

# PERCOLATING COSMIC STRING NETWORKS FROM KINATION

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String Pheno 2024  
Padua



(Talk based on JC, Copeland, Hardy, [Sanchez-Gonzales](#) 2406.12637,  
also see [Apers](#), JC, Copeland, [Mosny](#), [Revello](#) 2401.04064)

Parallel  
Thursday  
2.30pm



Parallel  
Thursday  
2.45pm

# WHERE IS THE CENTRE OF THE WORLD?

- N=0 supersymmetry
- Hierarchies
- Weak couplings

**Strong coupling**  $AdS_5 \times S^5$

$$g_s \sim 1, \text{Volume} \sim l_s^6$$

**Dualities** **Black hole entropy**  
 $\mathcal{N} \geq 2$  Supersymmetry

**Landau-Ginzburg models**

# OUR HOME, THE UNIVERSE

- Our universe is *filled* with hierarchies and small numbers

$$\frac{\Lambda_{EW}}{M_P} \sim 10^{-16}$$

$$\frac{\delta\rho_{CMB}}{\rho} \sim 10^{-5}$$

$$\Lambda_{cc} \sim 10^{-120} M_P^4$$

$$\alpha_{SU(3)} \sim \frac{1}{11}, \alpha_{SU(2)} \sim \frac{1}{30}, \alpha_{U(1)_Y} \sim \frac{1}{60}$$

$$y_e \sim 10^{-5}, y_\mu \sim 10^{-3}, y_\tau \sim 10^{-2}$$

$$m_\nu \sim 10^{-3} \text{eV}$$

$$\theta_{QCD} \lesssim 10^{-10}$$

# OUR HOME, THE UNIVERSE

- The true string vacuum is the vacuum of *this* universe
- It must contain a method to generate hierarchies, small couplings and small numbers
- This makes the asymptotic boundaries of moduli space appealing

# ASYMPTOTICS OF MODULI SPACE

Cosmology

Supersymmetry Breaking

AdS Vacua - LVS, DGKT

Moduli and Standard Model spectra

Axions

Dark Energy/Quintessence

Many exciting talks this week - Lüst, Halverson,  
Pedro, Vafa, Valenzuela, Blumenhagen,  
Marchesano, Parameswaran

# GETTING TO THE CENTRE OF THE WORLD

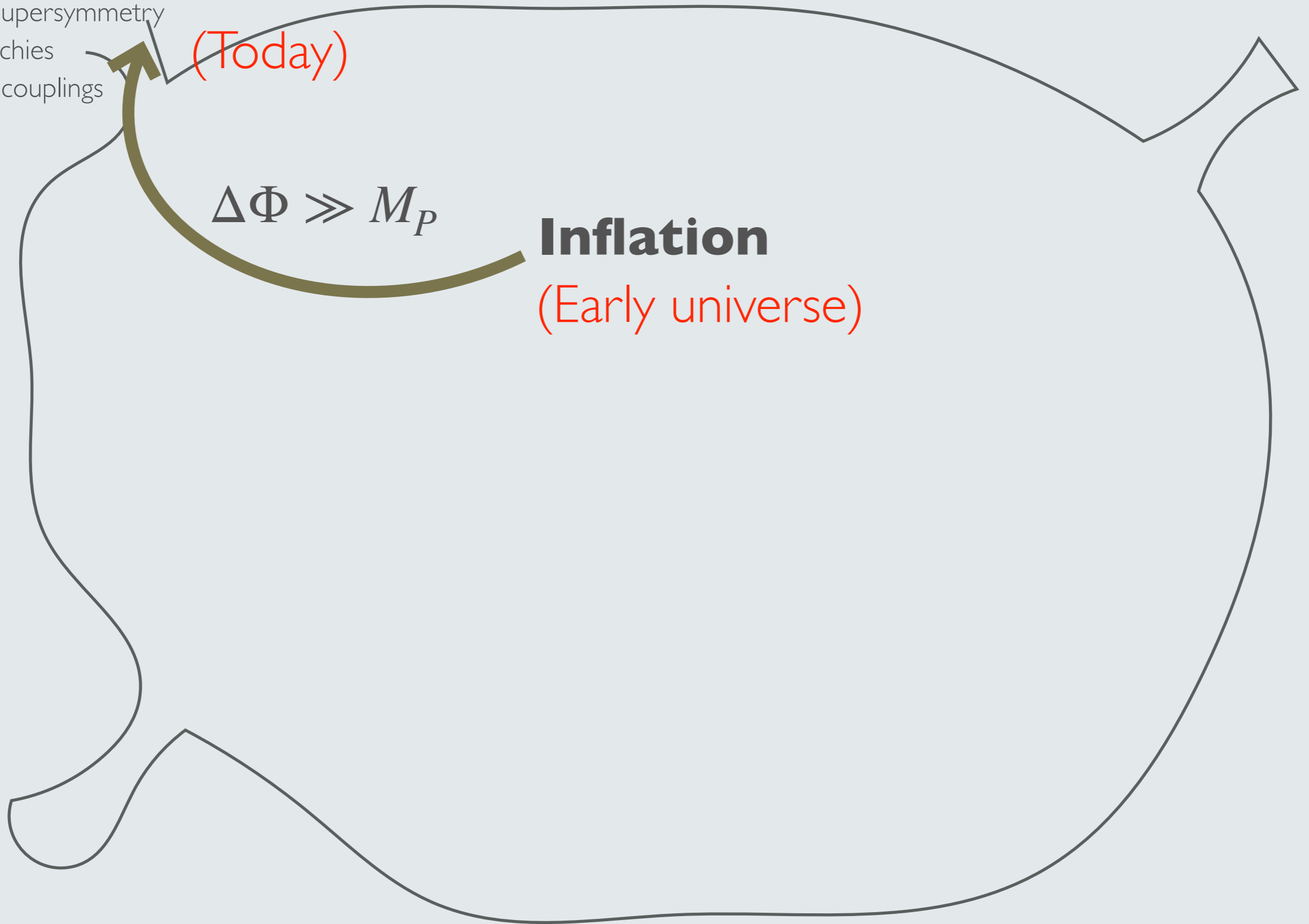
- N=0 supersymmetry
- Hierarchies
- Weak couplings

(Today)

$$\Delta\Phi \gg M_P$$

**Inflation**

(Early universe)



# GETTING TO THE END OF THE WORLD

- We assume we start with an inflationary model of <your favourite model> and end with a vacuum of <my favourite vacuum>
- Talk is about the middle: route from end of inflation to vacuum near the boundaries of moduli space

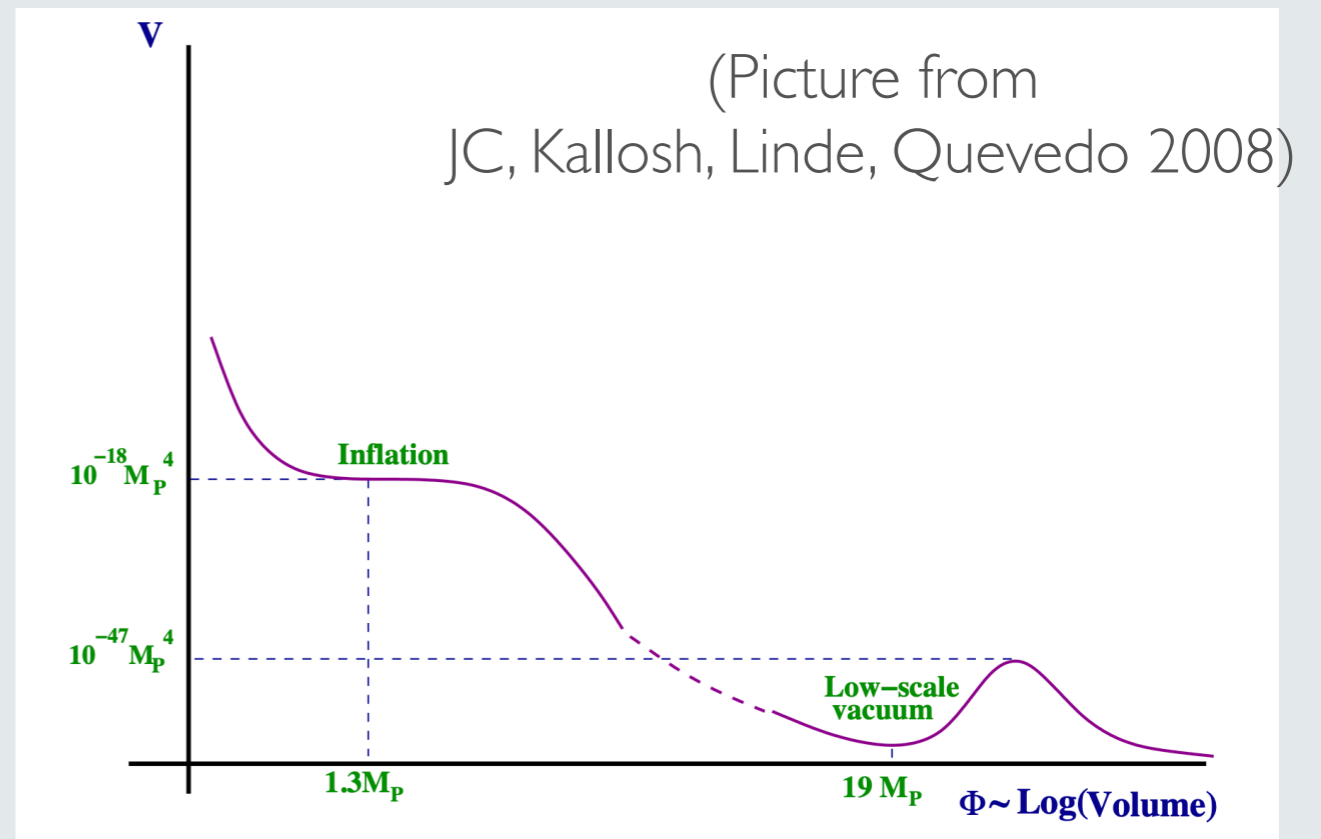
- After inflation, field starts rolling down exponential slope

$$V \sim V_0 \exp\left(-\sqrt{\frac{27}{2}} \frac{\Phi}{M_P}\right)$$

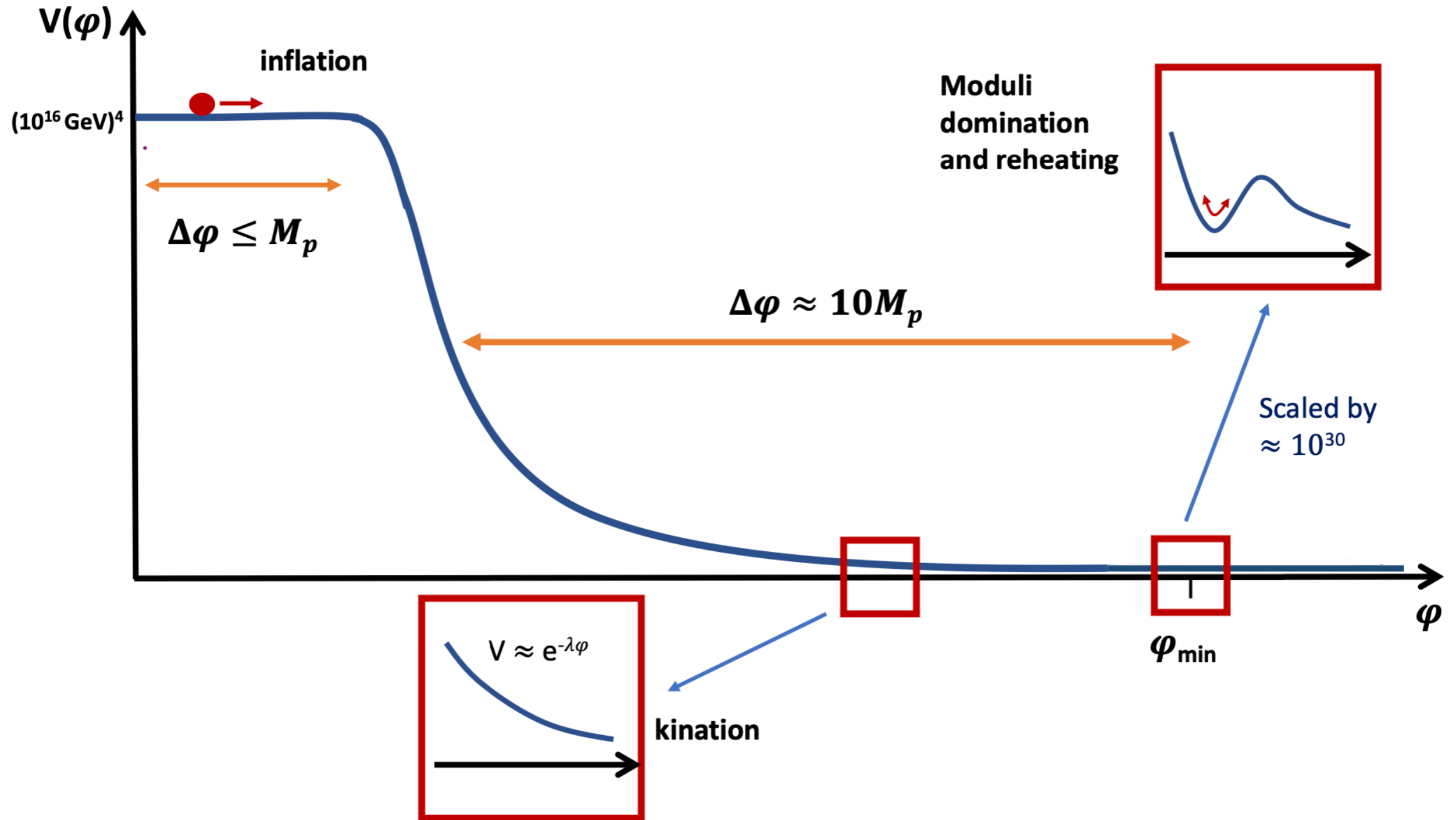
(exponent as in LVS, but not material)

- Universe enters a *kination* epoch

$$a(t) \sim t^{1/3}, \quad \rho_{KE} \sim \frac{1}{a(t)^6}, \quad \rho_\gamma \sim \frac{\epsilon}{a(t)^4}$$

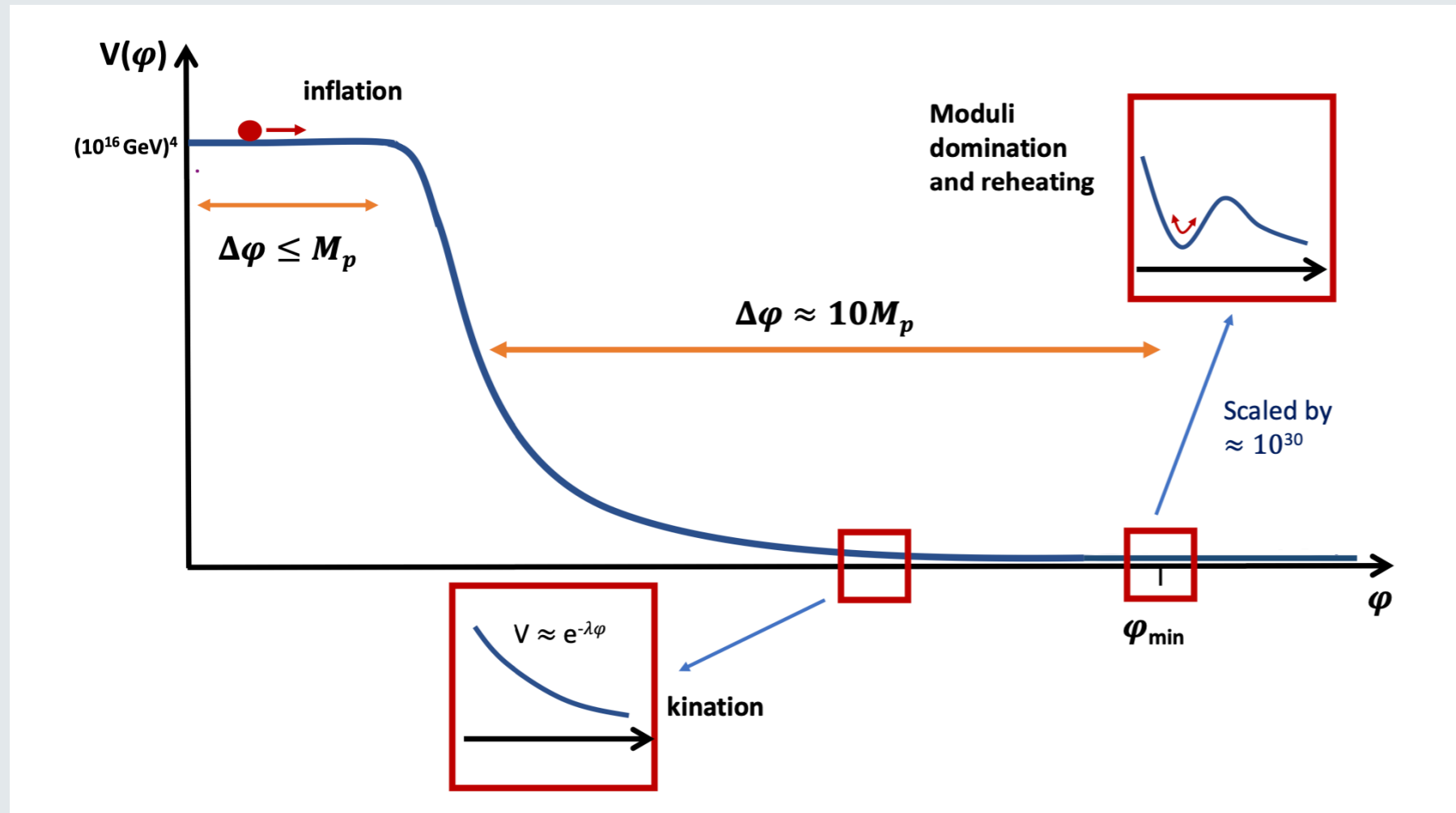


# GETTING TO THE END OF THE WORLD

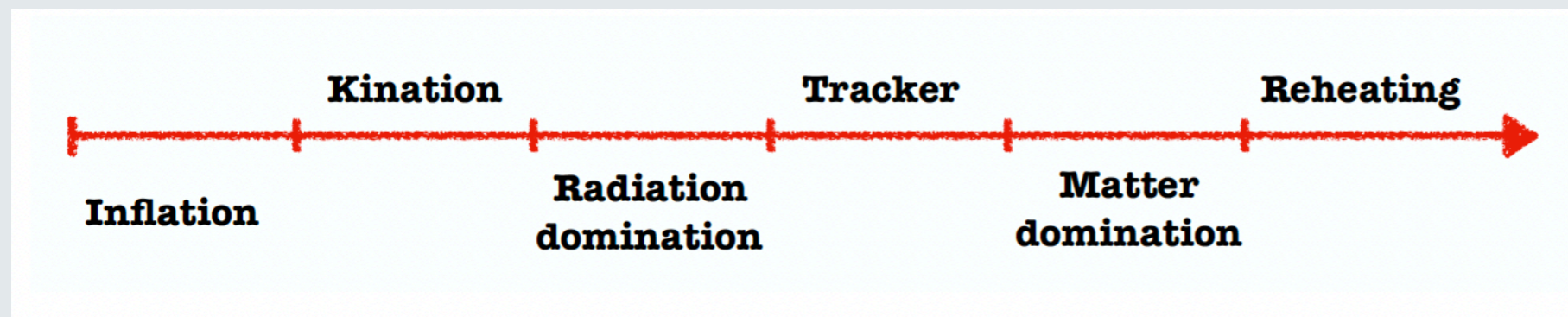




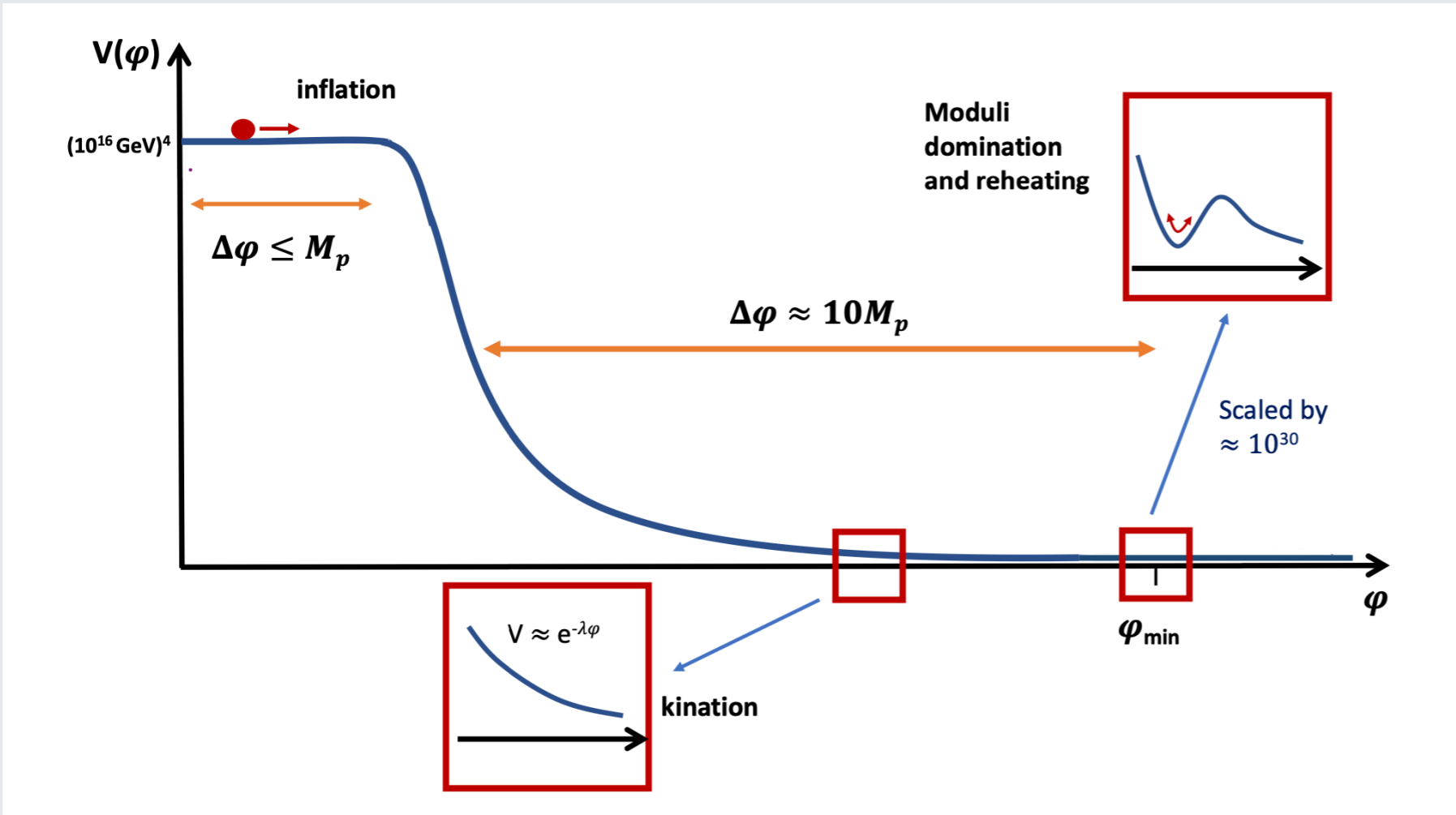
# NOVEL COSMOLOGICAL HISTORY



This motivates a distinctive 'stringy' cosmological history quite distinct from the normal assumption of radiation domination after inflation

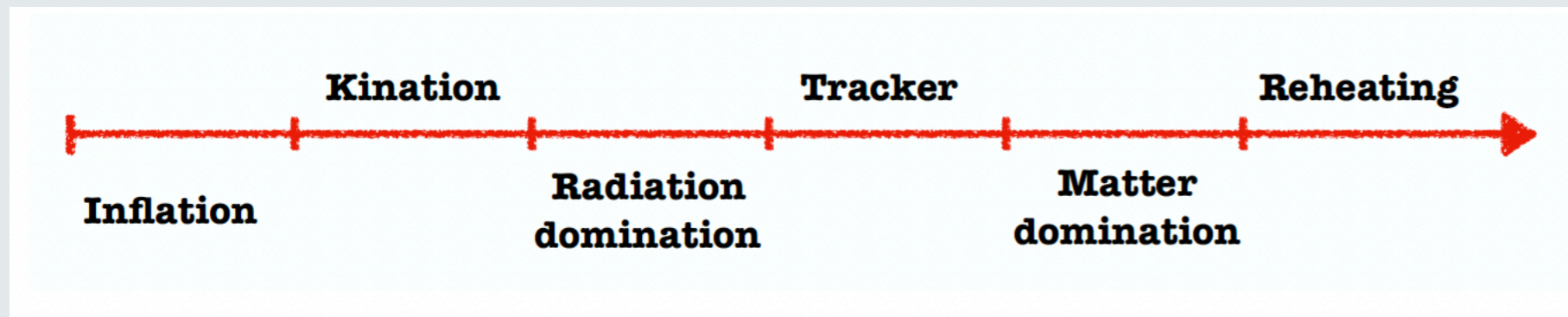


# NOVEL COSMOLOGICAL HISTORY



First half of the universe (on a log scale) is unconstrained: great opportunity for string theory

Many aspects discussed in Thursday parallel talks by **Fien Apers** and **Filippo Revello**



# KINATION EPOCHS

- During kination epoch, kinating field evolves as

$$\Phi(t) = \Phi_0 + \sqrt{\frac{2}{3}} M_P \ln \left( \frac{t}{t_0} \right)$$

- Field moves through  $\sim M_P$  in field space each Hubble time

Long kination epoch implies large transPlanckian field excursions

- String theorists should **care!** - trans-Planckian field excursions

$\Delta\Phi \gg M_P$  is home territory

- Novel cosmology: **real opportunities for string phenomenology**

My talk:  
Cosmic Strings

# DYNAMICS OF COSMIC (SUPER)STRINGS

- Cosmic (super)strings long-studied candidate for new stringy cosmologies

Kibble 76 Witten 85

Brandenberger+Vafa 86 Sarangi+Tye 02 Copeland+Polchinski 05

- Dynamics of closed strings set by Nambu-Goto action in fixed spacetime background,  $\mu$  is the string tension

$$S_{NG} = - \int d^2\xi \mu \sqrt{-\gamma}$$

- What are the dynamics? (assuming stability and FRLW metric)

$$ds^2 = dt^2 - a(t)^2 (dx^2 + dy^2 + dz^2)$$

# DYNAMICS OF COSMIC (SUPER)STRINGS (FIXED TENSION)

- Equations of motion follow from NG action

$$x_{,a}^{\nu;a} + \Gamma_{\beta\rho}^{\nu}(g)\gamma^{ad}x_{,d}^{\beta}x_{,a}^{\rho} = 0$$

- Focus on **circular string loops**

$$X^{\mu}(t, \sigma) = R(t)(\cos \sigma, \sin \sigma, 0)$$

- Study equations of motion in FLRW background (gauge choice identifies worldsheet and spacetime time)

# DYNAMICS OF COSMIC (SUPER)STRINGS (FIXED TENSION)

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$$X^\mu(t, \sigma) = R(t) (\cos \sigma, \sin \sigma, 0)$$

- Equations of motion are

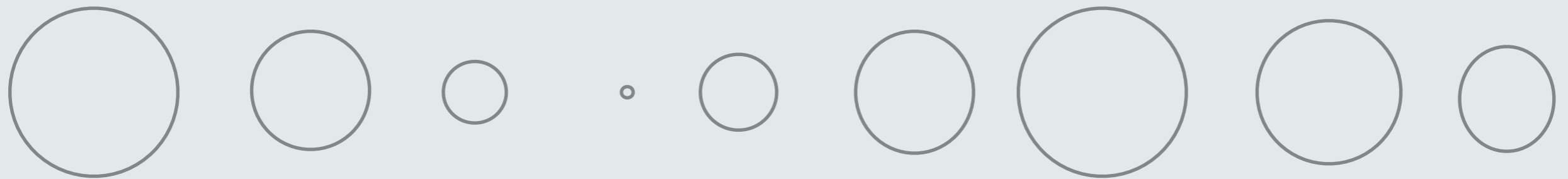
$$\left( \varepsilon = \sqrt{\frac{a^2 R^2}{1 - a^2 \dot{R}^2}} \equiv a R_{\max} \right)$$

$$\frac{\dot{\varepsilon}}{\varepsilon} = H - 2a^2 \dot{R}^2 H \quad \langle a^2 \dot{R}^2 \rangle = 1/2$$

$$\ddot{R} + H\dot{R} + \varepsilon^{-2}R + 2H(1 - a^2 \dot{R}^2)\dot{R} = 0$$

- Loops oscillate with a fixed maximum (physical) size  $R_{\max}$

# EVOLUTION OF A LOOP



Circular string loops oscillate in and out back on themselves at **constant physical radius** and **shrink in comoving coordinates**

More complicated exact solutions also exist (**Burden, Kibble+Turok**)

Loops are left behind as universe expands  
(and gradually decay by emission of gravitational waves)

# DYNAMICS OF COSMIC (SUPER)STRINGS (VARYING TENSION)

- Equations of motion follow from NG action

$$x_{,a}^{\nu;a} + \Gamma_{\beta\rho}^{\nu}(g)\gamma^{ad}x_{,d}^{\beta}x_{,a}^{\rho} + \frac{\mu_{,\rho}}{\mu}\gamma^{ab}x_{,a}^{\rho}x_{,b}^{\nu} - \frac{\mu^{,\nu}}{\mu} = 0,$$

- Focus on **circular string loops**

$$X^{\mu}(t, \sigma) = R(t)(\cos \sigma, \sin \sigma, 0)$$

- Study equations of motion in **kinating** FLRW background  
(gauge choice identifies worldsheet and spacetime time)



# DYNAMICS OF COSMIC (SUPER)STRINGS (VARYING TENSION)

- Focus on **circular string loops**

$$X^\mu(t, \sigma) = R(t) (\cos \sigma, \sin \sigma, 0)$$

- Equations of motion are

$$\varepsilon = \sqrt{\frac{a^2 R^2}{(1 - a^2 \dot{R}^2)}} \equiv a R_{\max}$$

$$\frac{\dot{\varepsilon}}{\varepsilon} = H - a^2 \dot{R}^2 \left( 2H + \frac{\dot{\mu}}{\mu} \right) \quad \langle a^2 \dot{R}^2 \rangle = 1/2$$

$$\ddot{R} + H\dot{R} + \varepsilon^{-2} R + \left( 2H + \frac{\dot{\mu}}{\mu} \right) (1 - a^2 \dot{R}^2) \dot{R} = 0$$

- High-frequency oscillation at (physical) amplitude  $R_{\max}$  **but....**

# DYNAMICS OF COSMIC (SUPER)STRINGS (VARYING TENSION)

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$$\varepsilon = \sqrt{\frac{a^2 R^2}{(1 - a^2 \dot{R}^2)}} \equiv a R_{\max}$$

$$\frac{\dot{\varepsilon}}{\varepsilon} = H - a^2 \dot{R}^2 \left( 2H + \frac{\dot{\mu}}{\mu} \right) \quad \langle a^2 \dot{R}^2 \rangle = 1/2$$

- Decreasing tension causes loops to grow with cosmic time
- Right hand side of equation determines precisely how loops compared to scale factor (cf  $\frac{\dot{a}}{a} = H$ )

# DYNAMICS OF COSMIC (SUPER)STRINGS (VARYING TENSION)

- Equations of motion are

$$\varepsilon = \sqrt{\frac{a^2 \dot{R}^2}{(1 - a^2 \dot{R}^2)}} \equiv a R_{\max}$$

$$\frac{\dot{\varepsilon}}{\varepsilon} = H - \cancel{a^2 \dot{R}^2 \left( 2H + \frac{\dot{\mu}}{\mu} \right)} \quad \langle a^2 \dot{R}^2 \rangle = 1/2$$

- Decreasing tension causes loops to grow with cosmic time
- If crossed out term vanishes, oscillating loops grow precisely with the scale factor (cf  $\frac{\dot{a}}{a} = H$ )

# DYNAMICS OF COSMIC (SUPER)STRINGS (VARYING TENSION)

- Equations of motion are

$$\varepsilon = \sqrt{\frac{a^2 R^2}{(1 - a^2 \dot{R}^2)}} \equiv a R_{\max}$$

$$\frac{\dot{\varepsilon}}{\varepsilon} = H - a^2 \dot{R}^2 \left( 2H + \frac{\dot{\mu}}{\mu} \right) \quad \langle a^2 \dot{R}^2 \rangle = 1/2$$

- When  $2H + \frac{\dot{\mu}}{\mu} < 0$ , oscillating loops grow faster than the scale factor **and will percolate given enough time**

# KINATION AND TIME-VARYING TENSION

- We want to make  $2H + \frac{\dot{\mu}}{\mu}$  as negative as possible
- This requires
  - (a)  $H \equiv \frac{\dot{a}}{a}$  as small as possible
  - (b)  $\frac{\dot{\mu}}{\mu}$  as large and negative as possible
- Kination epochs are ideal as
  - (a)  $a(t) \sim t^{1/3}$  and so growth is as slow as any other fluid
  - (b) All energy is in kinetic evolution of a modulus and so maximises rate of change of tension vev

# KINATION AND TIME-VARYING TENSION

- During volume modulus kination, volume grows with time

$$\frac{\mathcal{V}}{\mathcal{V}_0} = \frac{t}{t_0}$$

- For superstrings,  $G\mu \sim m_s^2$  and so  $\mu \propto t^{-1}$  using standard relationship

$$m_s \sim \frac{M_P}{\sqrt{\mathcal{V}}}$$

- It follows that

$$2H + \frac{\dot{\mu}}{\mu} = -H < 0$$

and so loops of fundamental strings grow faster than the scale factor!

- Loops of fundamental strings **grow in comoving coordinates** and can percolate!

# PERCOLATING SUPERSTRINGS

- During kination, scale factor and loop radius grow as

$$a(t) \sim t^{1/3}$$
$$R_{max}(t) \sim t^{1/2}$$

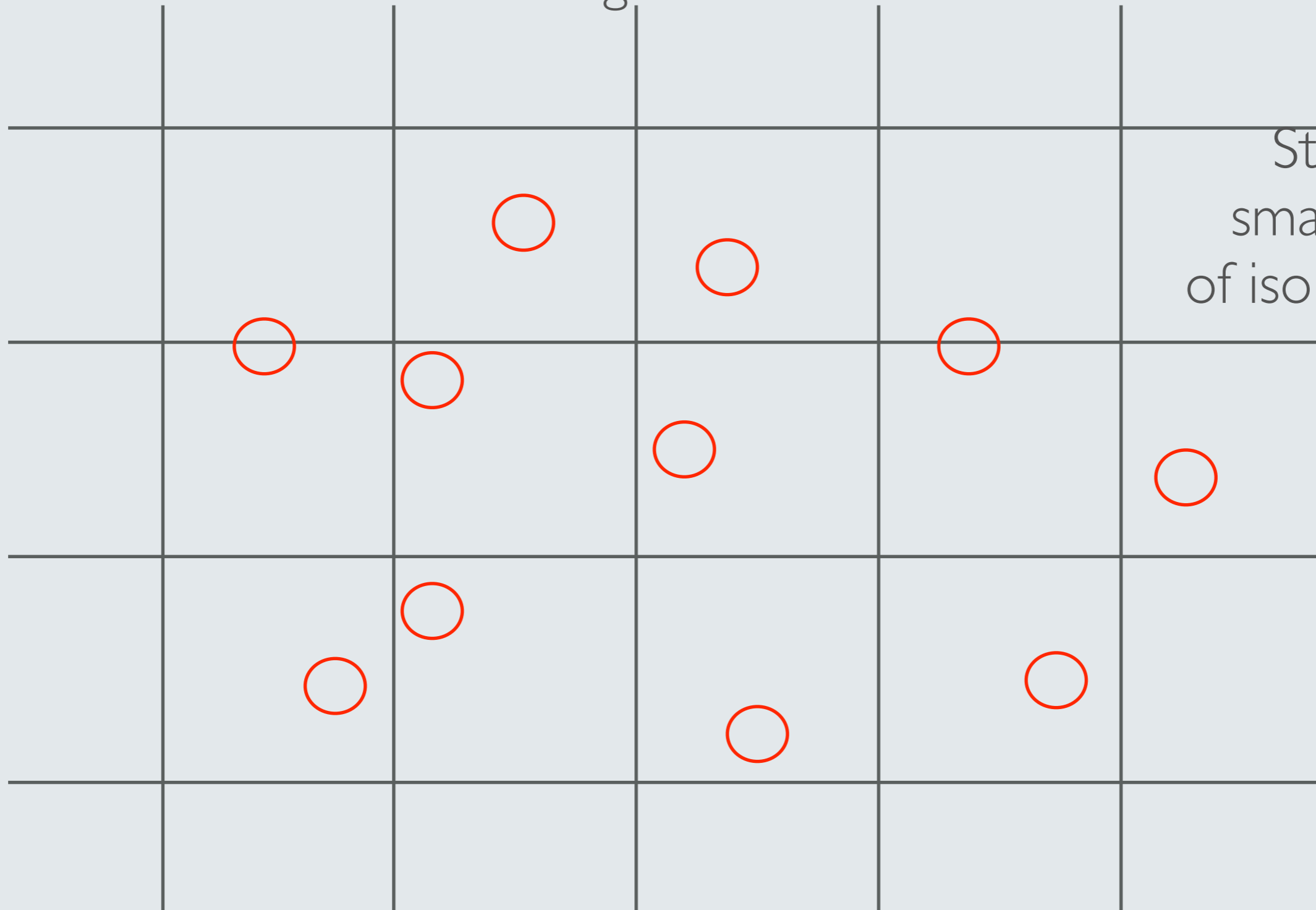
- In comoving coordinates,

$$R_{max,comoving} \sim t^{1/6} \sim \left( \frac{\mathcal{V}_f}{\mathcal{V}_i} \right)^{1/6}$$

- Long kination epochs (closely tie to vacua in asymptotic region of moduli space) essential to give percolation

# PERCOLATING SUPERSTRINGS

Fixed comoving coordinates

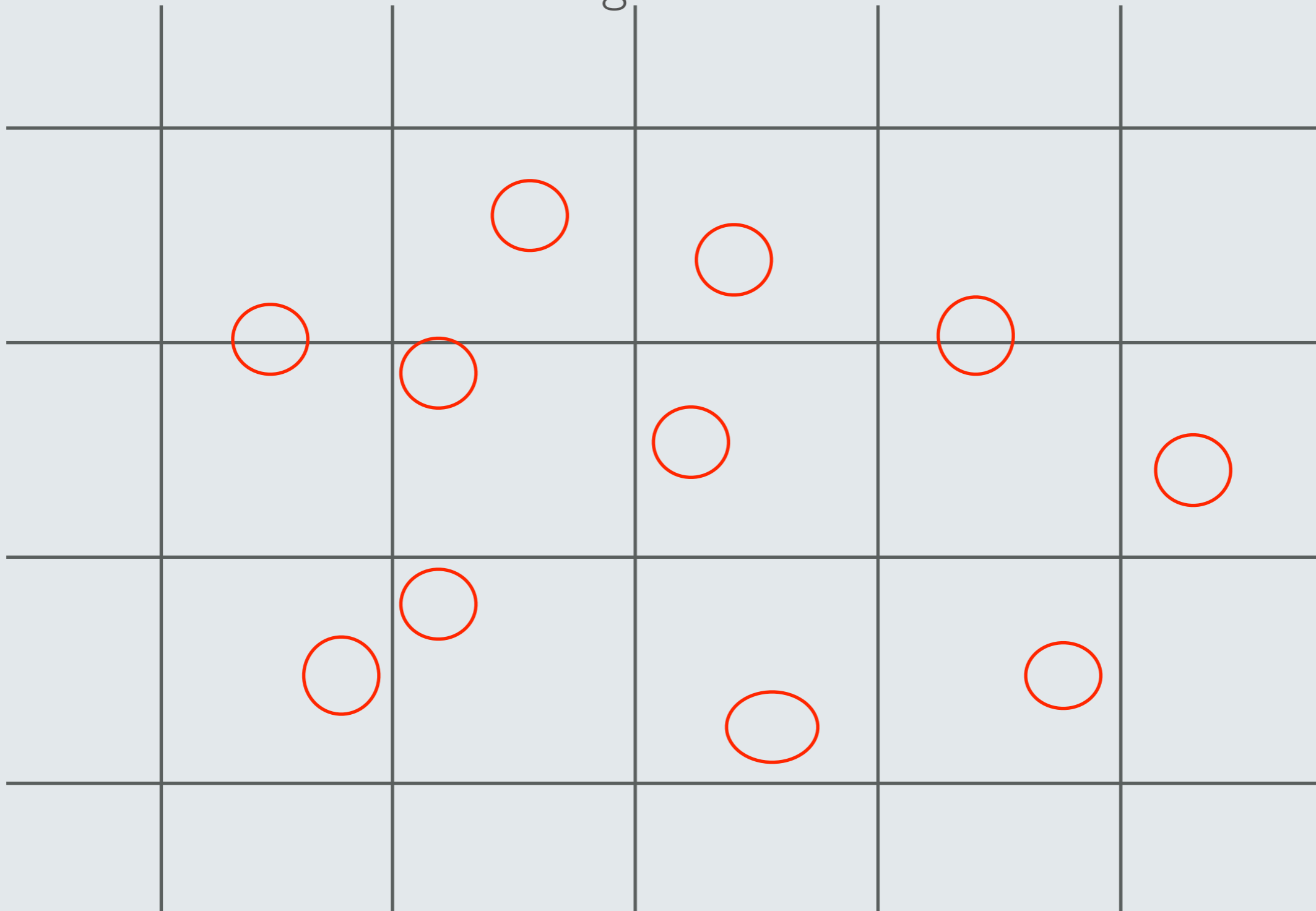


Start with  
small number  
of isolated strings



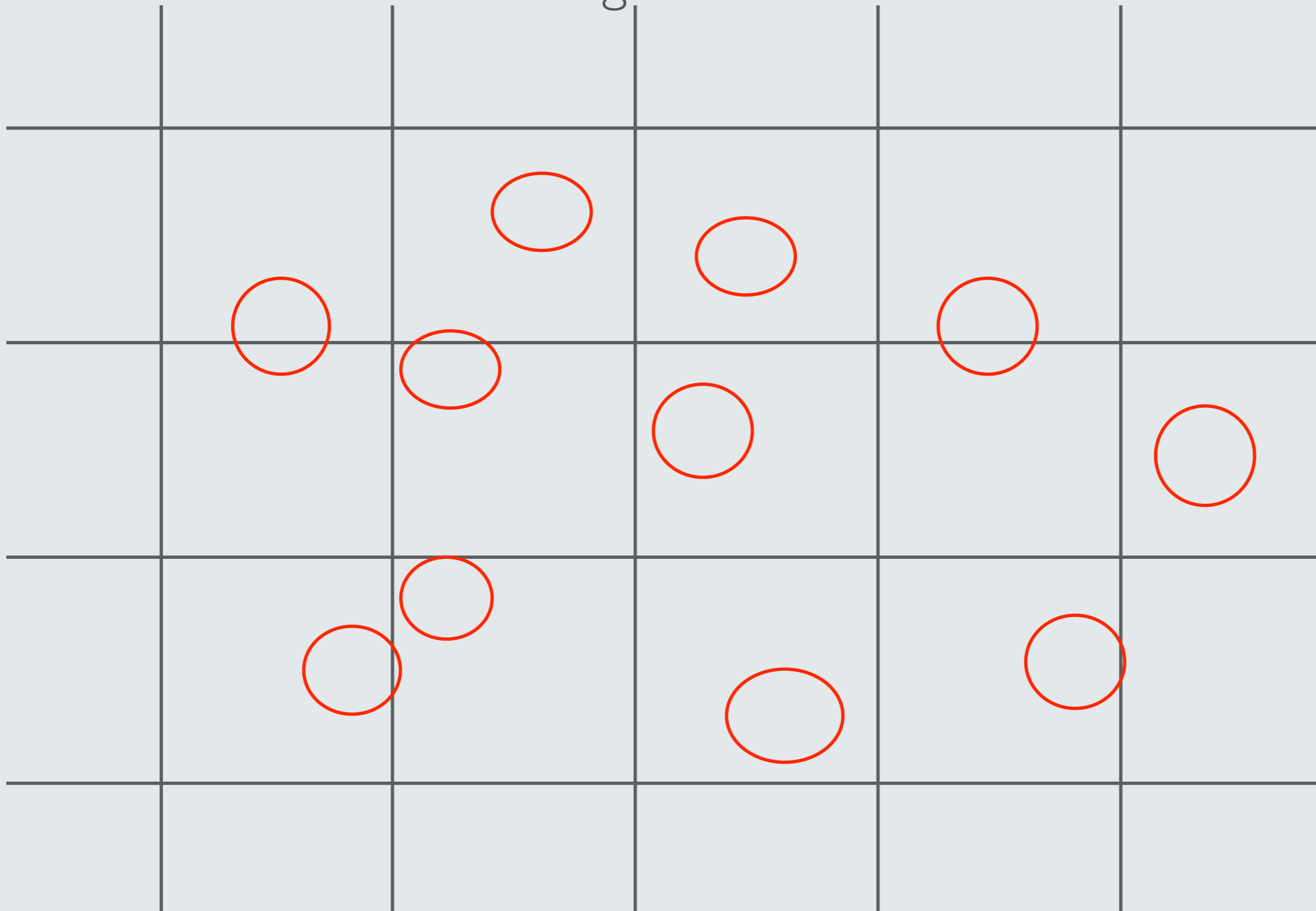
# PERCOLATING SUPERSTRINGS

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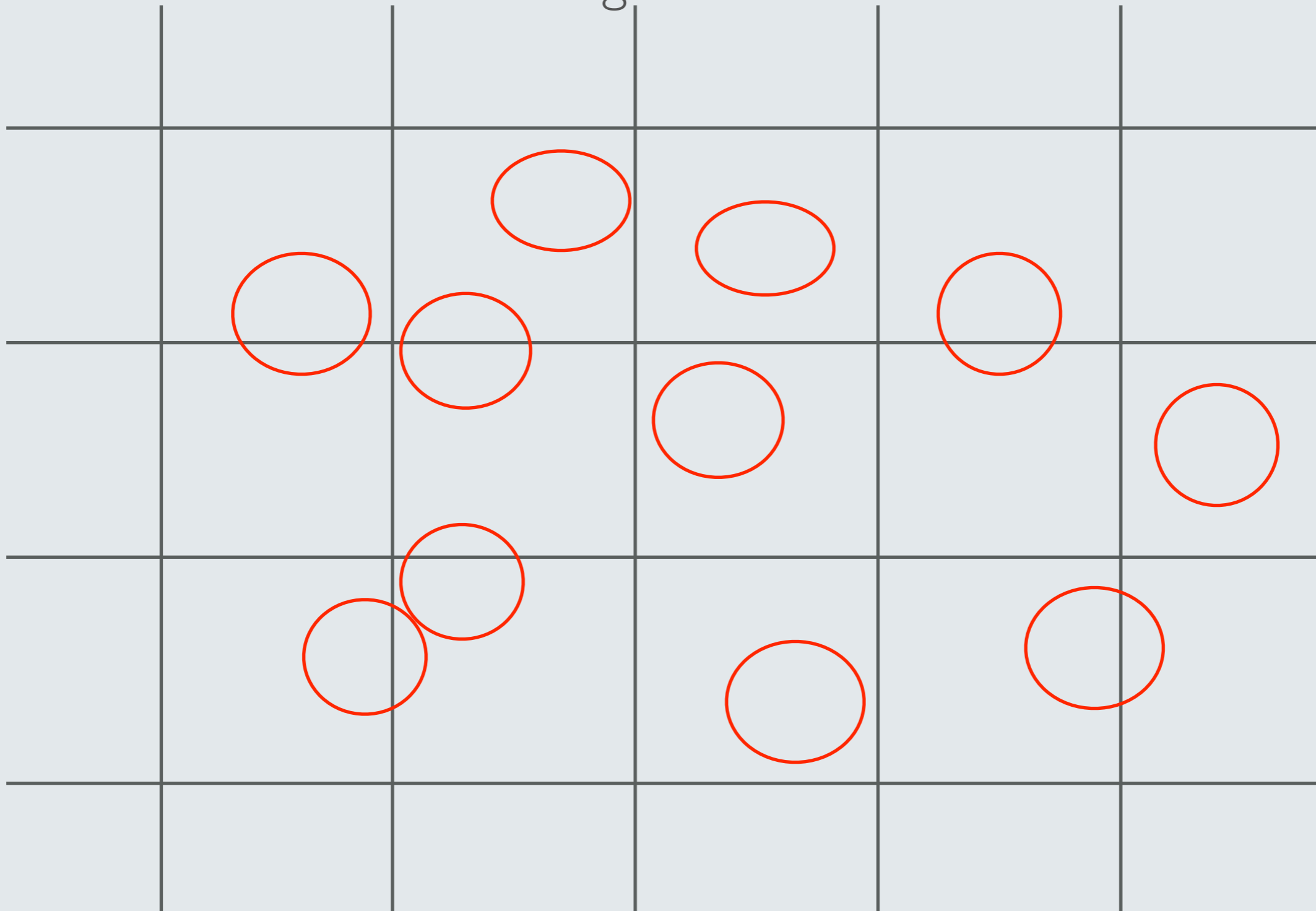
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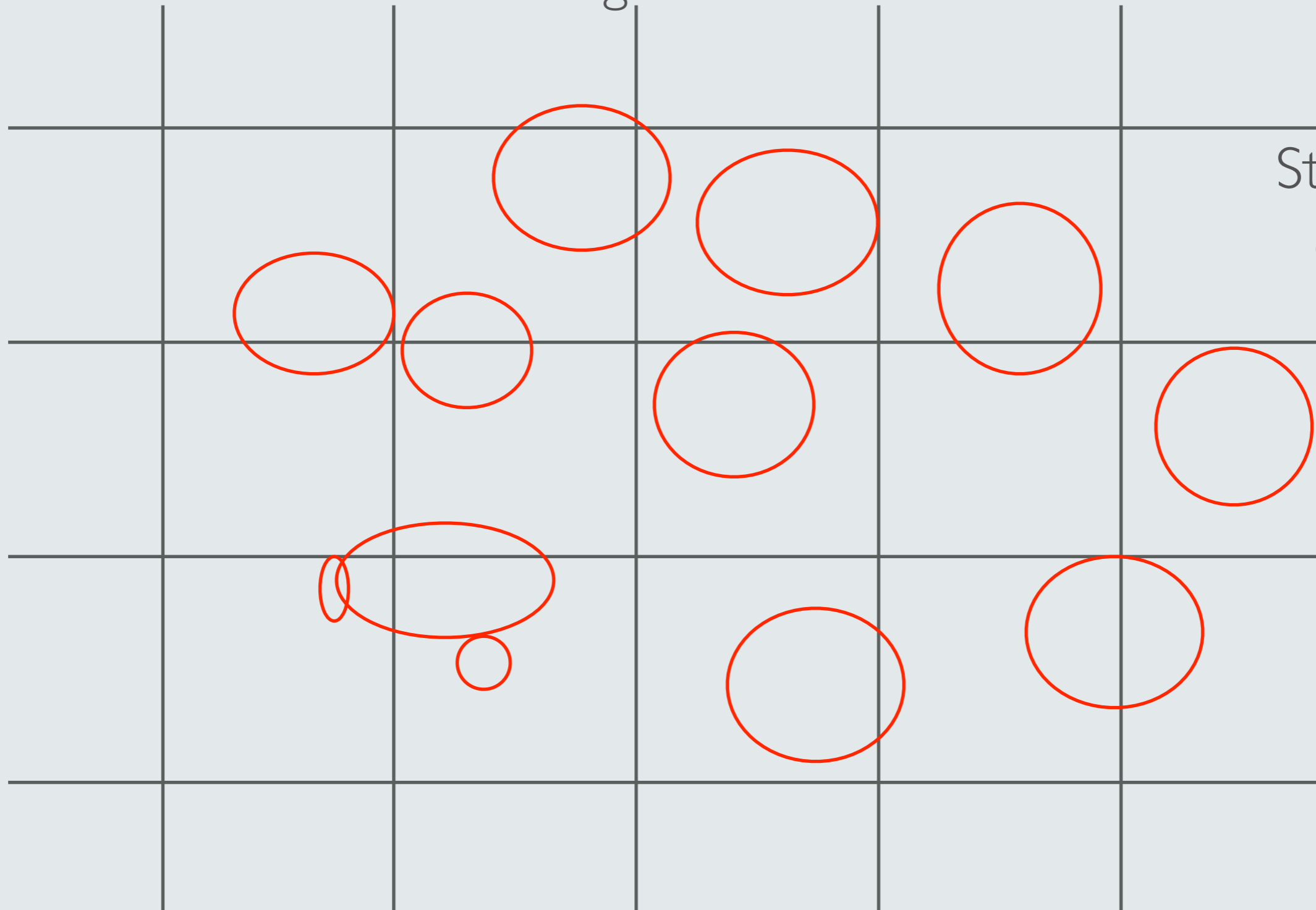
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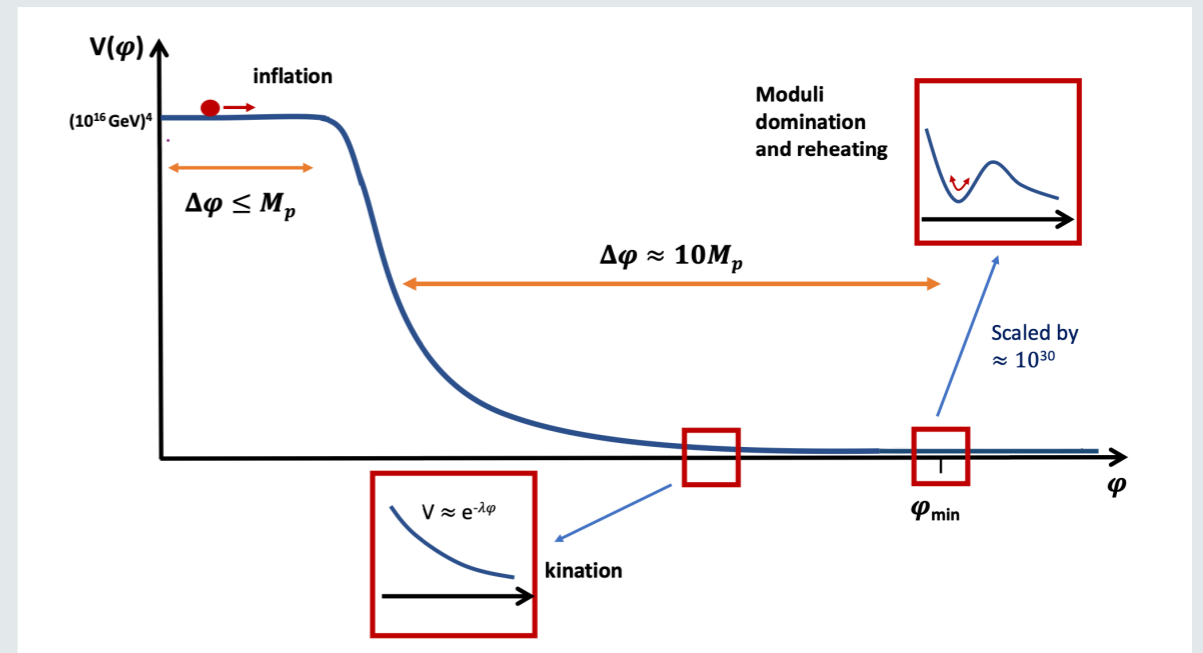
Strings meet  
and form  
network

# PERCOLATING SUPERSTRINGS

- Percolation ends once kination ends

(a) background reaches tracker solution as radiation catches up with kination

(b) kinating modulus settles down in its final minimum



- String networks enter a scaling regime  $\rho_{strings,init} \sim \mu H^2$  and lose energy via emission of gravitational radiation

- Final tension has  $\mu \sim m_s^2$  with  $G\mu \sim \frac{m_s^2}{M_p^2} \sim 10^{-10}$  in for

phenomenologically appealing vevs for volume.

# WILL NOT TALK ABOUT

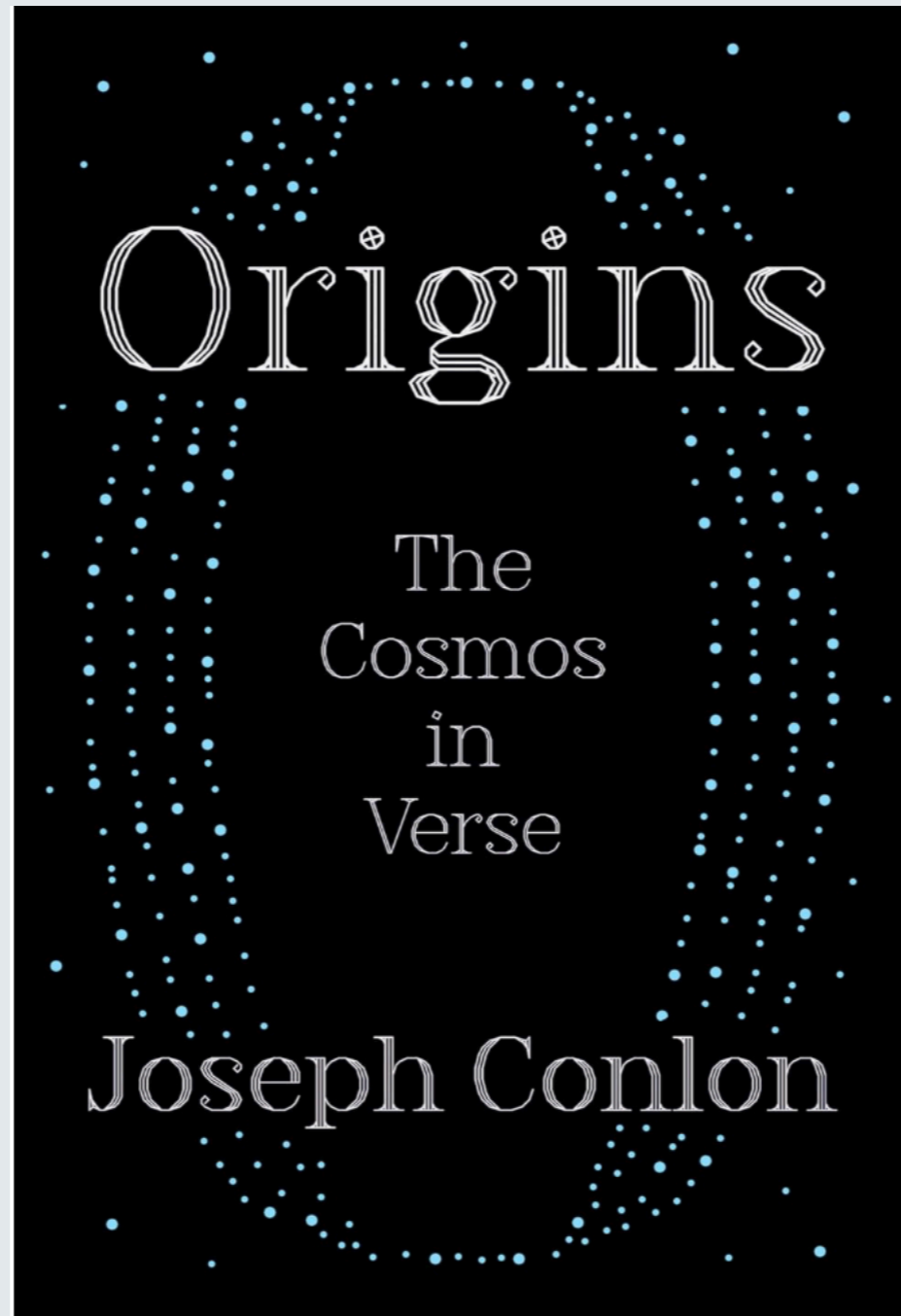
- Conditions for overall stability of cosmic strings against immediate fragmentation (Copeland + Polchinski 2005)
- Initial conditions: we need a starting population of isolated small loops (from quantum nucleation? Brane inflation?)
- Gravitational wave emission from closed loops (see paper; too slow to affect loop growth)
- Numerical details of endpoint of percolation
- Effects of  $2H + \frac{\dot{\mu}}{\mu} < 0$  and loop growth on any existing string network



# CONCLUSIONS

- If *String Phenomenology* means understanding this universe, the boundaries of moduli space are natural places to live: contains interesting vacua (LVS, DGKT)
- Reaching such vacua involves **long kination epochs**
- During volume kination, fundamental string loops **grow and can percolate** to form a cosmic string network
- New mechanism (distinct from Kibble mechanism) to form string networks
- With LVS final vacuum, string network today with  $G\mu \sim 10^{-10}$
- Such a fundamental cosmic string network with  $10^{-7} \lesssim G\mu \lesssim 10^{-11}$  - in reach of upcoming experiments (cf NANOGrav)

# ADDENDUM: CHRISTMAS PRESENT



Oneworld Publications  
November 7th 2024