

# Tearing and Mirror Instability

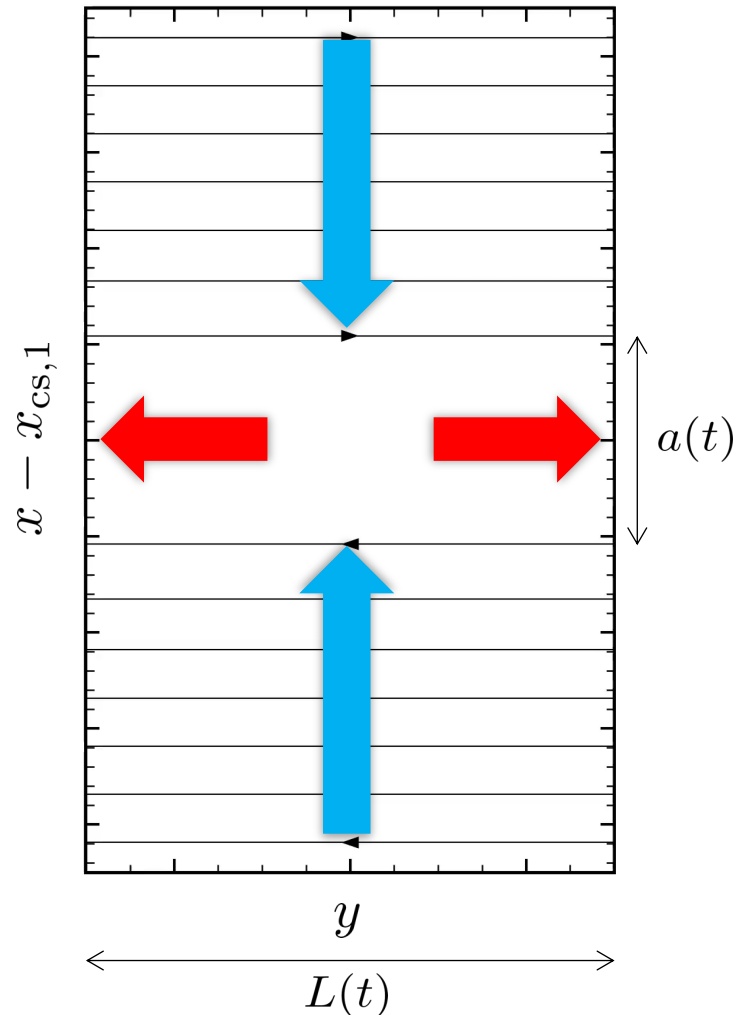
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**UNIVERSITY**

# Current Sheet (CS) Formation



Reconnection does not immediately start in nature

Consider field reversal with Harris equilibrium

$$\mathbf{B}_r(t, x) = B_r(t) \tanh \left[ \frac{x}{a(t)} \right] \hat{\mathbf{y}}$$

Impose background incompressible fluid flow:

$$\mathbf{u}(t, x, y) = \frac{1}{\Gamma(t)\tau_{cs}} (-x\hat{\mathbf{x}} + y\hat{\mathbf{y}})$$
$$\Gamma(t) \doteq 1 + \frac{t}{\tau_{cs}}$$

# Hyper-resistivity Tearing Instability

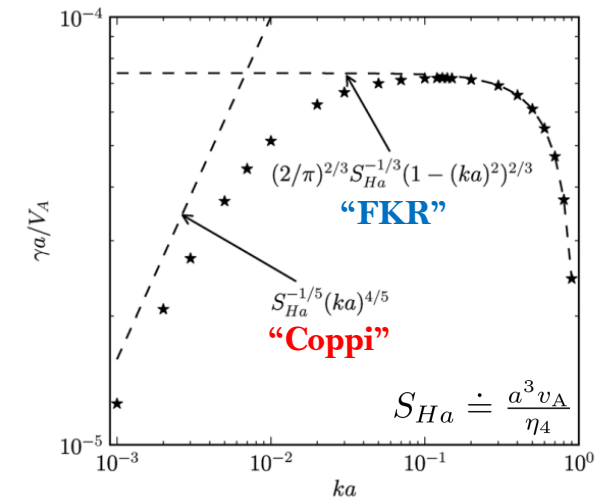
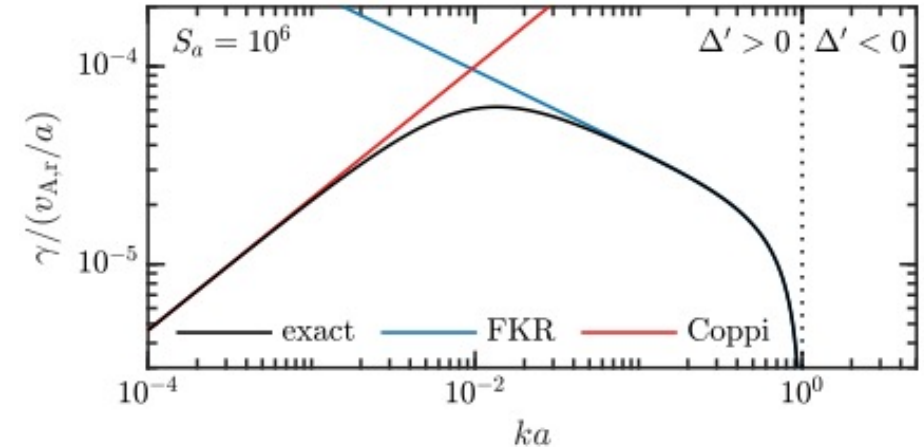
→ Reconnection needs CS to go tearing unstable, for Harris sheet:

$$\Delta'(t, N) = \frac{2N}{L(t)} \left[ \frac{1}{N^2} \frac{L^2(t)}{a^2(t)} - 1 \right]$$

→ In Ohmic tearing, there are two asymptotic regimes:

- Coppi : Multiple magnetic islands
- FKR : Single magnetic island

→ In hyper-resistive tearing has flat “FKR” regime, thus tearing mode with single island is always dominant.



Huang+(2013)

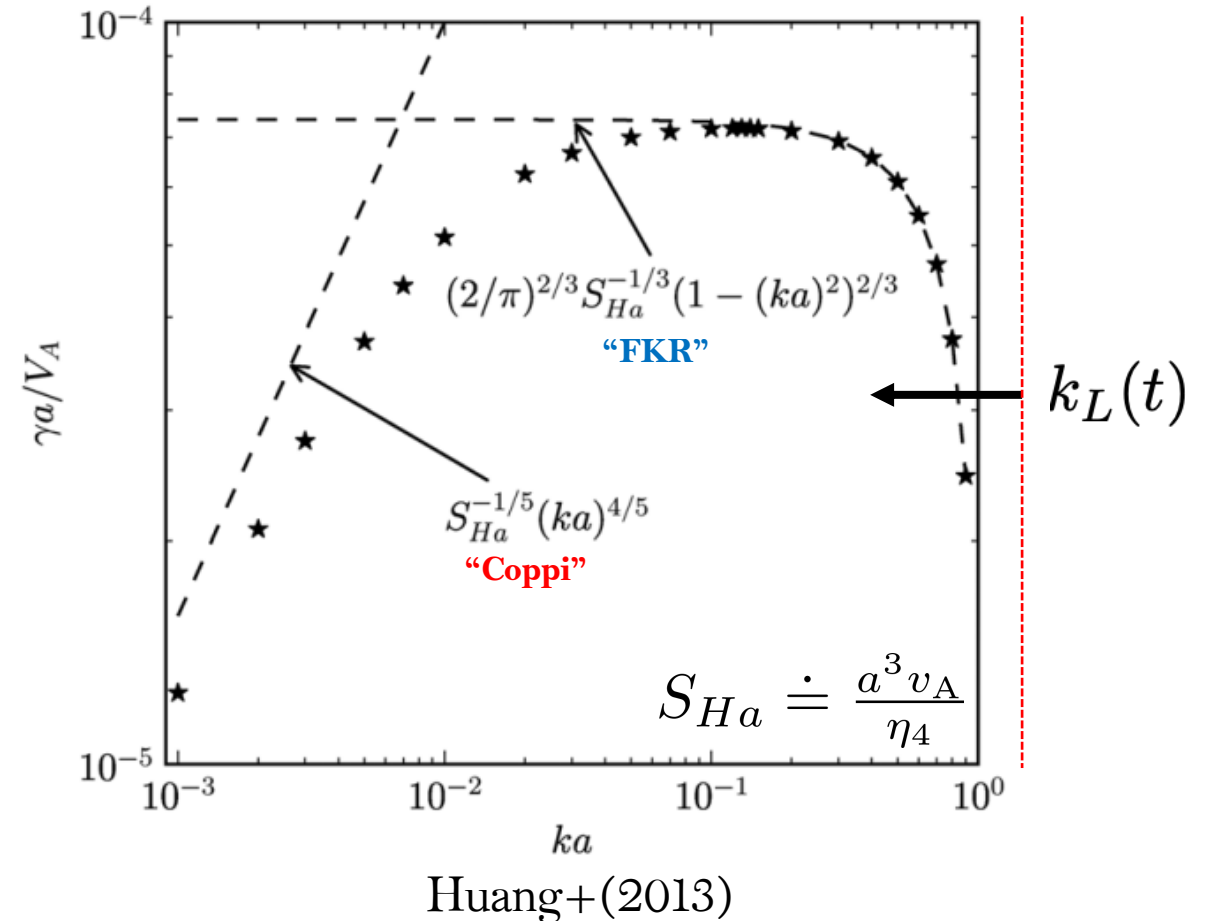
# Hyper-resistivity Tearing Instability

→ As CS evolves, aspect ratio changes and longer tearing wavelengths become available.

→ Estimate of tearing onset is from the first unstable mode of  $N \sim 1$ .

→ Growth rate must satisfy  $\gamma_t \tau_{cs} \gtrsim 1$  and  $k_t a \ll 1$  giving

$$\left(1 + \frac{t_{cr}}{\tau_{cs}}\right)^{2/3} \frac{a_0}{L_0} \ll S_{a0}^{1/3} M_{A0} \lesssim \left(1 + \frac{t_{cr}}{\tau_{cs}}\right)^{8/3}.$$



# Magnetic Accumulation Causes Mirror Instability

→ CS compression cause magnetic amplification

$$B_r(t) = B_{r0}\Gamma(t)$$

→ In collisionless system, the double adiabatic conservations gives pressure anisotropy

$$\Delta_p(t, \xi(t)) = [\Gamma(t)]^3 - 1 \approx \frac{3t}{\tau_{cs}}$$

→ Large pressure anisotropy can be unstable to mirror instability given that

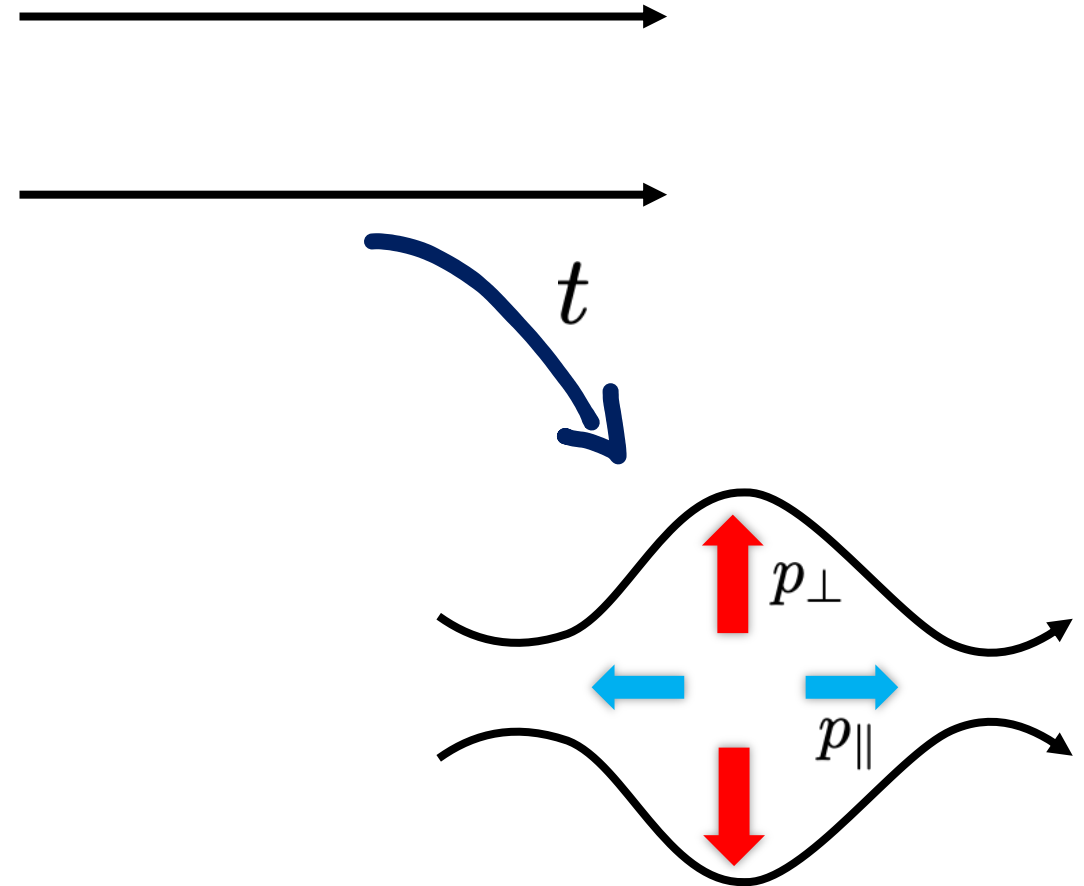
$$\Lambda_m \doteq \Delta_p - \frac{1}{\beta_{\perp}} > 0$$

→ Bulk region with stronger initial field will be unstable first with mirror time satisfying

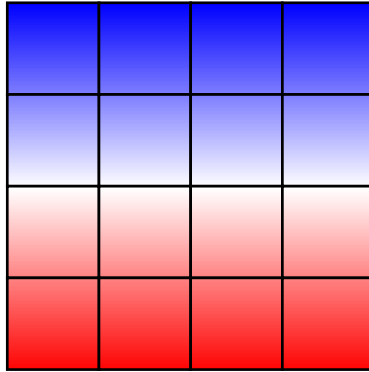
$$\frac{t_m}{\tau_{cs}} \approx \frac{1}{3\beta_{\perp}(0, \xi_0)} \ll 1$$

→ The pressure anisotropy will then be regulated at mirror regulation time

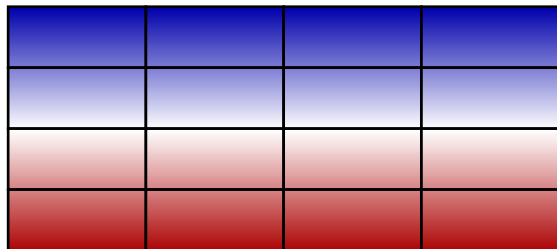
$$\frac{t_{m,reg} - t_m}{\tau_{cs}} \approx \frac{C_m}{3} (\Omega_i \tau_{cs})^{-1/2}$$



# Numerical Framework



**Code**



**Lab**

→ Hybrid-kinetic code **Pegasus++** with kinetic ions and neutralizing electron background fluid.

- Hyper-resistivity acts as the electron inertia

→ Compression is modeled using continuous frame transformation

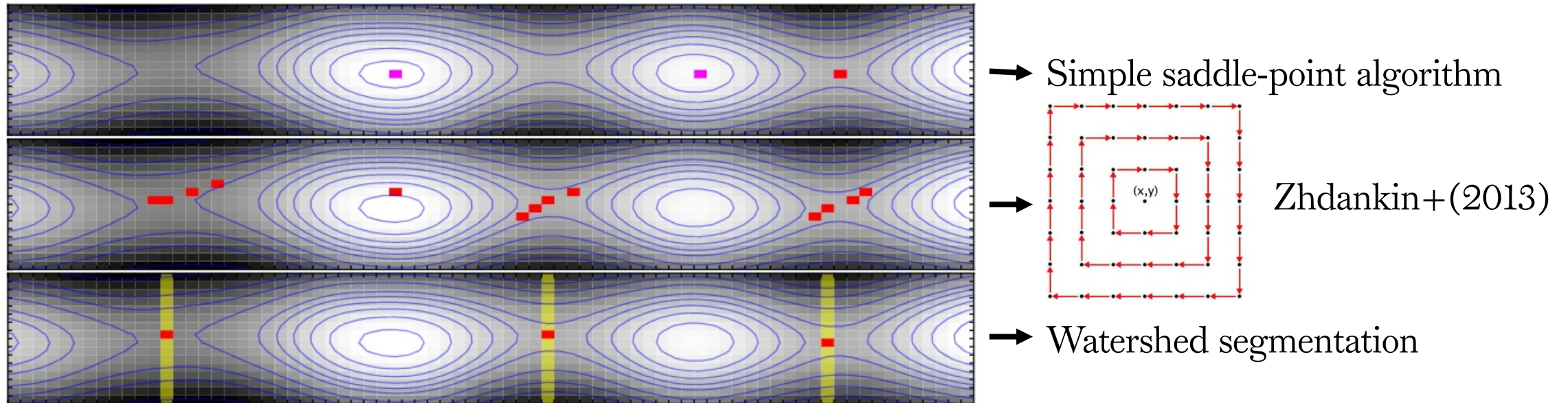
$$\Lambda(t) = \Gamma^{-1}(t)\hat{x}\hat{x} + \Gamma(t)\hat{y}\hat{y} \quad \longrightarrow \quad \begin{aligned} E' &= \Lambda E \\ B' &= \lambda \Lambda^{-1} B \\ n' &= \lambda n \\ u' &= \Lambda^{-1} u \end{aligned}$$

$$\Gamma(t) \doteq 1 + \frac{t}{\tau_{cs}}$$

→ Initialize with double Harris-sheet with periodic boundary conditions

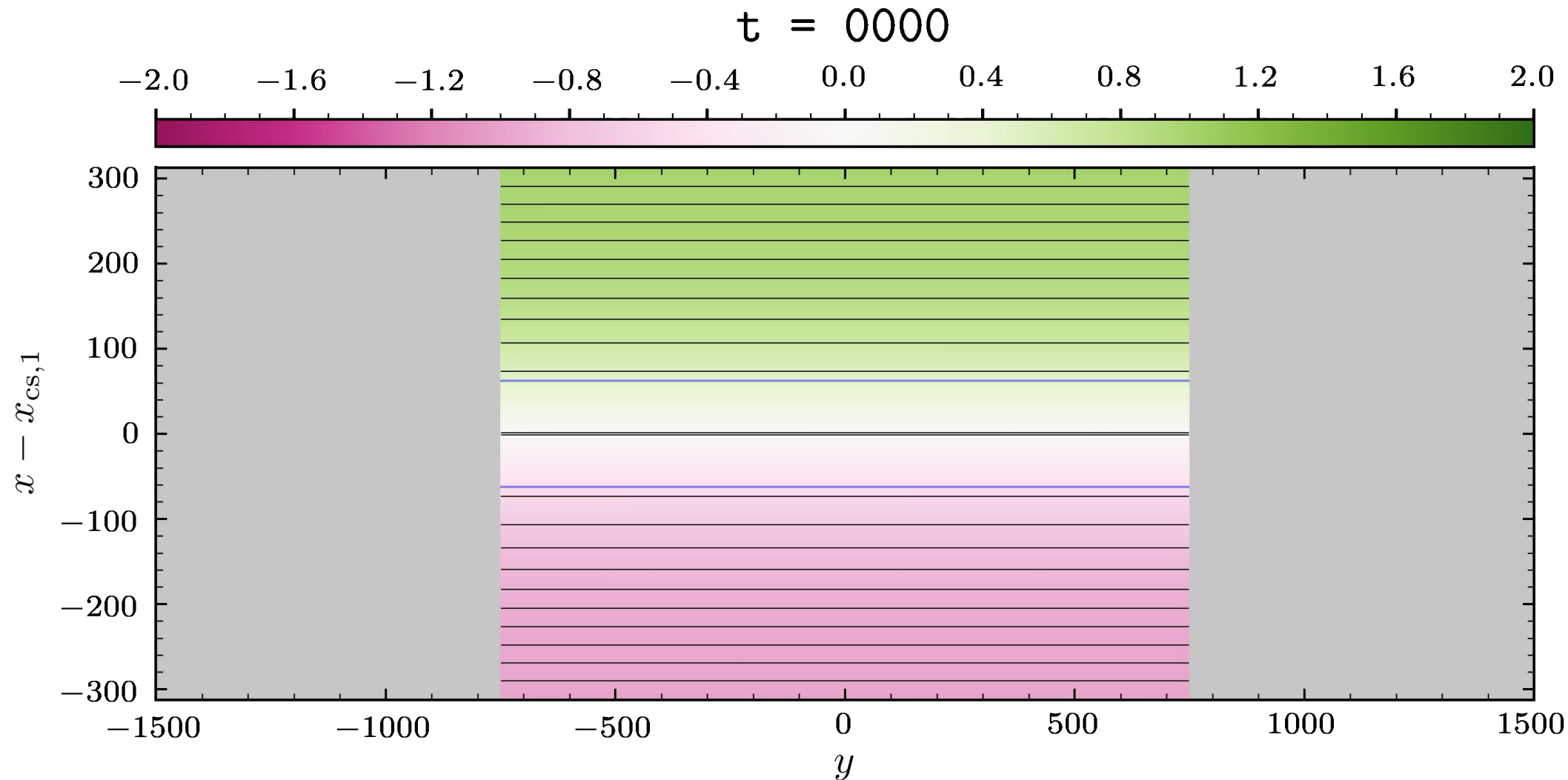
$$B_r(t=0) = B_{r0} \left[ \tanh\left(\frac{x-x_{cs,1}}{a_0}\right) - \tanh\left(\frac{x-x_{cs,2}}{a_0}\right) - 1 \right] \hat{y}$$

# Watershed Segmentation to Identify X-Points



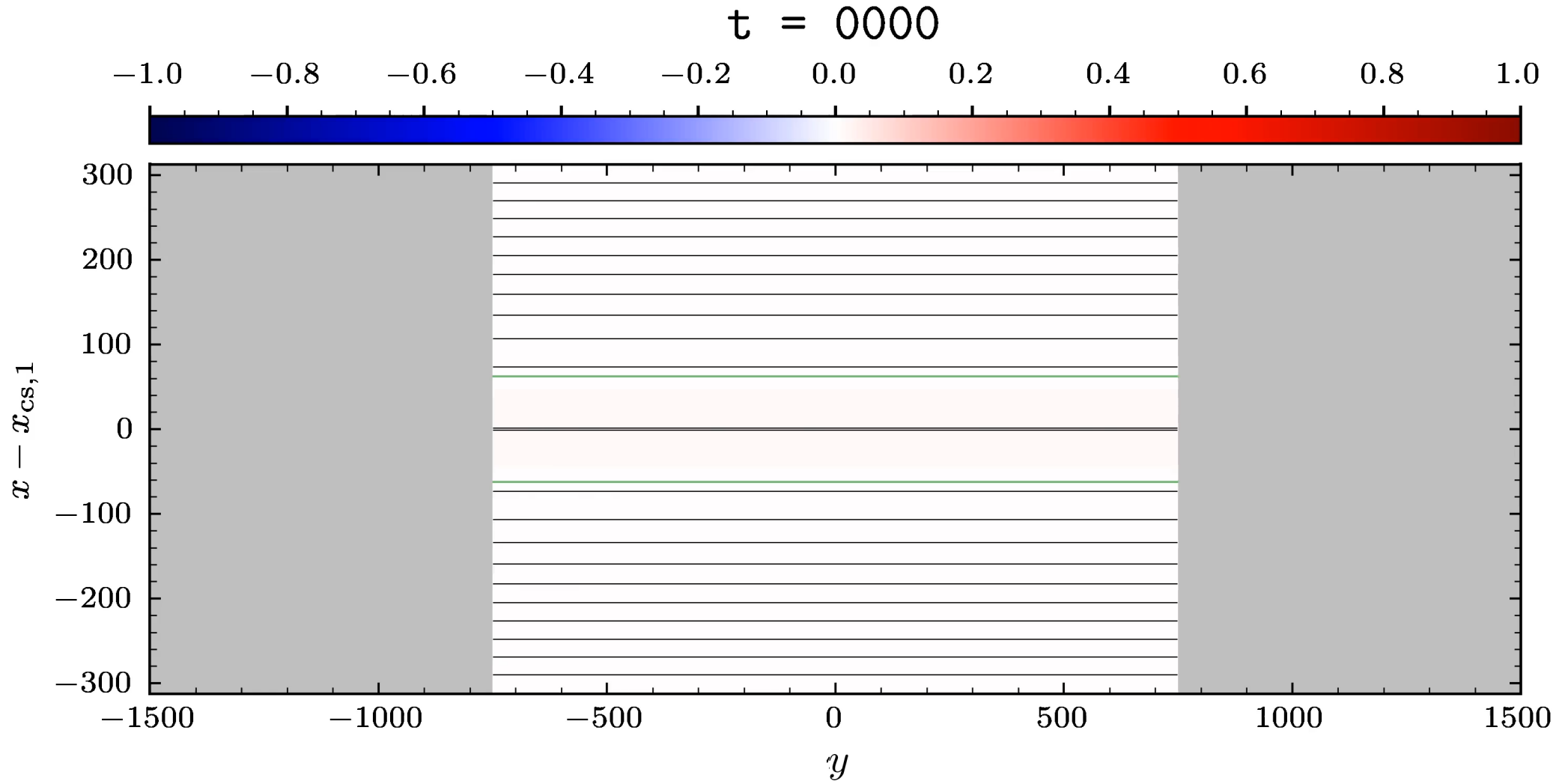
Magnetic island acts as flood basins with X-points as the local minimum point along the boundaries

# Fiducial Evolution $\tau_{CS} = 1000, a_0 = 125$





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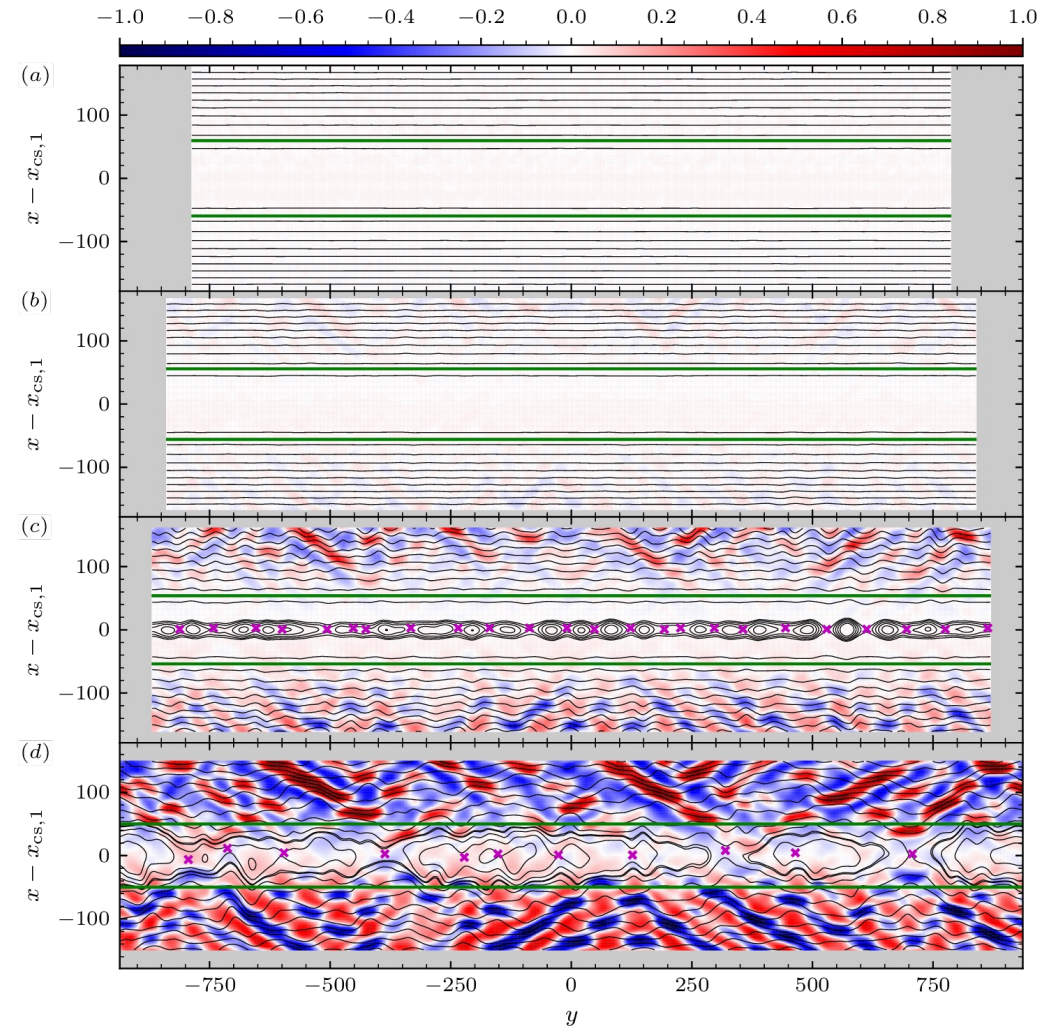
→ Focus at the four stages of evolution:

- (a) :  $t = 50$
- (b) :  $t = 120$
- (c) :  $t = 160$
- (d) :  $t = 250$

→ Higher field strength in Bulk region starts the mirror instabilities early as shown in figure (b).

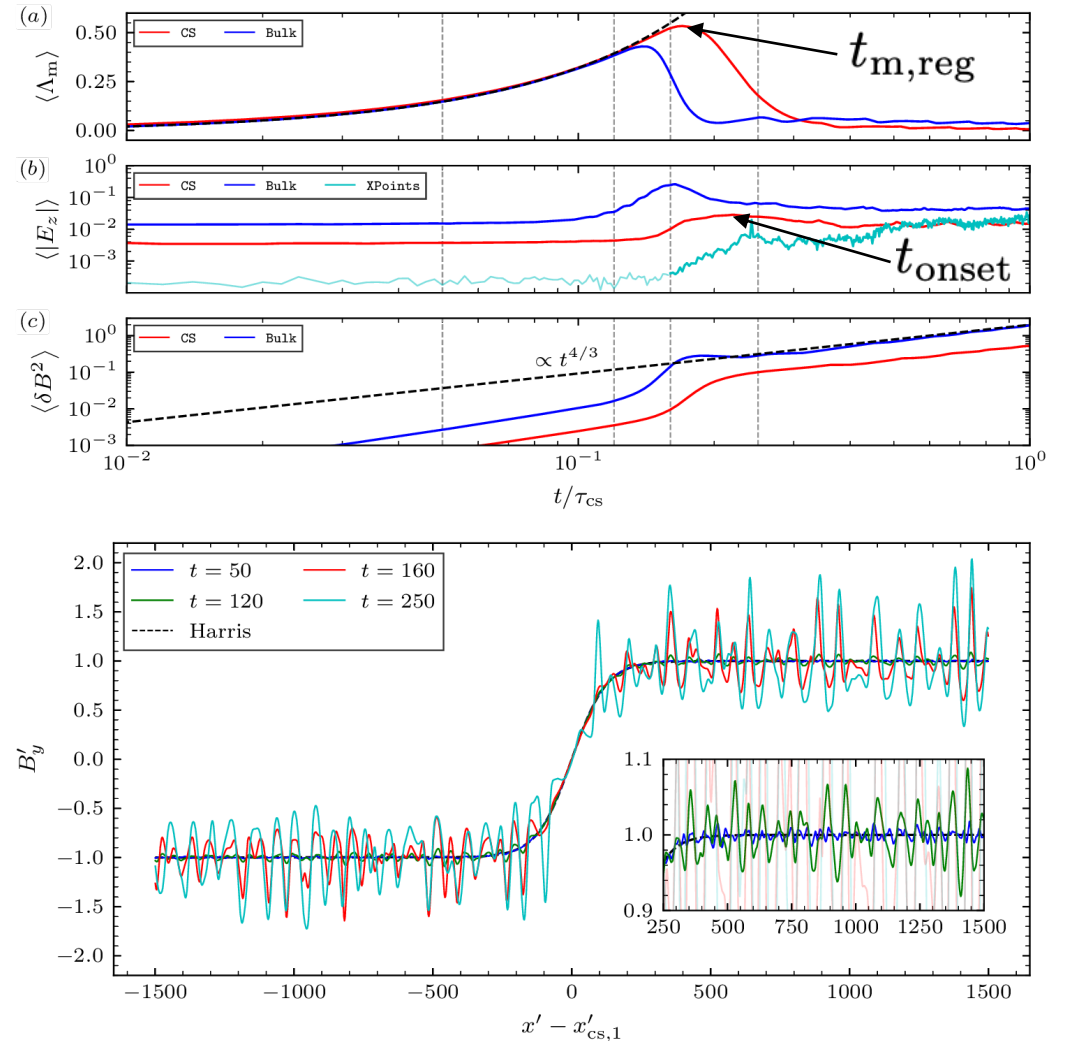
→ In figure (c), the CS should still be **tearing stable**, but we start seeing magnetic mirror.

→ In figure (d), the reconnection is well underway with non-linear tearing as indicated by **merging magnetic islands**.



# Fiducial Evolution $\tau_{CS} = 1000, a_0 = 125$

- Average quantities based on the **Bulk** vs **CS** region as well as the averaged values over the location of X-points as **XPoints**.
- Confirmed secular growth of magnetic fluctuations due to mirror instabilities in the **Bulk**.
- Mirror fluctuations get progressively closer to the neutral line as time progress.
- The value of average  $E_z$  in **CS** region decrease slowly as the magnetic islands grow.
- Measured average  $E_z$  over the **XPoints** matches the previously reported value in prior hyper-resistive reconnection publications (e.g., Huang, 2013).



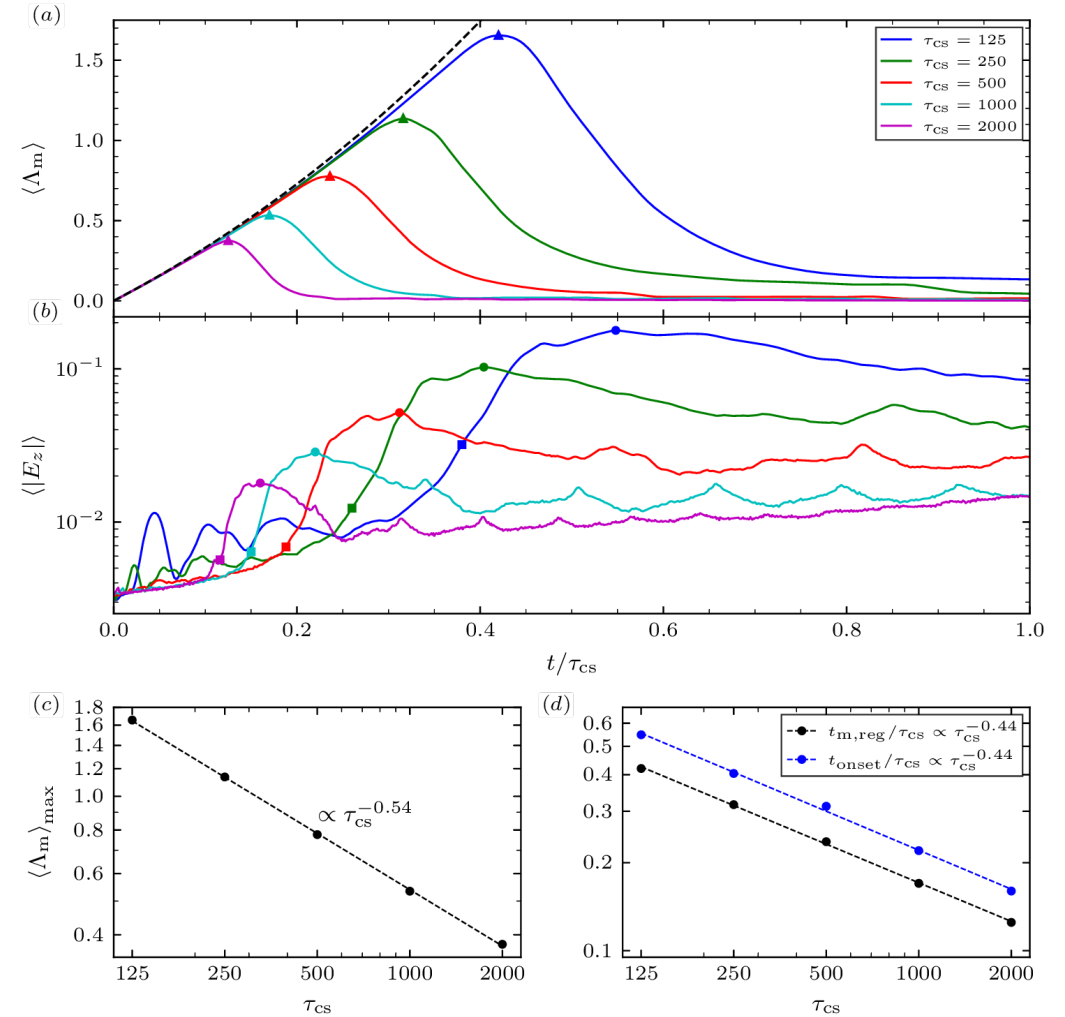
# Variation of $\tau_{CS}$

→ Initial mirror growth in CS follows double adiabatic condition, with peak values of  $\Lambda_m$  satisfies theoretical prediction

$$\Lambda_{m,\max} = C_m (\Omega_i \tau_{CS})^{-1/2}$$

$$\frac{t_{m,\text{reg}} - t_m}{\tau_{CS}} \approx \frac{C_m}{3} (\Omega_i \tau_{CS})^{-1/2}$$

→ The dominant linear tearing modes can be obtained from time averaging between the peak of mirror instability parameters and the second  $E_z$  peak, i.e. between  $t_{m,\text{reg}}$  and  $t_{\text{onset}}$ .



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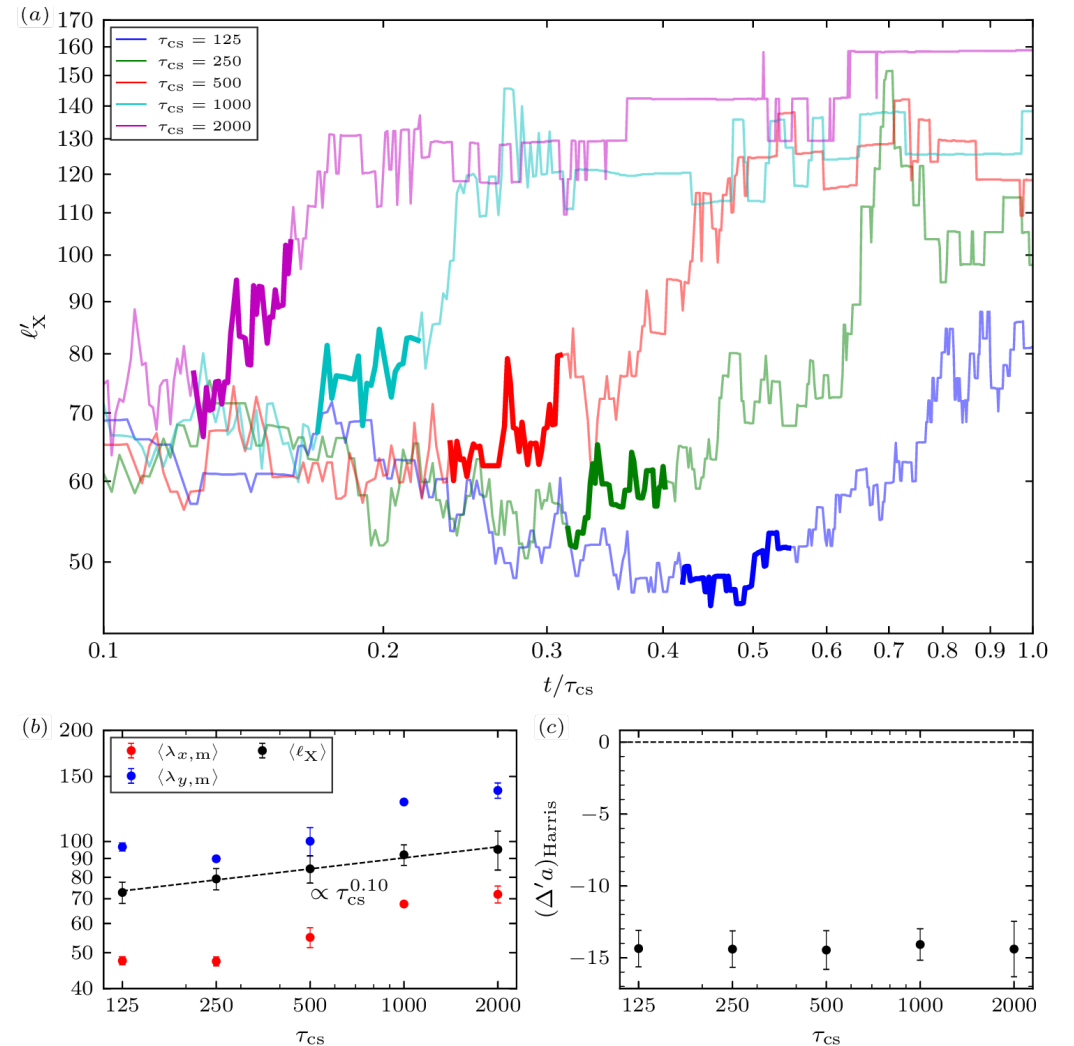
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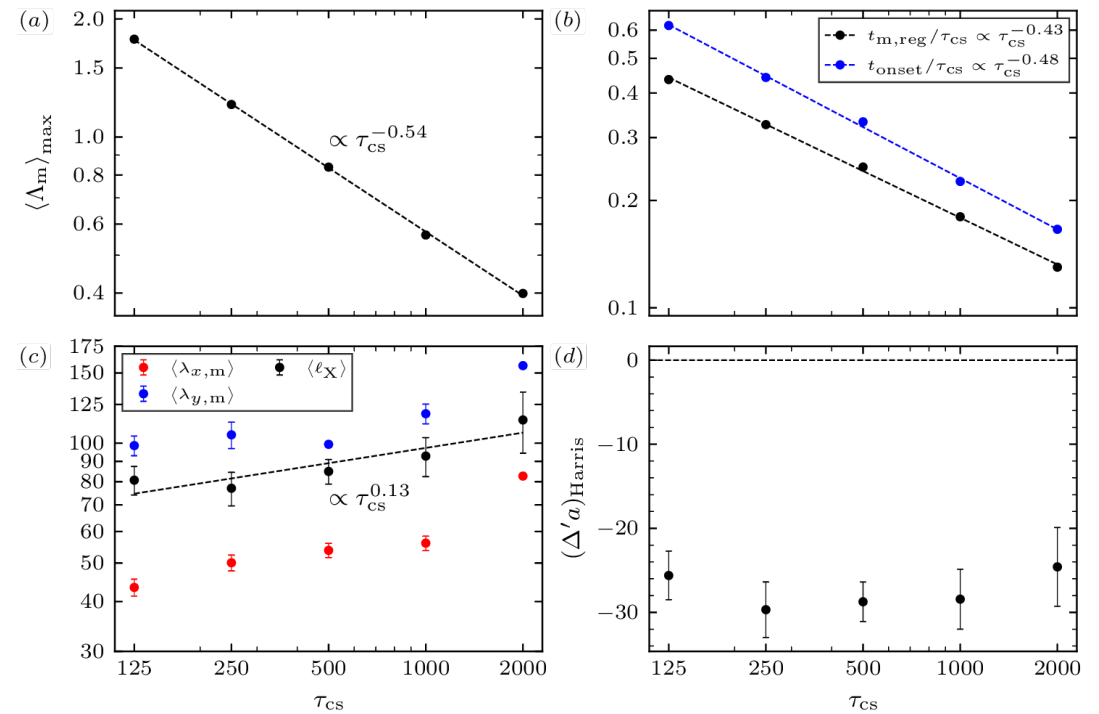
→ The most dominant mode scale similar to mirror modes indicating strong interaction between the two.

- Classically these tearing modes should still be stable.



# Wide Sheet with Variation of $\tau_{CS}$

- The reconnection onset time does not differ significantly from the thinner sheet, while the wider initial sheet should still be much more stable to tearing instability.
- Consistent scaling for CS with double initial thickness confirms that the mirror-instability is seeding the tearing modes.



# Summary

- During CS formation in high-beta collisionless plasma, the pressure anisotropy build up can trigger **mirror instability** disrupting the sheet with ion Larmor scale perturbations.
- The rapid growth of the **mirror fluctuations stimulates tearing modes** by wrinkling the current sheet, effectively reducing CS thickness and changing the value of  $\Delta'$ .
- The resulting tearing instability **onsets at earlier time** and **on smaller scales** than it would have without the mirrors.
  - Onset time scales at approximately  $\propto \tau_{CS}^{1/2}$ .
  - Tearing modes are at intermediate scales between the parallel and perpendicular wavelengths of the mirror modes.
- In turbulence context, this result puts the importance of **kinetic instabilities** that can disturb magnetic folds, which locally can be seen as CS.
- \* **Watershed segmentation** can be used to robustly determine locations of X-points.