

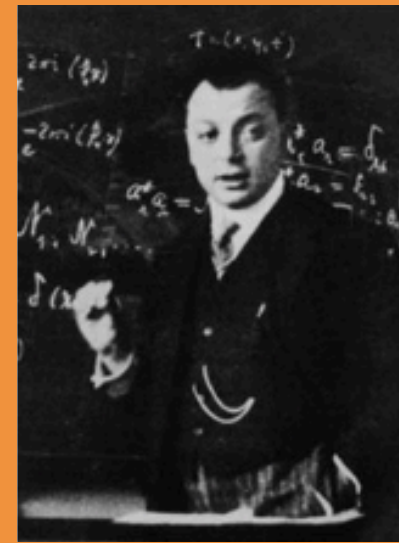
[Wolfgang Pauli Institute, Vienna](#)

14th Plasma Kinetics Working Meeting

24 July - 4 August, 2023

**Wrapped
is here.**

#WPIWRAPPED



[Wolfgang Pauli Institute, Vienna](#)

14th Plasma-Kinetics Working Meeting

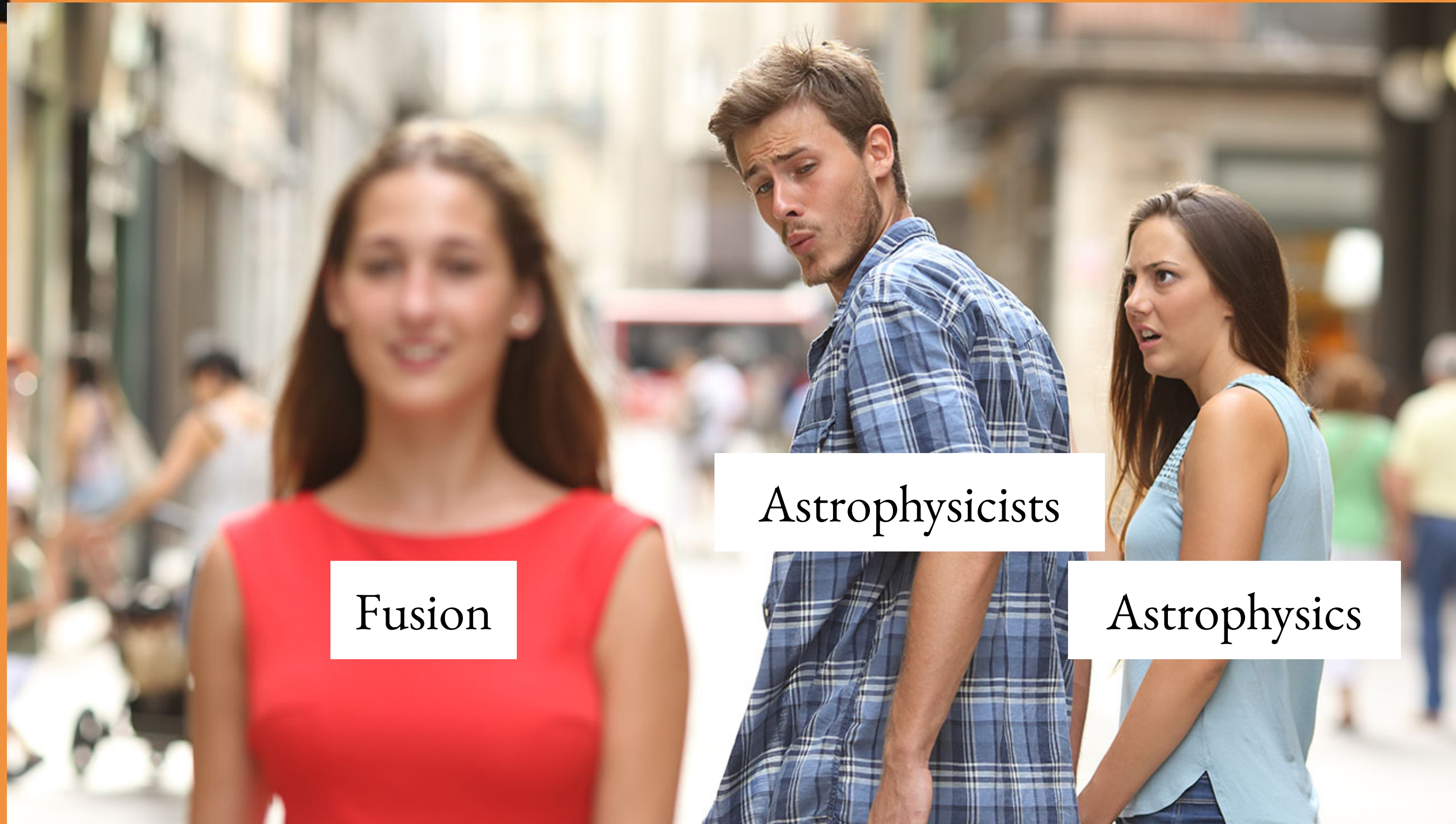
24 July - 4 August, 2023

**Wrapped
is here.**

#WPIWRAPPED



Ladies and gentlemen, it has been a very good year for the Society



Fusion

Astrophysicists

Astrophysics

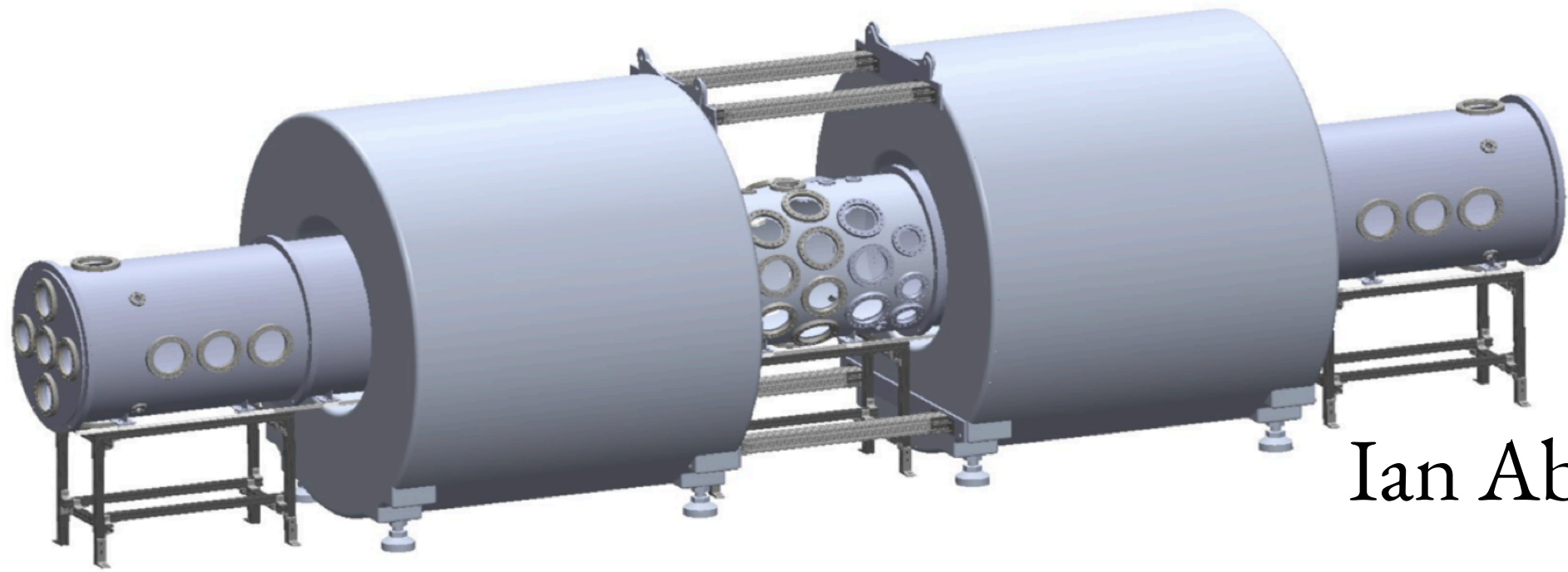




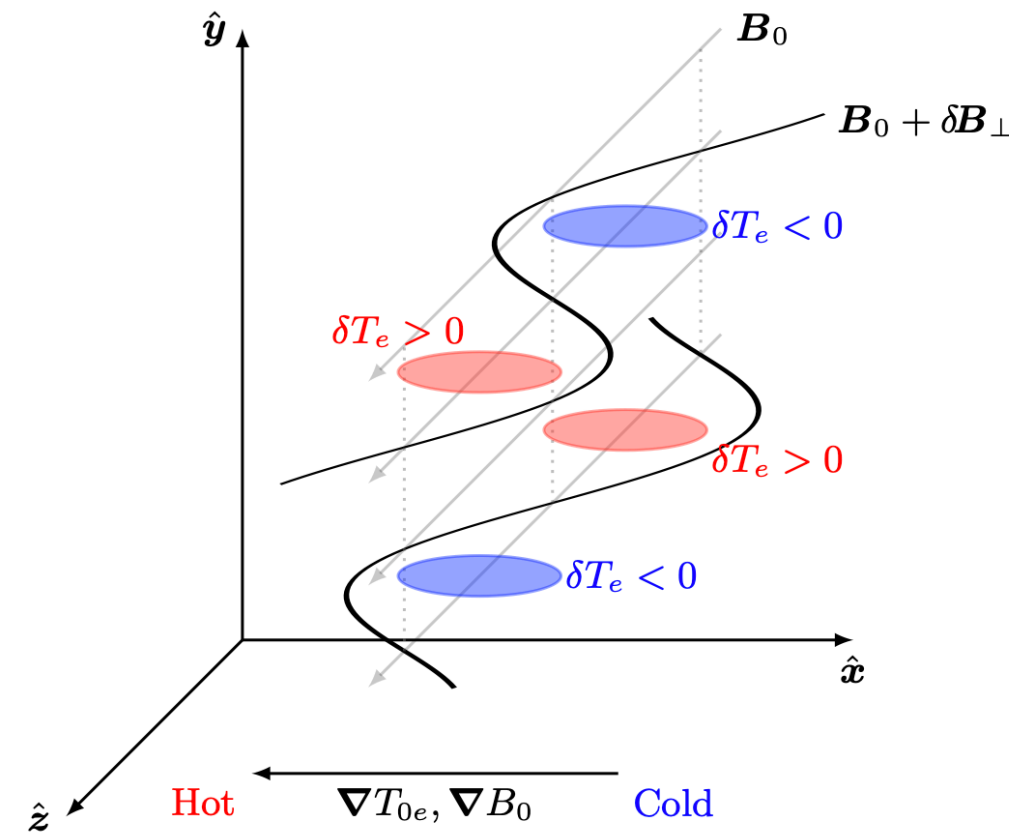
Recurring themes

You played **30 talks**
but these played again
and again

And again.



Ian Abel



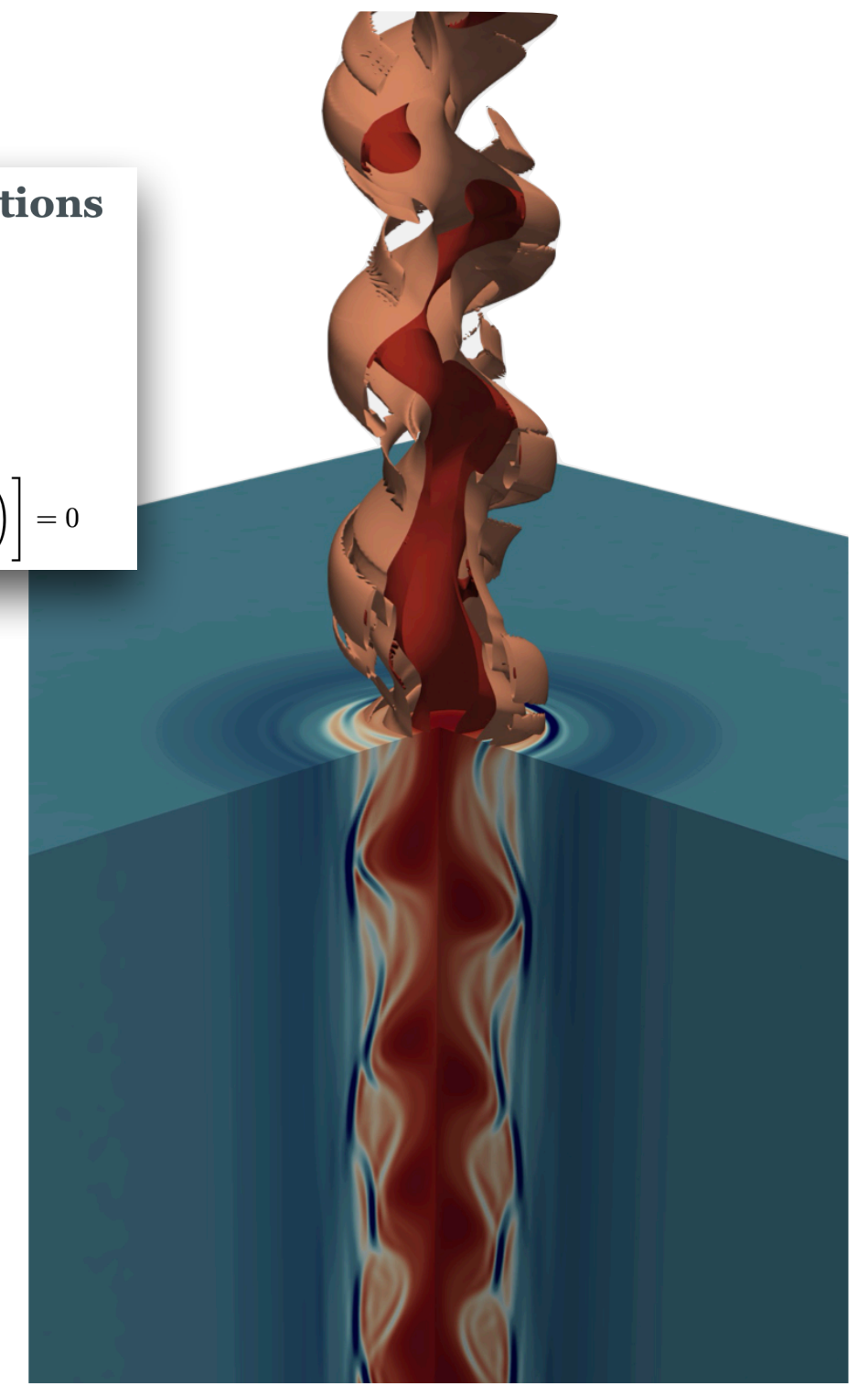
Toby Adkins

Summary of rational surface conditions

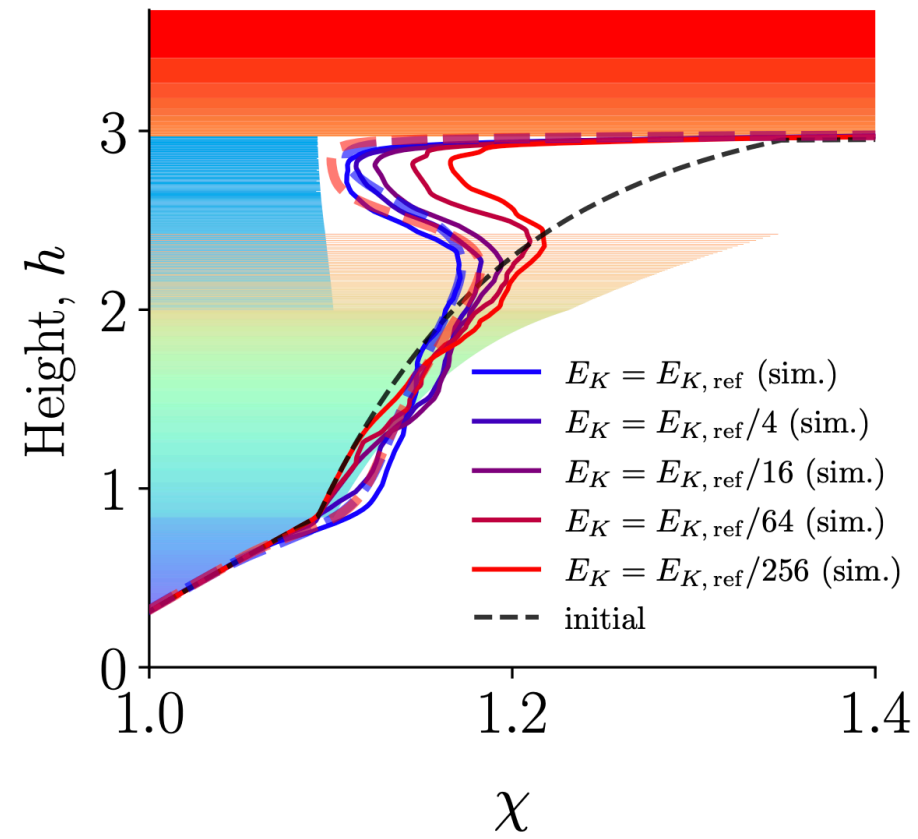
- Hamada condition: $P' \frac{d}{d\alpha} \left(\oint \frac{dl}{B} \right) = 0$
- New condition

$$\frac{d}{d\alpha} \left[\left(\oint \frac{B}{|\nabla\rho|^2} dl \right)^{-1} \left(\frac{c\Psi_t'}{4\pi} \oint dv - \oint \frac{J_{PS}}{|\nabla\rho|^2} dl + \frac{c}{4\pi} \oint \frac{[\mathbf{B} \cdot \nabla \hat{\mathbf{n}} \cdot (\mathbf{B} \times \hat{\mathbf{n}}) + (\mathbf{B} \times \hat{\mathbf{n}}) \cdot \nabla \hat{\mathbf{n}} \cdot \mathbf{B}]}{B|\nabla\rho|^2} dl \right) \right] = 0$$

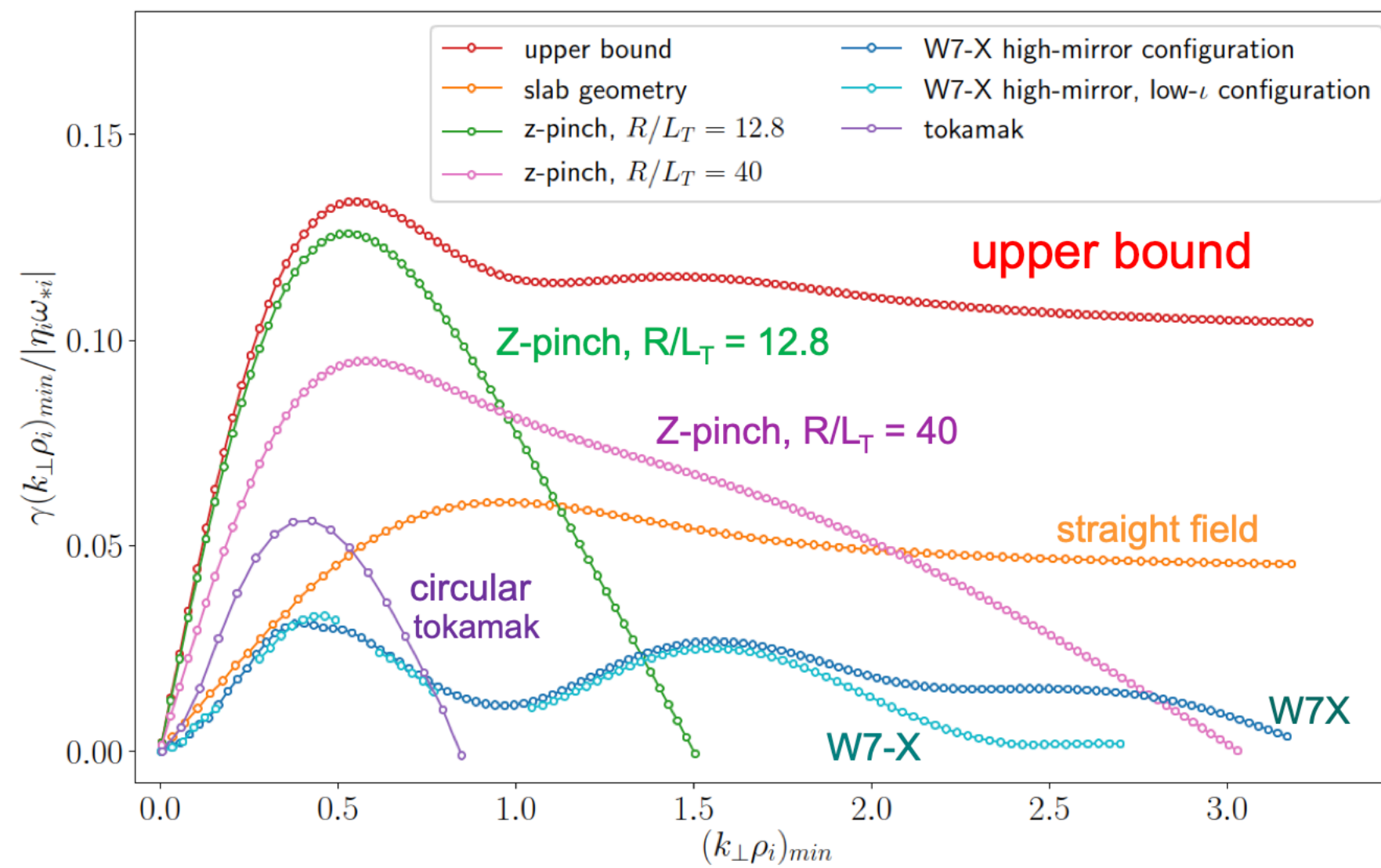
Felix Parra



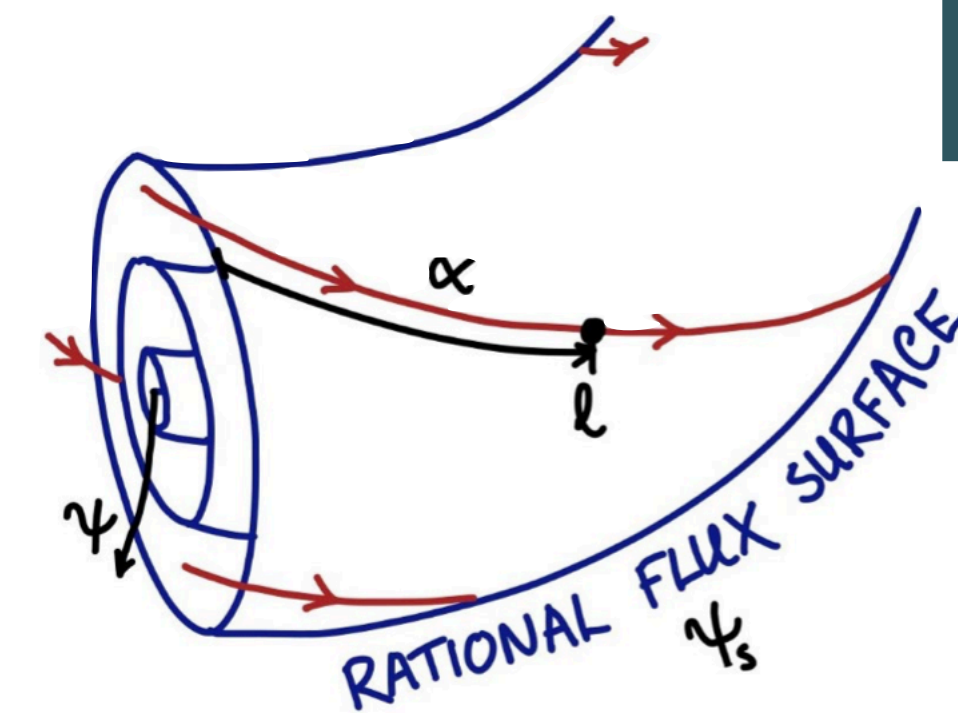
Alex Velberg



David Hosking

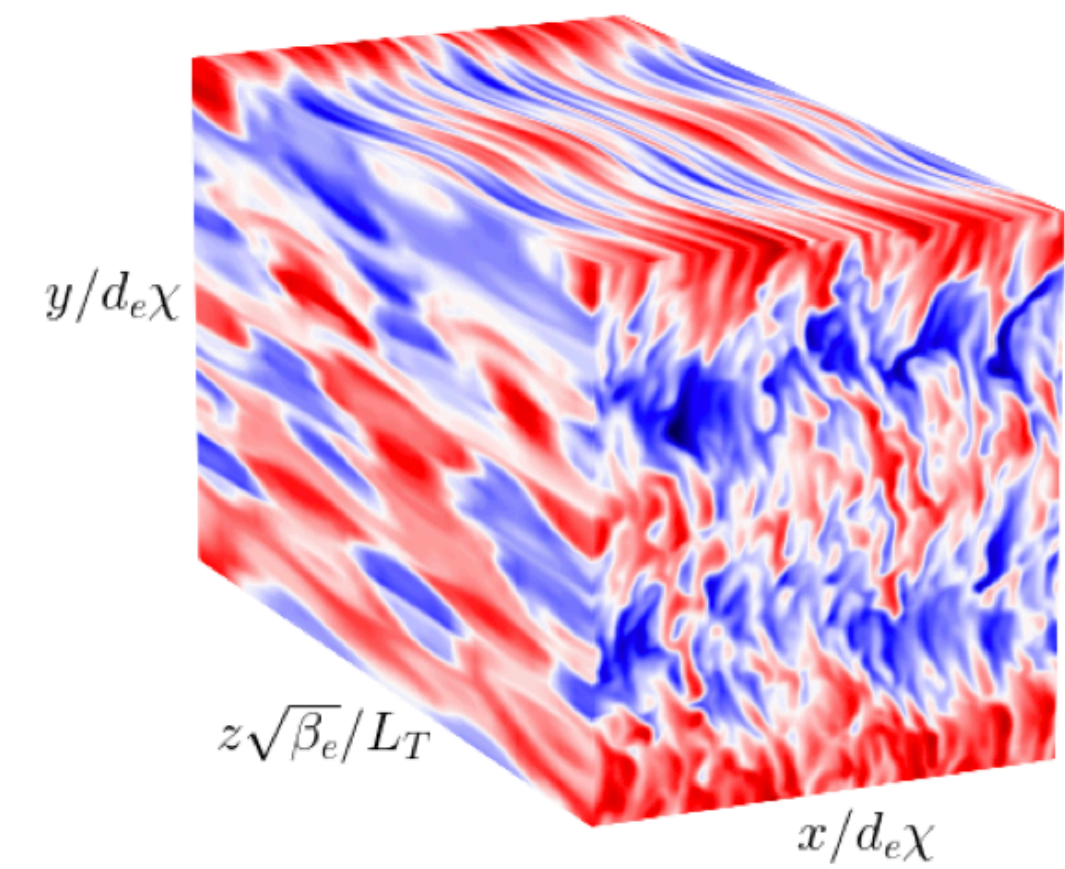


Per Helander

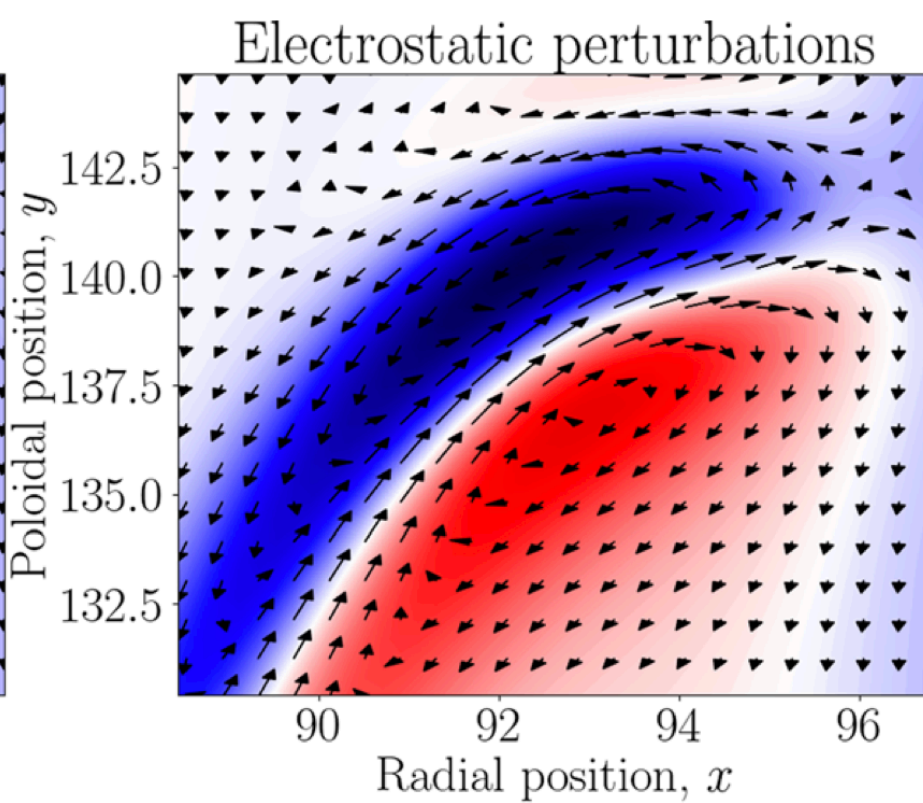
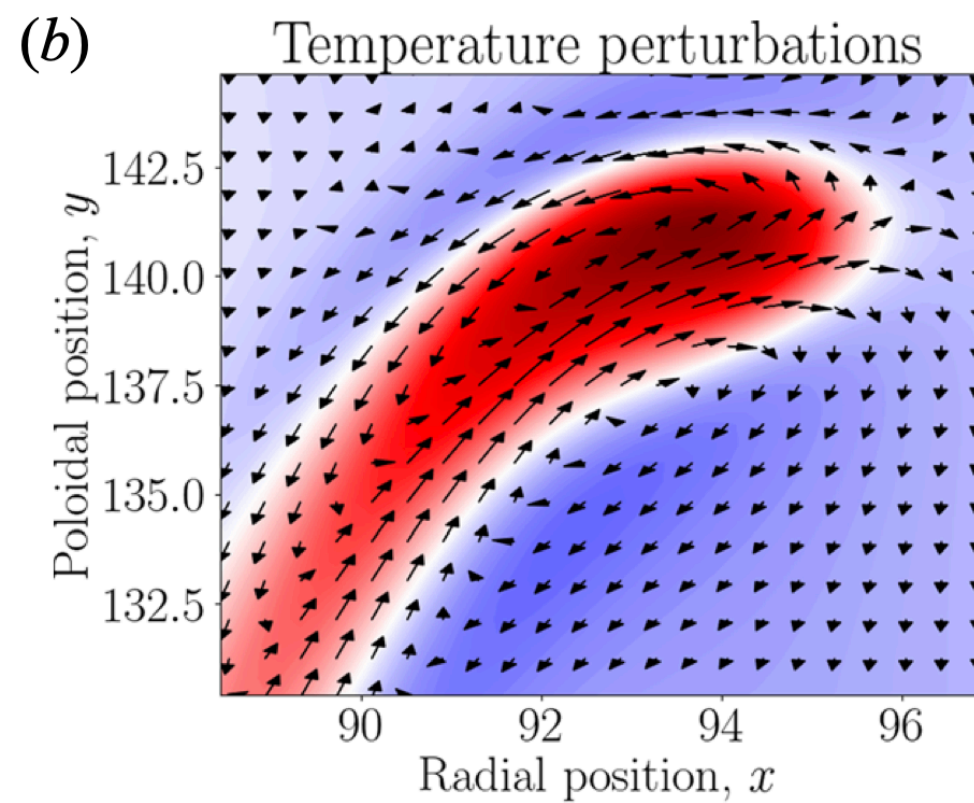


Thomas Foster

Equilibria and stability



Toby Adkins

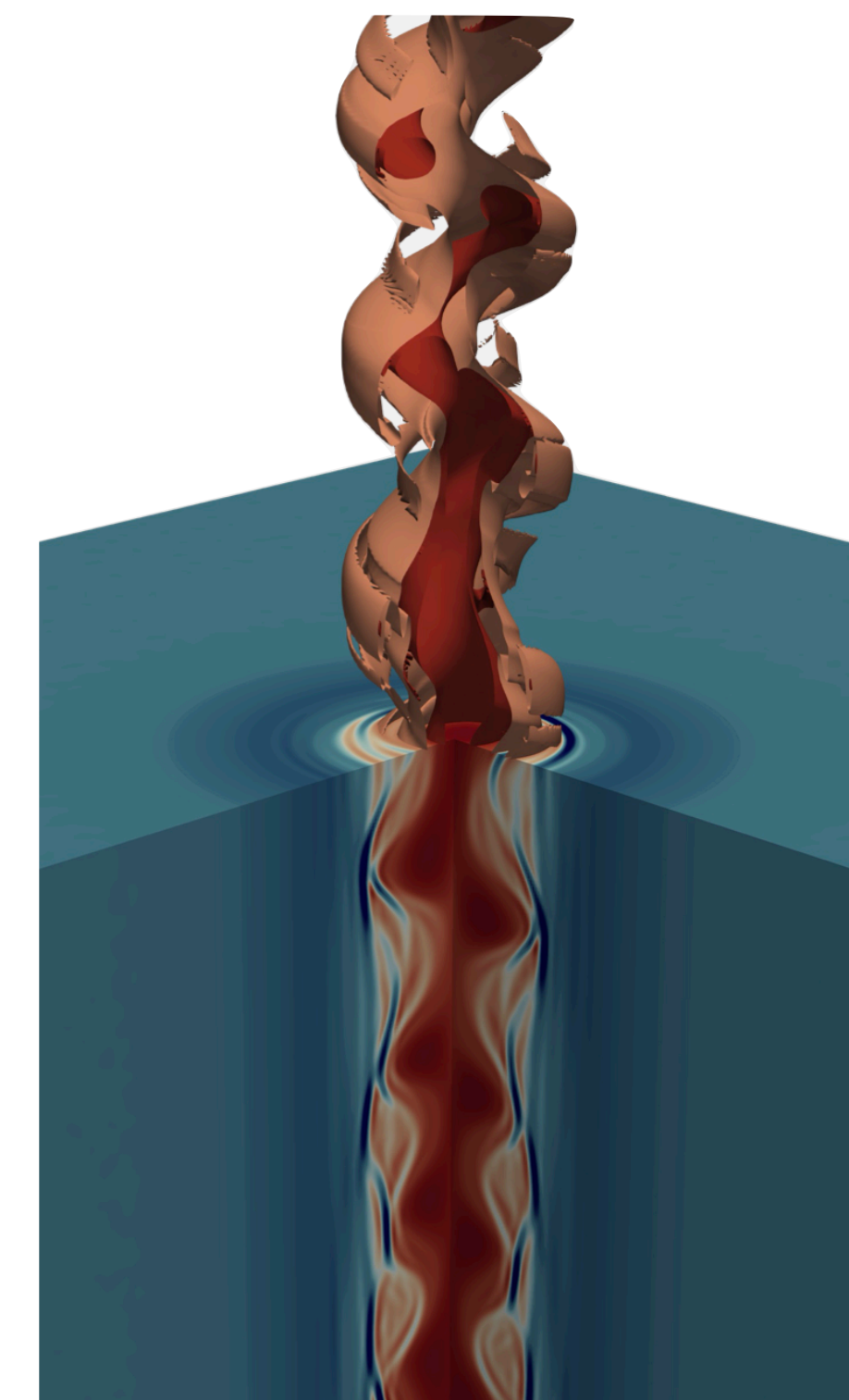


Plamen Ivanov

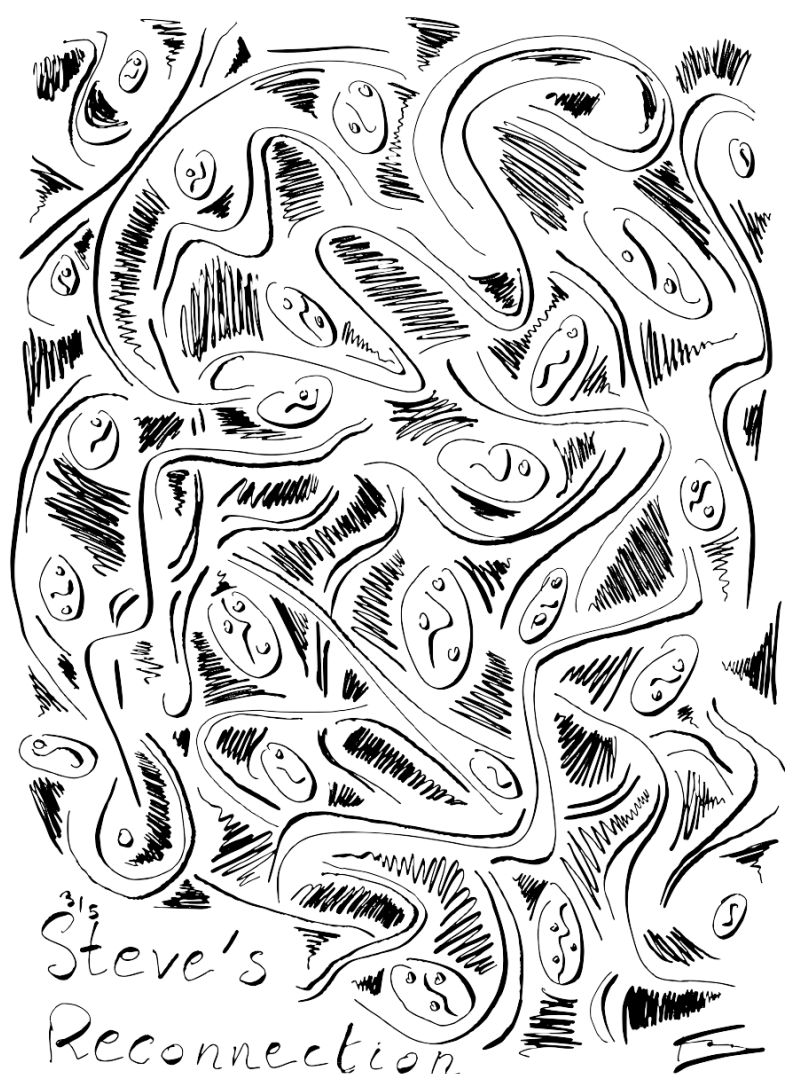


Nuno's
Reconnection

Nuno Loureiro

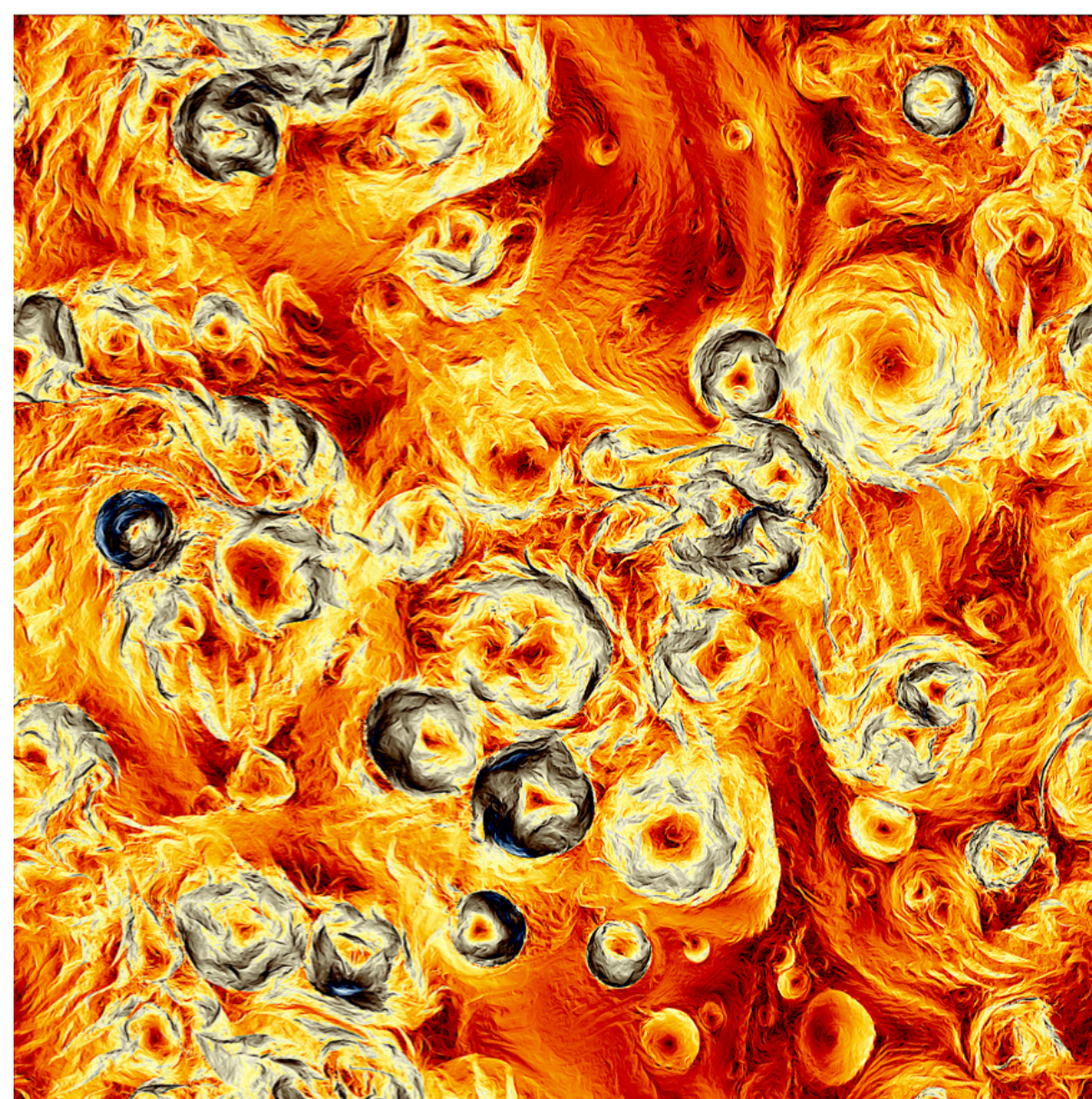


Alex Velberg

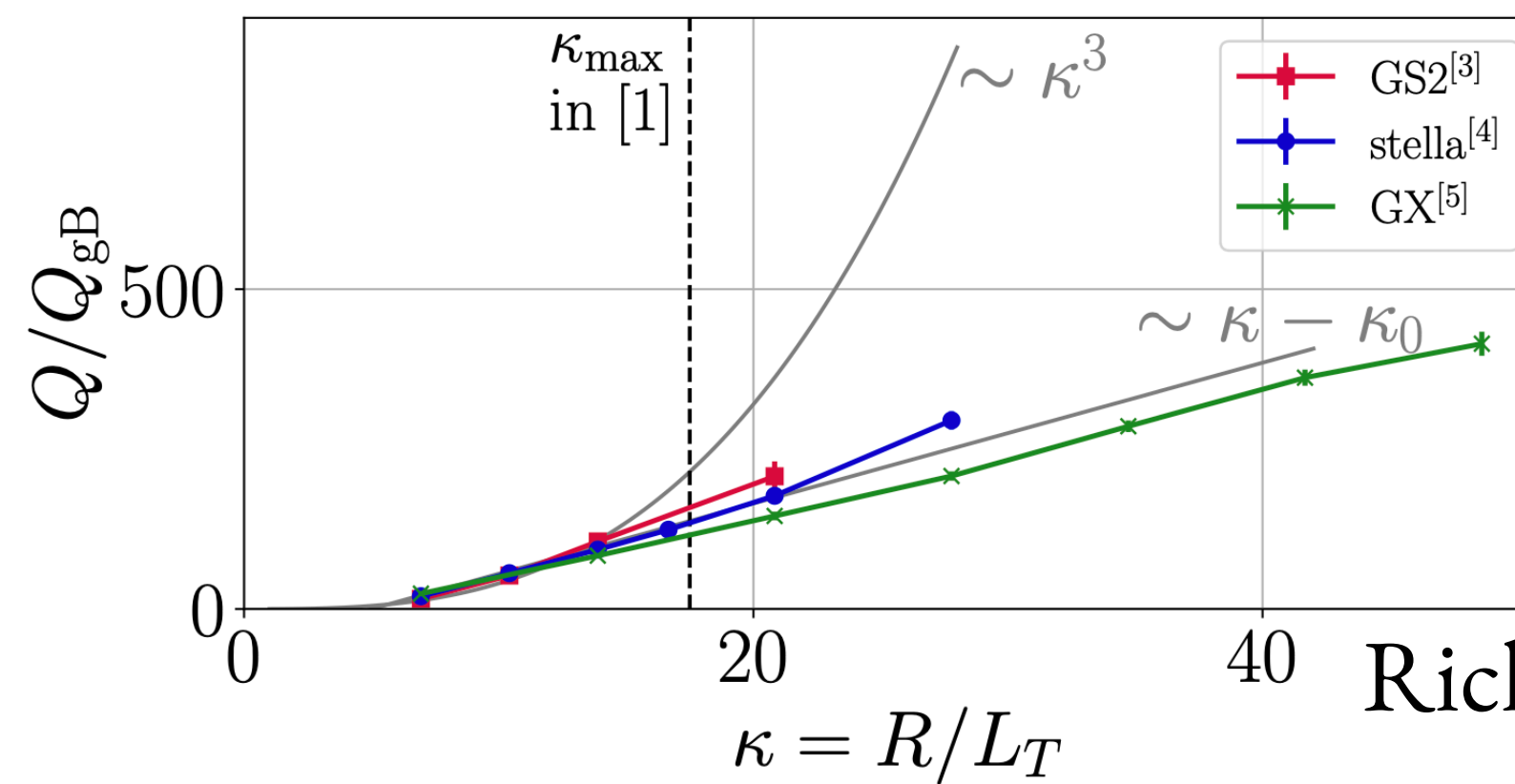


Steve's
Reconnection

Steve Cowley



Romain Meyrand

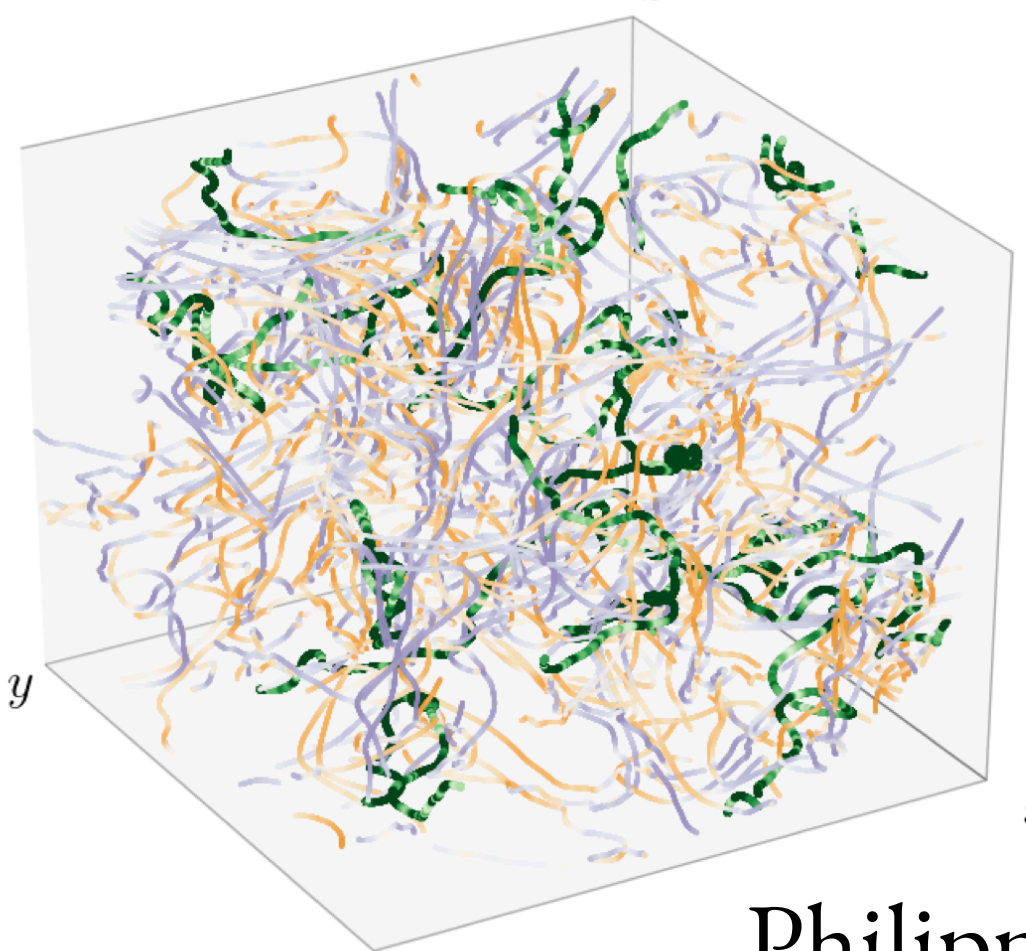


Richard Nies

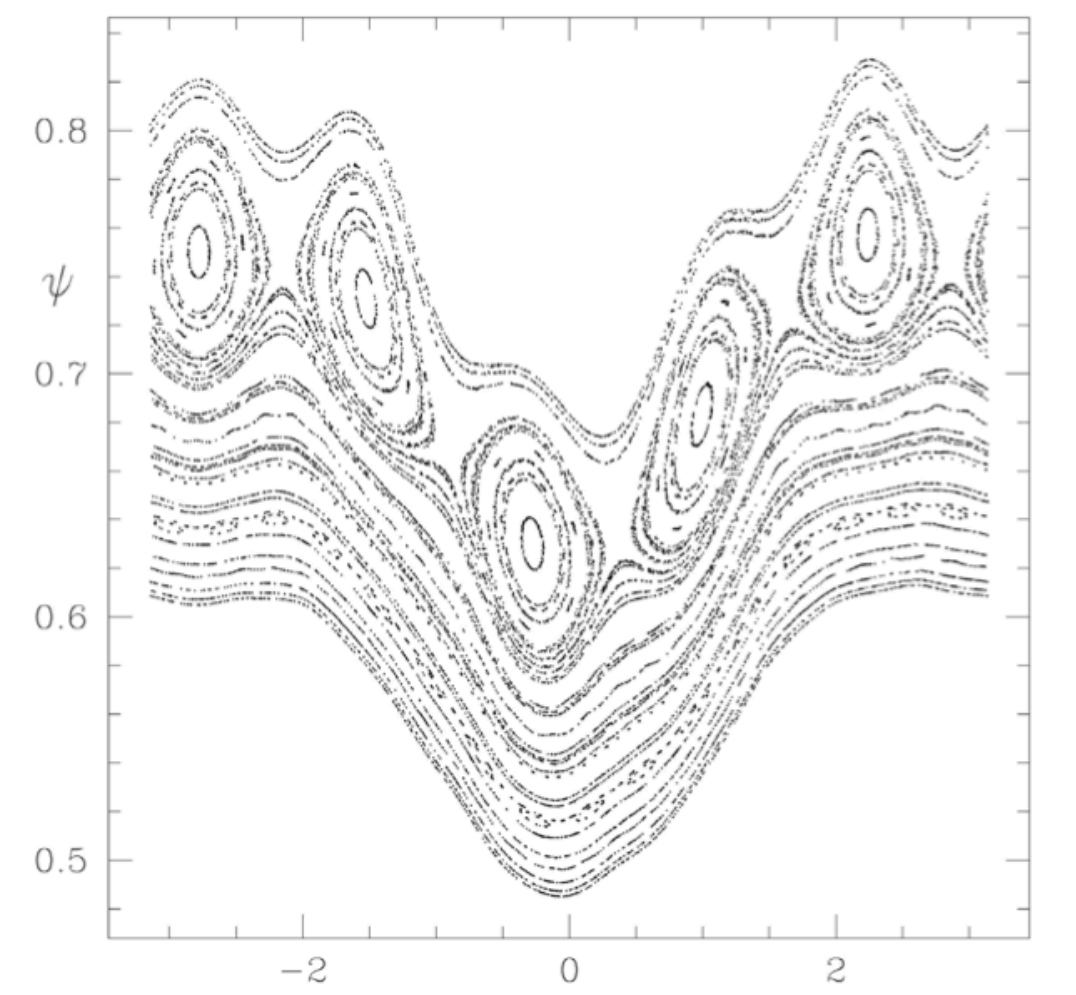
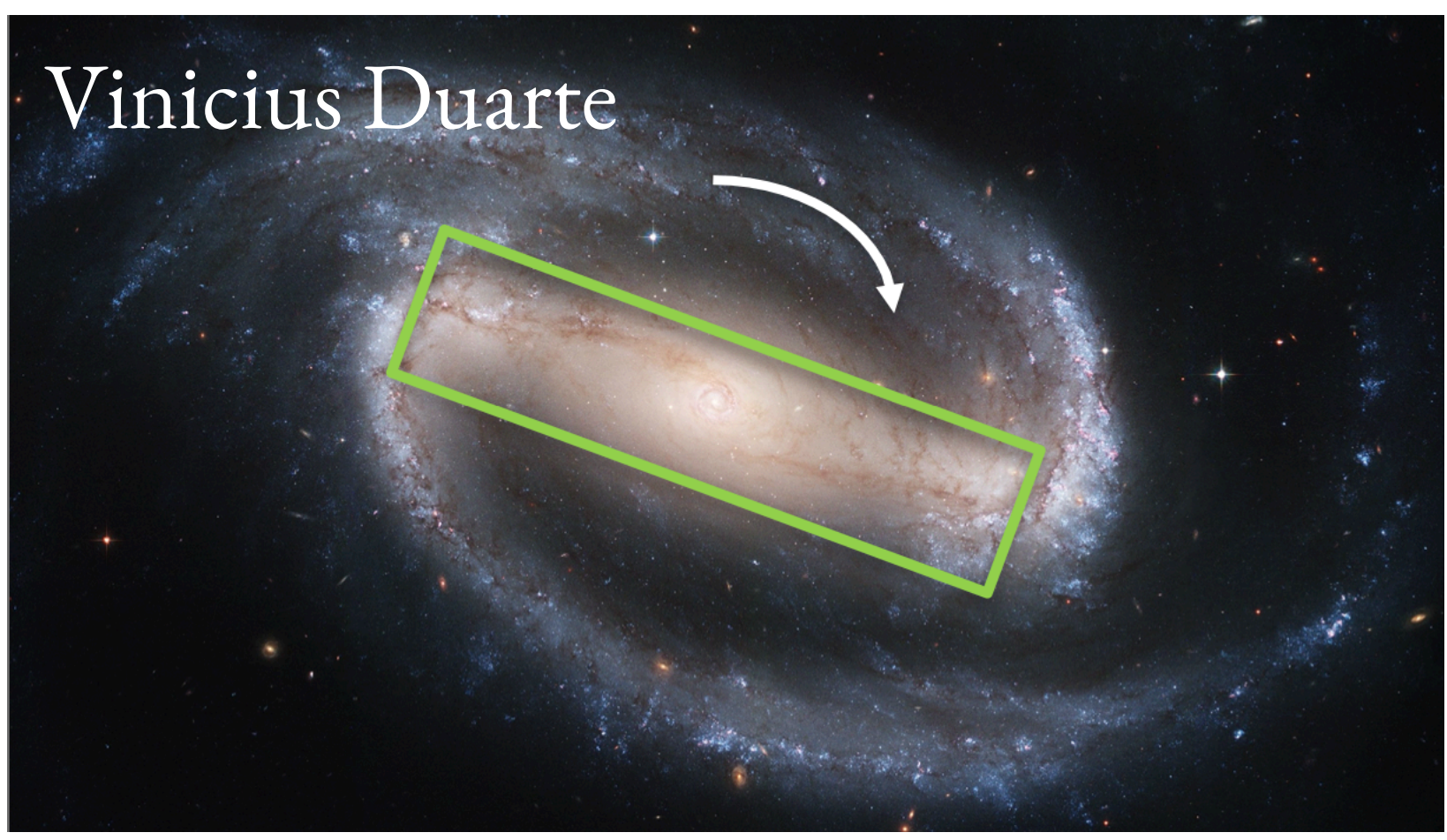
$$\Delta\beta \sim 1 \rightarrow \frac{\delta B_{\parallel}}{B_0} \sim \frac{\delta B_{\perp}}{B_0} \sim \frac{1}{\beta} \sim \Delta \sim \frac{\delta\rho}{\rho_0} \sim \frac{u_{\perp}}{v_A} \sim \frac{u_{\parallel}}{v_{th}} \sim \epsilon$$

Steve Majeski

Magnetic reconnection and turbulence



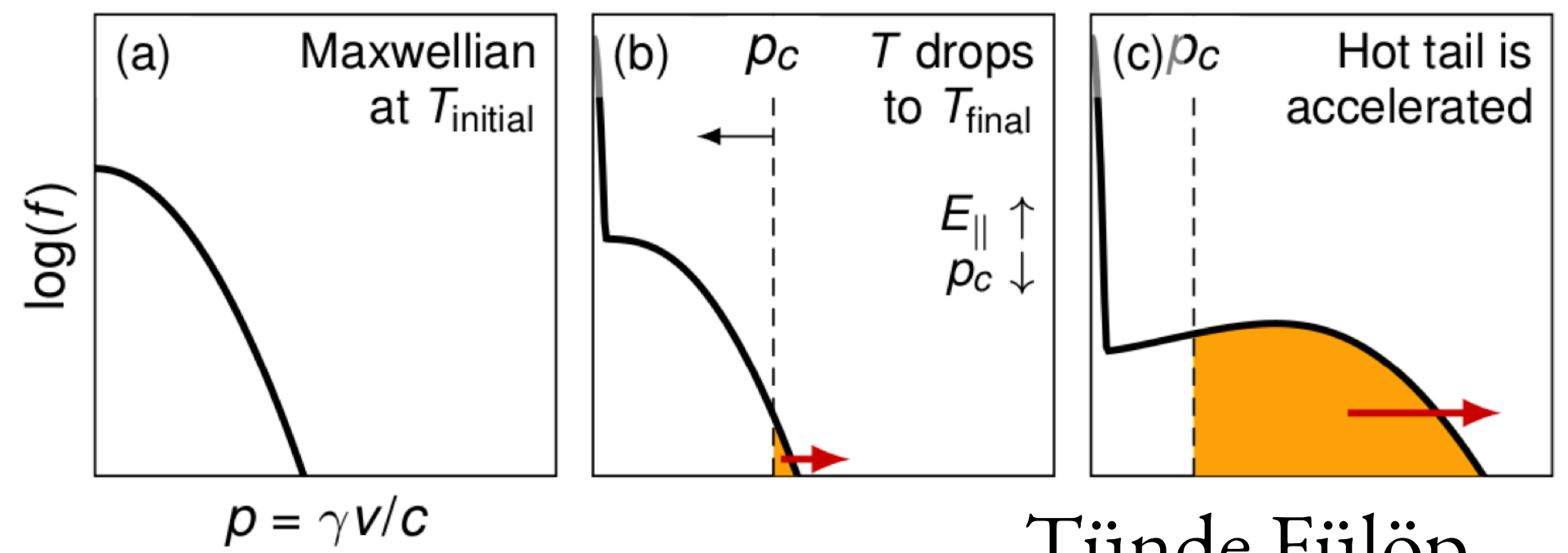
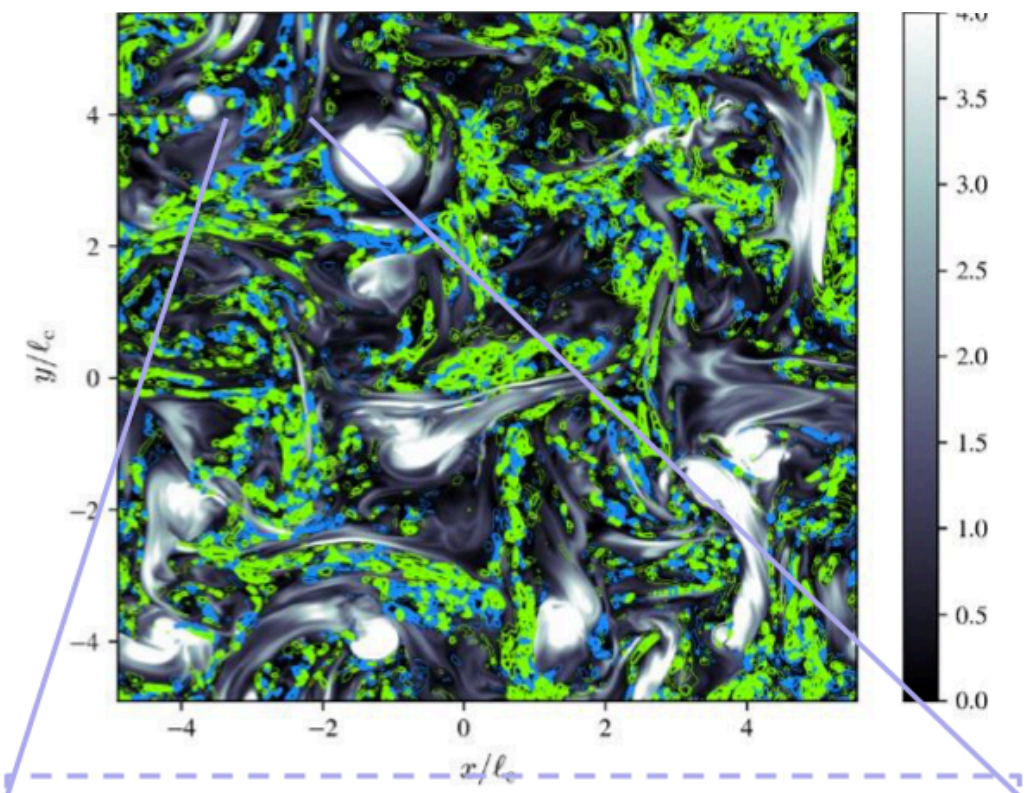
Patrick Reichherzer



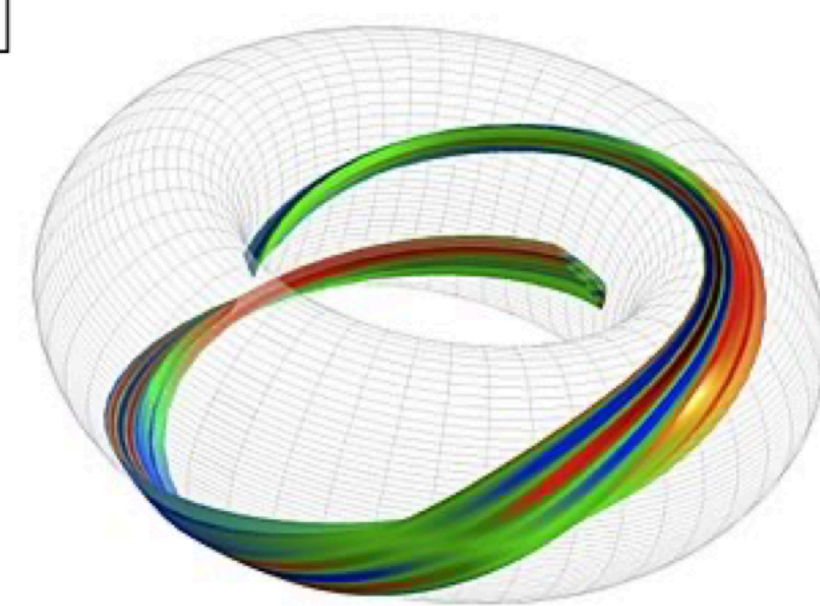
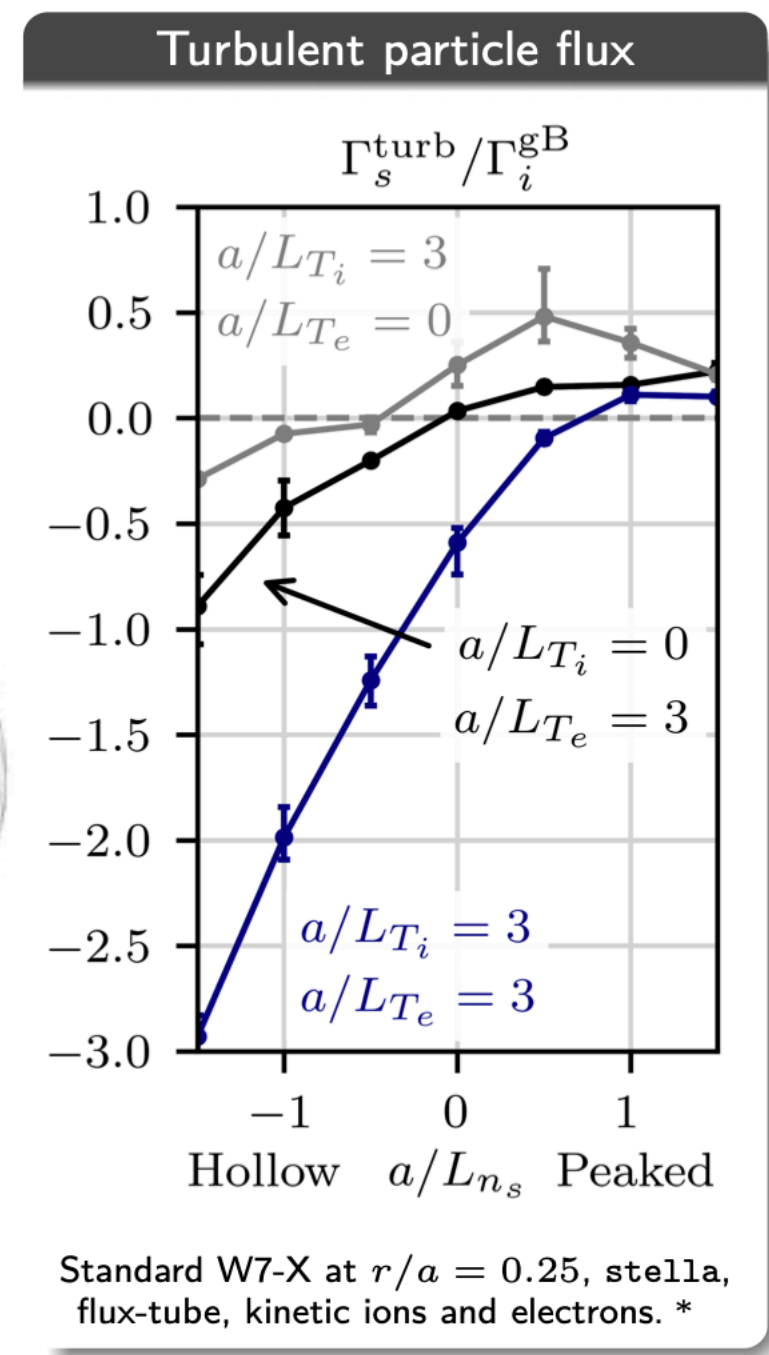
Thomas Foster

Philipp Kempskii

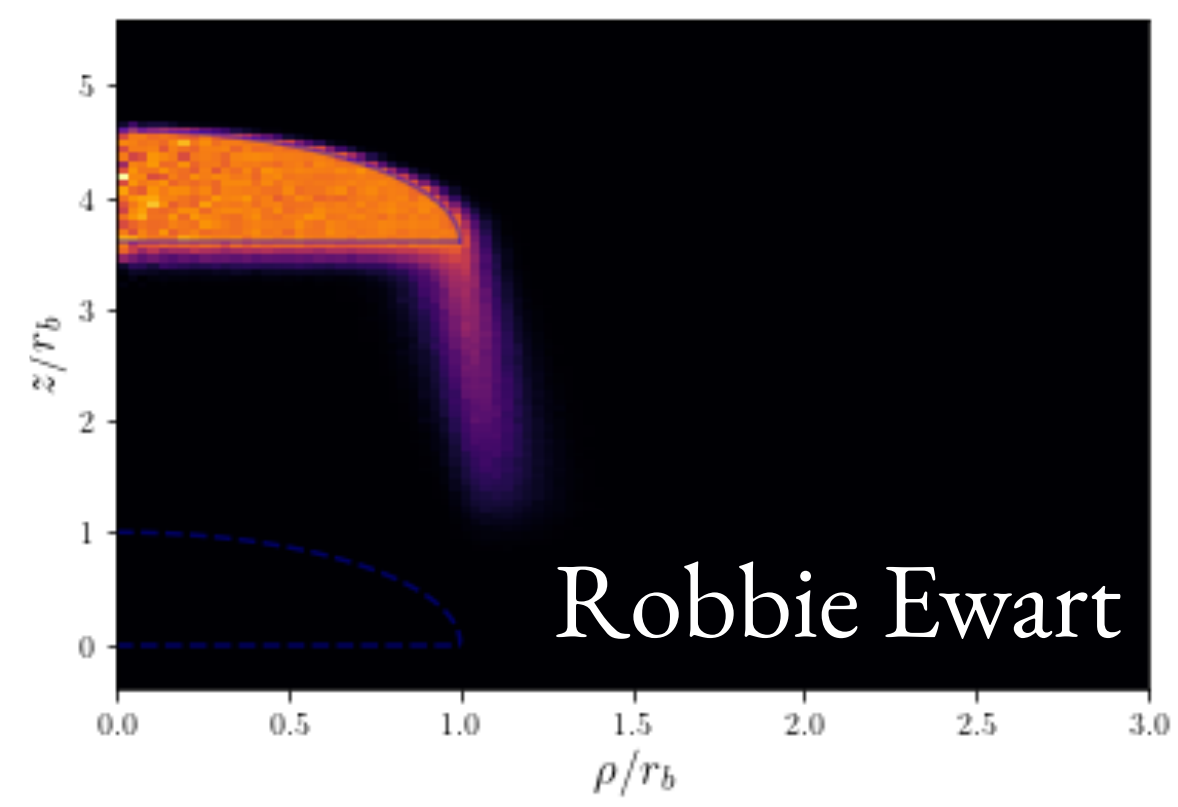
Martin Lemoine



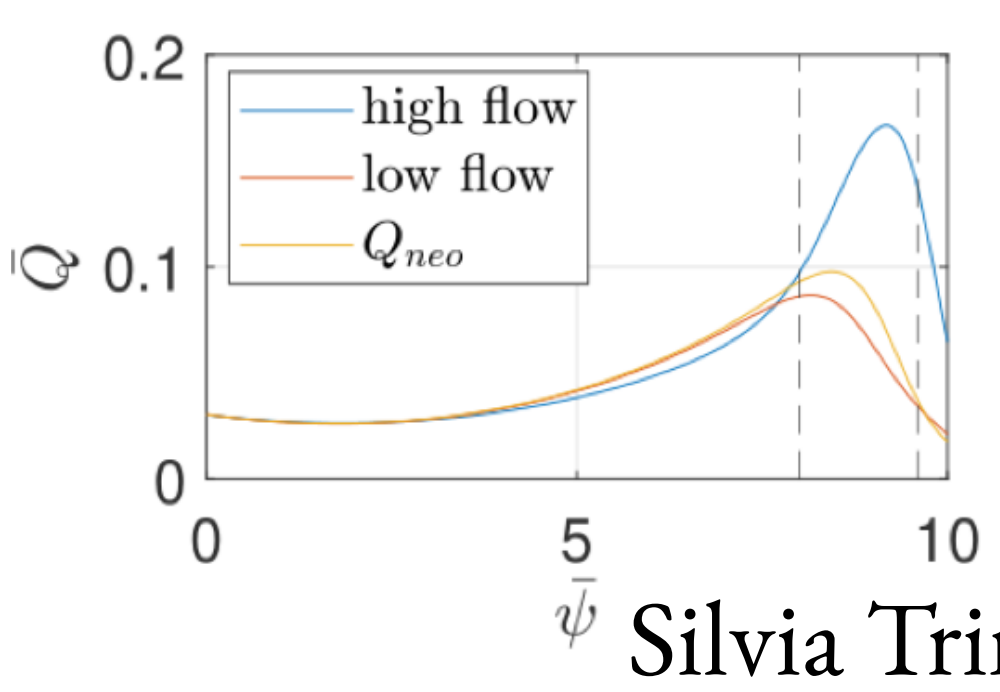
Tünde Fülöp



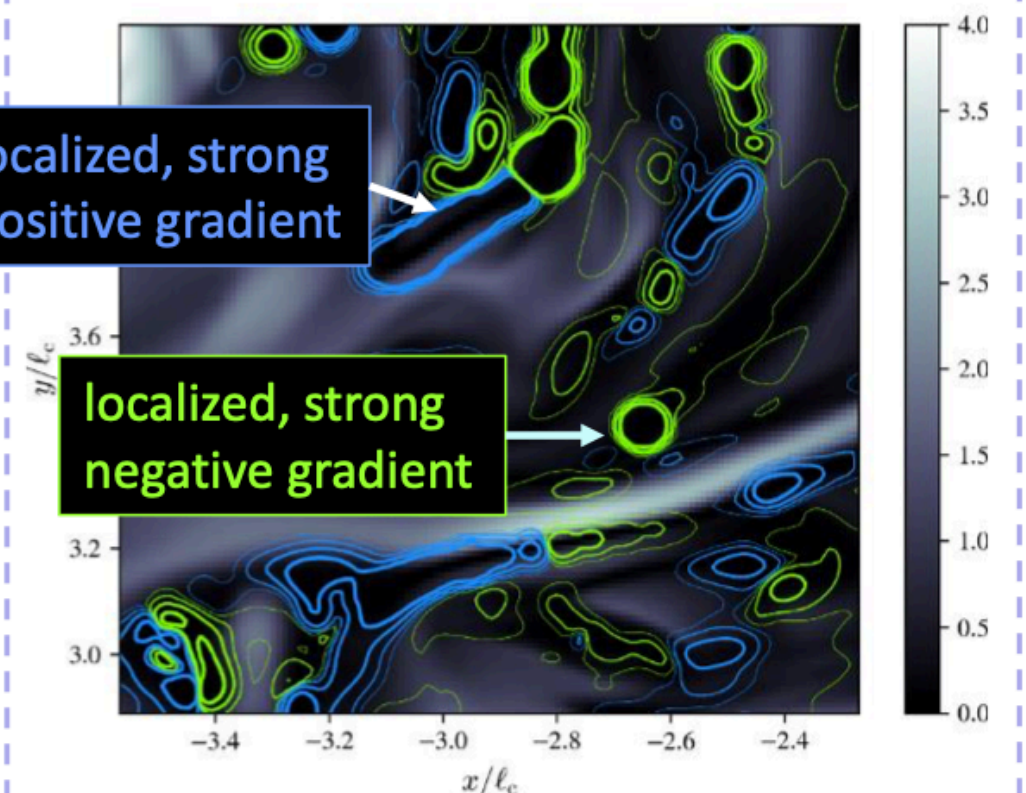
Georgia Acton



Robbie Ewart

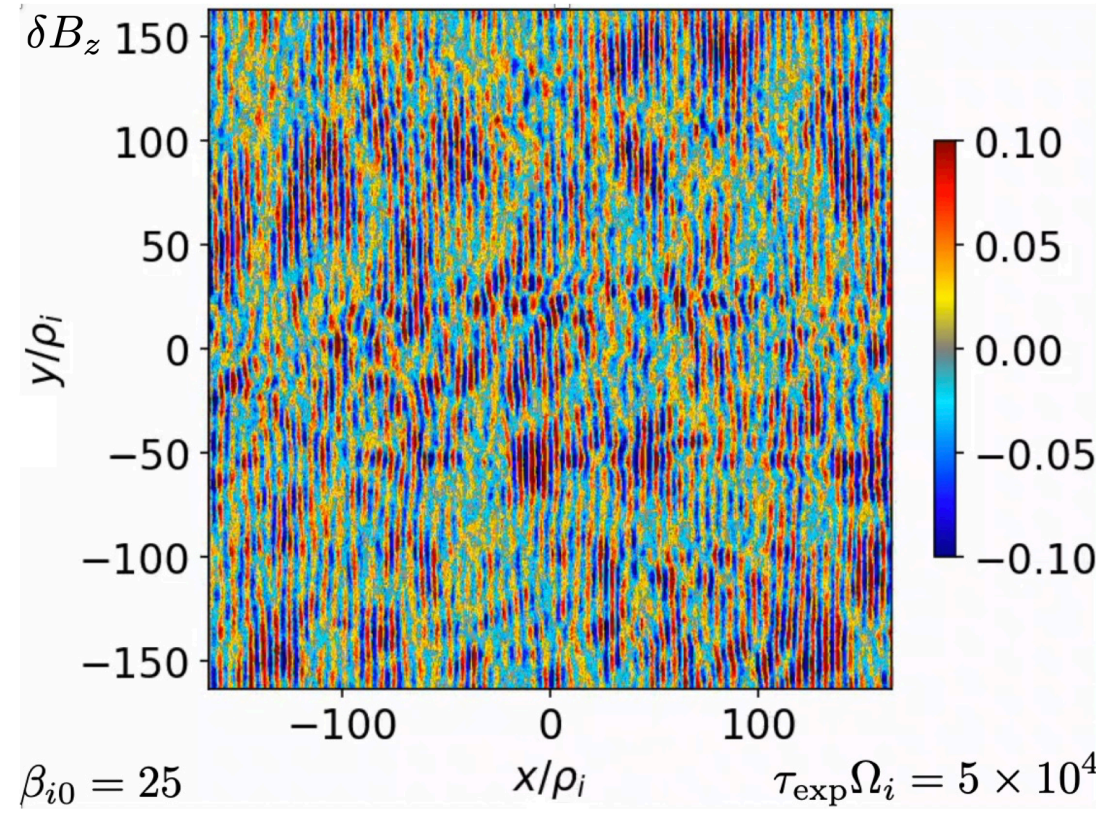


Silvia Trinczek

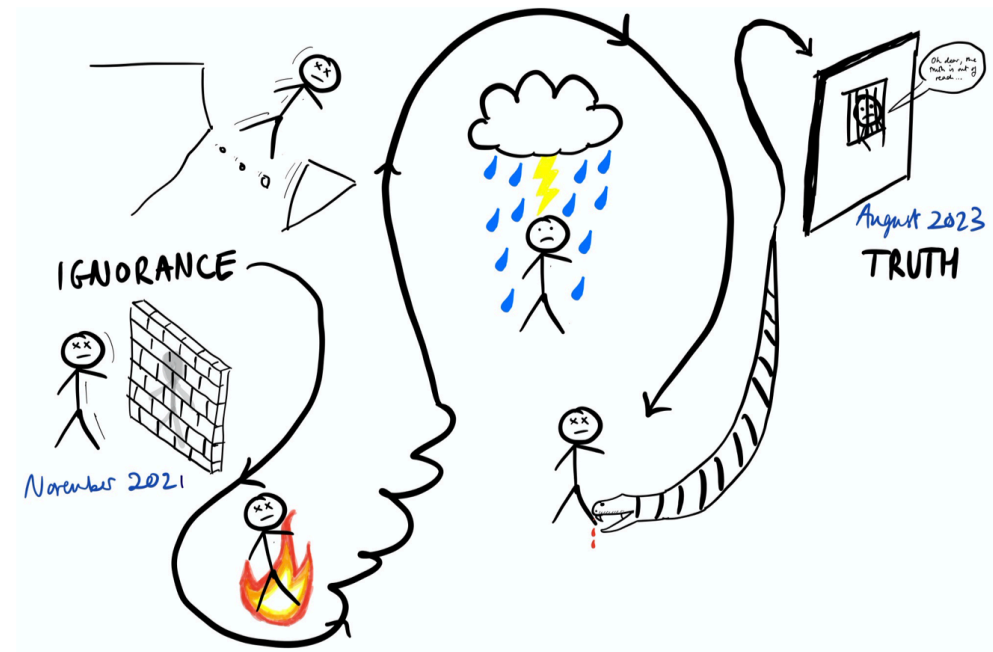


Transport — particles + heat

Hanne Thienpondt



Archie Bott



$$\mathfrak{e}_f[f] = \frac{1}{2} \frac{\partial}{\partial \xi} \left\{ (1 - \xi^2) \nu_{\text{eff,pl}}(v\xi) \left[\frac{\partial f_i}{\partial \xi} - \frac{2v v_{\text{wv,pl}}(v\xi)}{v_{\text{thi}}^2} f_M \right] + (1 - \xi^2) \nu_{\text{eff,ob}}(v\xi) \frac{\partial f_i}{\partial \xi} \right\}$$

Lemma

The gyro-moment model writes

$$\partial_t \tilde{U} + \sum_{i=1}^3 \partial_{x_i} (\tilde{A}_i(\mathbf{x}, t) \tilde{U}(\mathbf{x}, t)) = \tilde{B}(\mathbf{x}) \tilde{U}(\mathbf{x}, t) \quad (1)$$

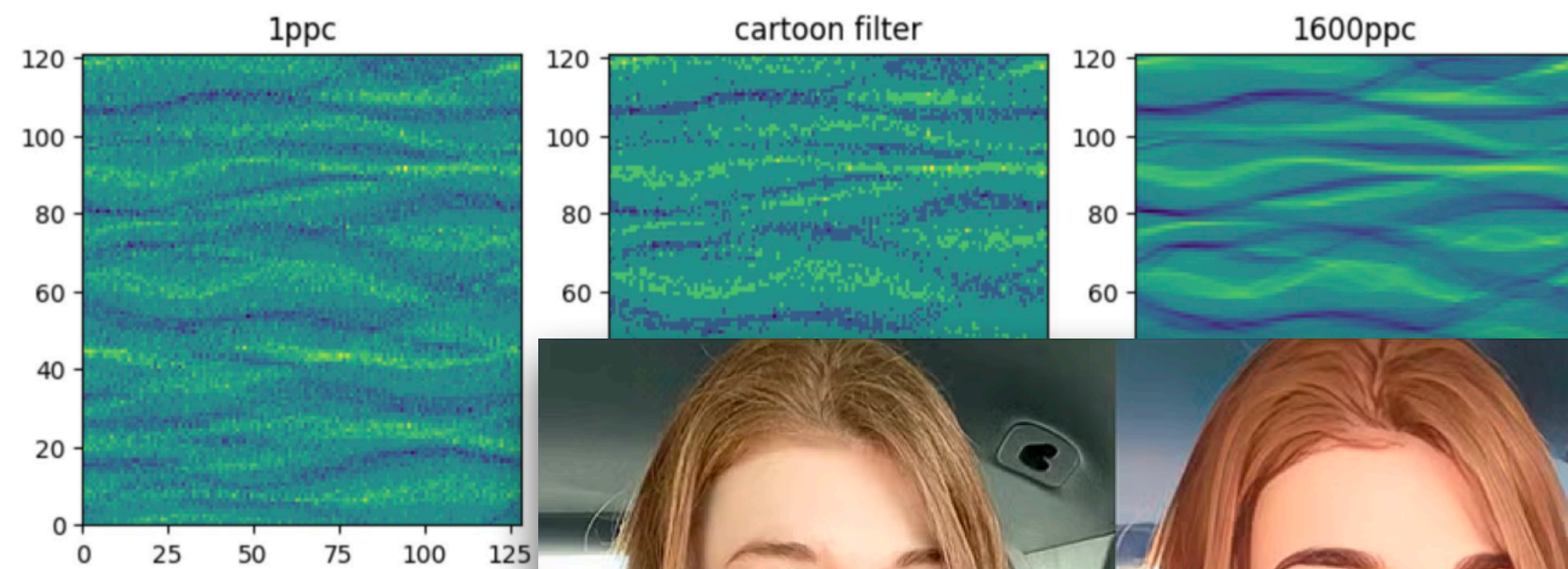
where

$$\tilde{A}_i(\mathbf{x}) = P^t A_i(\mathbf{x}) P, \quad \tilde{B}(\mathbf{x}) = P^t B(\mathbf{x}) P.$$

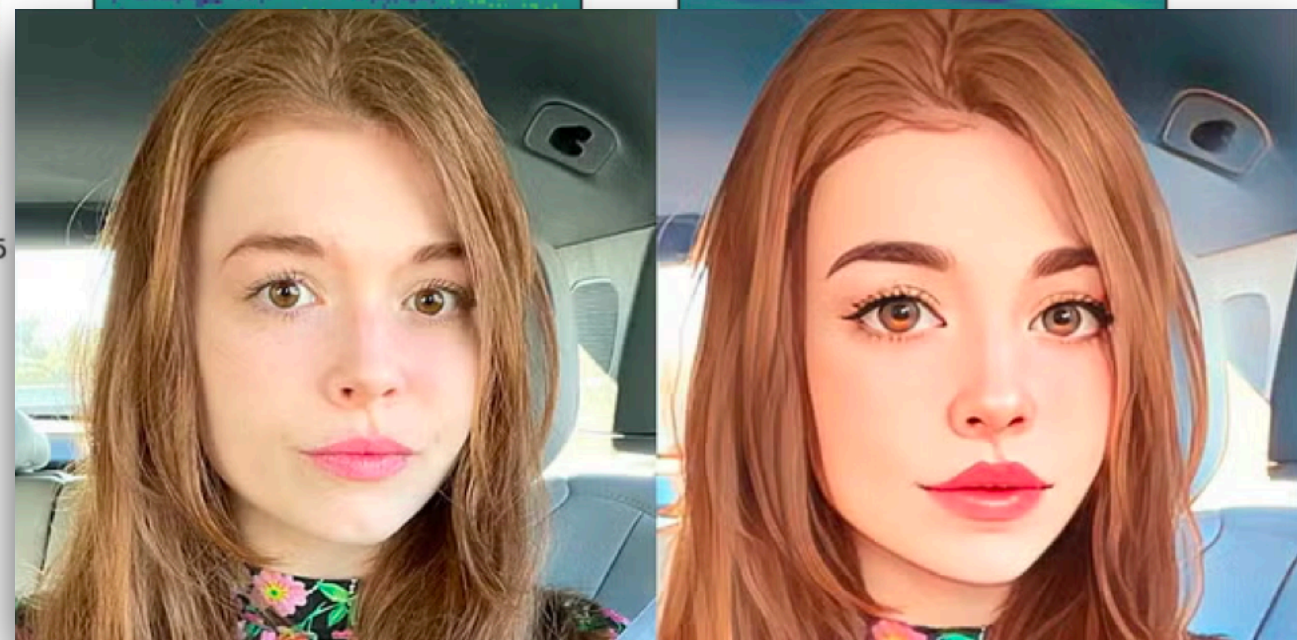
This numerical scheme works fine, but it is physically disgusting. Bruno Despres

Bruno Després

Nuno + Loureiro



Anatoly Spitkovsky



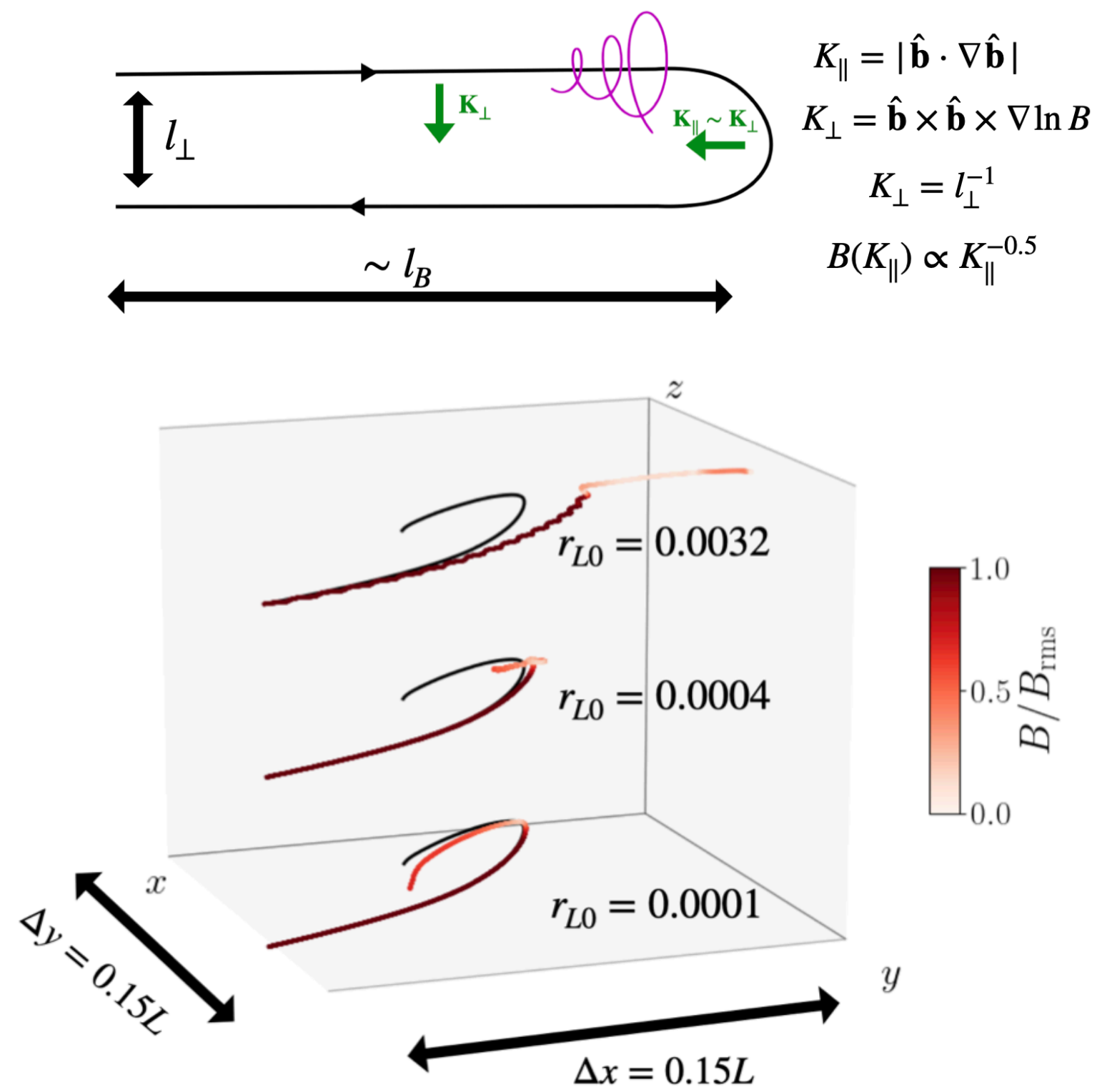
$$\Delta\beta \sim 1 \quad \rightarrow \quad \frac{\delta B_{\parallel}}{B_0} \sim \frac{\delta B_{\perp}}{B_0} \sim \frac{1}{\beta} \sim \Delta \sim \frac{\delta\rho}{\rho_0} \sim \frac{u_{\perp}}{v_A} \sim \frac{u_{\parallel}}{v_{th}} \sim \epsilon$$

Ordering 1D CGL-MHD...

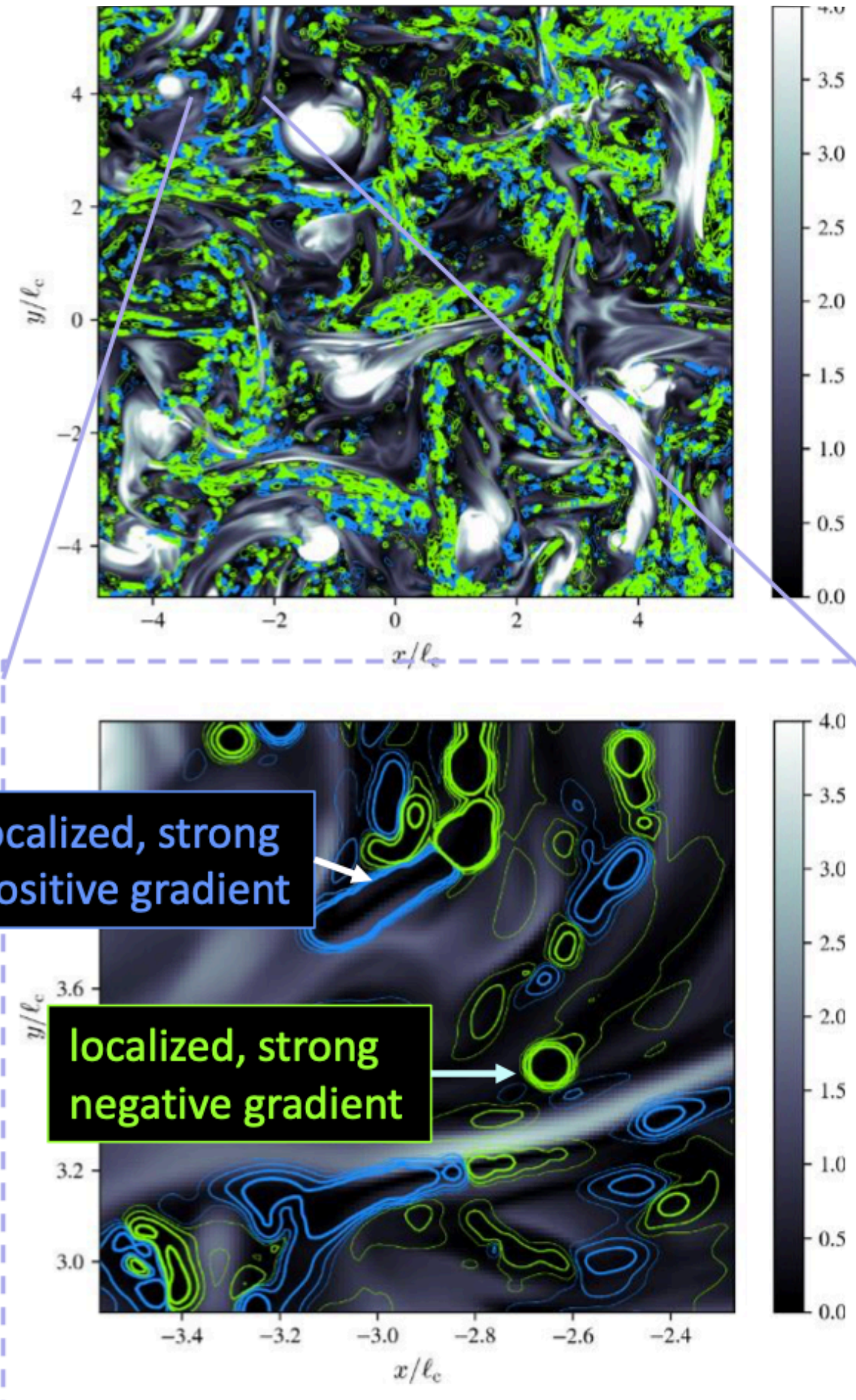
$$\frac{\partial z^{\pm}}{\partial t} \mp v_A \frac{\partial z^{\pm}}{\partial x} = v_A \frac{\beta}{4} \frac{\partial}{\partial x} [(z^+ - z^-) \Delta] \quad + \quad \text{Linear ion acoustic waves}$$

Steve Majeski

Subgrid models and closures

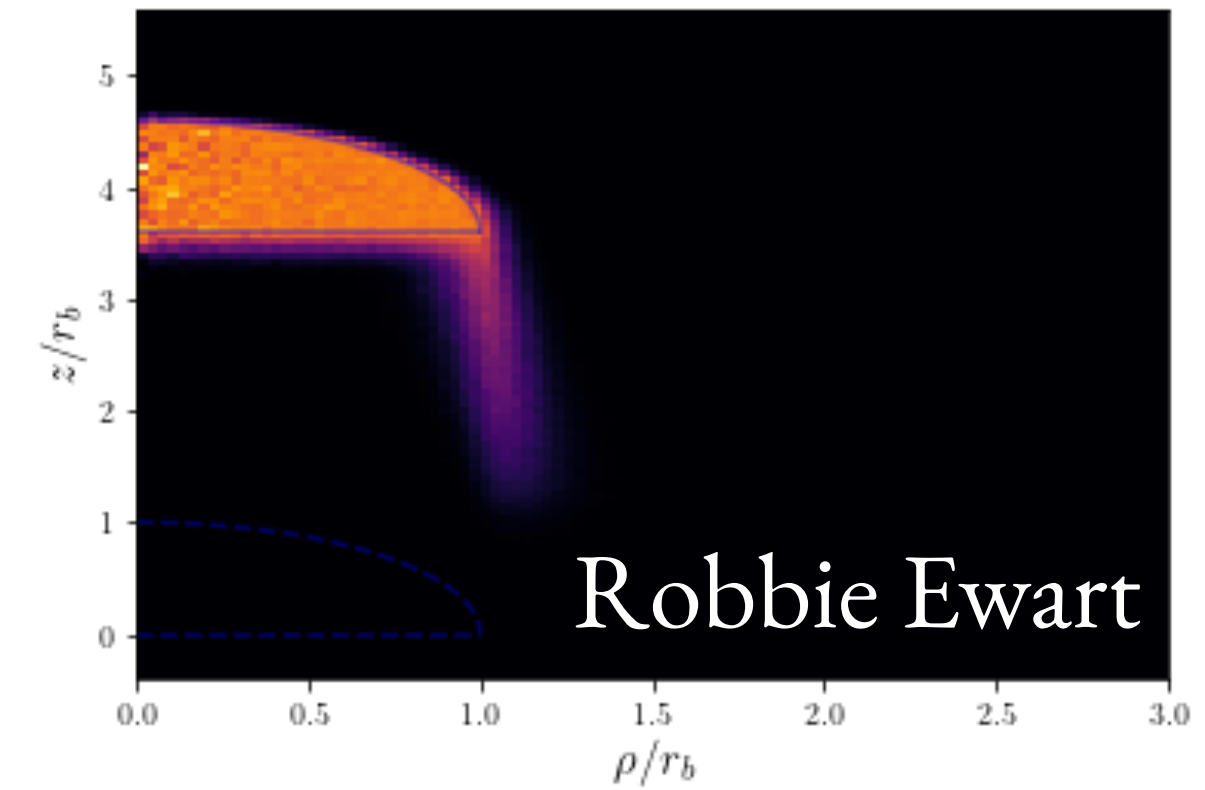


Philipp Kempskiii



Martin Lemoine

Nuno + Loureiro



Robbie Ewart



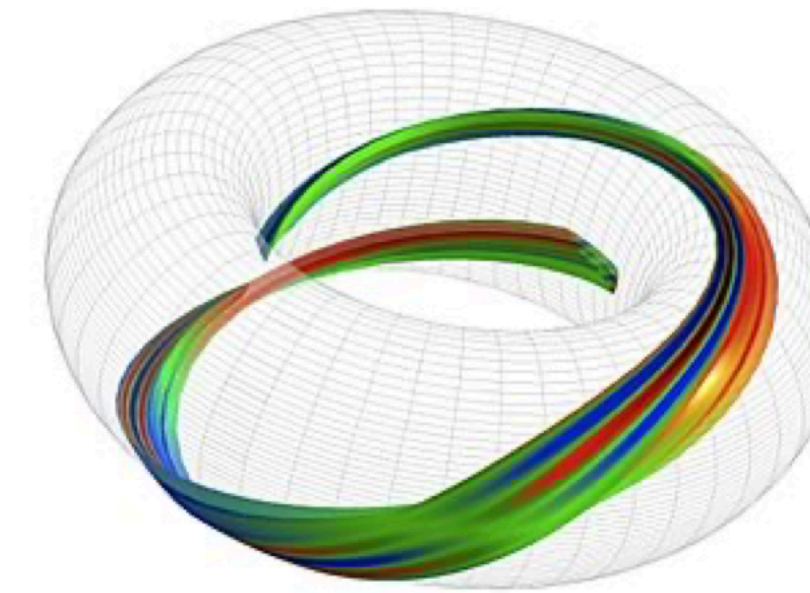
Patrick Reichherzer

Intermittent structures

Where there is geometry there are problems

Geometry is at the root of all evil...

$$\frac{\partial g}{\partial t} = \underbrace{(\text{geometric factors})}_{\text{e.g. } \hat{\mathbf{b}} \cdot \nabla z} \cdot (\nabla g + \nabla \underbrace{\langle \phi \rangle_{\mathbf{R}}}_{J_{0,k} \hat{\phi}_k})$$

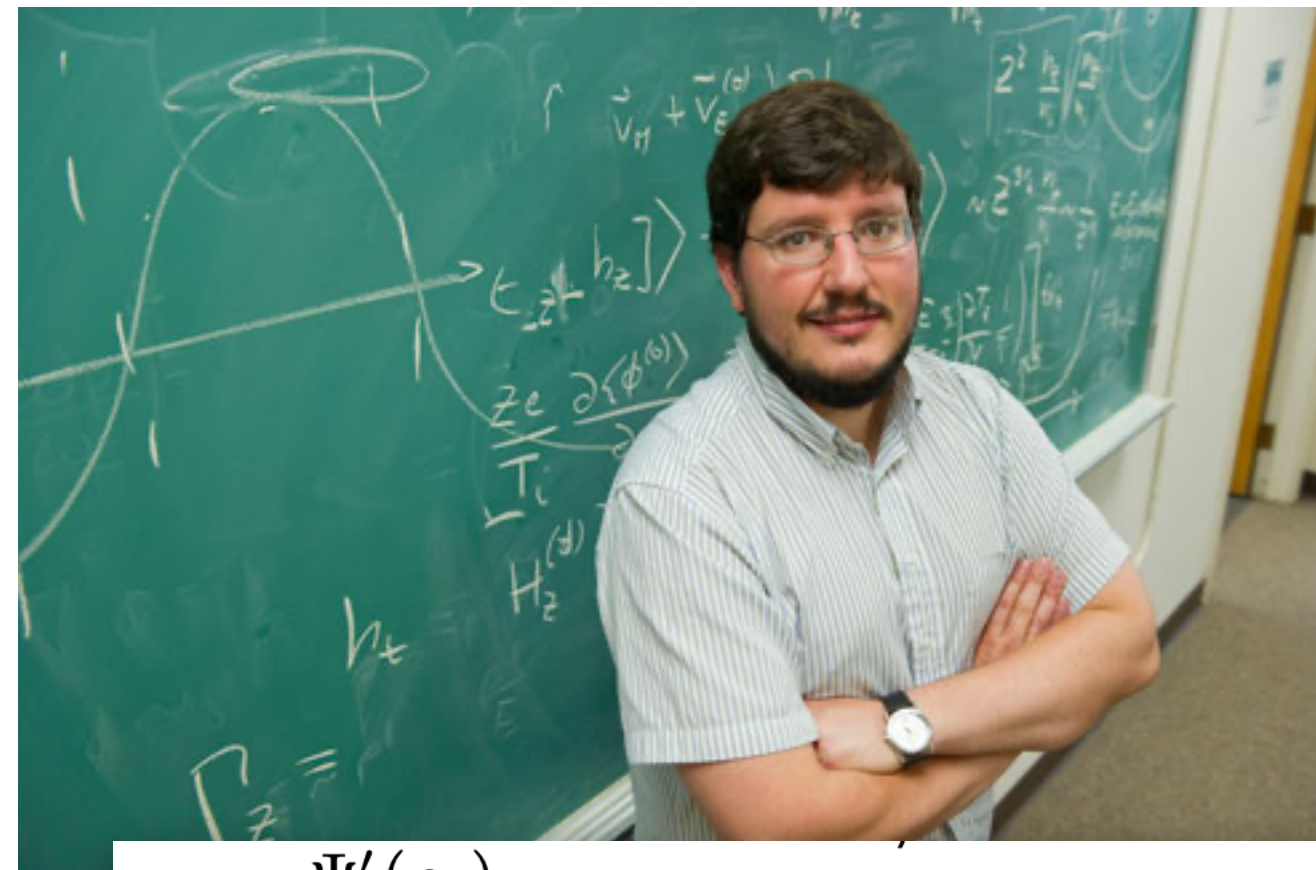


Georgia Acton

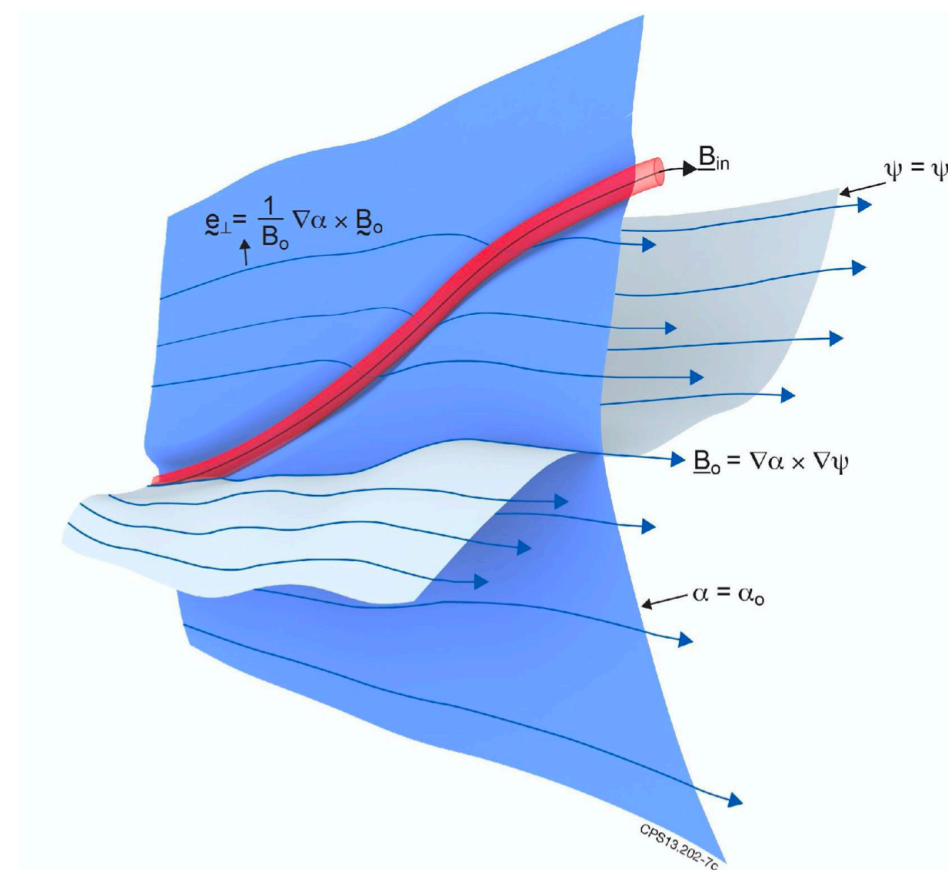
Do we really need the torus? Lessons learned from the humble slab

- ▶ It appears that the TAI instability mechanism appears to survive the transition to toroidicity. **Another win for the slab?**

Toby Adkins

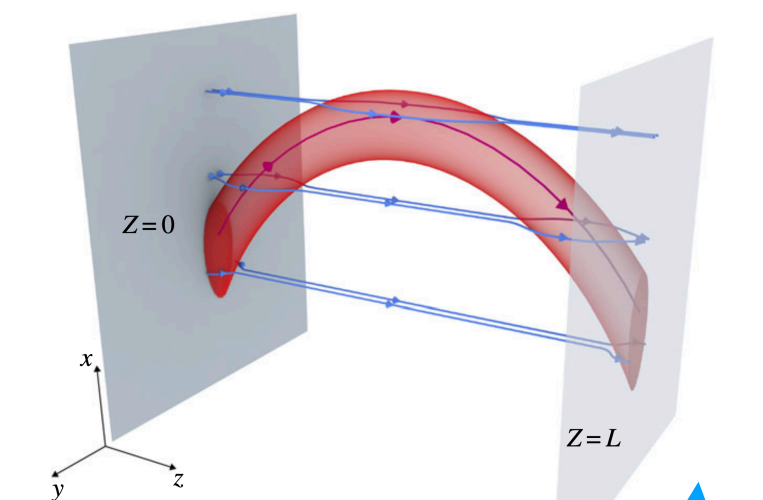


Felix Parra

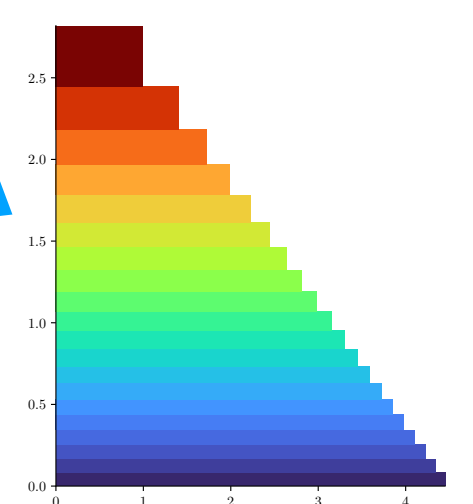


Steve Cowley

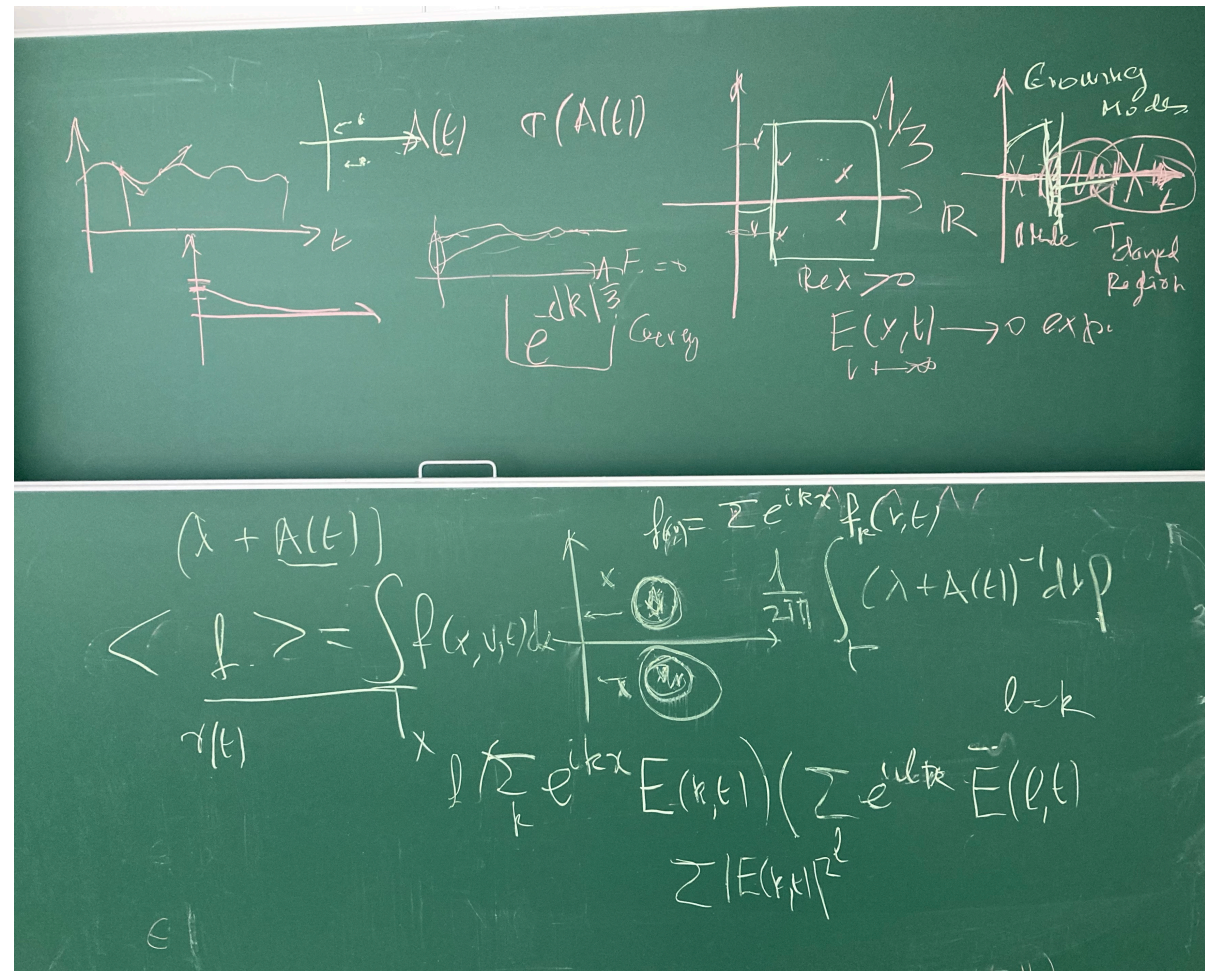
$$\mathbf{B}'_a = \frac{\Psi'_t(\rho_a)}{\mathcal{J}_a} [\partial_u \nu'_a \partial_v \mathbf{x}_a + (l'(\rho_a) - \partial_v \nu'_a) \partial_u \mathbf{x}_a] + \left(\frac{\Psi''_t(\rho_a)}{\Psi'_t(\rho_a)} - \frac{\mathcal{J}'_a}{\mathcal{J}_a} \right) \mathbf{B}_a + \mathbf{B}_a \cdot \nabla_S \left(\frac{\mathcal{J}_a}{|\partial_u \mathbf{x}_a \times \partial_v \mathbf{x}_a|} \hat{\mathbf{n}}_a \right)$$



David Hosking



Geometry and whether it matters



Claude Bardos

Summary of rational surface conditions

- Hamada condition: $P' \frac{d}{d\alpha} \left(\oint \frac{dl}{B} \right) = 0$

- New condition

$$\frac{d}{d\alpha} \left[\left(\oint \frac{B}{|\nabla\rho|^2} dl \right)^{-1} \left(\frac{c\Psi'_t}{4\pi} \oint dv - \oint \frac{J_{PS}}{|\nabla\rho|^2} dl + \frac{c}{4\pi} \oint \frac{[\mathbf{B} \cdot \nabla \hat{\mathbf{n}} \cdot (\mathbf{B} \times \hat{\mathbf{n}}) + (\mathbf{B} \times \hat{\mathbf{n}}) \cdot \nabla \hat{\mathbf{n}} \cdot \mathbf{B}]}{B|\nabla\rho|^2} dl \right) \right] = 0$$

Felix Parra

Lemma

The gyro-moment model writes

$$\partial_t \tilde{U} + \sum_{i=1}^3 \partial_{x_i} (\tilde{A}_i(\mathbf{x}, t) \tilde{U}(\mathbf{x}, t)) = \tilde{B}(\mathbf{x}) \tilde{U}(\mathbf{x}, t) \quad (1)$$

where

$$\tilde{A}_i(\mathbf{x}) = P^t A_i(\mathbf{x}) P, \quad \tilde{B}(\mathbf{x}) = P^t B(\mathbf{x}) P.$$

Bruno Després

$$\tilde{h}_s^{\text{even}} = \lambda^2 h_s^{\text{even}}(x/\lambda^2, y/\lambda^2, z/\lambda^{2/\alpha}, t/\lambda^2),$$

$$\tilde{h}_s^{\text{odd}} = \lambda^{2/\alpha} h_s^{\text{odd}}(x/\lambda^2, y/\lambda^2, z/\lambda^{2/\alpha}, t/\lambda^2),$$

$$\tilde{\phi} = \lambda^2 \phi(x/\lambda^2, y/\lambda^2, z/\lambda^{2/\alpha}, t/\lambda^2),$$

Toby Adkins

Mathematics and plasma physics

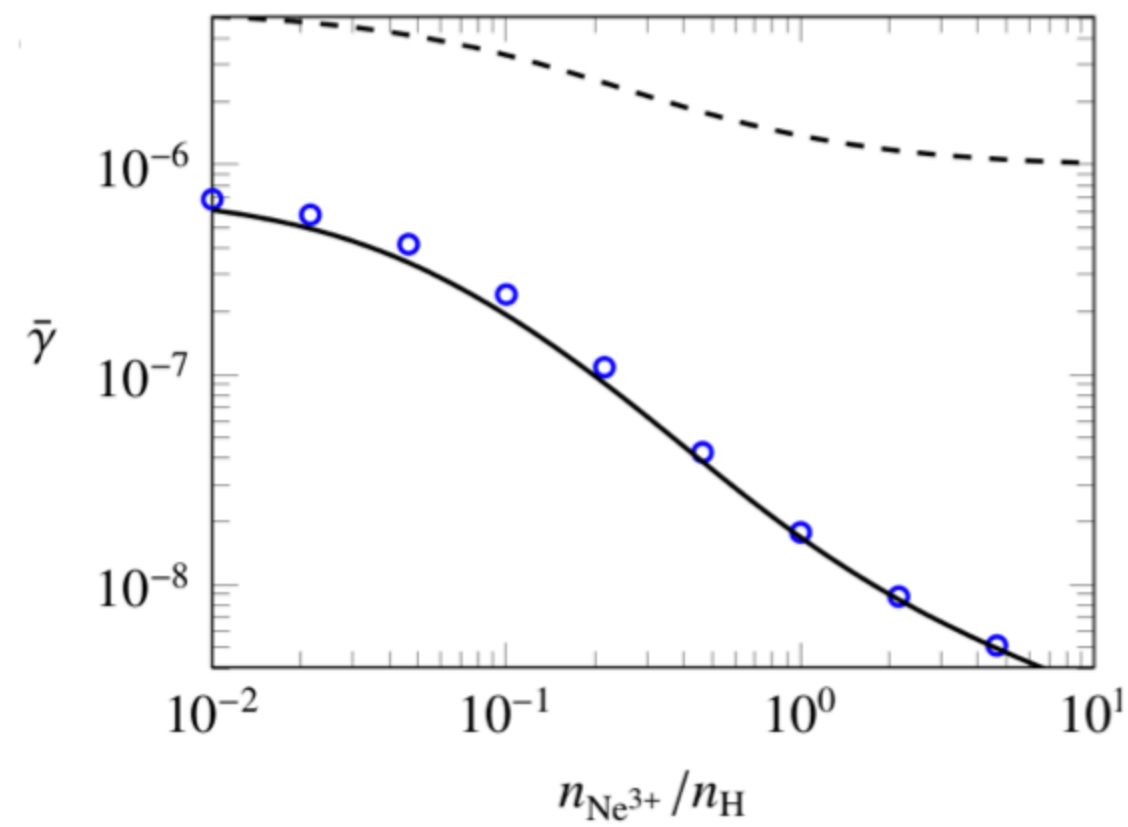
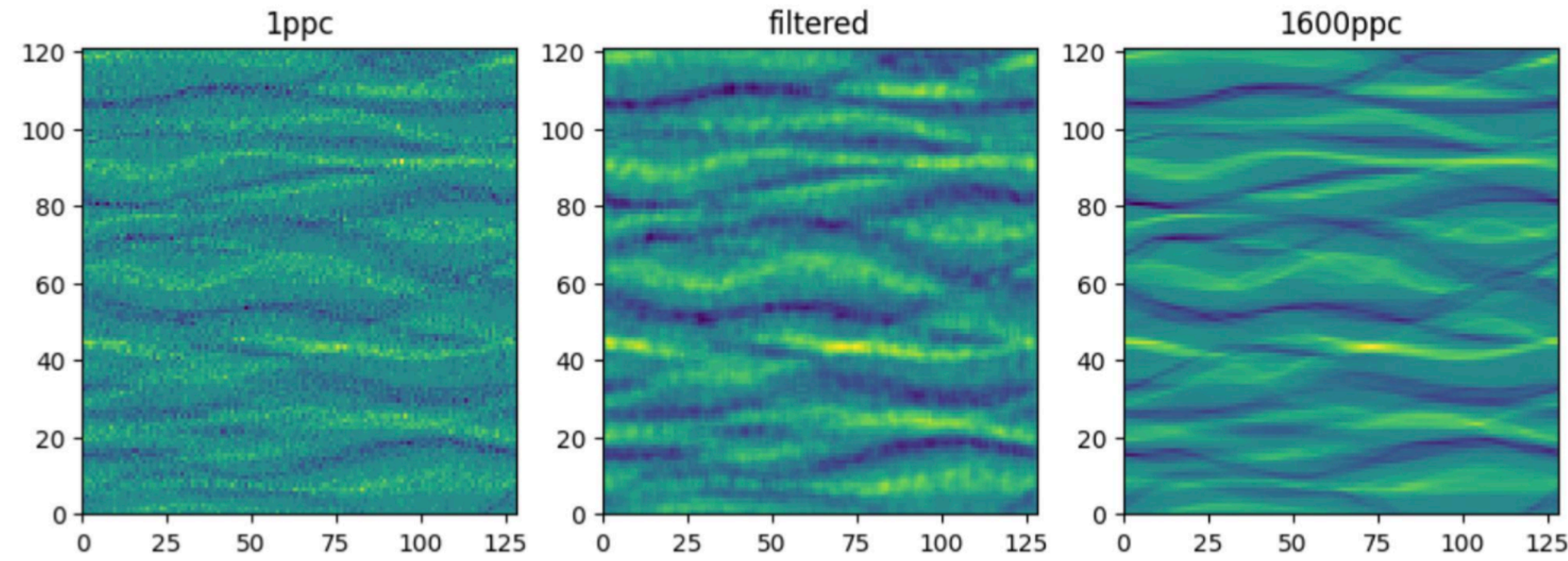


Figure shows the Dreicer generation rate in the presence of Ne^{3+} , obtained by neural network trained on kinetic simulations (solid), kinetic simulations (blue circles) and the Connor-Hastie formula (dashed)

Tünde Fülöp

Micromirrors mediating multiscale motions in magnetised megastructures

Patrick Reichherzer



Anatoly Spitkovsky

Per Helander stepped up with rigorous bounds,
Gyrokinetic instabilities, no room for playgrounds.
The growth rates constrained, like children on a leash,
The plasma particles danced, within limits, they'd reach,
And the audience clapped, with their hands making sounds.

Matt Kunz



François Rincon

Artificial intelligence



David Hosking

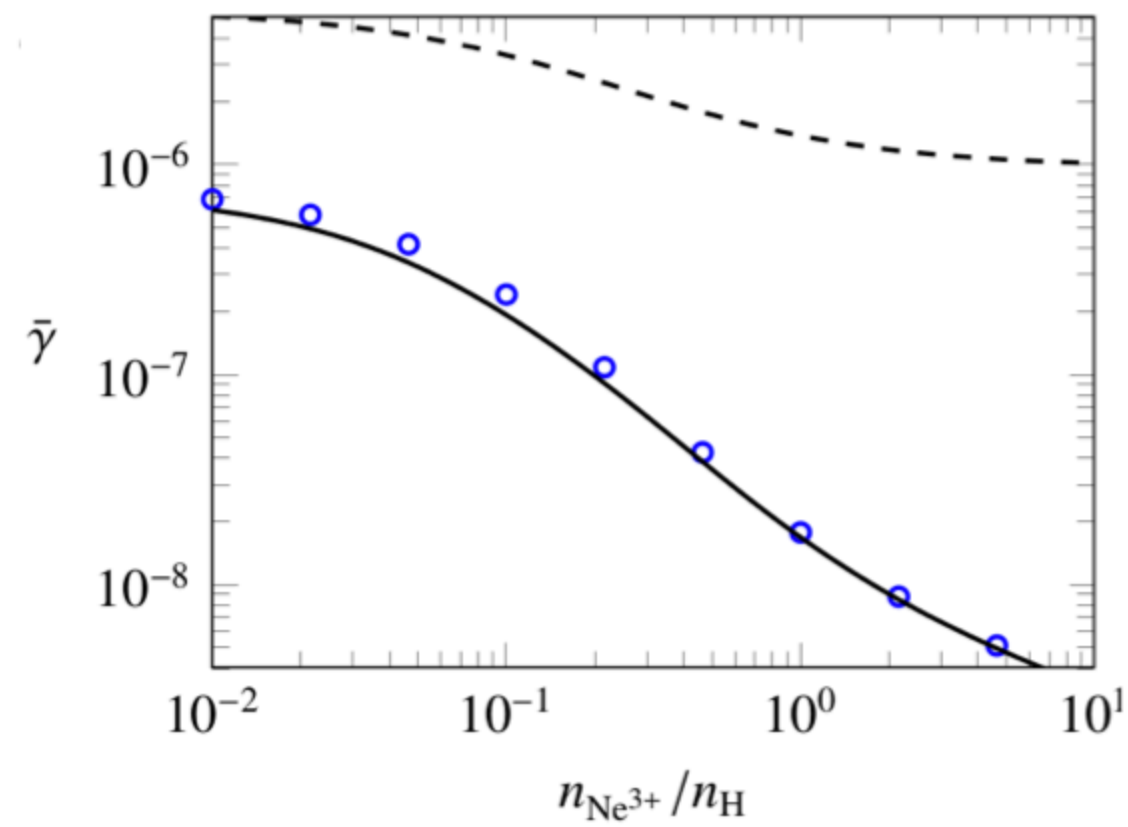
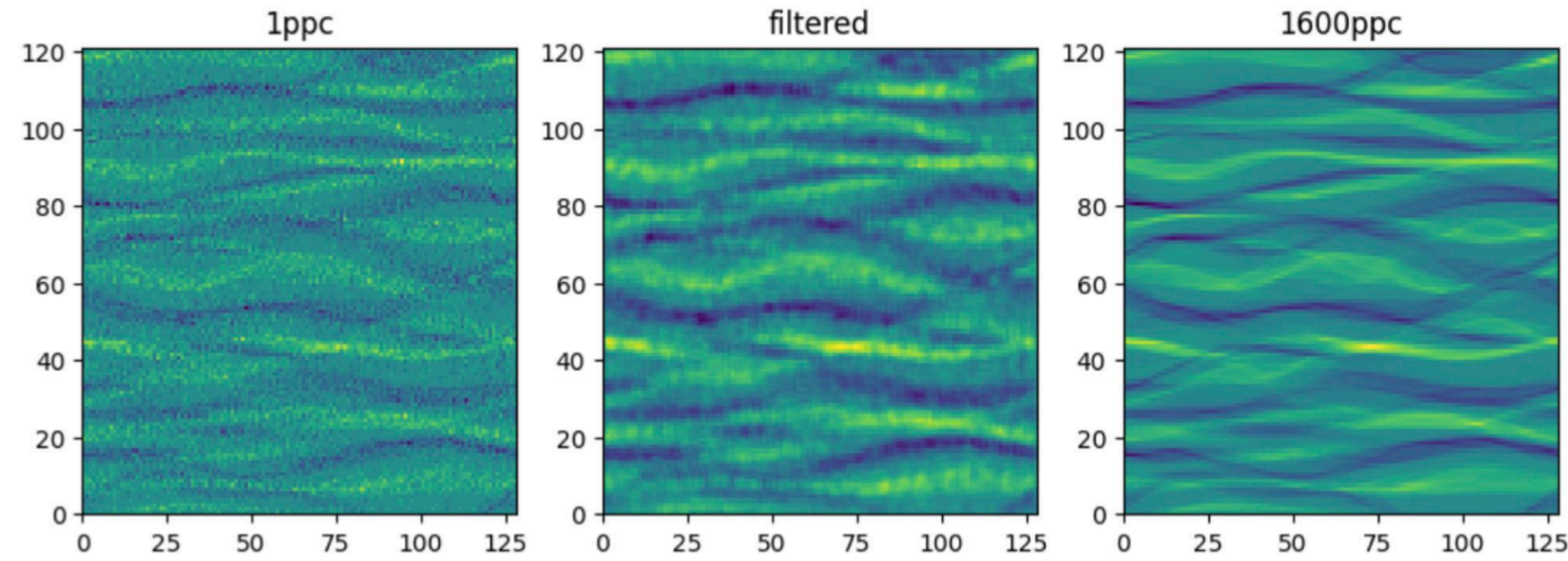


Figure shows the Dreicer generation rate in the presence of Ne^{3+} , obtained by neural network trained on kinetic simulations (solid), kinetic simulations (blue circles) and the Connor-Hastie formula (dashed)

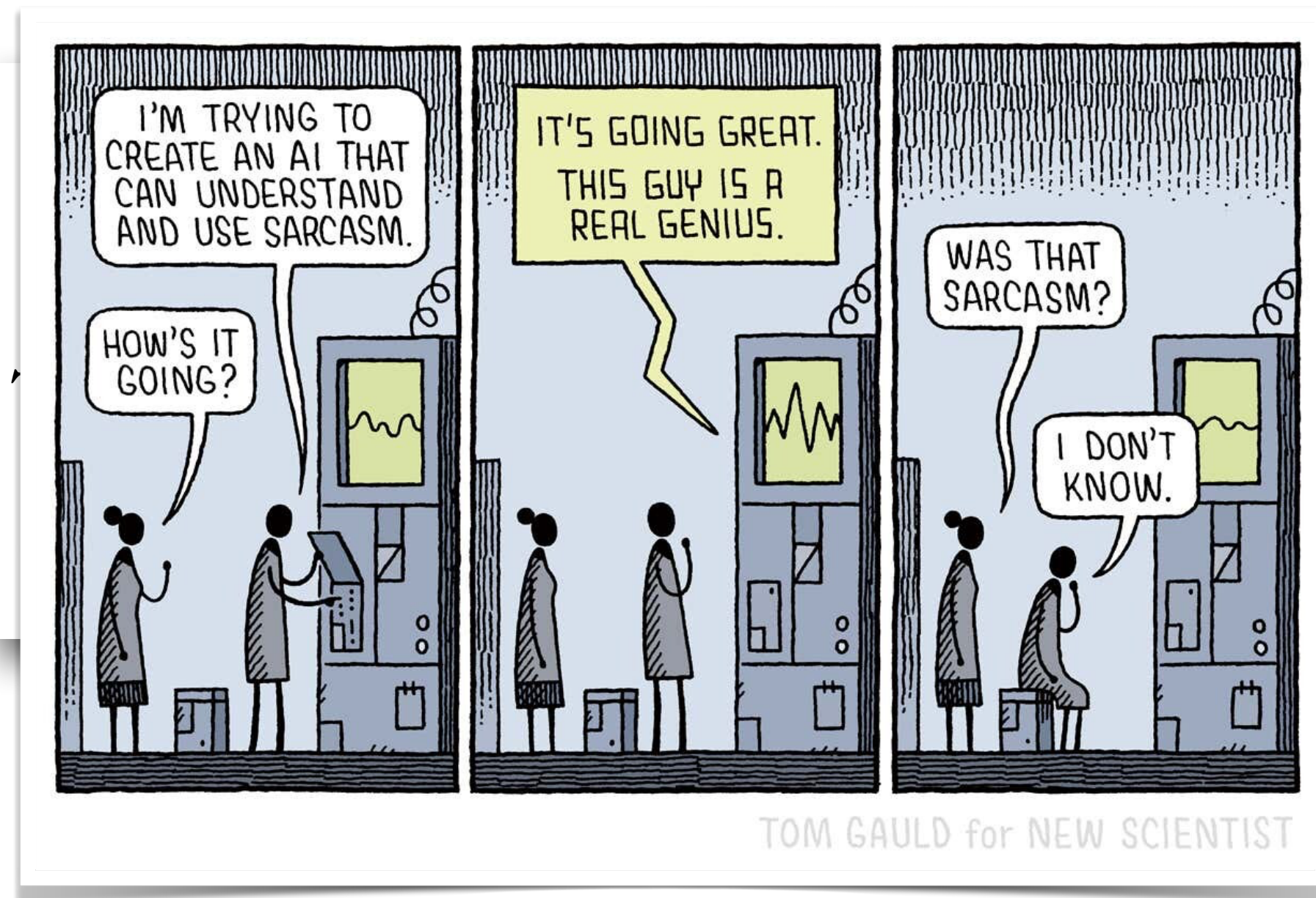
Tünde Fülöp

Micromirrors mediating multiscale motions in magnetised megastructures

Patrick Reichherzer

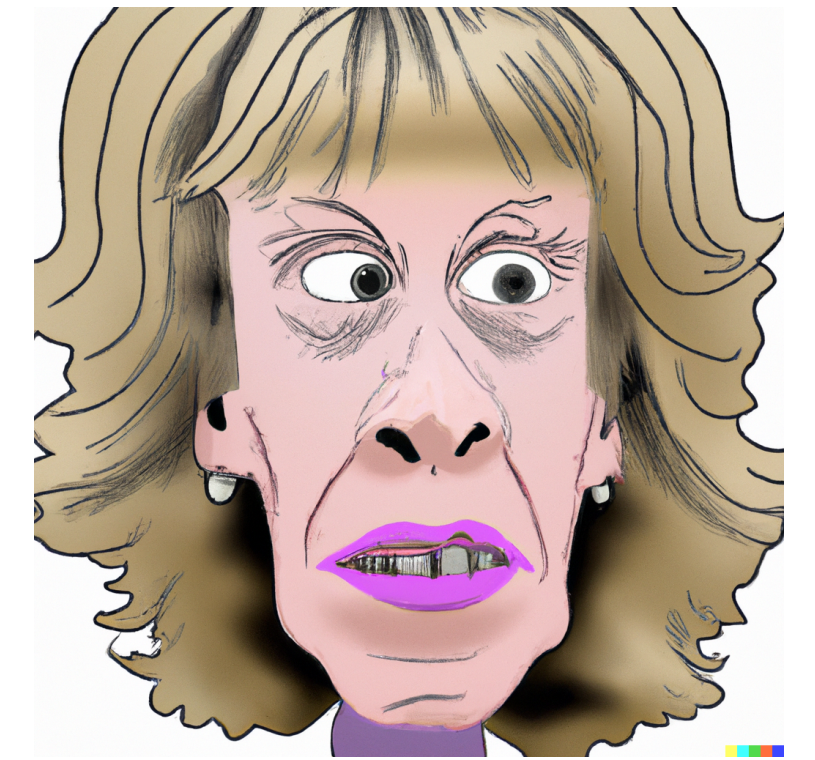


Anatoly Spitkovsky



ands,
ounds.
a leash,
'd reach,
making

Matt Kunz



David Hosking

Artificial intelligence



François Rincon