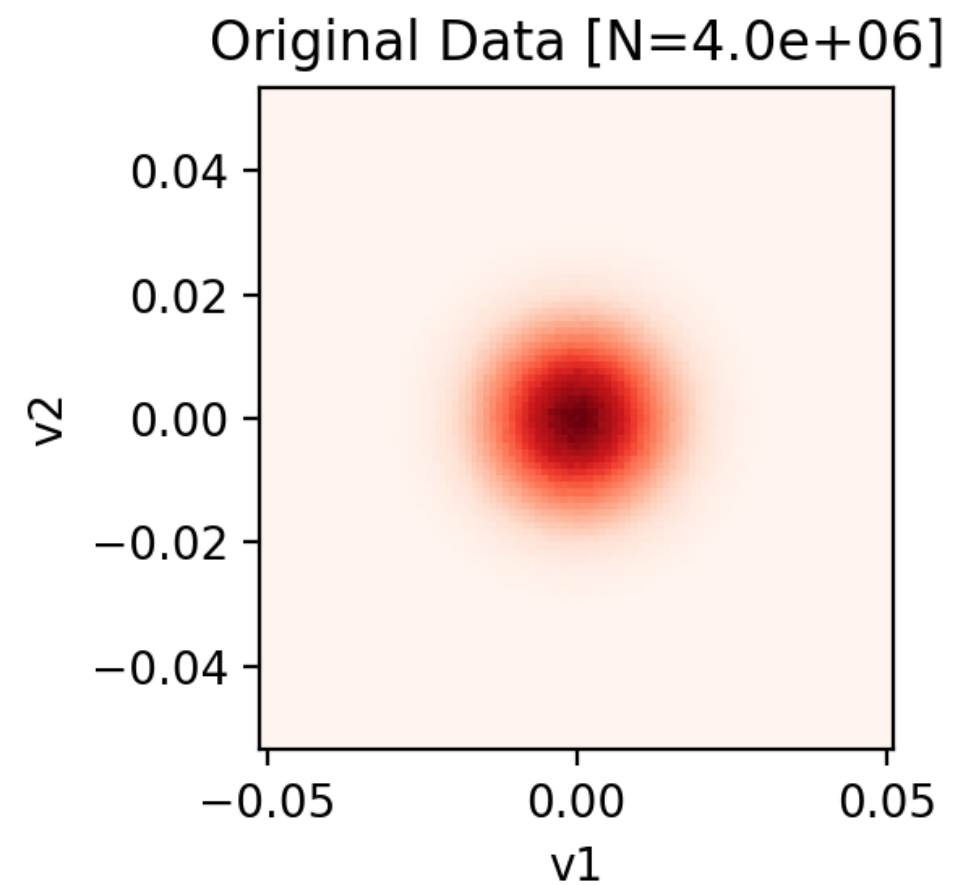
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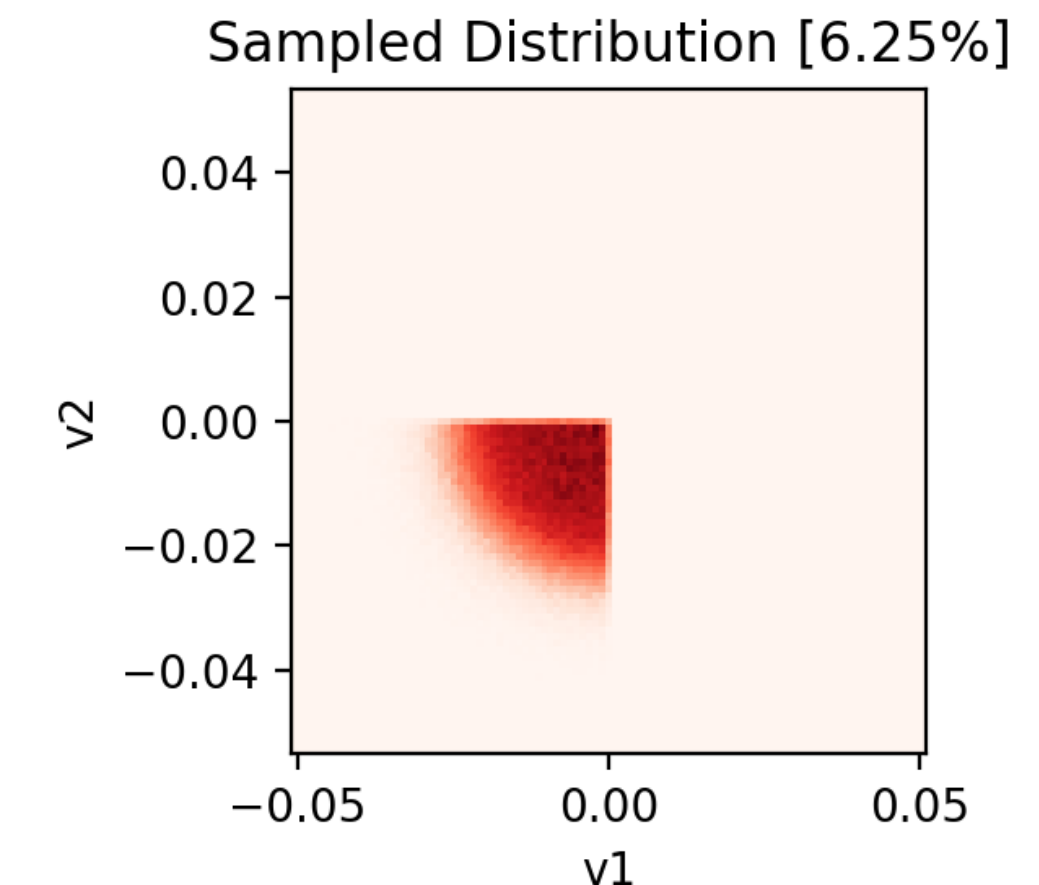
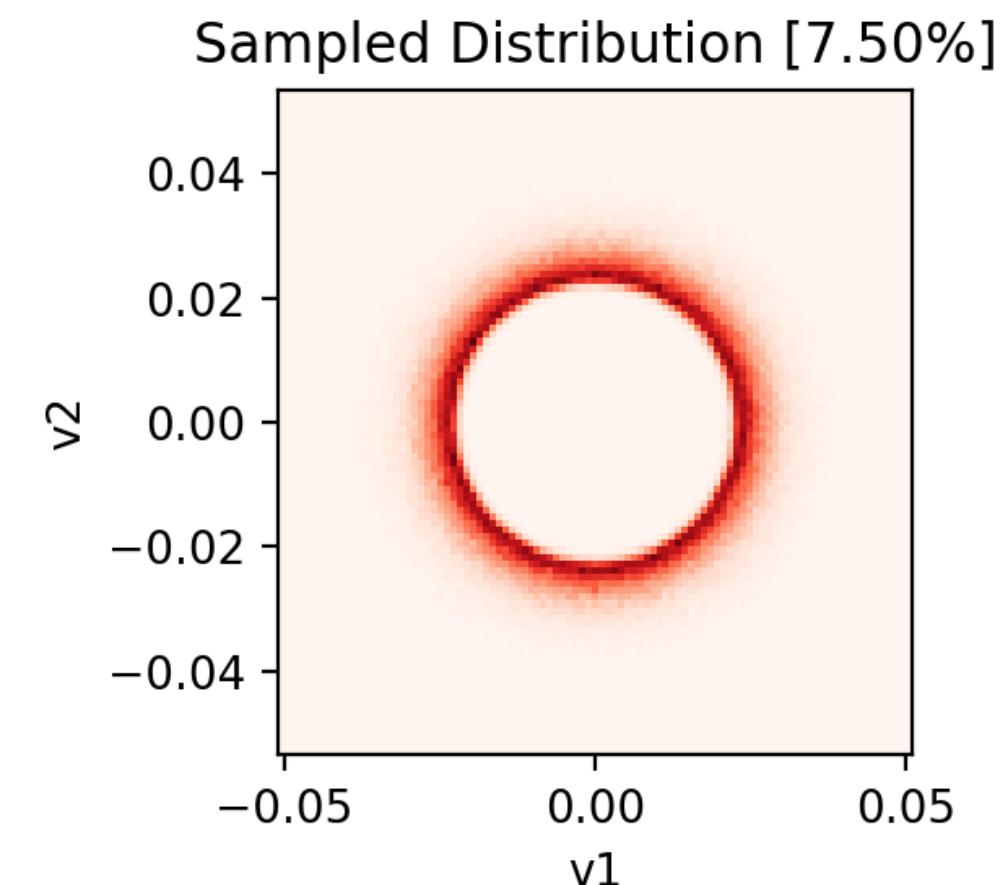
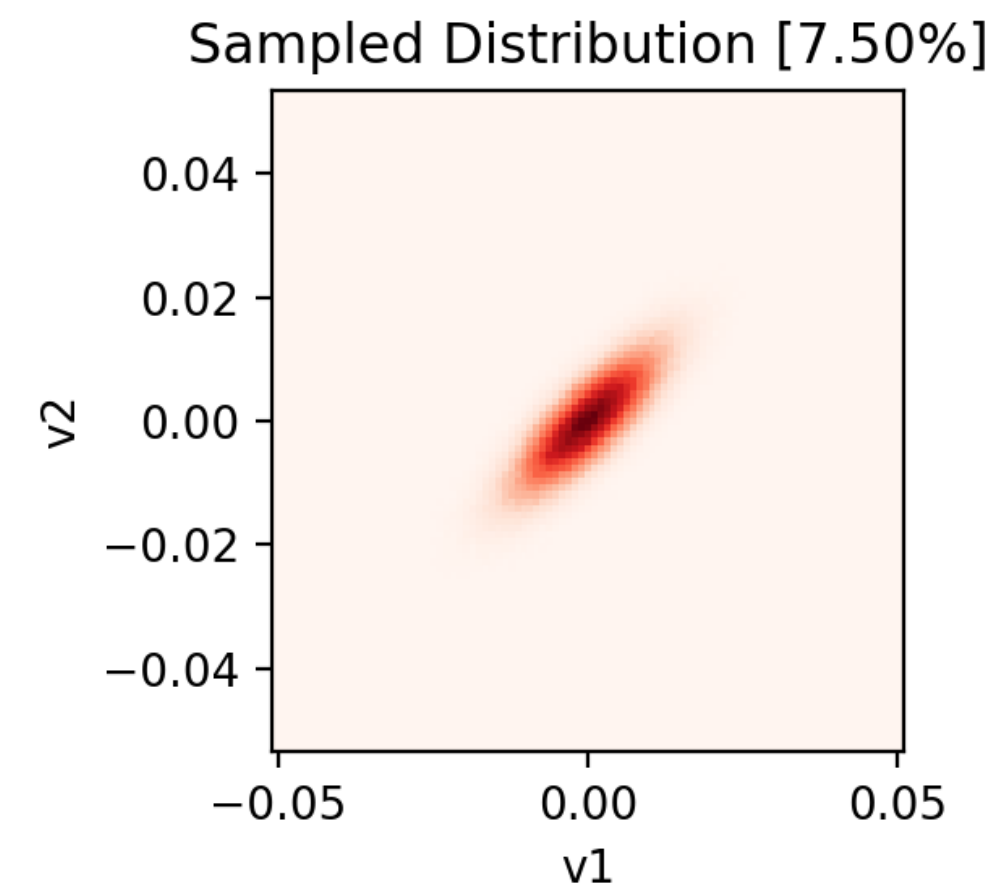
This is an ill-posed problem (there exists a family of solutions) \longrightarrow Train with multiple sub-populations

We use different sets of training and test sub-populations

Train (9x)

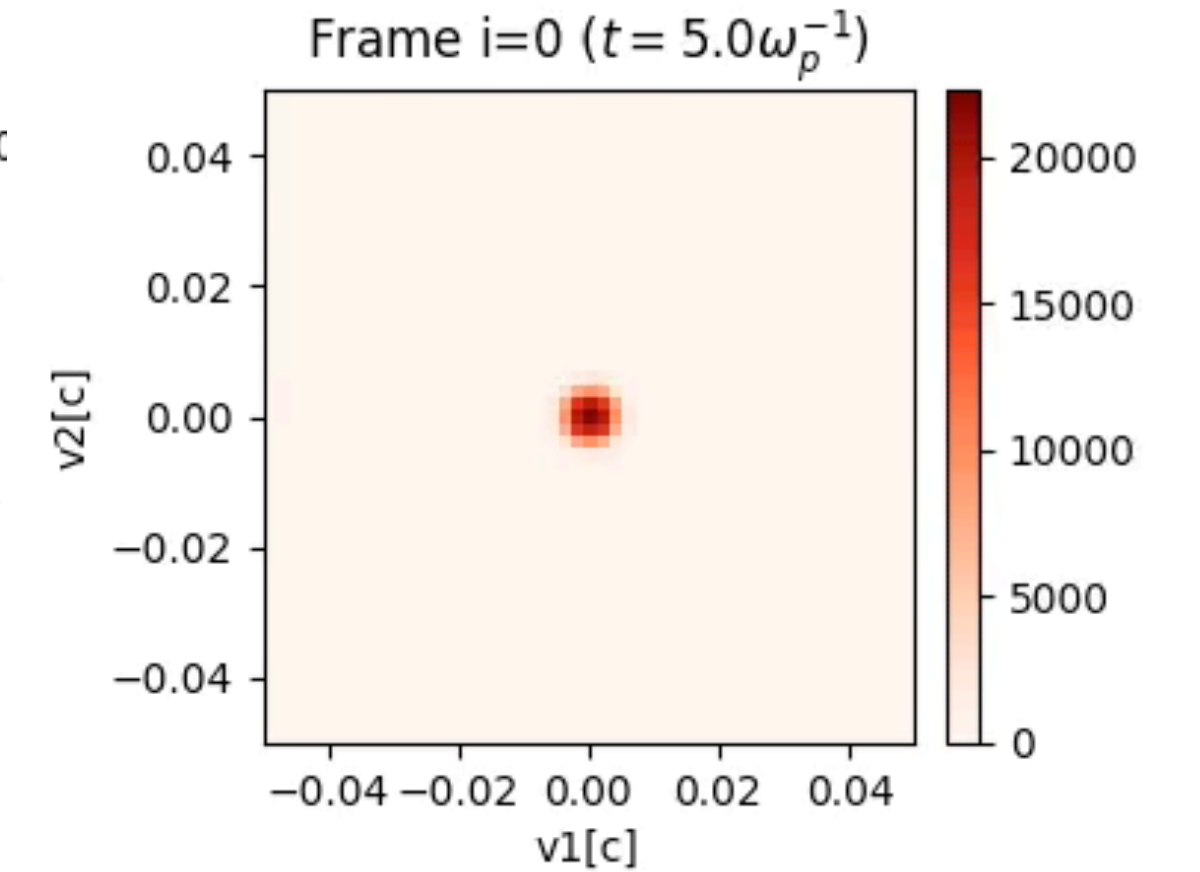
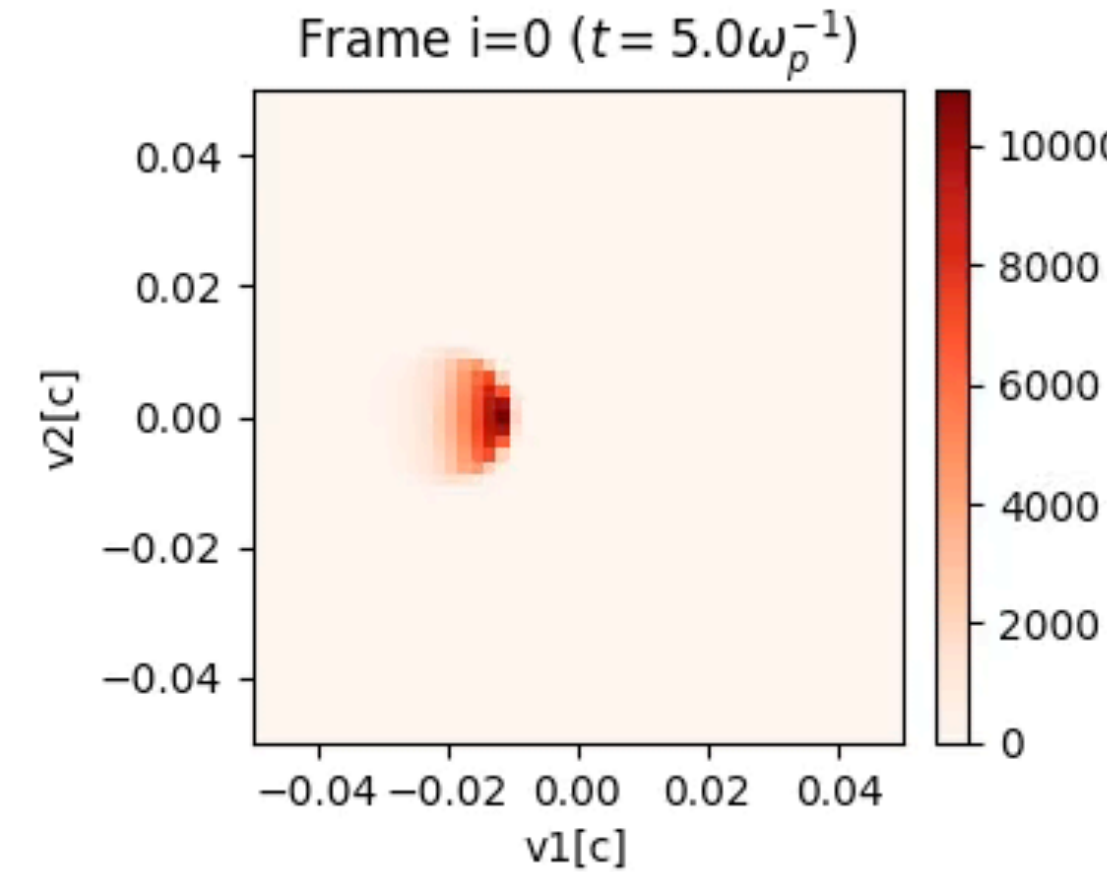
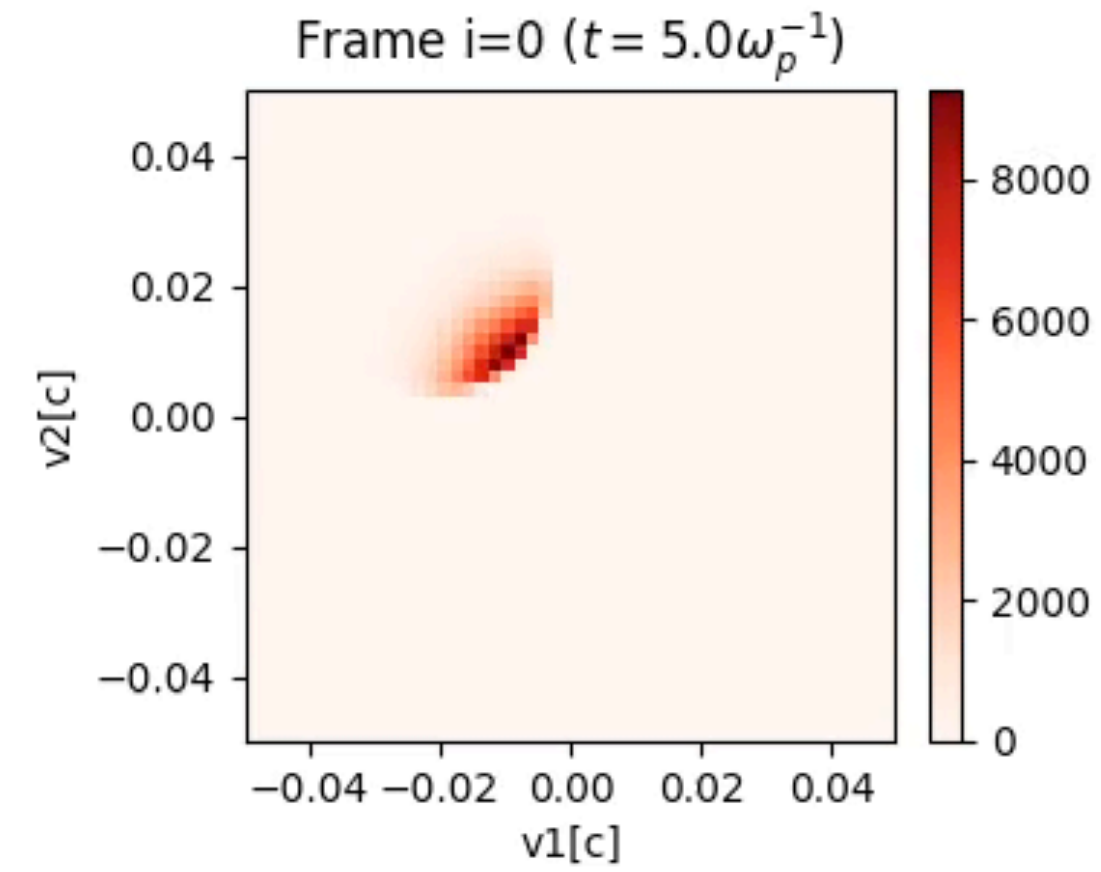
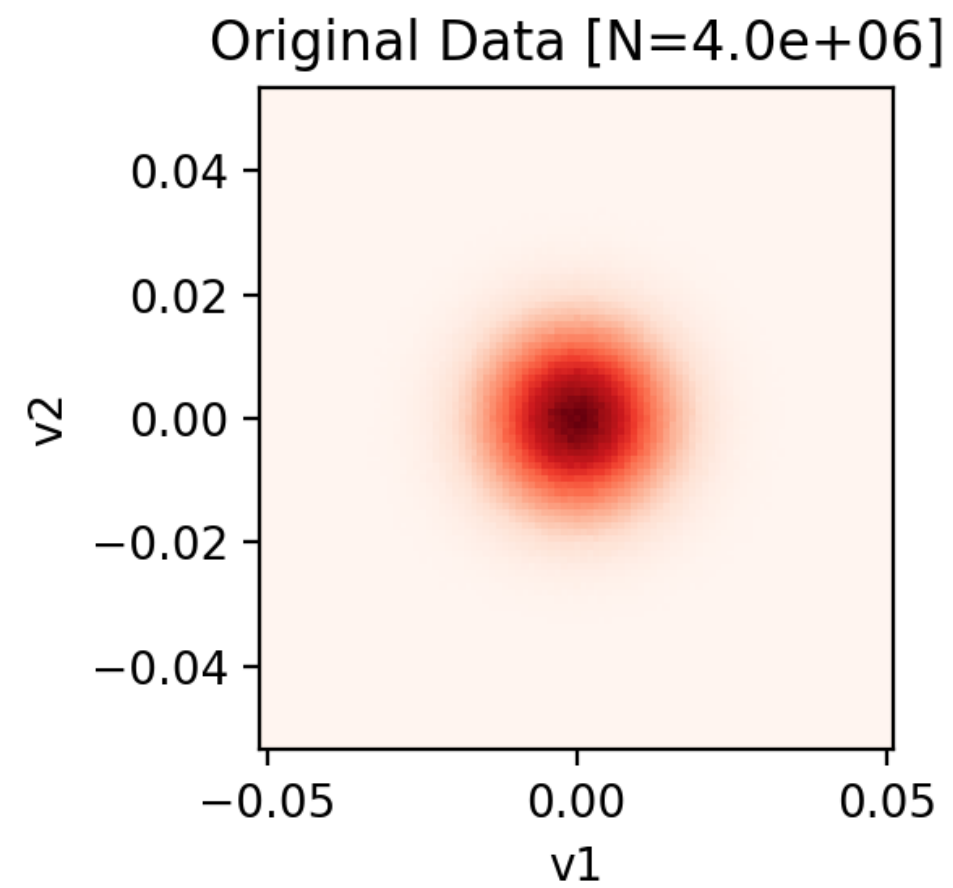


Test (20x)

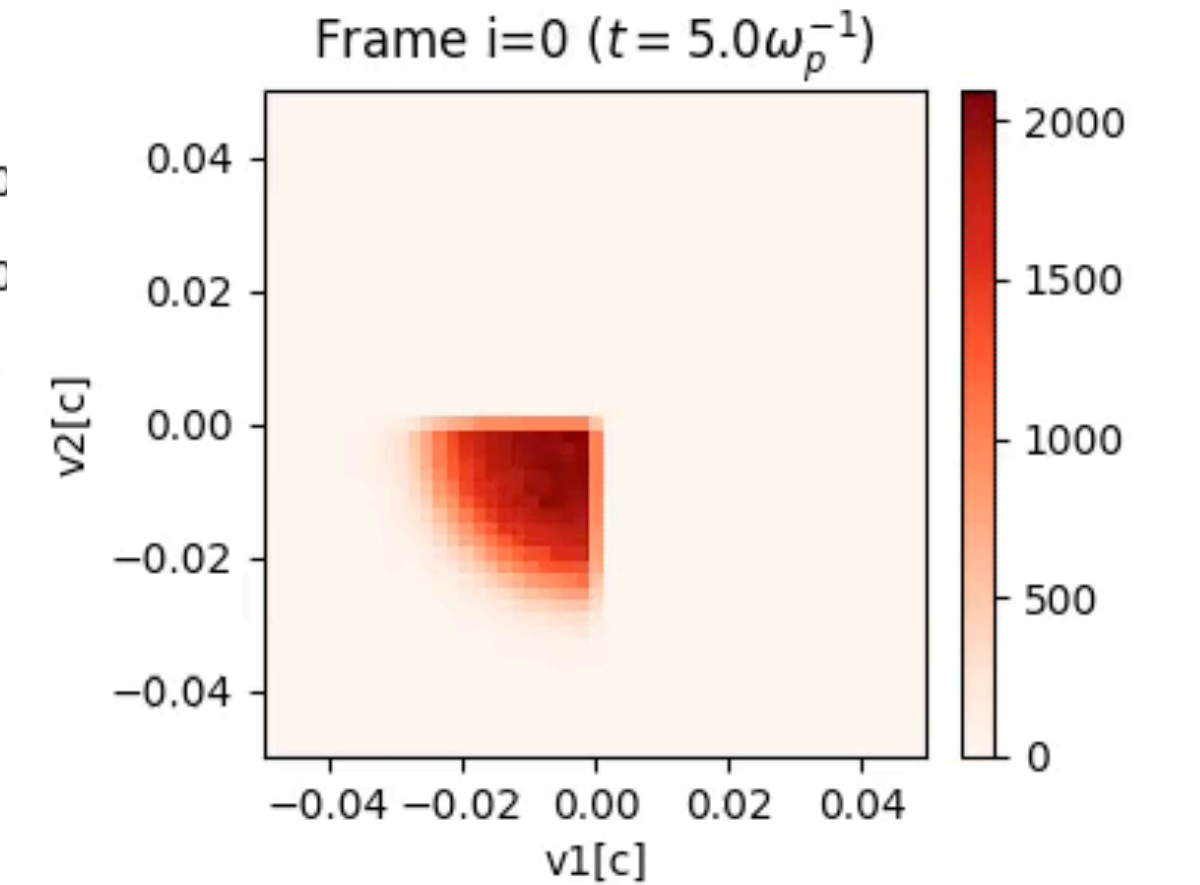
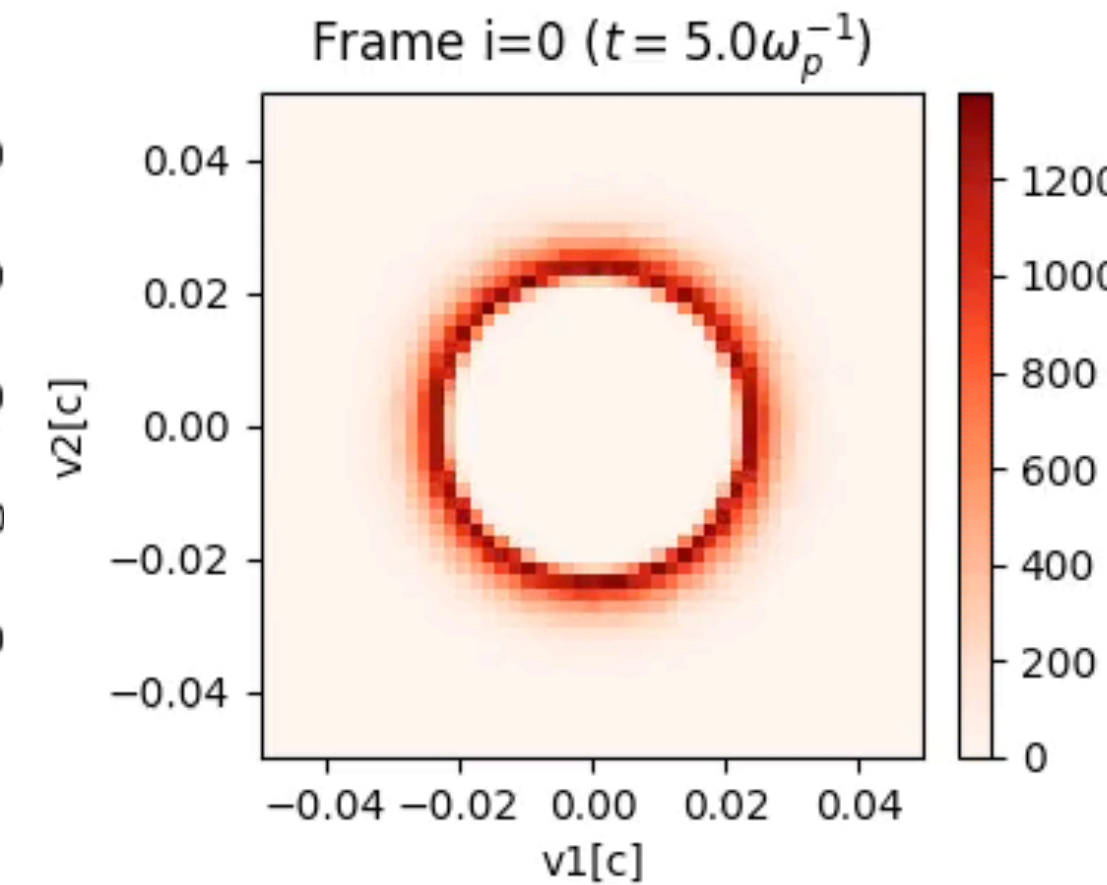
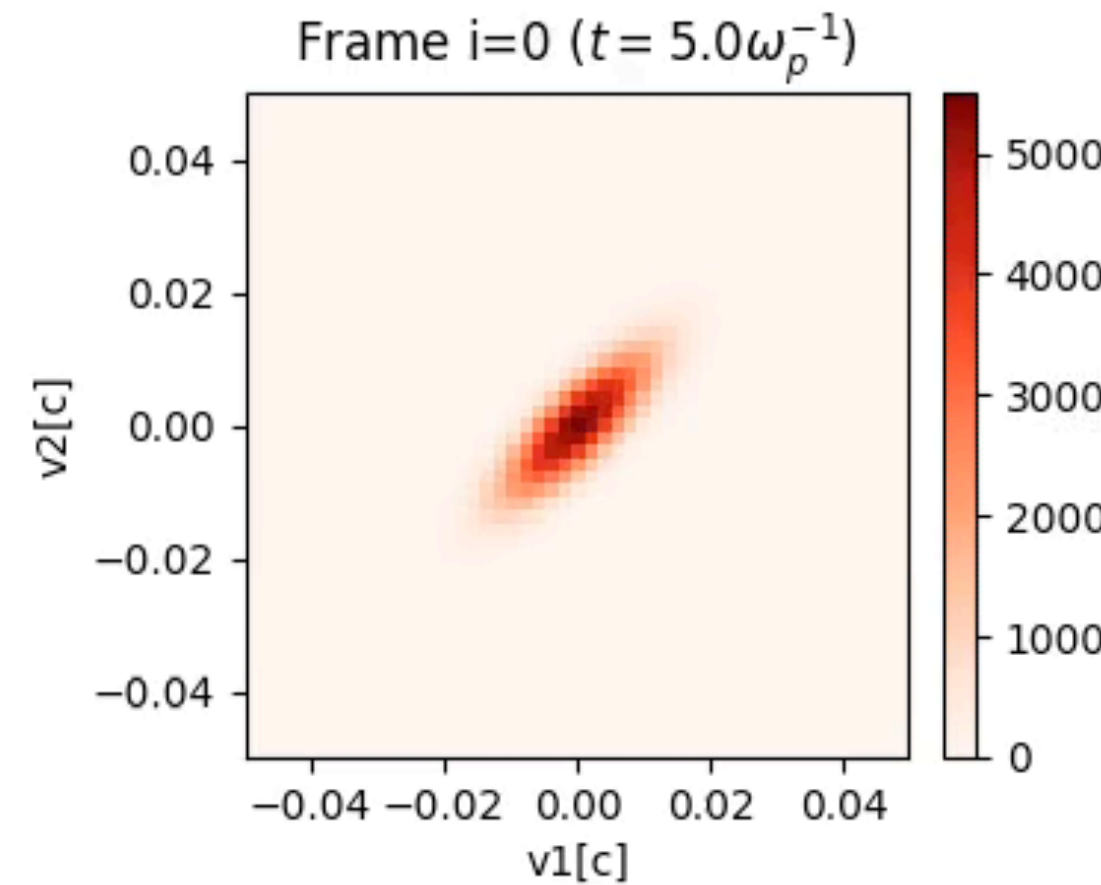


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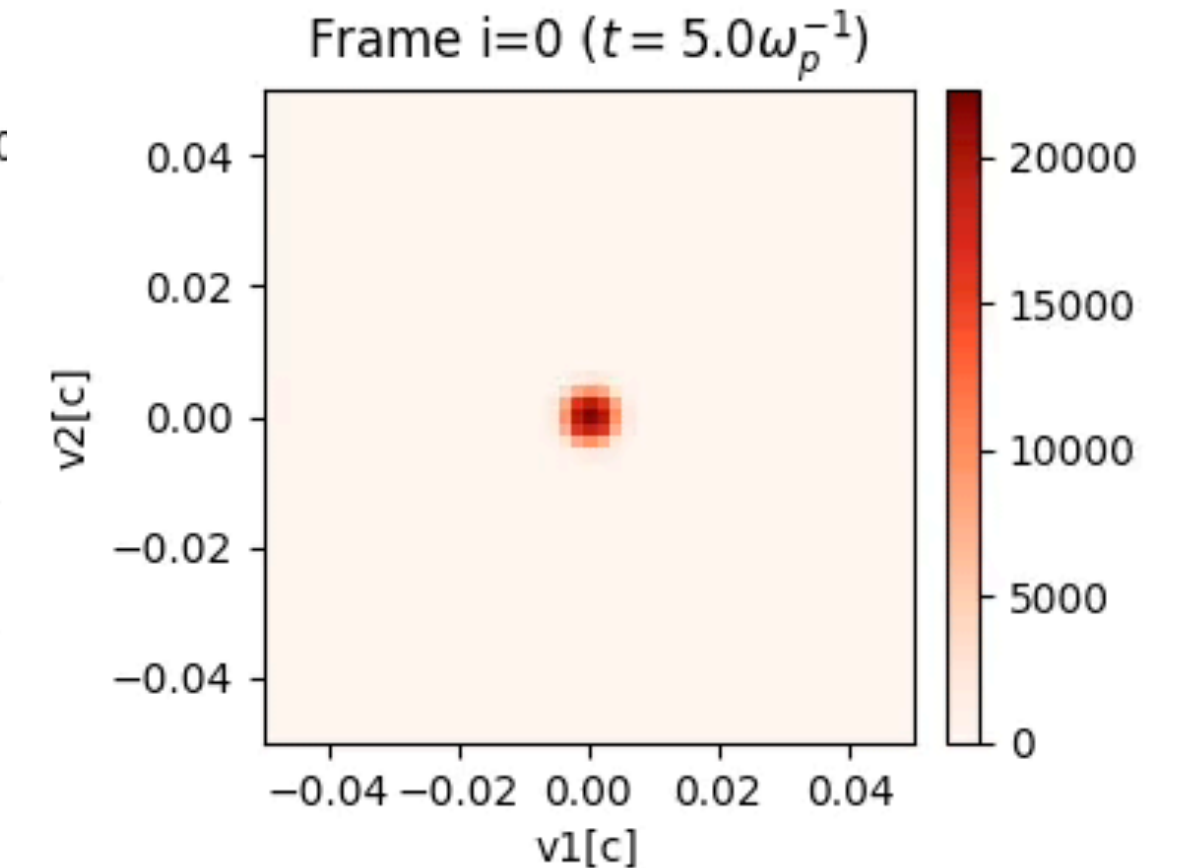
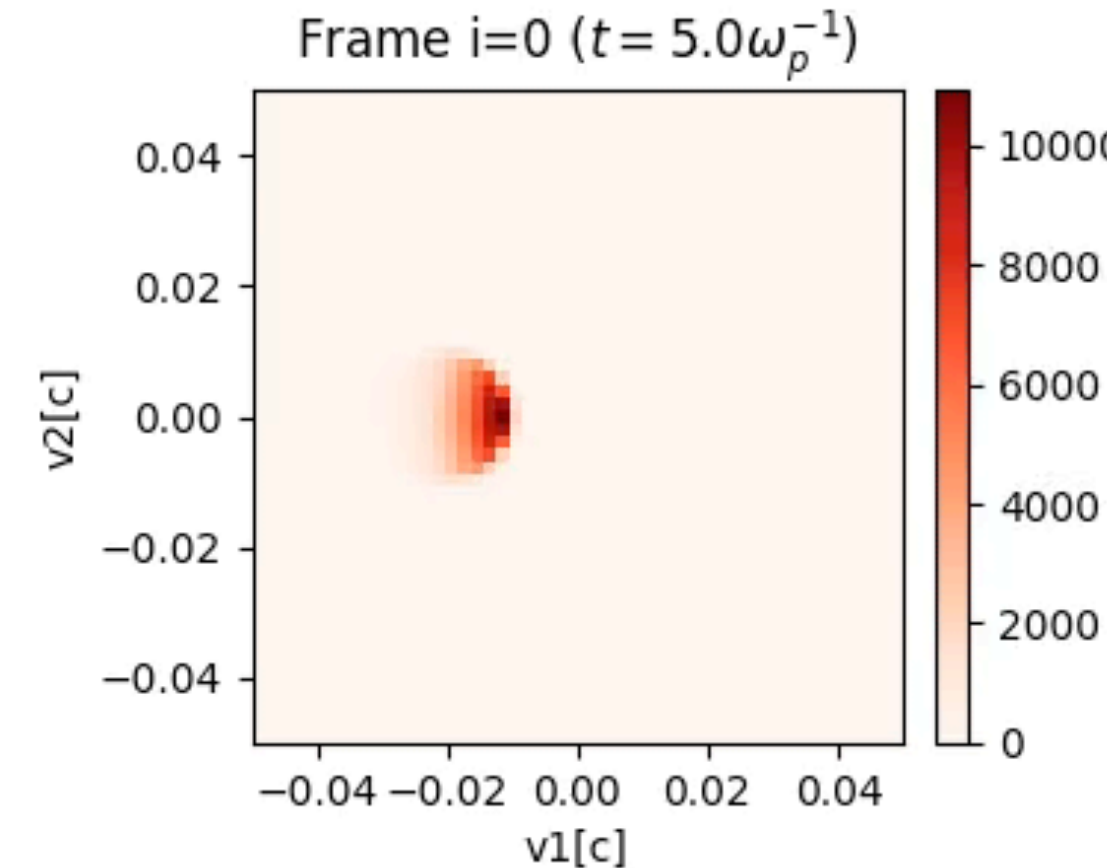
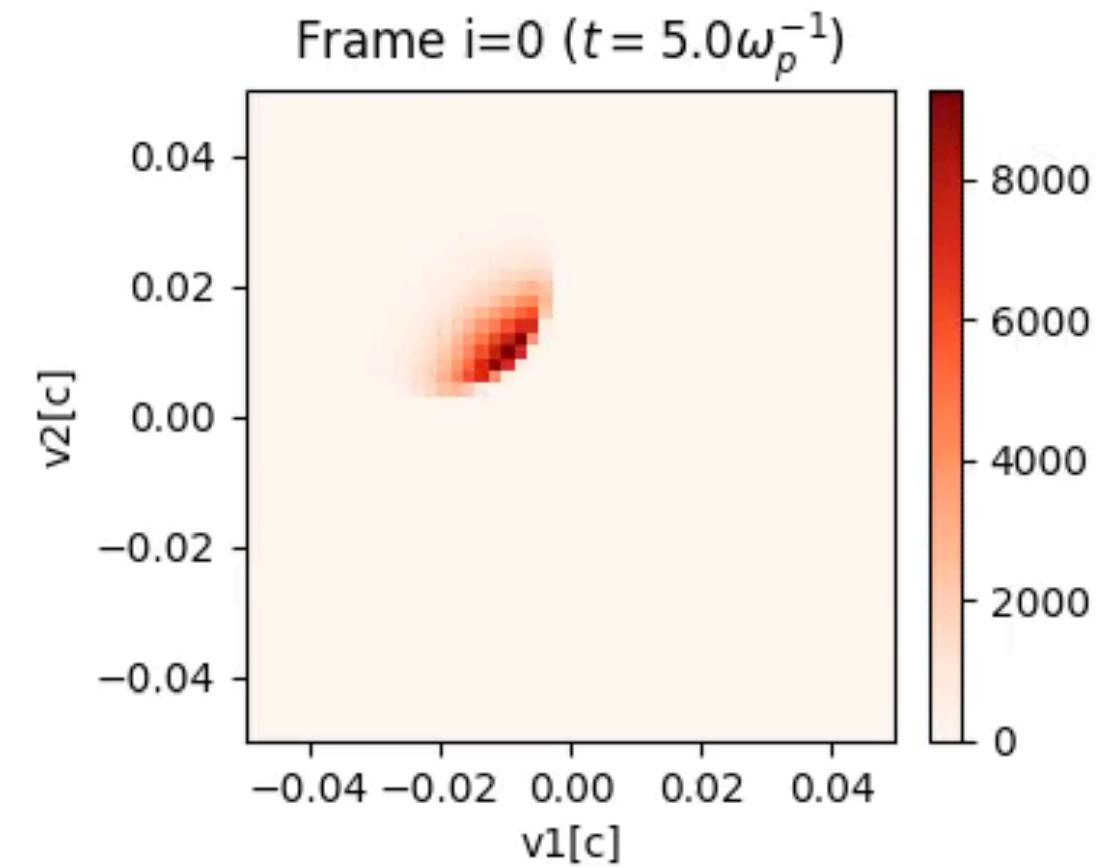
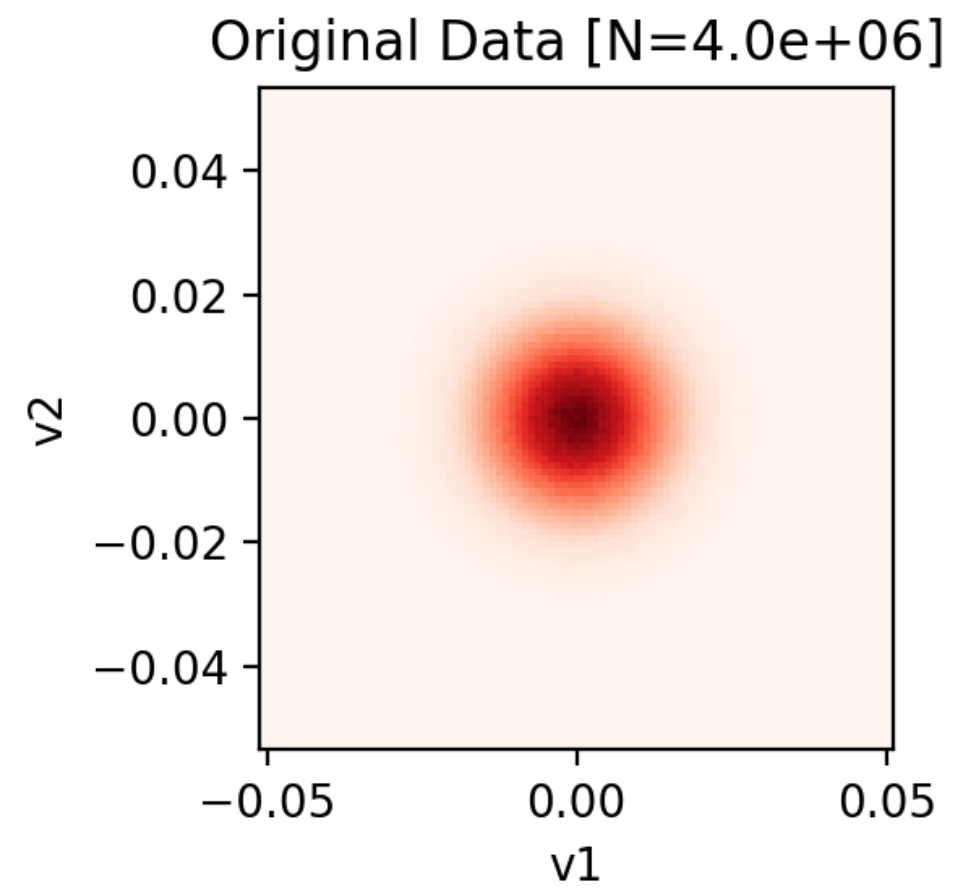


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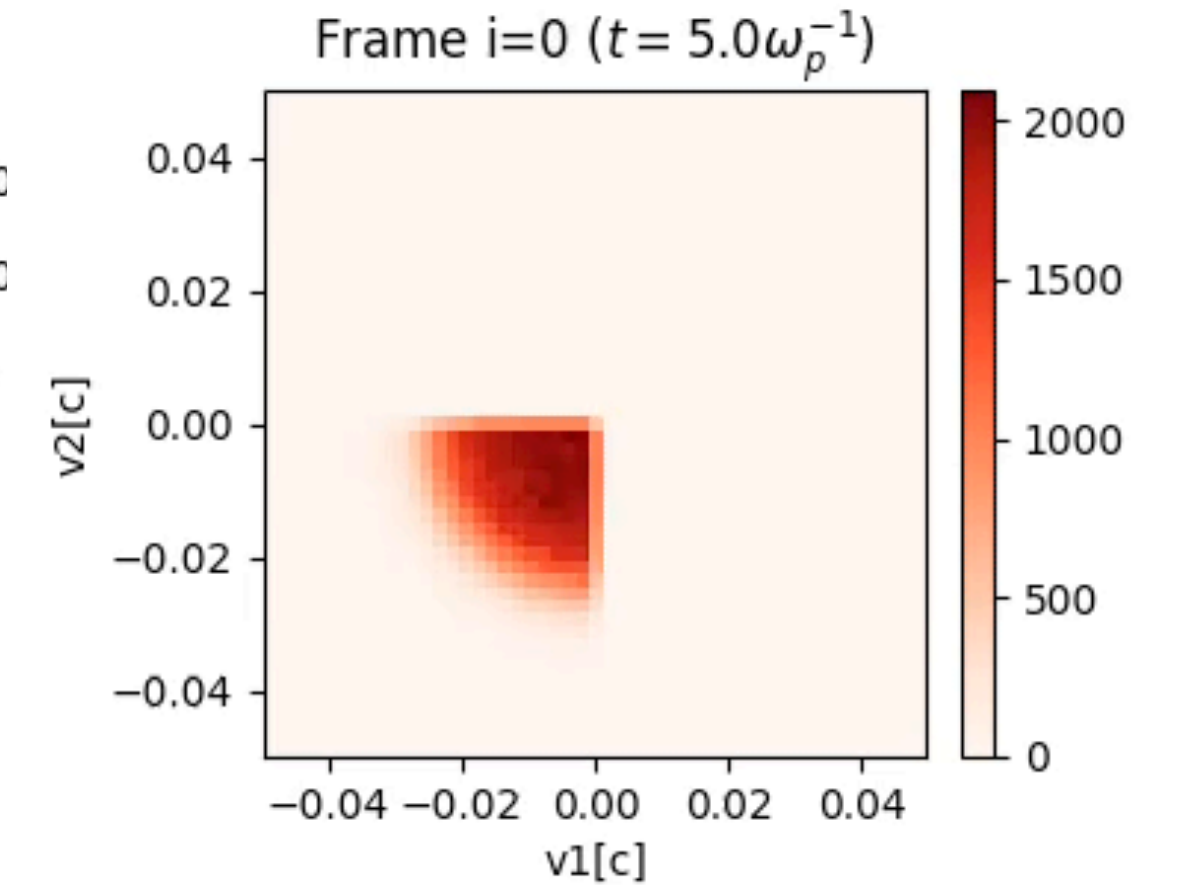
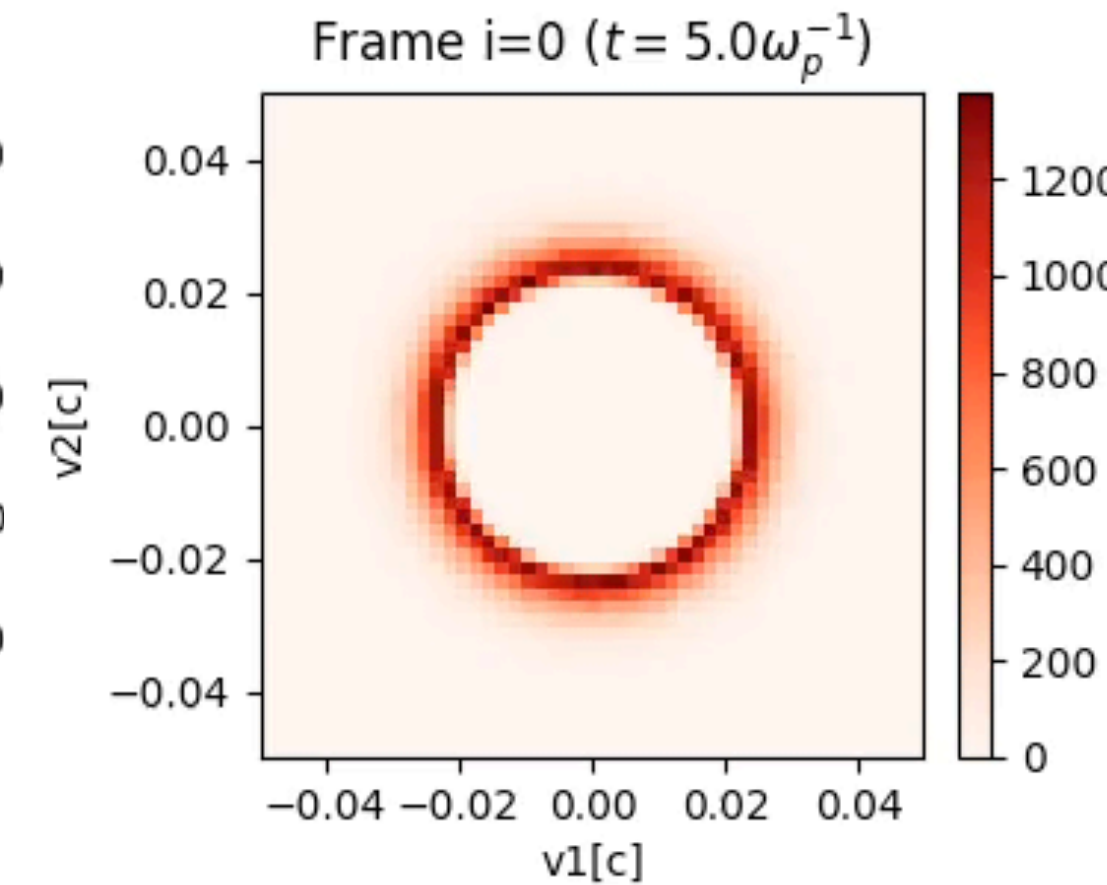
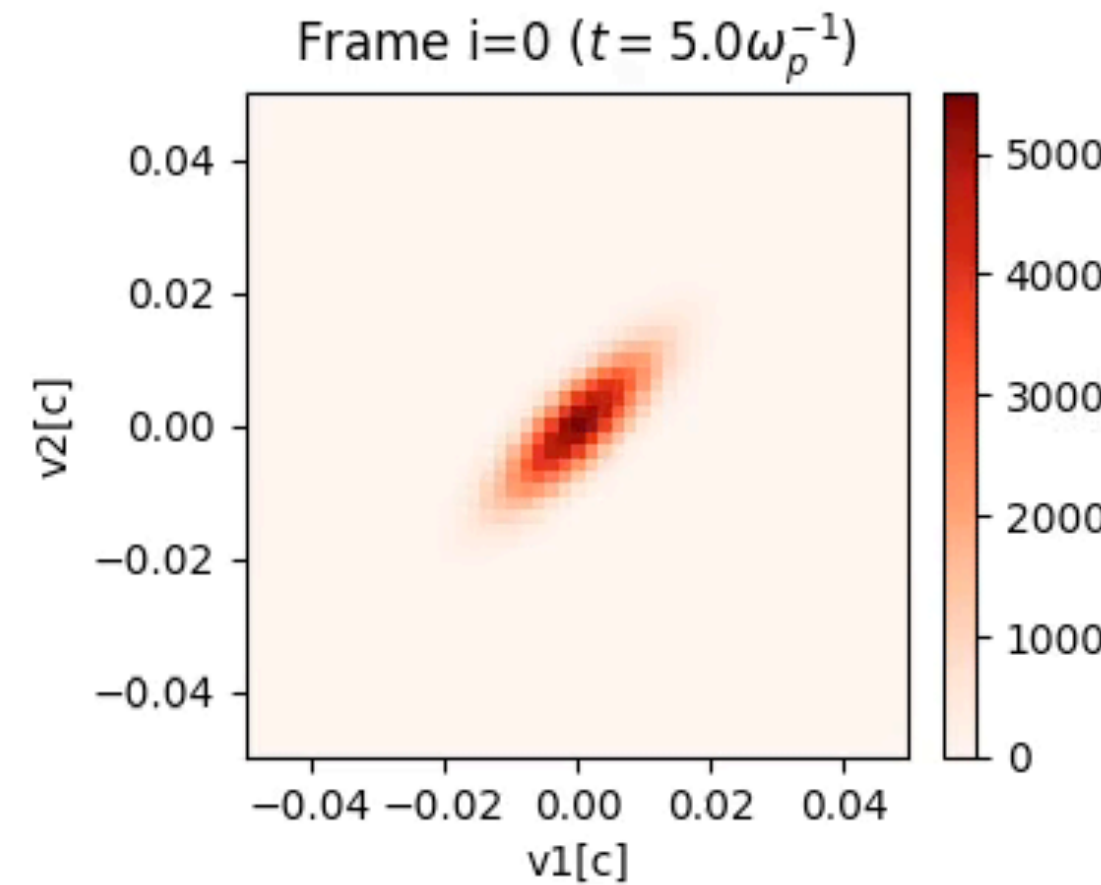


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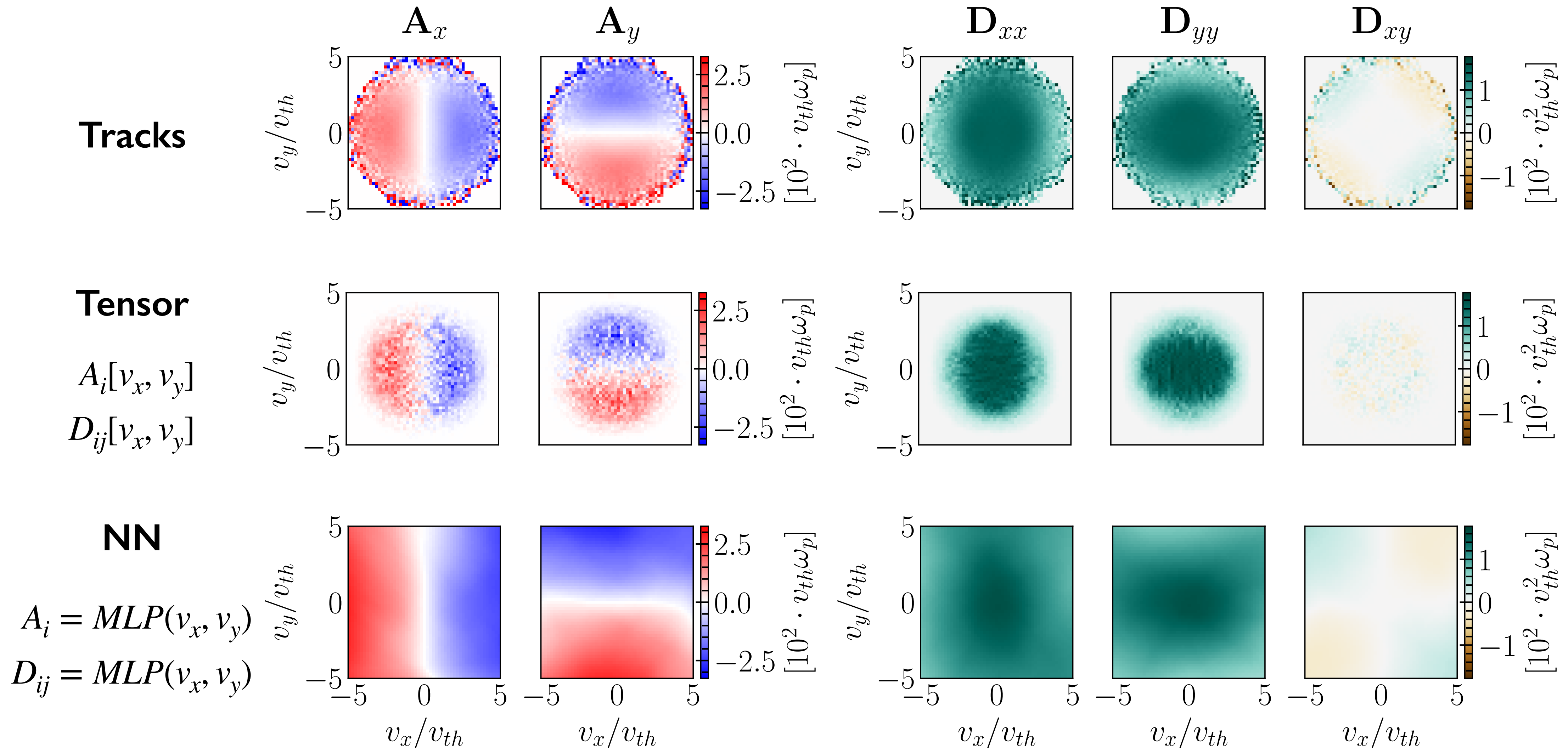
Train (9x)



Test (20x)



We can parameterise A/D using a Tensor (discrete) or a NN (continuous)

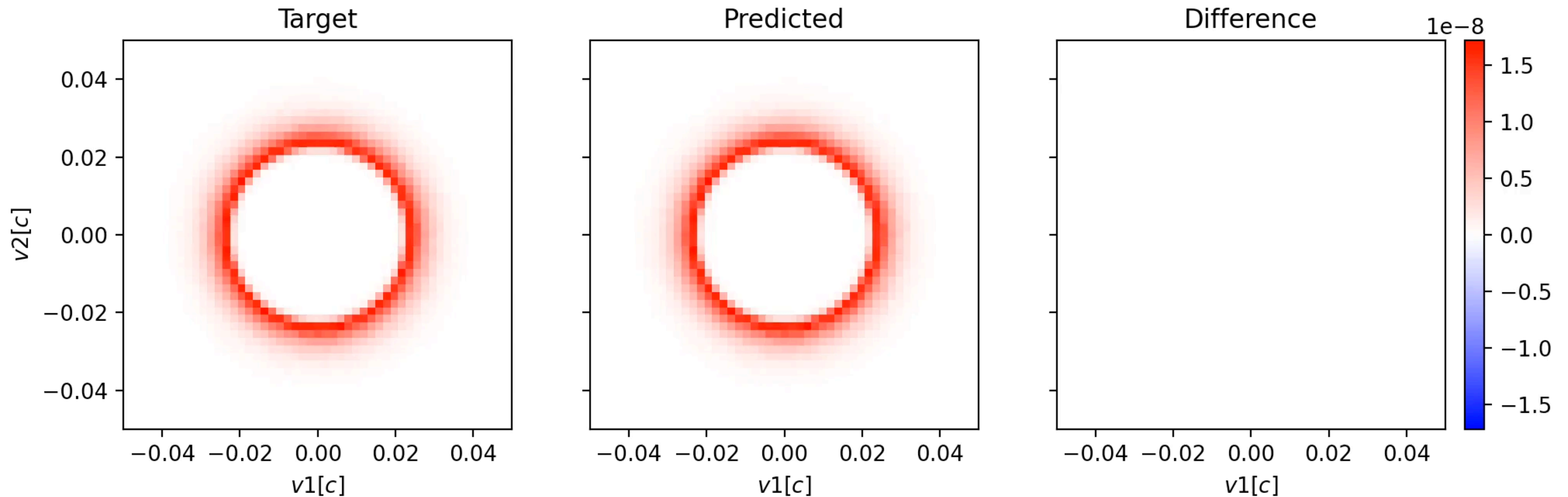


Can these operators reproduce dynamics?

$$f^{t+1} = f^t + \Delta t \left(-\nabla_{\mathbf{v}} \cdot (\mathbf{A} f^t) + \frac{1}{2} \nabla_{\mathbf{v}} \nabla_{\mathbf{v}} \cdot (\vec{\mathbf{D}} f^t) \right)$$

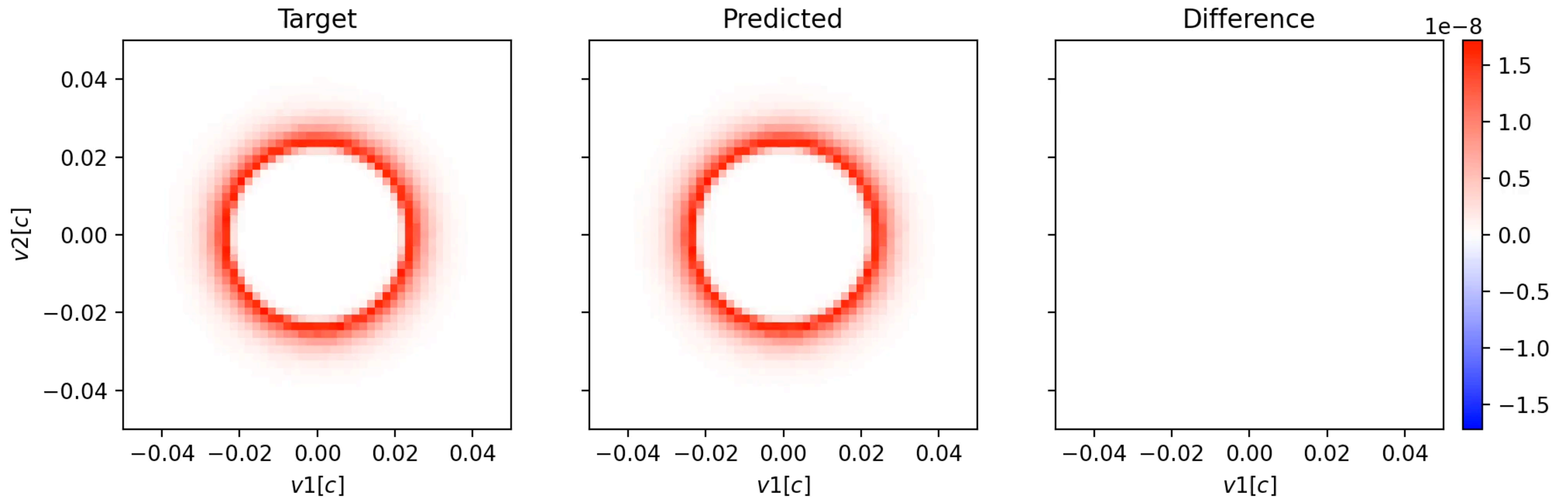
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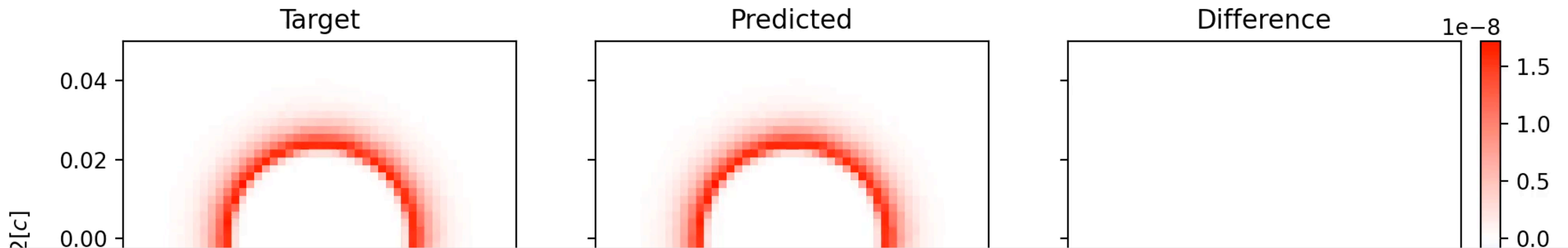
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Next steps

General purpose library: from phase space data, retrieve **A** and **D** (to be inserted on Fokker-Planck codes) for varying plasma conditions, and from different sources of data

Sub-module to capture (PIC or other) collisions for mesoscale simulations

Meta analysis: use different A and D for different plasma conditions (n, B, T) to learn more general behaviour e.g. via sparse regression



MC models in PIC simulations

New simulator models - GNN collisional plasma model

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The (ground) truth? - collisions in PIC codes

What is the ground truth?

Klimontovich + Maxwell's equations

$$\frac{\partial N}{\partial t} + \mathbf{v} \cdot \nabla_{\mathbf{x}} N - \frac{q}{m} (\mathbf{E}^m + \mathbf{v} \times \mathbf{B}^m) \cdot \nabla_{\mathbf{v}} N = 0$$

This is the particle-in-cell algorithm (with finite-size particles): statistical mechanics is well-known (e.g. H. Okuda and C. Birdsall (1970), R. Hockney (1971), M.Touati et al. (2022), S.Jubin et al. (2024)) + Born-Infeld electrodynamics

What if the cell/particle size is shorter than the classical electron radius?

What are the challenges of running $\ll 1$ ppc?

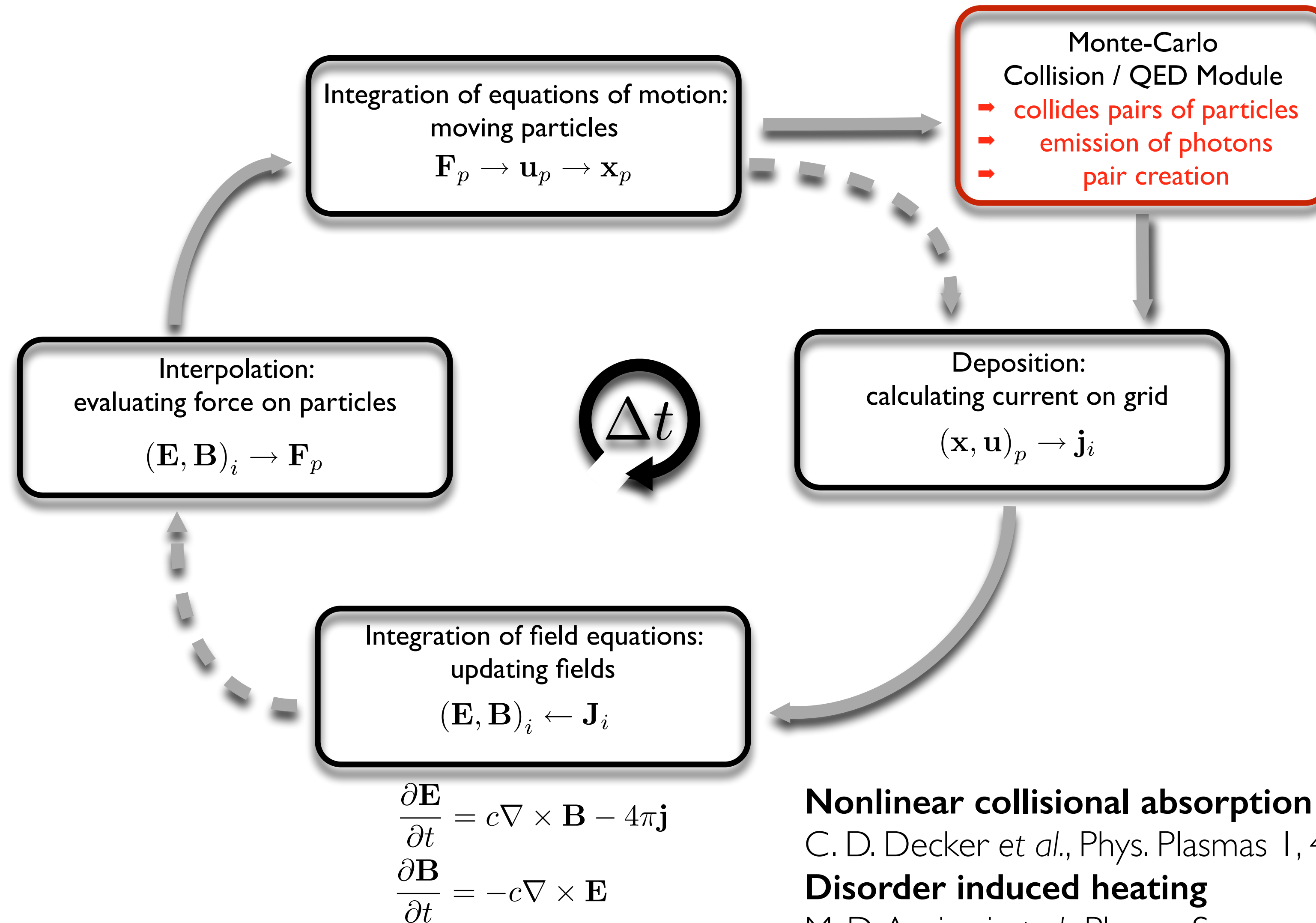
Field initialization becomes critical + Computation determined by grid (N^3 or N^2)

Numerical heating (still need to resolve Debye length) + Very small time steps (CFL) + numerical transition radiation

Validation against theory (but theory is very limited - only 2D) or computational models (MD non relativistic)

Shape functions to capture quantum effects?

The PIC loop by itself should be able to model collisional dynamics



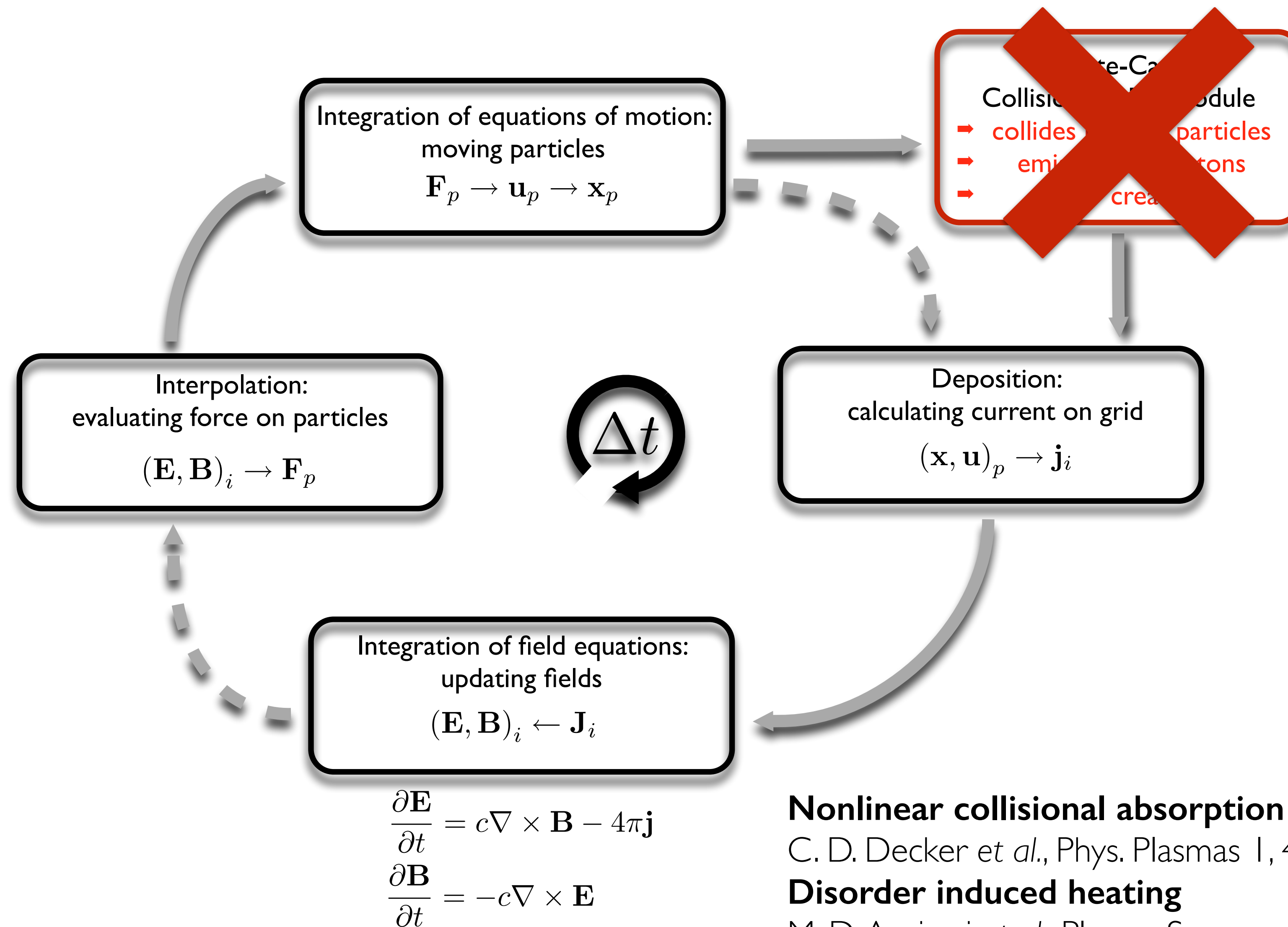
Nonlinear collisional absorption in laser-driven plasmas

C. D. Decker *et al.*, Phys. Plasmas 1, 4043–4049 (1994)

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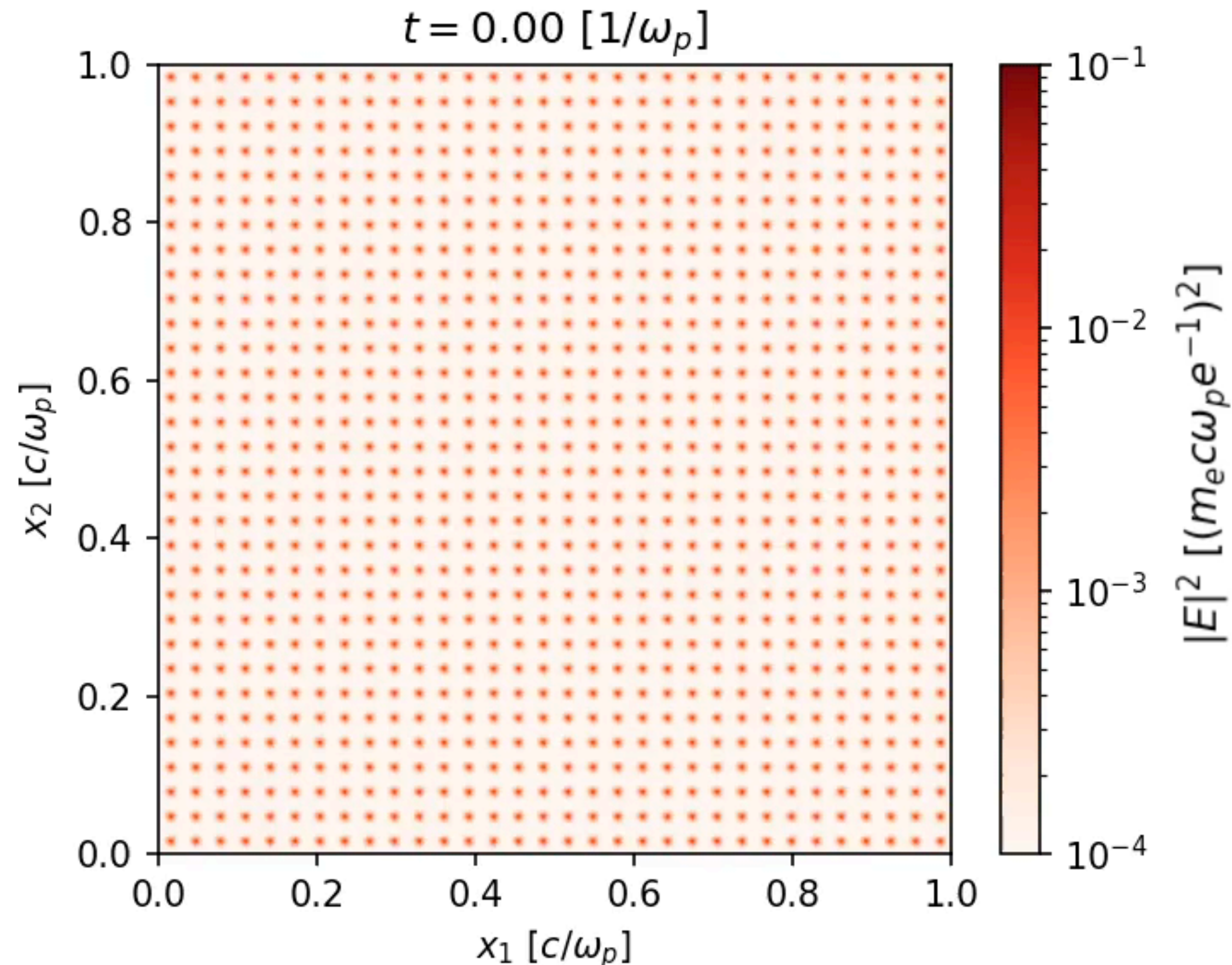
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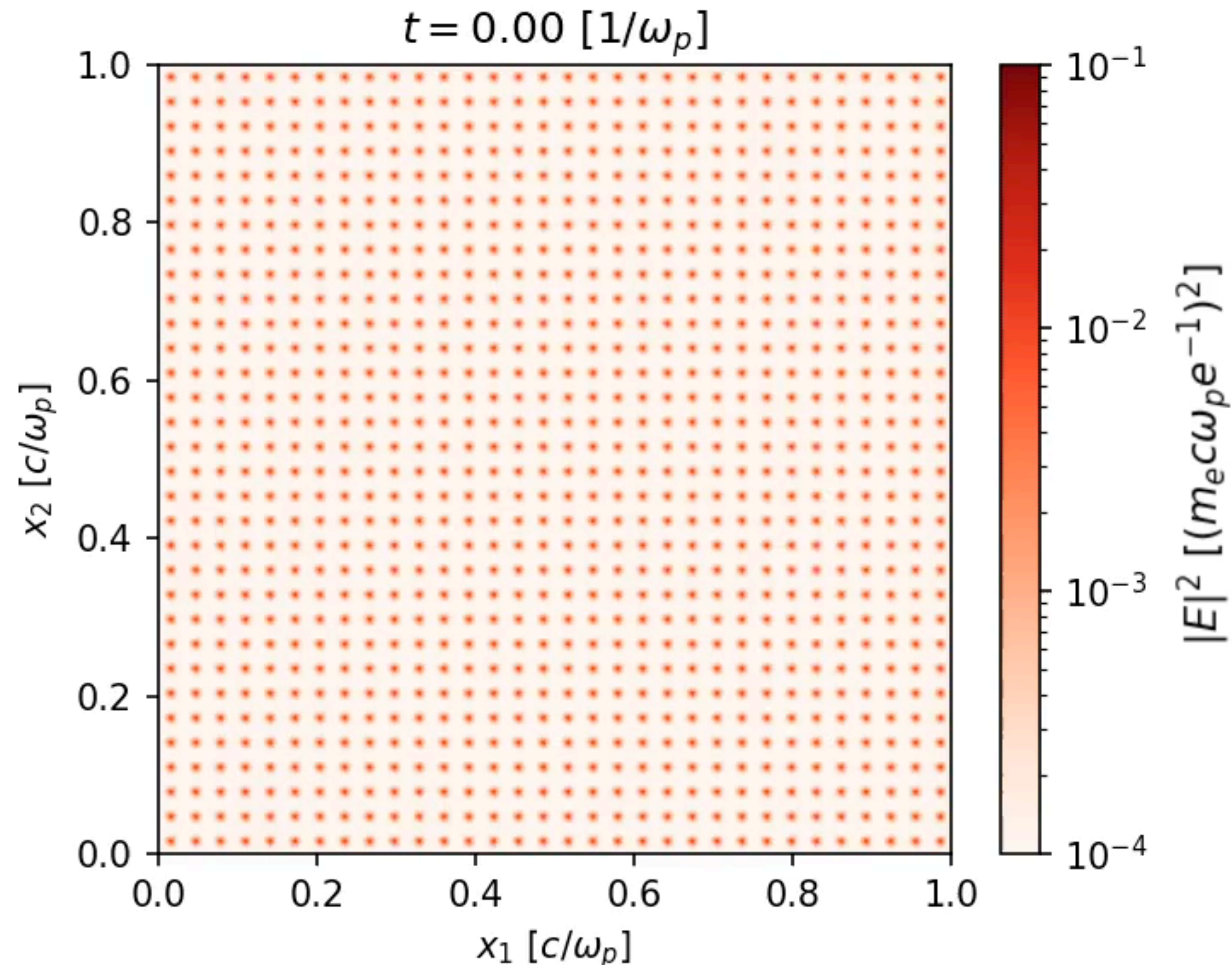
Simulating self-consistent collisions with PIC

$$n\lambda_D^2 \simeq 0.1 \quad (\text{Average interparticle distance} \gg \lambda_D)$$

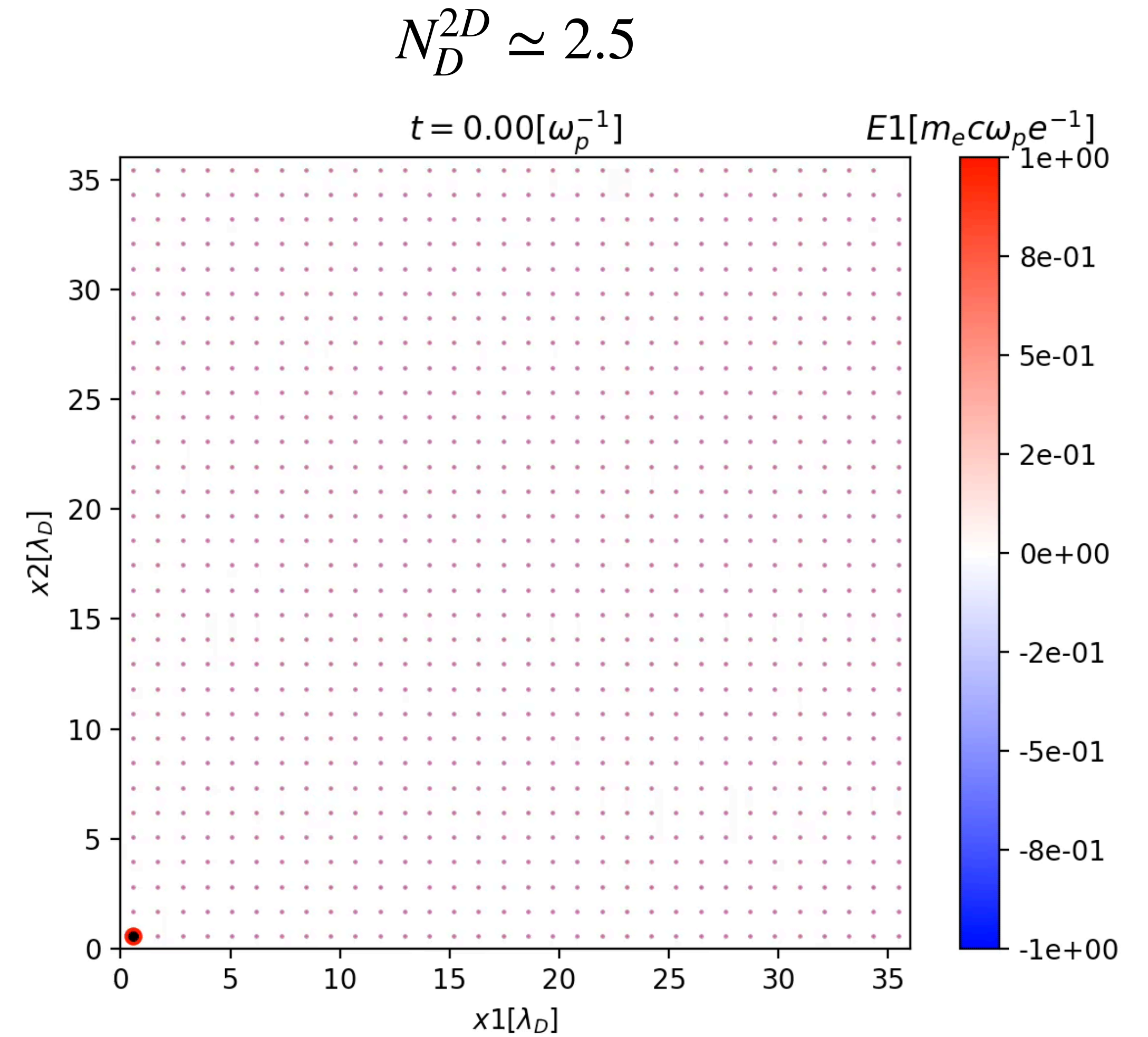
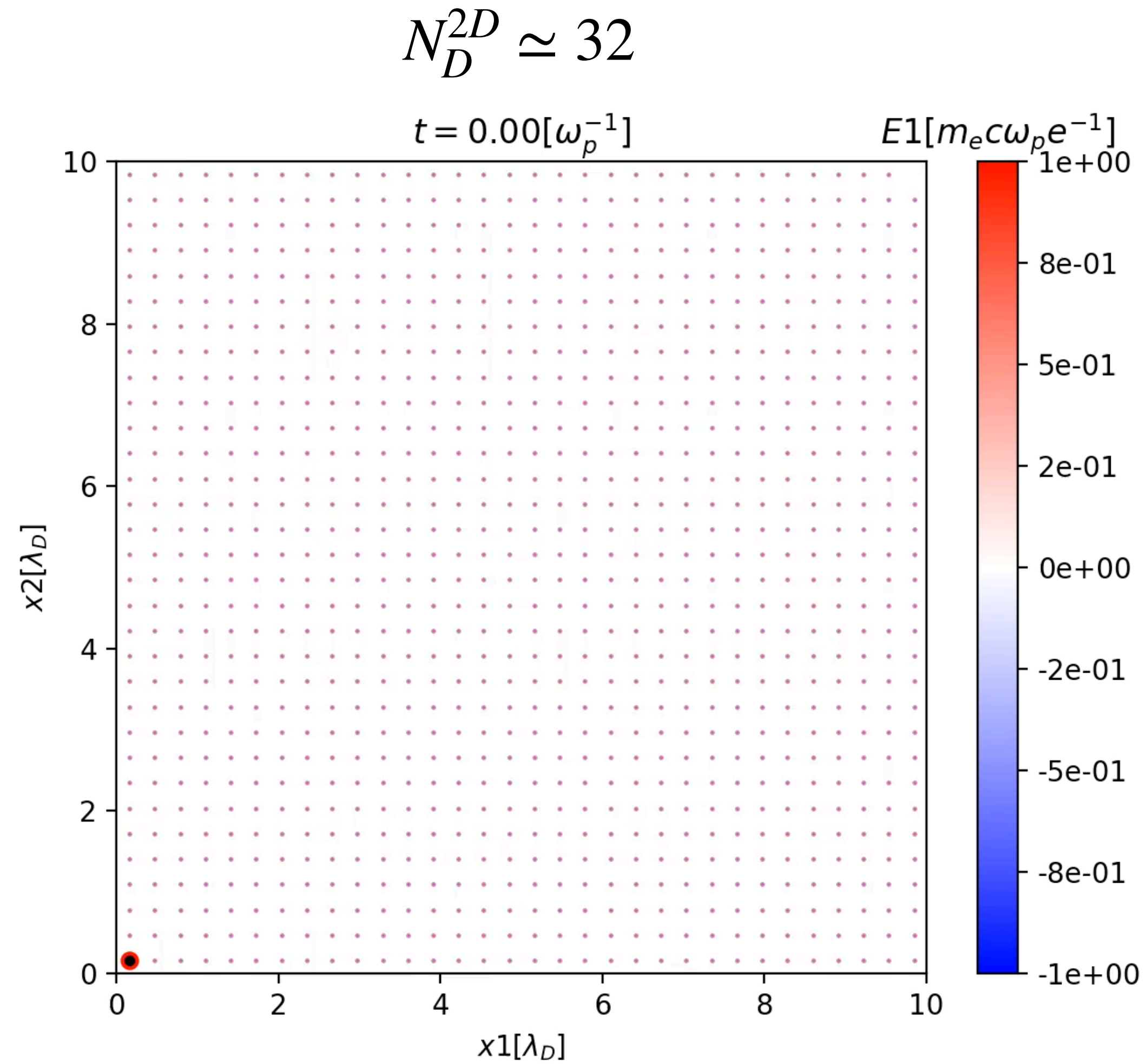


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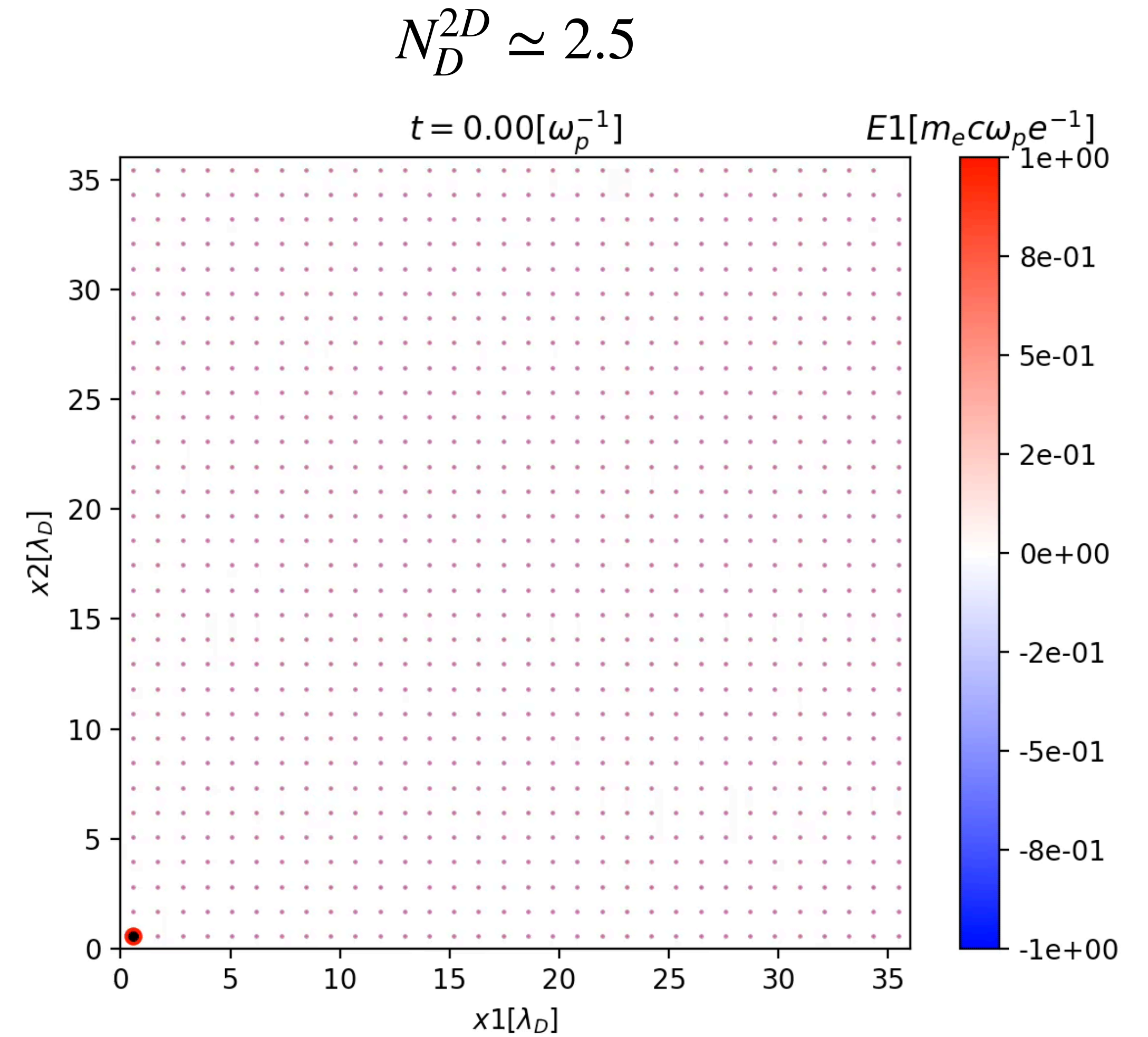
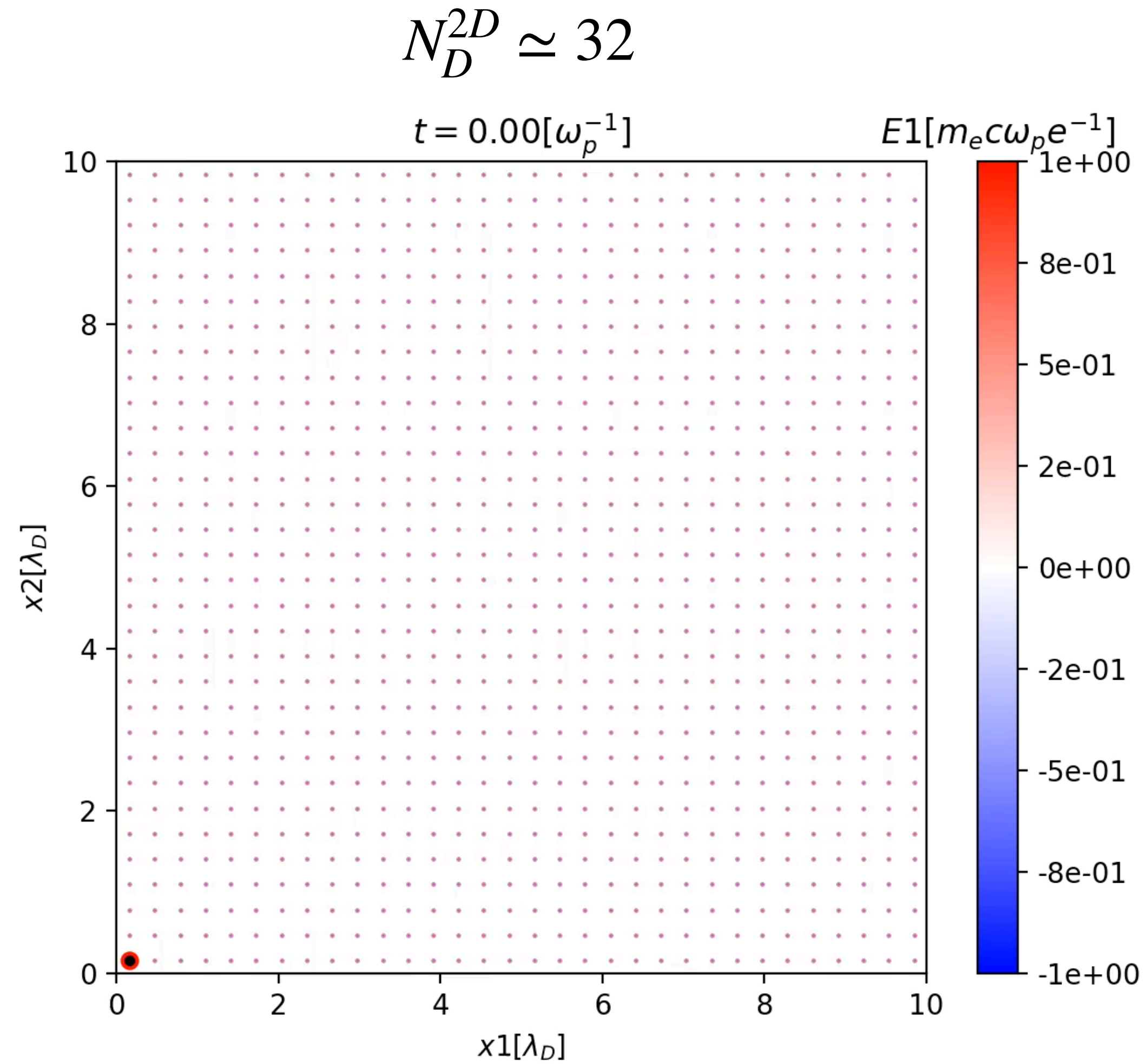
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Can we learn operators that describe the collisional dynamics?

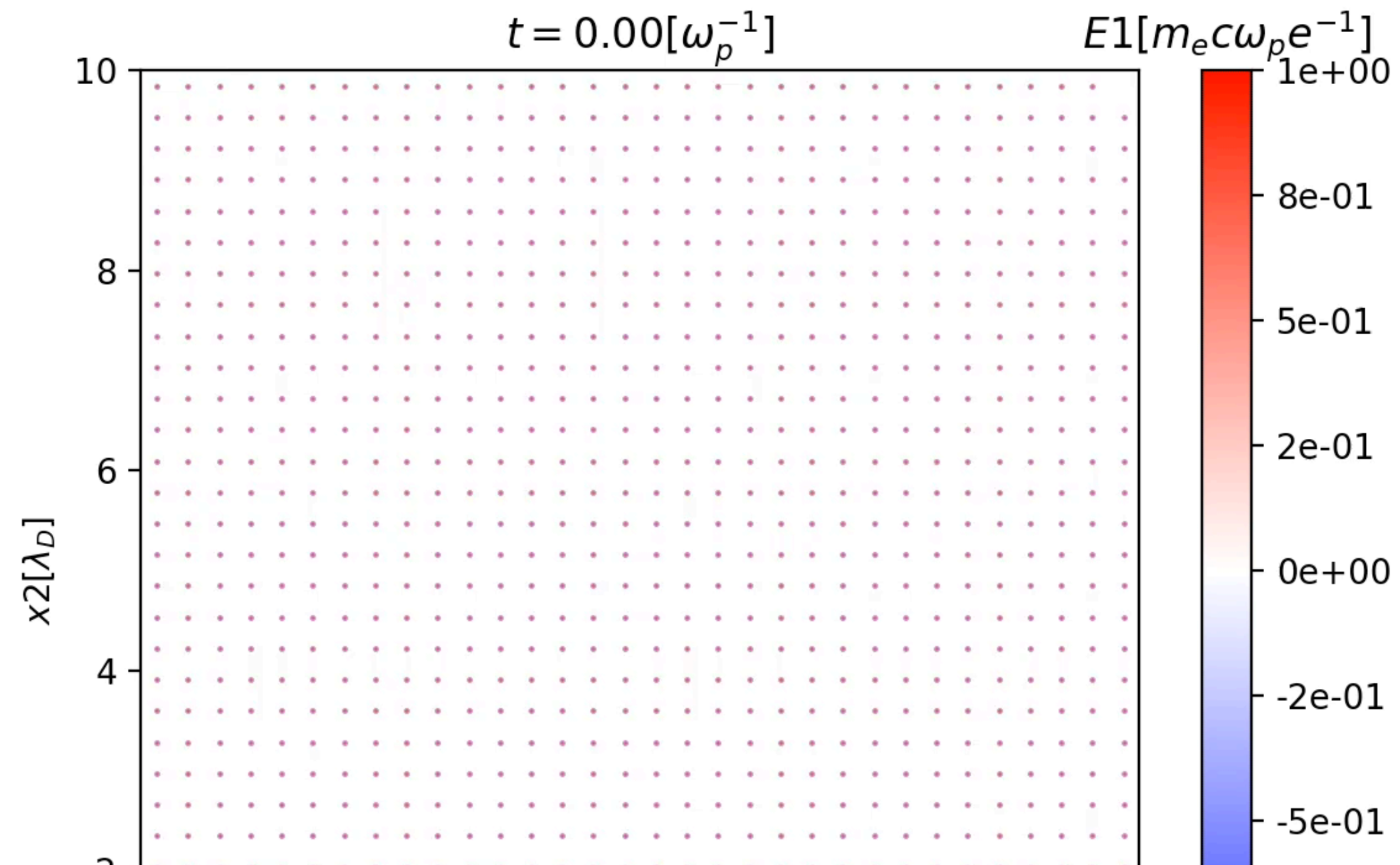


Can we learn operators that describe the collisional dynamics?

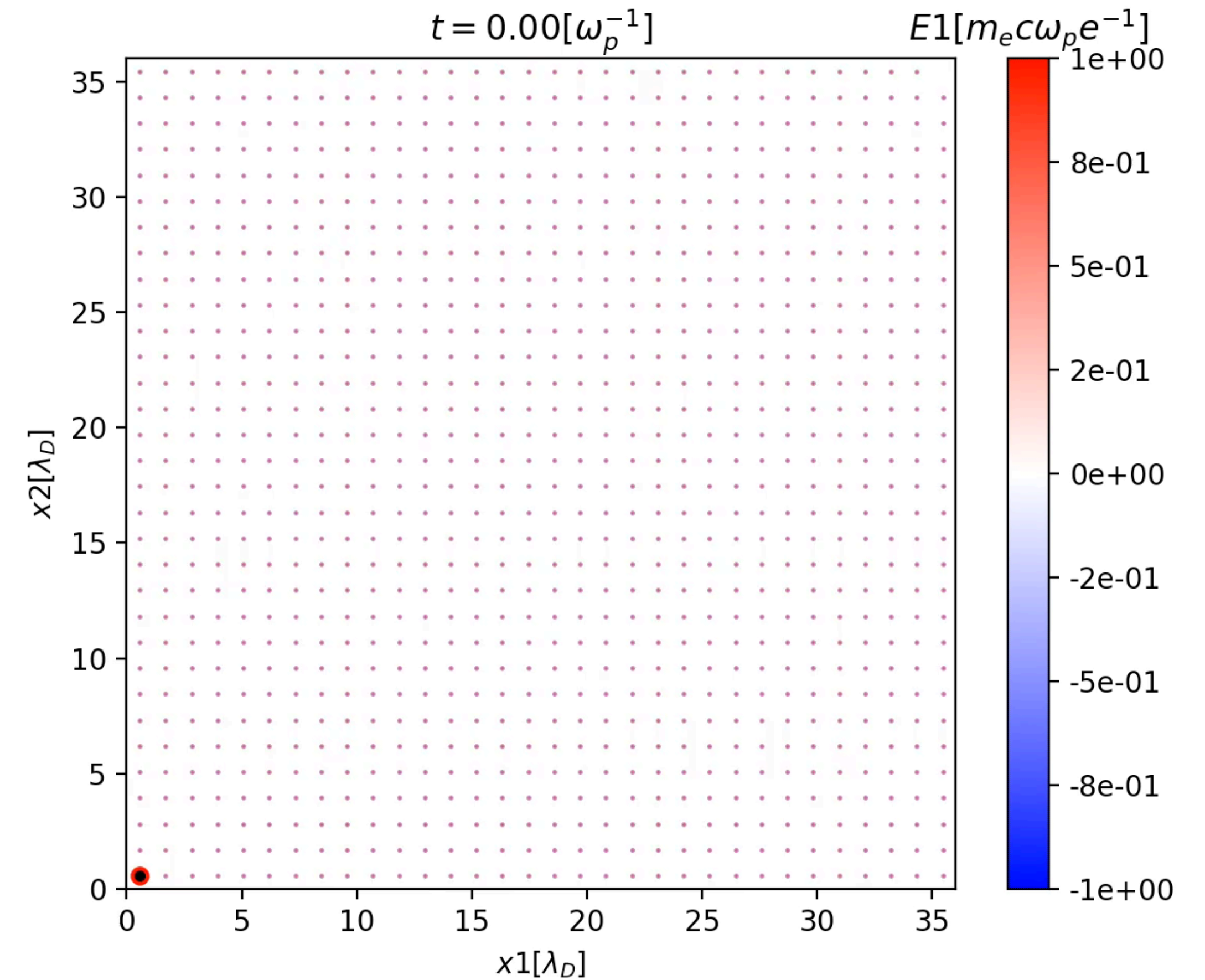


Can we learn operators that describe the collisional dynamics?

$$N_D^{2D} \simeq 32$$



$$N_D^{2D} \simeq 2.5$$



Work in progress

Comparison with theory: collision frequencies + **A** and **D** and 2D Fokker-Planck model (Morales et al.)

What about B fields?

Can ML help us speed up standard plasma simulators? No speed up but huge memory gains

Can we build faster ML based simulators? Yes, but with significant modifications to the algorithms/structure/philosophy

What can we learn from data-driven approaches + ML? It looks like we can learn a lot (and the community is learning how to do it): e.g. collision operators

Can standard plasma simulators provide “high quality data” for data-driven discovery? Yes, pushing for additional developments in HPC simulations

Interplay between HPC and AI is just starting: “There are (many) unknown unknowns” (which is great for science!)