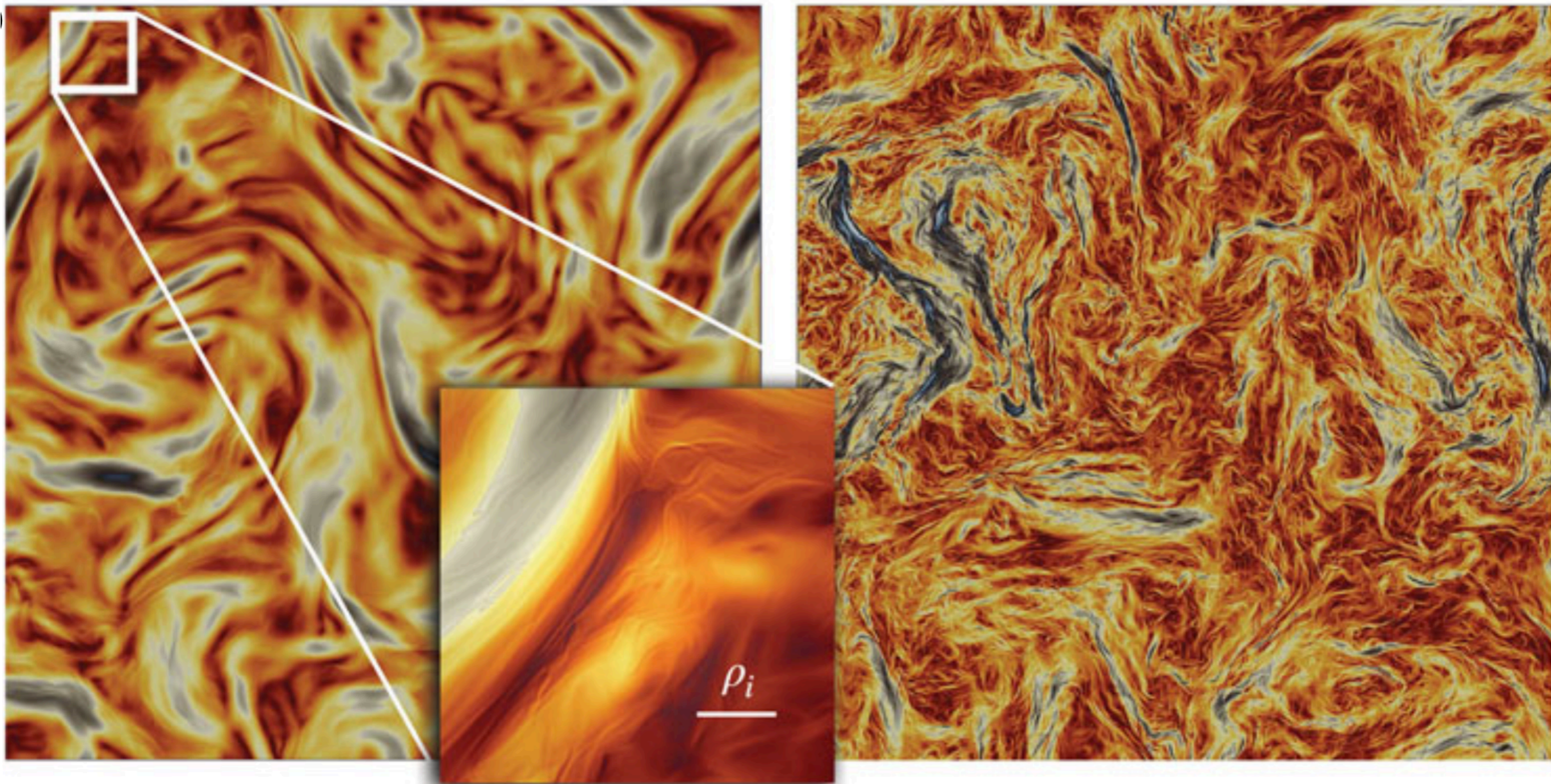


The helicity barrier



Heating and turbulence with imbalance

Model

FLR-MHD – low- β gyrokinetics without electron physics

$$\left(\frac{\partial}{\partial t} + \mathbf{u}_{\perp} \cdot \nabla_{\perp} \right) \frac{\delta n_e}{n_{0e}} = -\frac{c}{4\pi e n_{0e}} \left(\frac{\partial}{\partial z} + \mathbf{b}_{\perp} \cdot \nabla_{\perp} \right) \nabla_{\perp}^2 A_{\parallel} + \mathcal{D}_{6v} \frac{\delta n_e}{n_{0e}},$$

$$\left(\frac{\partial}{\partial t} + \mathbf{u}_{\perp} \cdot \nabla_{\perp} \right) A_{\parallel} = -c \frac{\partial \varphi}{\partial z} + \frac{c T_{0e}}{e} \left(\frac{\partial}{\partial z} + \mathbf{b}_{\perp} \cdot \nabla_{\perp} \right) \frac{\delta n_e}{n_{0e}} + \mathcal{D}_{6v} A_{\parallel},$$

$$\frac{\delta n_e}{n_{0e}} = -\frac{Z}{\tau} \left(1 - \hat{\Gamma}_0 \right) \frac{e\varphi}{T_{0e}},$$

$$\Theta_k^{\pm} = -\Omega_i \frac{v_{\text{ph}}(k_{\perp})}{k_{\perp}^2} \frac{\delta n_e}{n_{0e}} \mp \frac{A_{\parallel}}{\sqrt{4\pi m_i n_{0i}}}.$$

$$v_{\text{ph}}(k_{\perp}) = \frac{k_{\perp} \rho_i}{\sqrt{2}} \left(\frac{1}{1 - \hat{\Gamma}_0} + \frac{Z}{\tau} \right)^{1/2} \approx \begin{cases} 1 & k_{\perp} \rho_i \ll 1, \\ \left(\frac{1}{2} + \frac{Z}{2\tau} \right)^{1/2} k_{\perp} \rho_i & k_{\perp} \rho_i \gg 1. \end{cases}$$

Model

Invariants

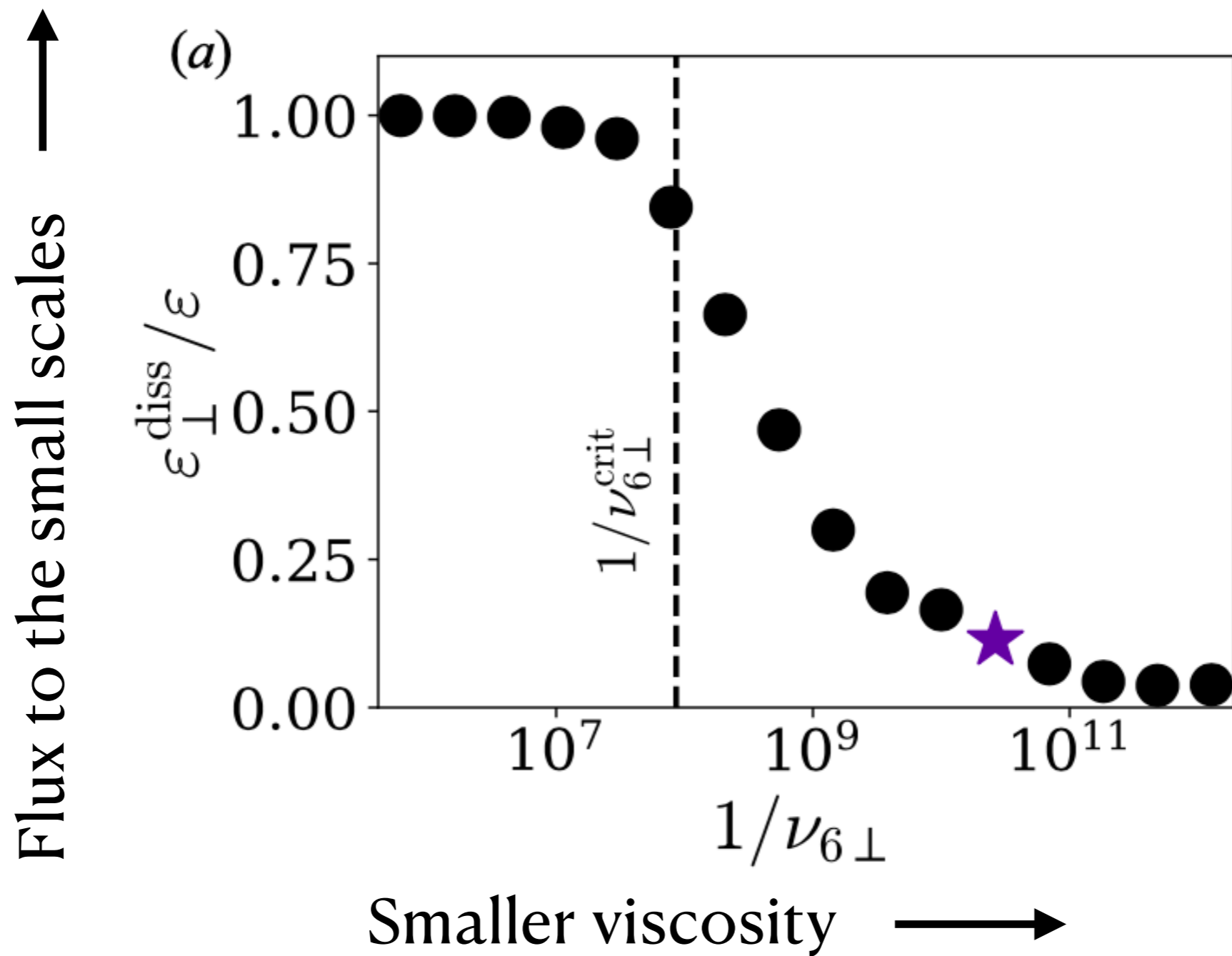
$$E = \frac{1}{4} \sum_k (|k_{\perp} \Theta_k^+|^2 + |k_{\perp} \Theta_k^-|^2) \quad \mathcal{H} = \frac{1}{4} \sum_k \frac{|k_{\perp} \Theta_k^+|^2 - |k_{\perp} \Theta_k^-|^2}{v_{\text{ph}}(k_{\perp})}$$



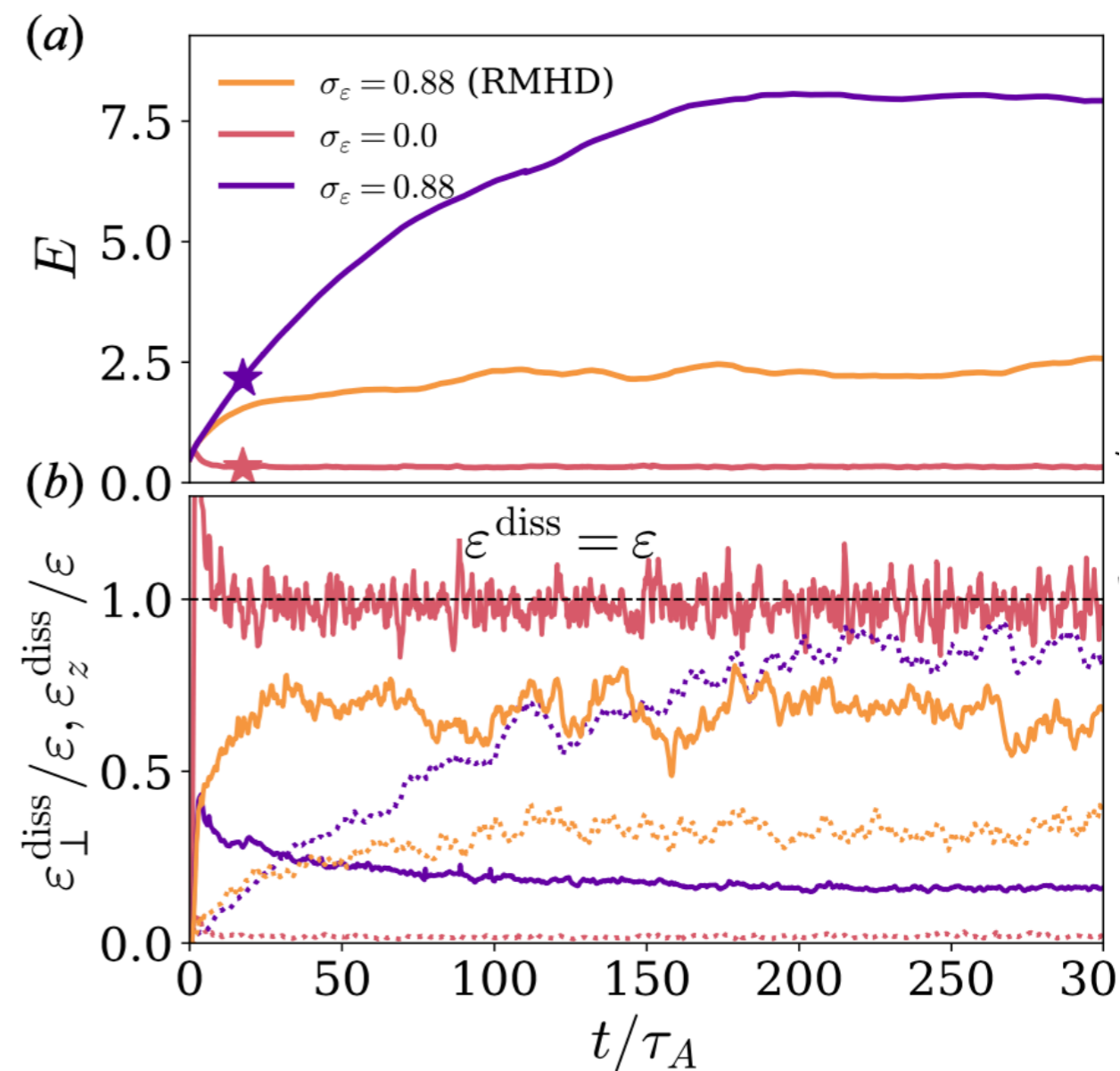
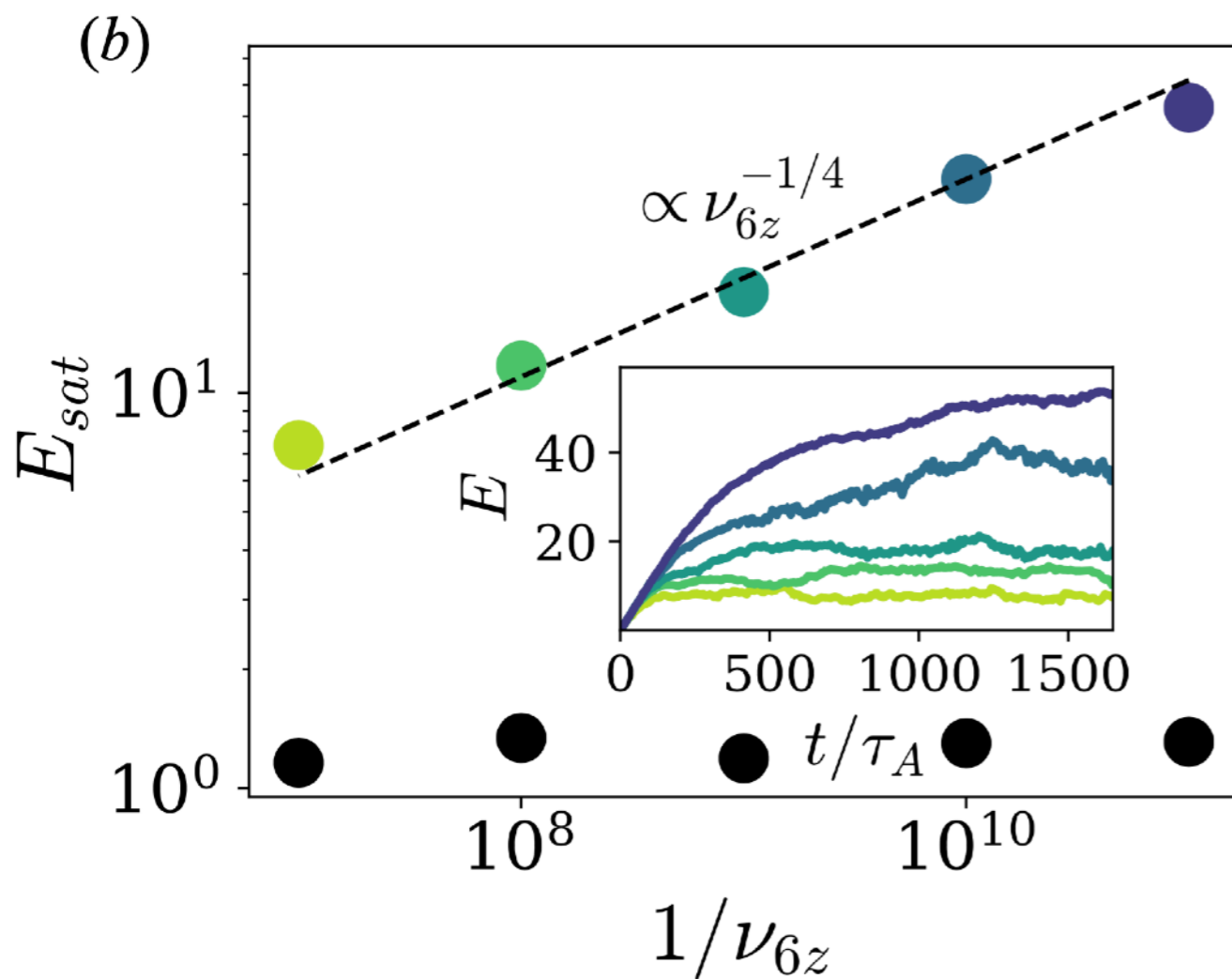
$$\sigma_{\varepsilon} = \frac{|\Pi_{\mathcal{H}}(k_{\perp})|}{\Pi(k_{\perp})} \leq \frac{1}{v_{\text{ph}}(k_{\perp})}$$

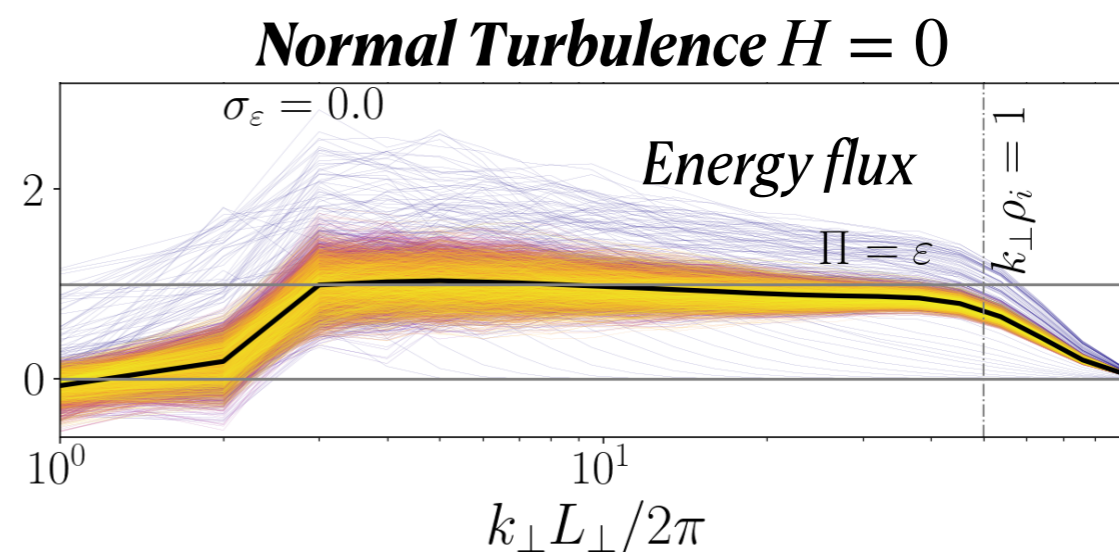
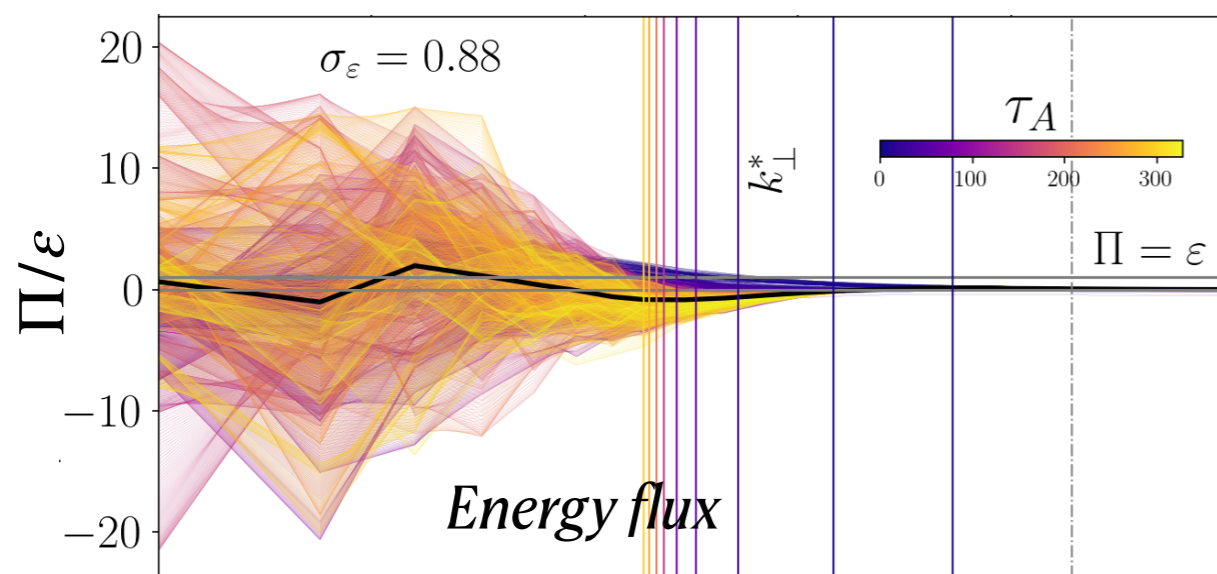
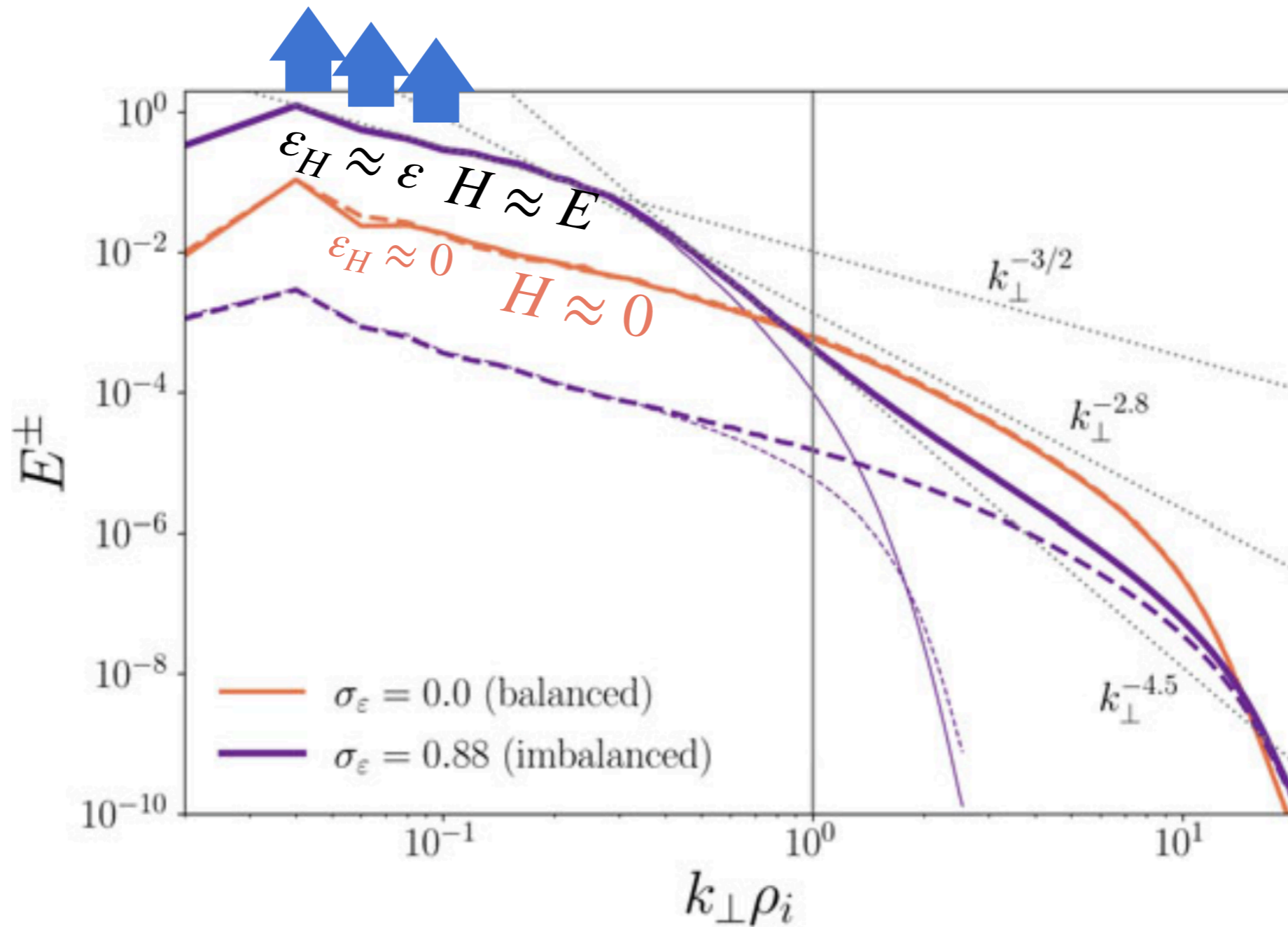
FLR effects strongly modify cascade at $k_{\perp} \rho_i \ll 1$

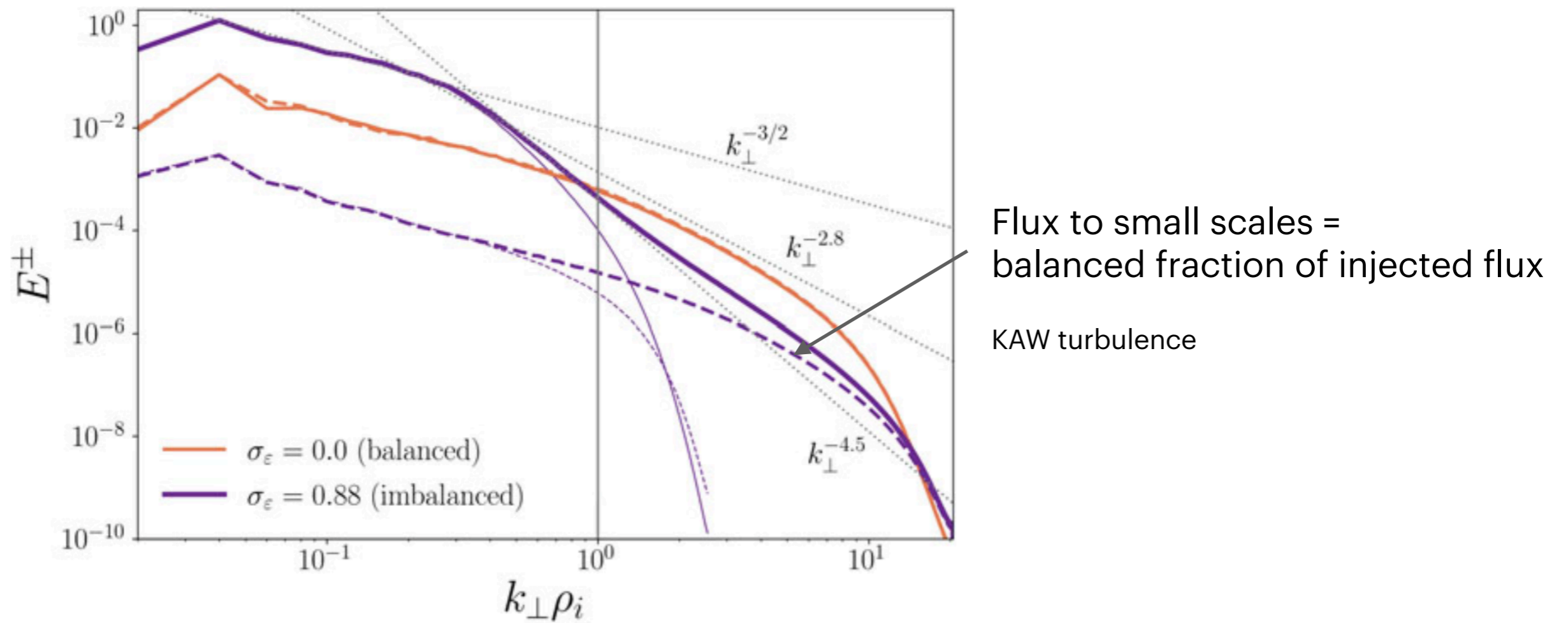
Violation of the zeroth law of turbulence



Parallel dissipation matters – “breaks” gyrokinetics



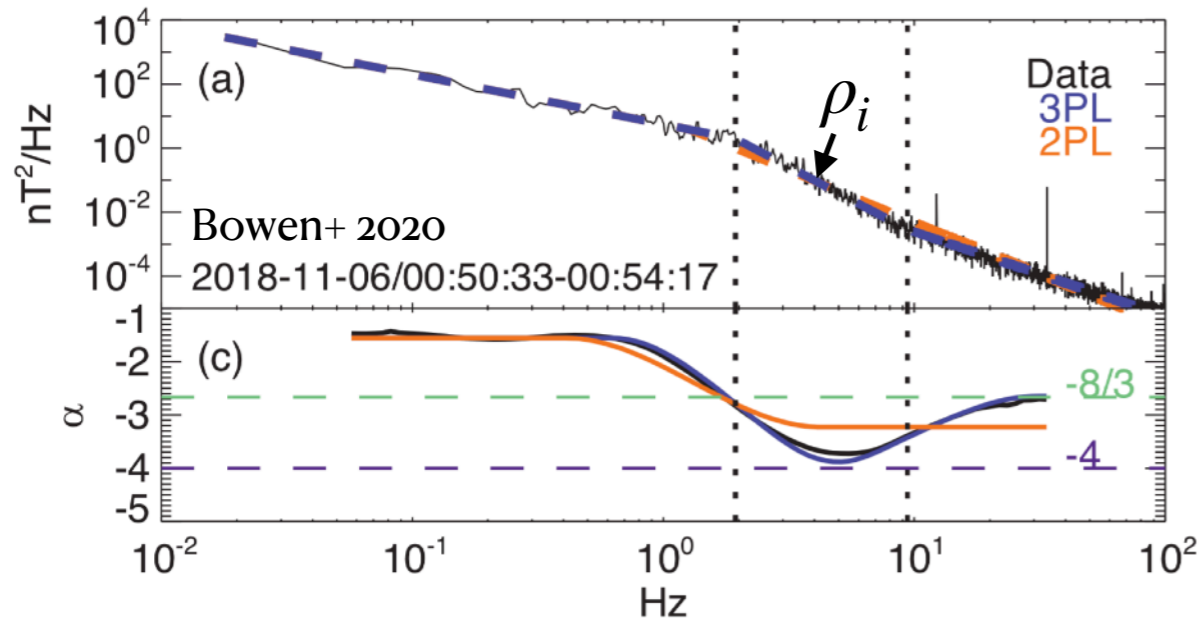




$$\frac{Q_i}{Q_e} \approx \frac{\text{Imbalanced fraction}}{\text{Balanced fraction}} = \frac{\varepsilon_H}{\varepsilon - \varepsilon_H} \gg 1 \quad (\text{if } \varepsilon \sim \varepsilon_H)$$

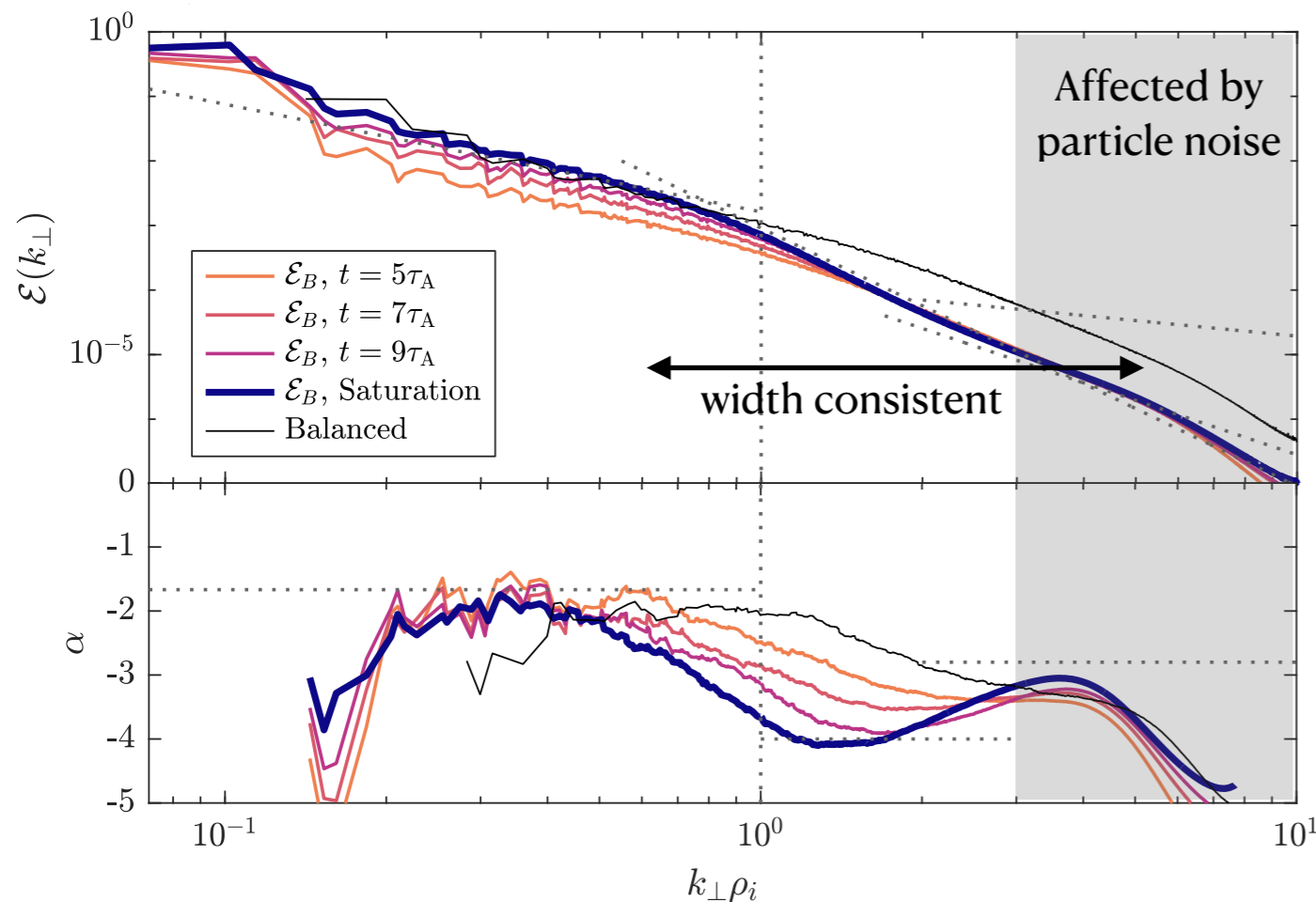
Observational Evidence

Magnetic spectra – transition range

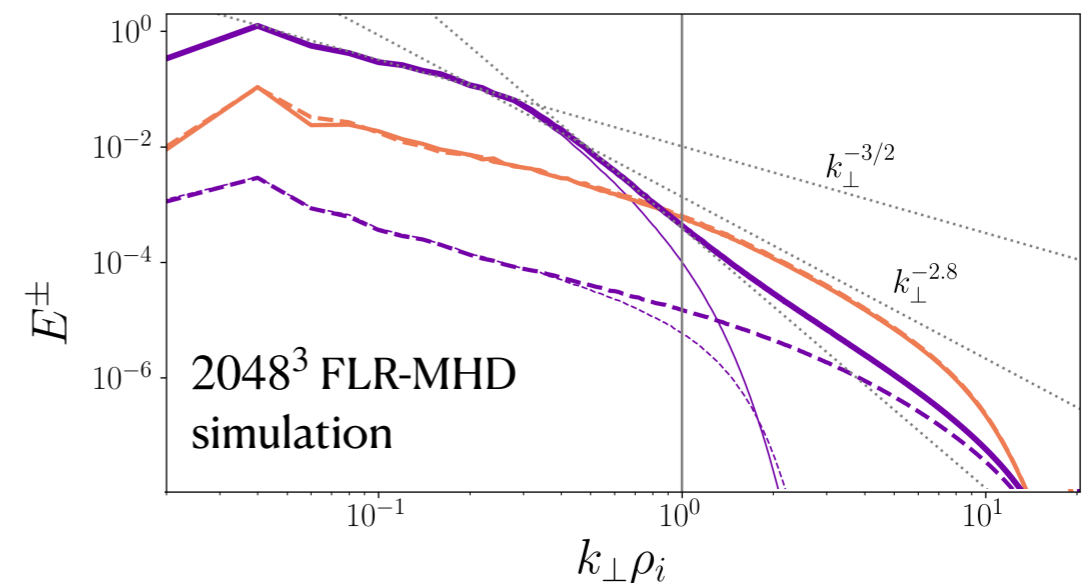


e.g., Leamon+ 1998, Smith+ 2006, Sahraoui+ 2009, Alexandrova+ 2013, Kiyani+ 2015, Bruno+ 2014, Vech+ 2018, Bowen+2020, Zhao+ 2020, Duan+ 2021, Huang+ 2021

Previous theories, e.g., Schekochihin+ 2009, Sahraoui+ 2010, Meyrand+ 2010, Lion+2010, Voitenko+ 2016, Mallet+ 2017



Physically: spectrum re-flattens because the balanced KAW cascade leaks through the barrier.

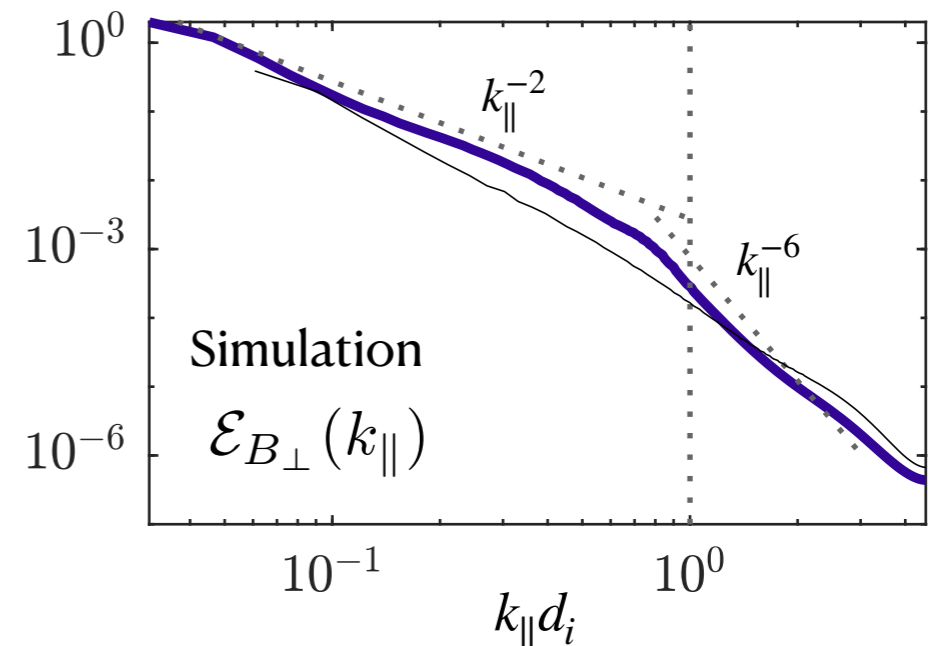
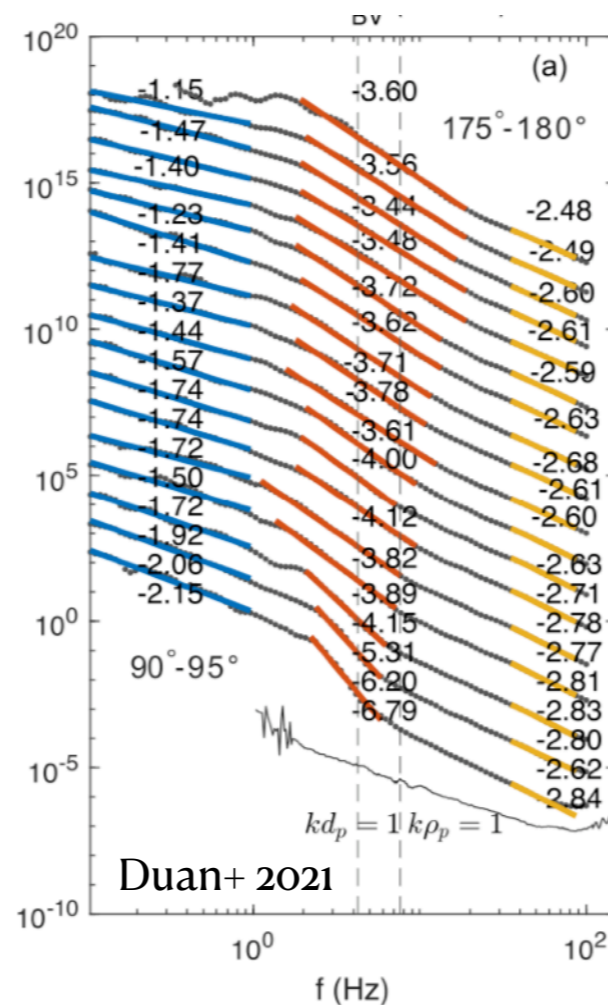
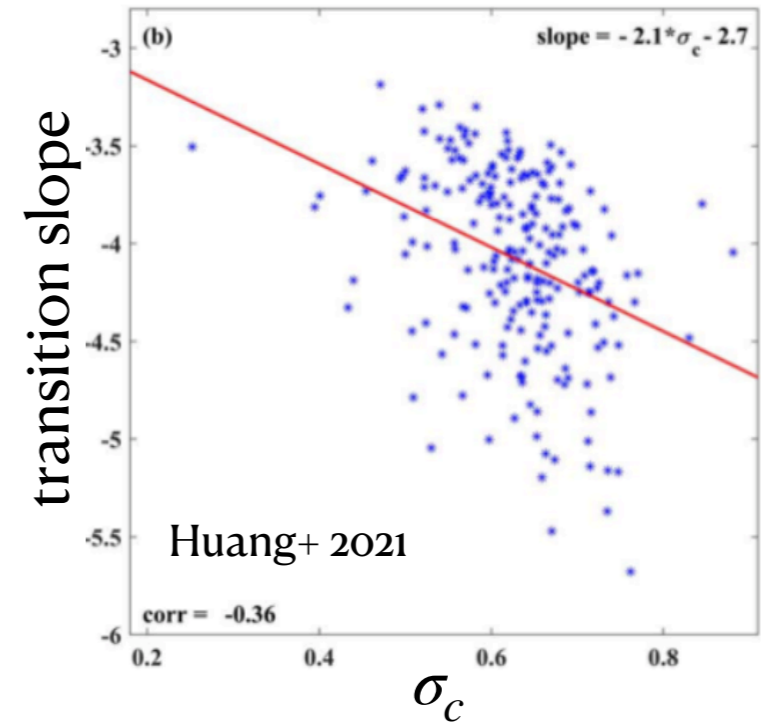


Observational Evidence

Magnetic spectra – transition range

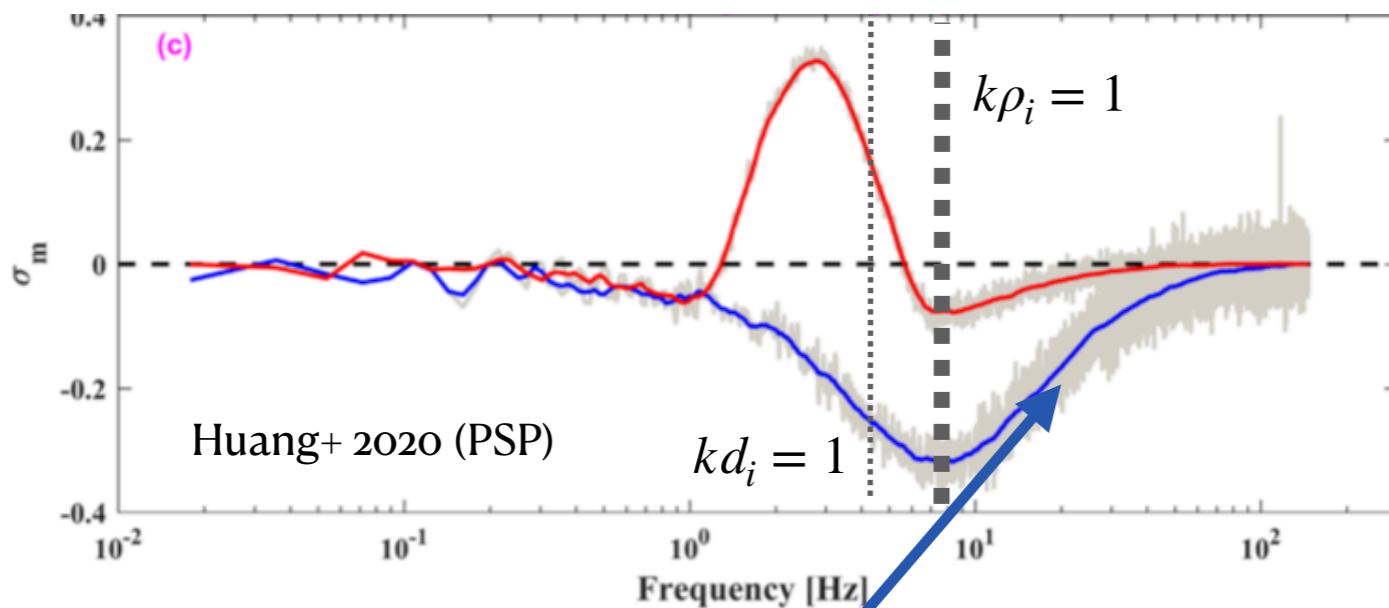
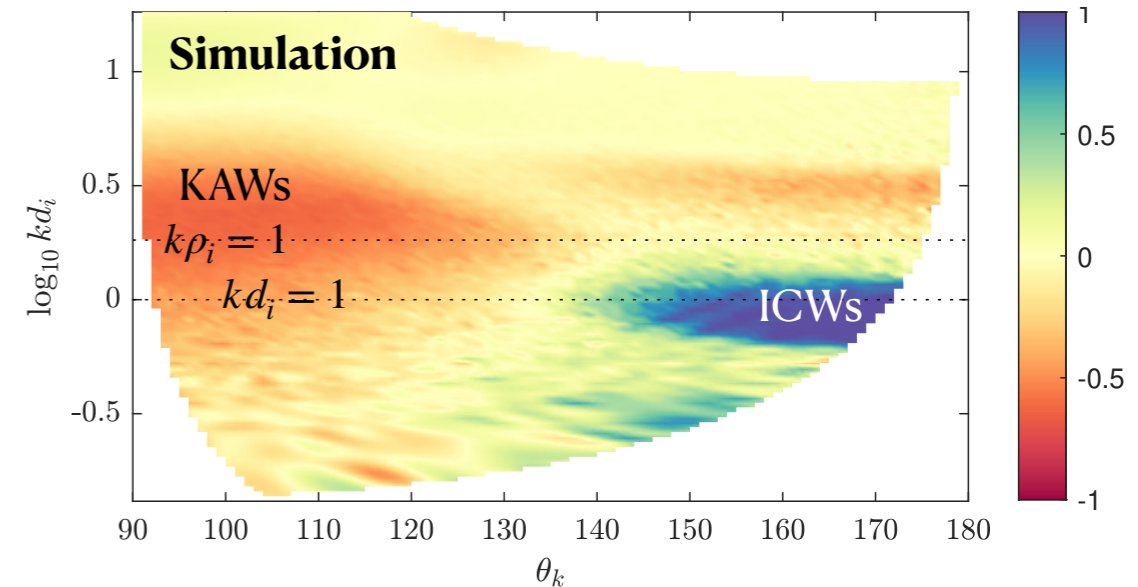
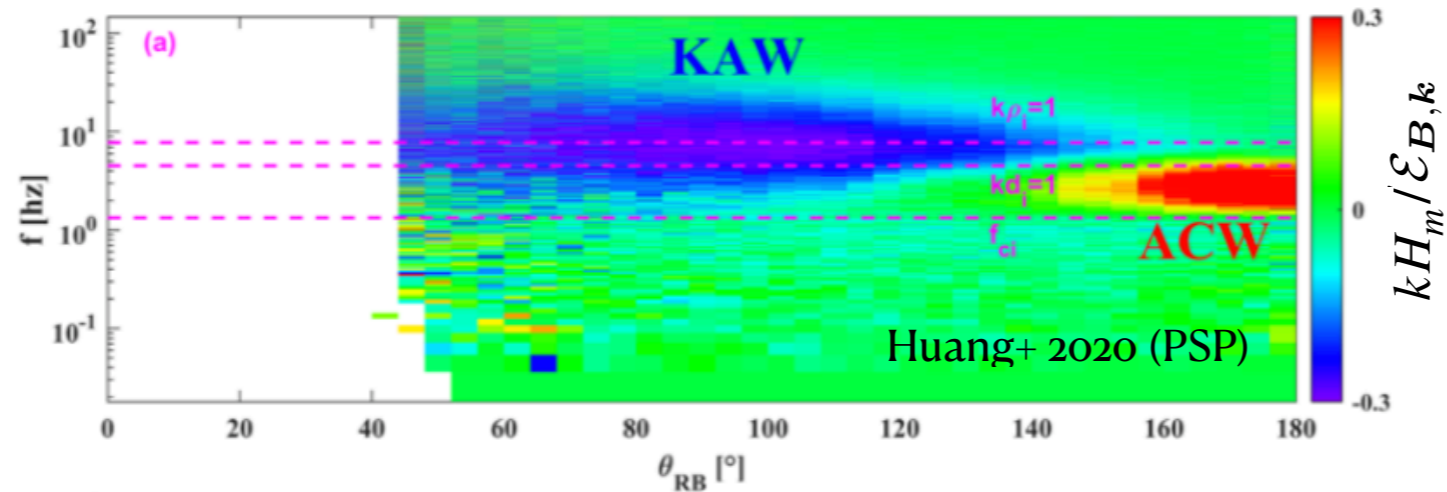
- Steeper transition range in larger amplitude, lower β , fast wind (higher imbalance)
- **Predictions:** *no transition range in balanced turbulence*
correlation of break position with imbalance

- Parallel spectra show extremely steep transition range
- Anisotropy scaling similar

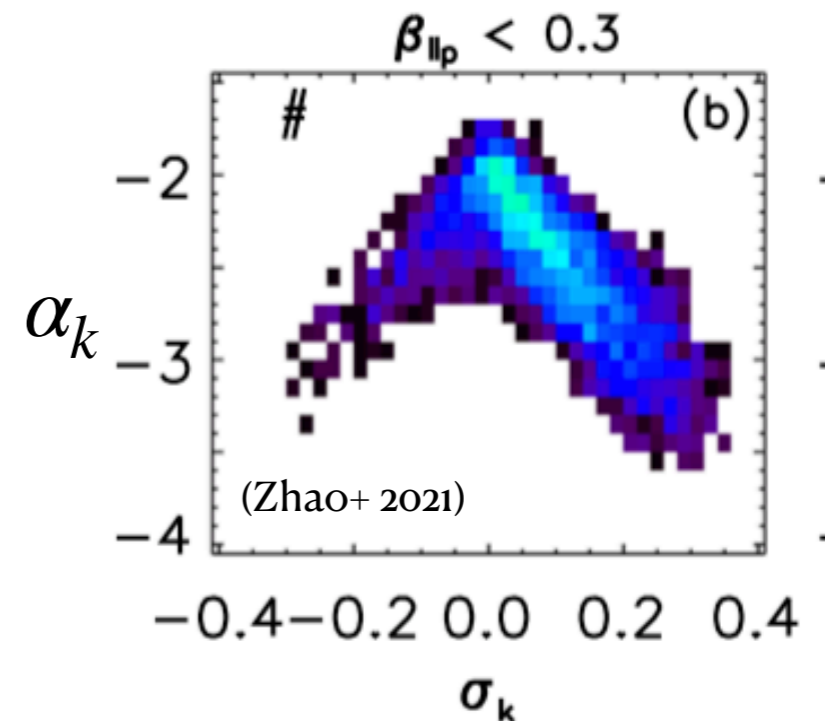


Observational Evidence

Magnetic spectra – helicity



Helicity correlated with transition-range slope

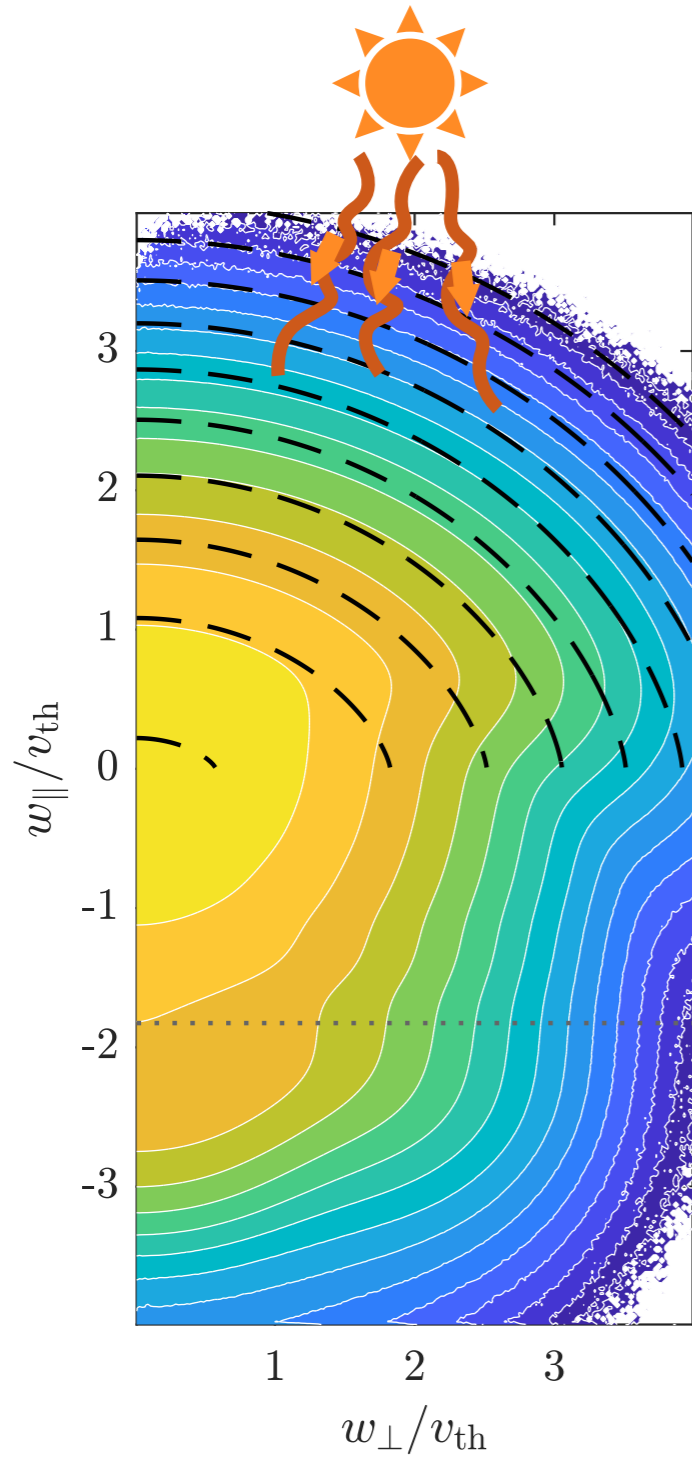


Natural explanation for
 $H_m \rightarrow 0$ at $k_{\perp}\rho_i > 1$
 e.g., Huang+2020, Woodham+2021 SH45A-2355

Physically: **balanced small-scale cascade**

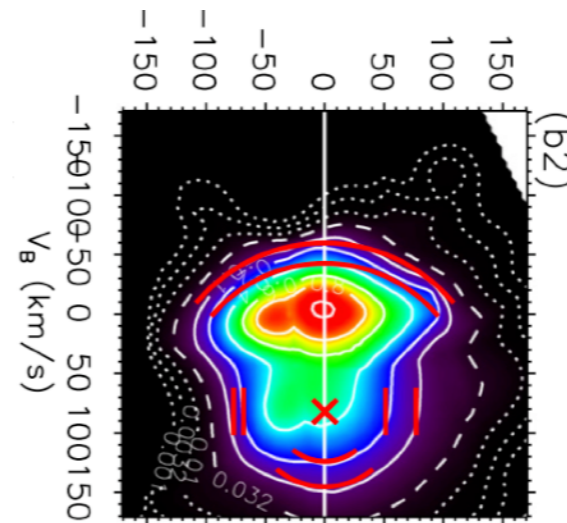
Observational Evidence

Distribution function

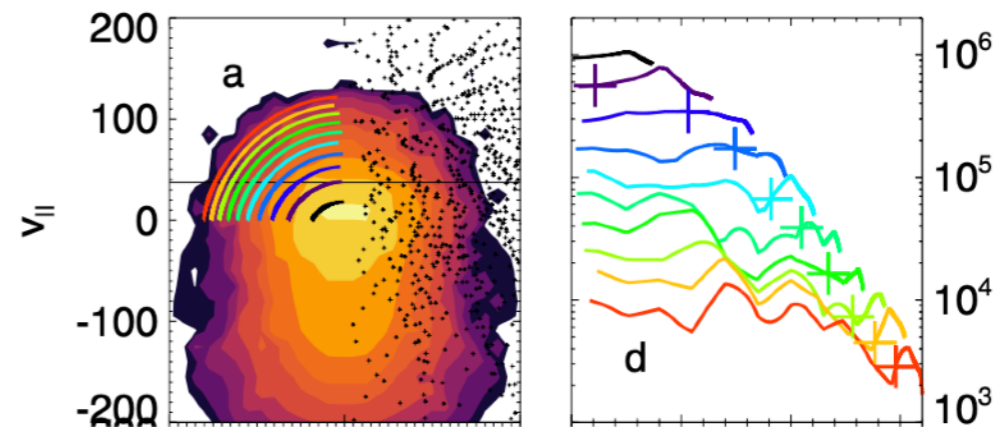


Helios 0.3AU (Marsch+ 1982)

WIND (He+ 2015)



PSP (Bowen+ 2021)



Beam: velocity $\simeq 1.2 - 1.5v_A$
 direction matches AW direction (He+ 2015b)

