

# Quasiparticle anisotropic hydrodynamics

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Kent State University

**Primary References:** M. Alqahtani, M. Nopoush, R. Ryblewski, and MS  
[1703.05808](#) (to appear in PRL) and [1705.10191](#)

Canterbury Tales of Hot QFTs in the LHC Era  
St John's College, Oxford  
July 11, 2017



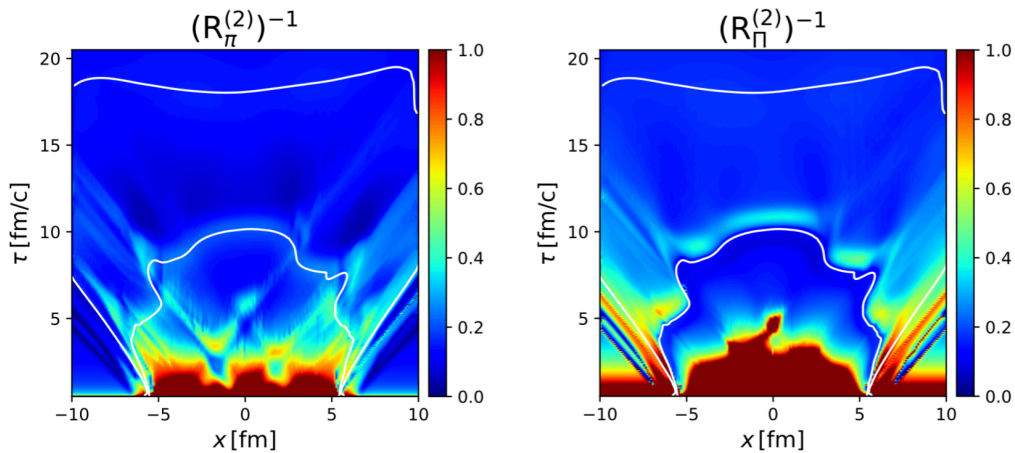
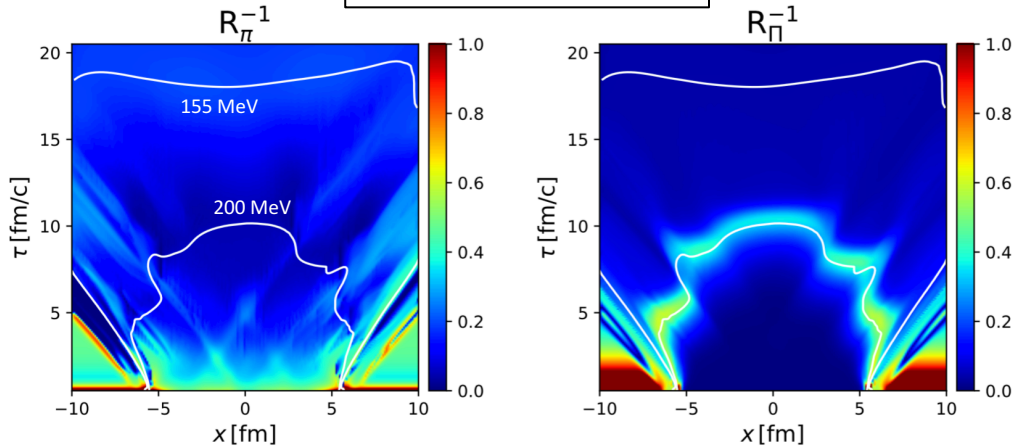
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# Motivation

- Viscous hydrodynamics is phenomenologically quite successful, however, the extreme environment generated in HICs presents a bit of a challenge to the standard formalism
- The **QGP** is born into a state of rapid longitudinal expansion which drives the system **out of equilibrium**
- There are many groups now focused on improving viscous hydrodynamics itself in order to better describe systems that are out of equilibrium, e.g. **anisotropic hydrodynamics (aHydro)**
- The goal of the aHydro program is to provide an optimized hydrodynamics(-like) framework that is **more accurate out of equilibrium**

# Pb-Pb @ 2.76 TeV - Don't worry, be happy?

$$R_{\pi}^{-1} \equiv \frac{\sqrt{\pi^{\mu\nu}\pi_{\mu\nu}}}{\mathcal{P}_0}, \quad R_{\Pi}^{-1} \equiv \frac{|\Pi|}{\mathcal{P}_0}.$$



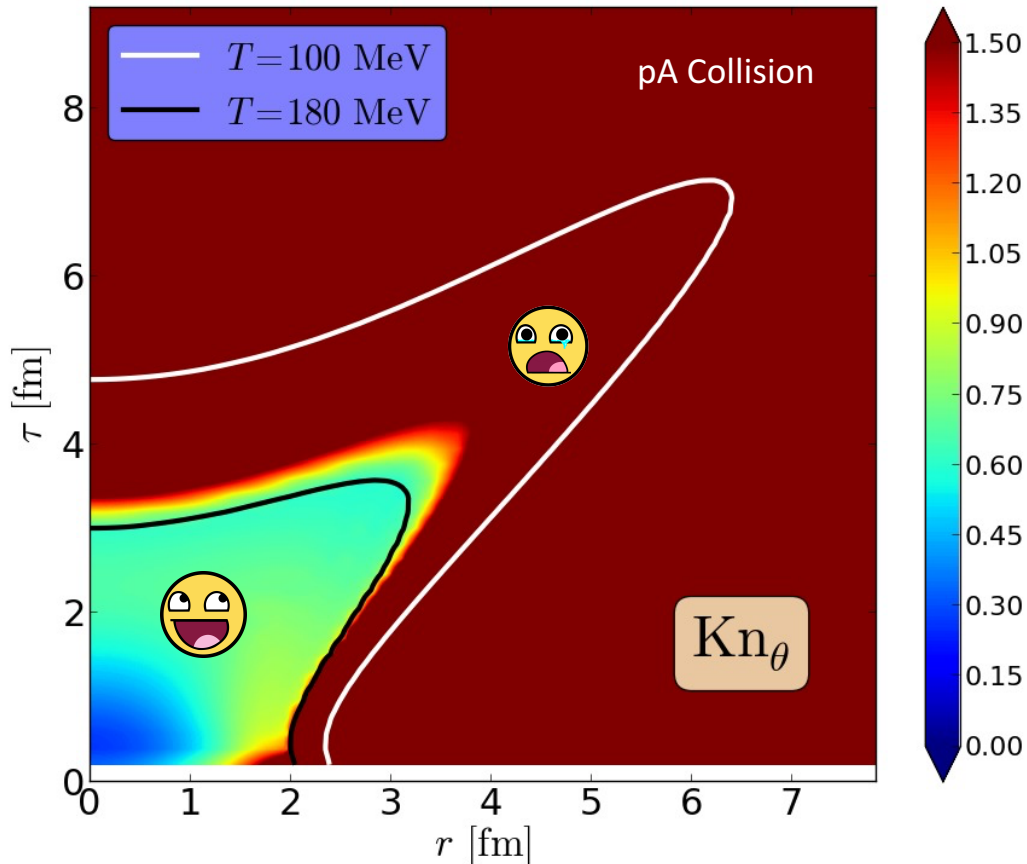
$$(R_{\pi}^{(2)})^{-1} \equiv \frac{\sqrt{\mathcal{J}^{\mu\nu}\mathcal{J}_{\mu\nu}}}{2\eta\sqrt{\sigma^{\mu\nu}\sigma_{\mu\nu}}}, \quad (R_{\Pi}^{(2)})^{-1} \equiv \frac{|\mathcal{J}|}{\zeta|\theta|}.$$

$$\begin{aligned} \tau_{\Pi}\dot{\Pi} + \Pi &= -\zeta\theta + \mathcal{J} + \mathcal{K} + \mathcal{R}, \\ \tau_n\dot{n}^{(\mu)} + n^{\mu} &= \kappa I^{\mu} + \mathcal{J}^{\mu} + \mathcal{K}^{\mu} + \mathcal{R}^{\mu}, \\ \tau_{\pi}\dot{\pi}^{(\mu\nu)} + \pi^{\mu\nu} &= 2\eta\sigma^{\mu\nu} + \mathcal{J}^{\mu\nu} + \mathcal{K}^{\mu\nu} + \mathcal{R}^{\mu\nu}. \end{aligned}$$

- $\mathcal{J}, \mathcal{J}^{\mu},$  and  $\mathcal{J}^{\mu\nu}$  are  $O(\text{Kn} R^{-1})$
- $\mathcal{K}, \mathcal{K}^{\mu},$  and  $\mathcal{K}^{\mu\nu}$  are  $O(\text{Kn}^2)$
- $\mathcal{R}, \mathcal{R}^{\mu},$  and  $\mathcal{R}^{\mu\nu}$  are  $O(R^{-2})$
- DNMR derivation assumes that  $\text{Kn} \sim R^{-1}$
- For this to be a reasonable approx, the 2<sup>nd</sup>-order terms should be smaller than the  $O(\text{Kn})$  Navier-Stokes terms
- **Secret:** In order for code to run stably, it is necessary to “dynamically regulate” the viscous corrections

# p-A @ 2.76 TeV - Don't be happy, worry!

Figure (sans emoticons): H. Niemi and G. Denicol, 1404.7327

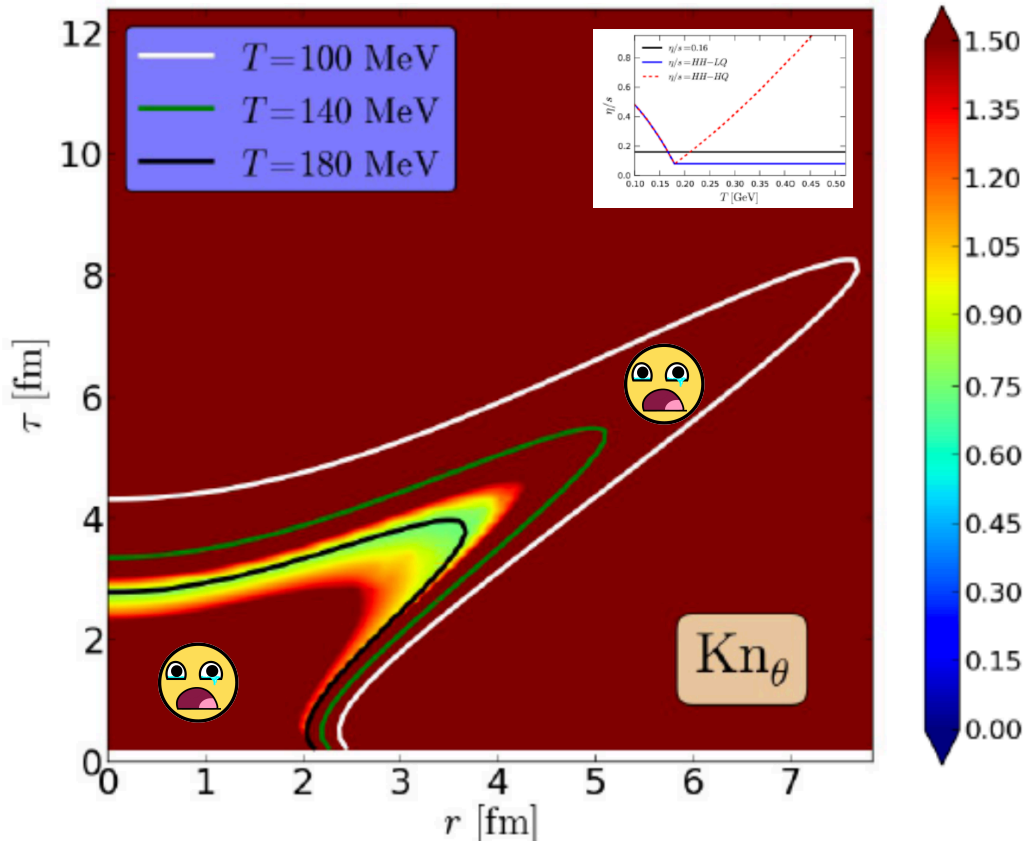


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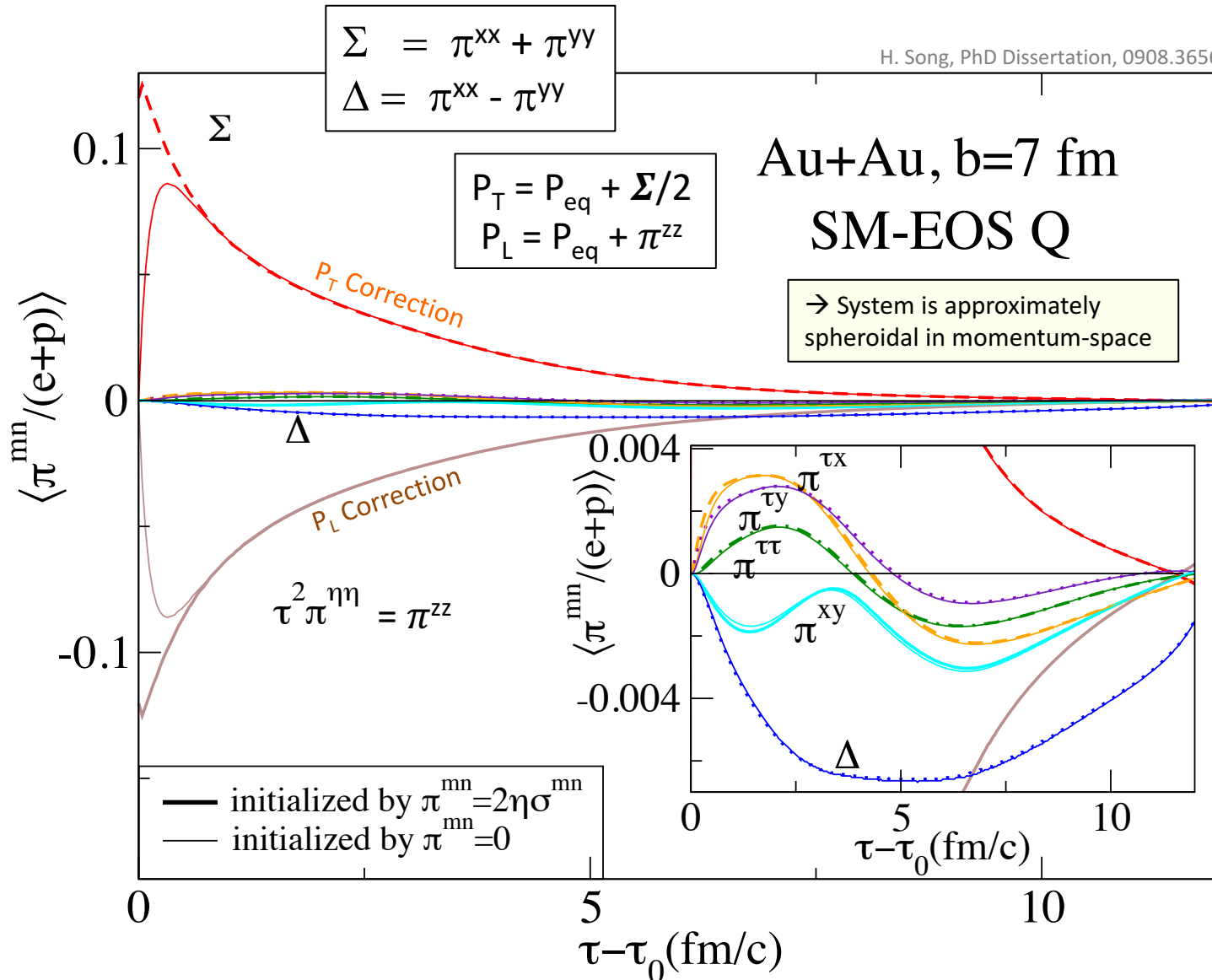


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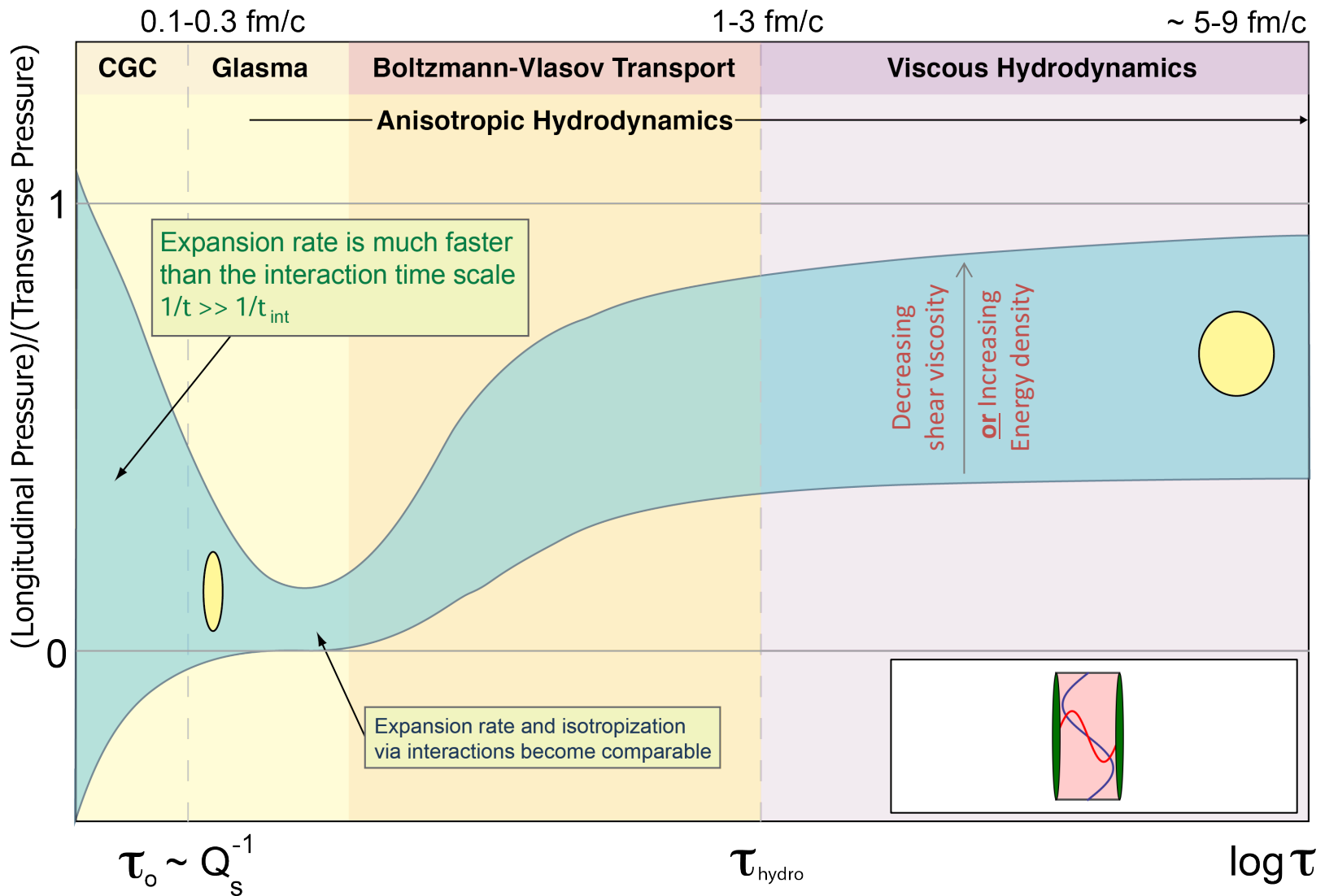
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# What are the largest viscous corrections?

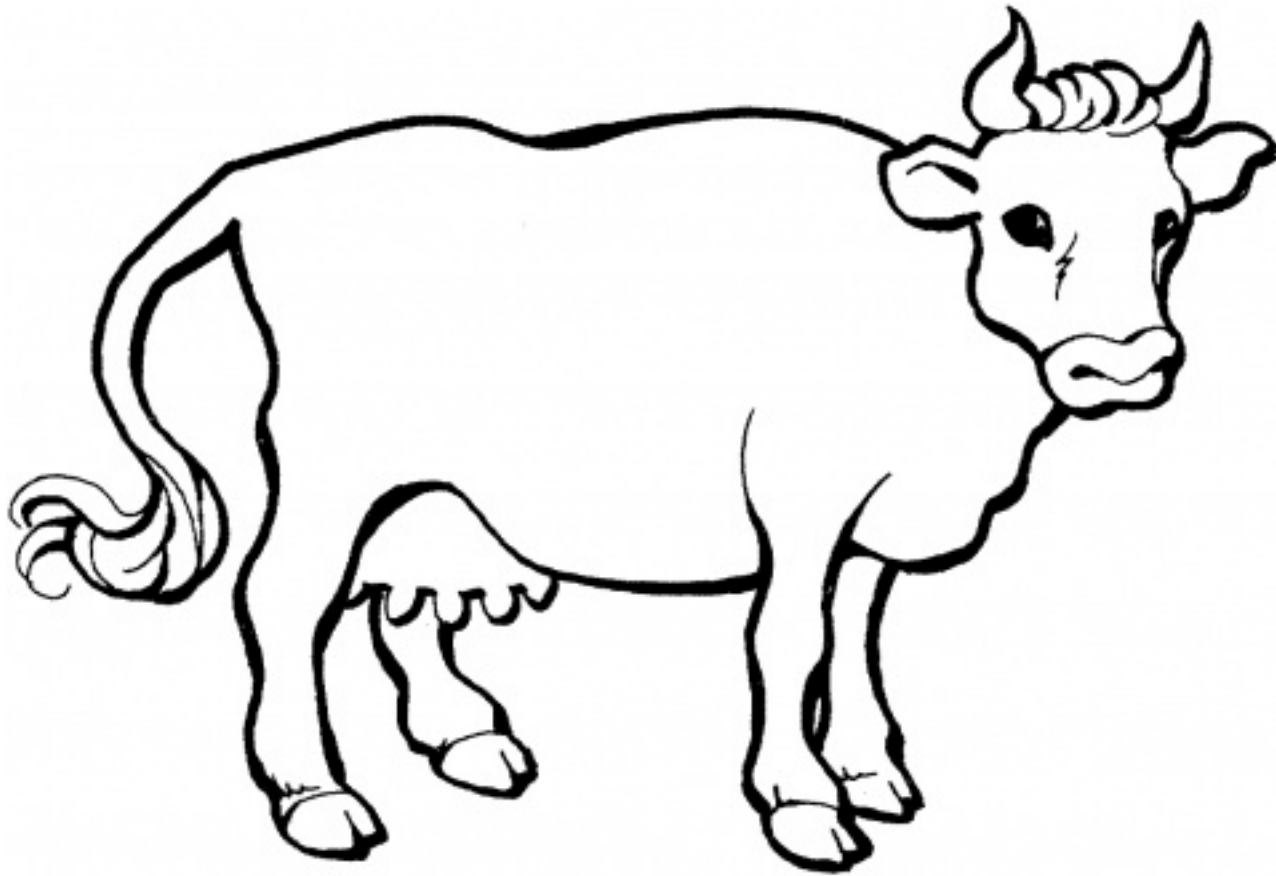
H. Song, PhD Dissertation, 0908.3656



# QGP momentum anisotropy cartoon

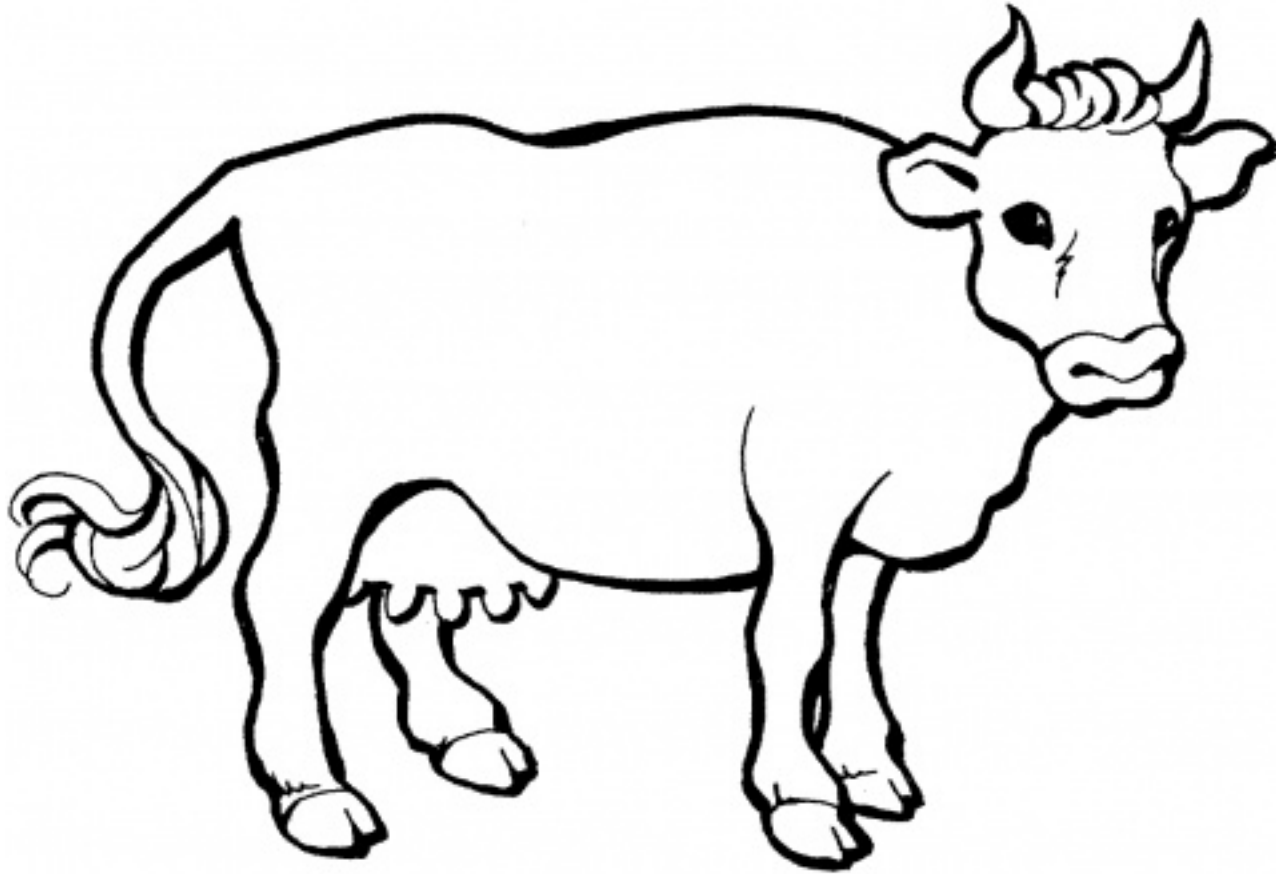


# Physics 101

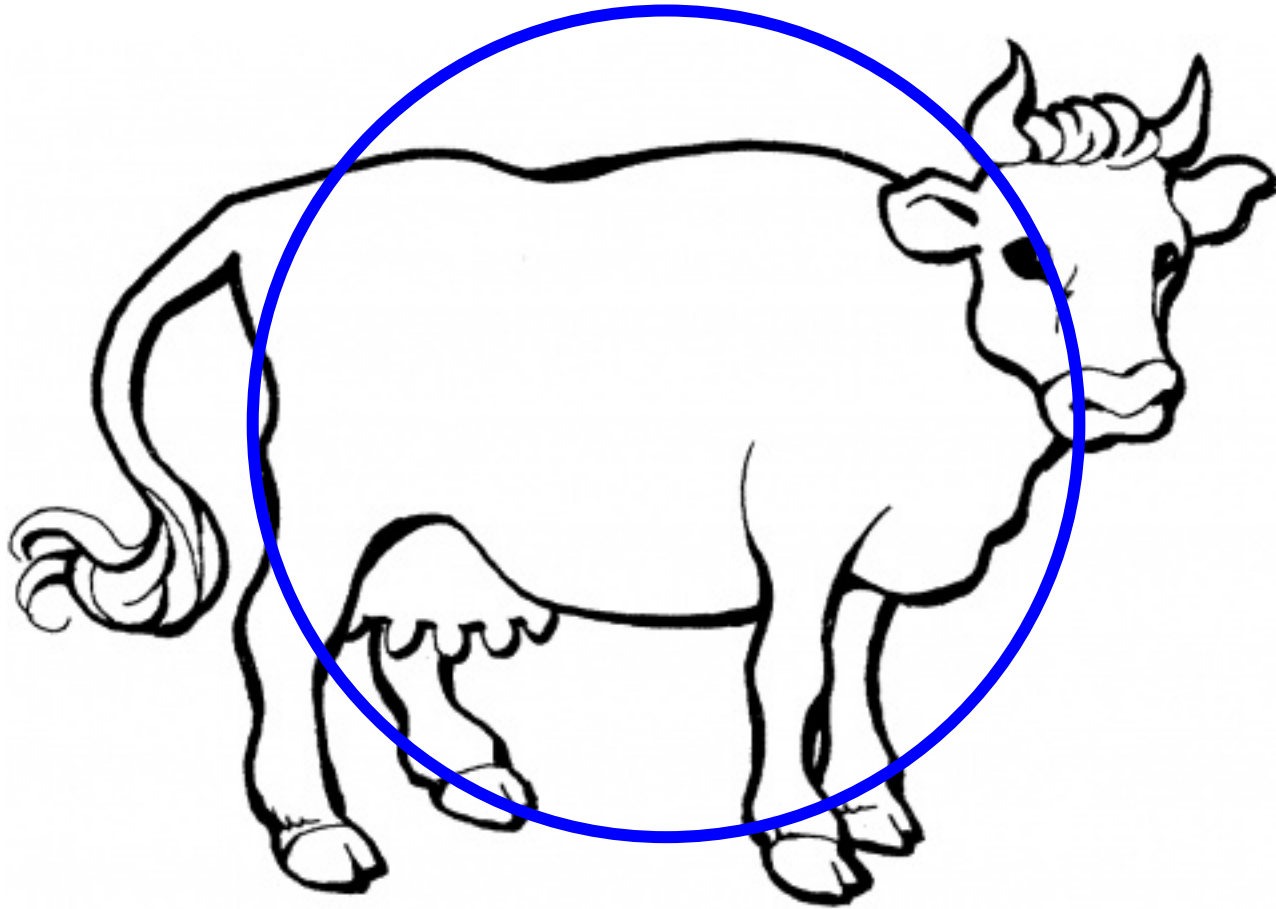




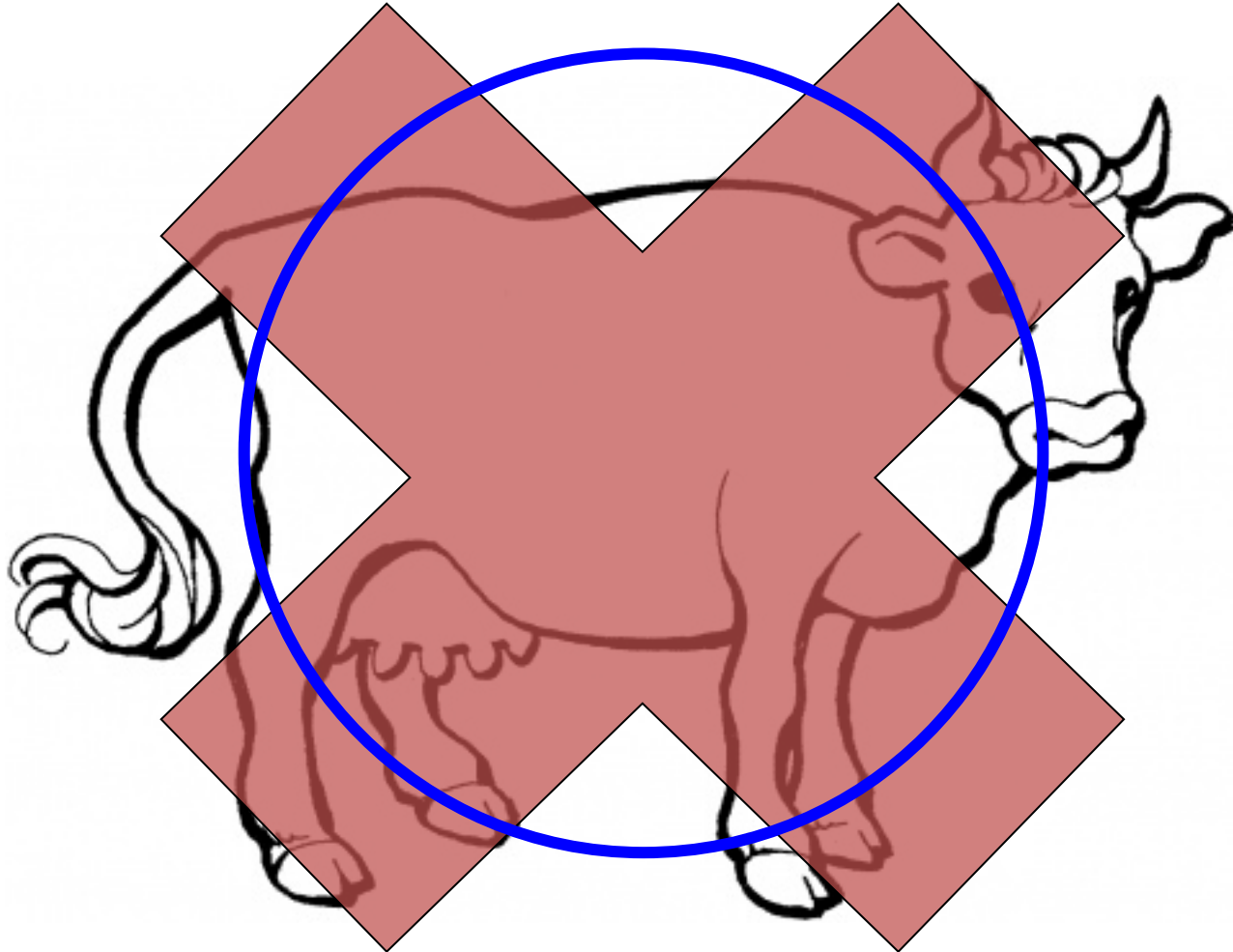
# Cows are spheres?



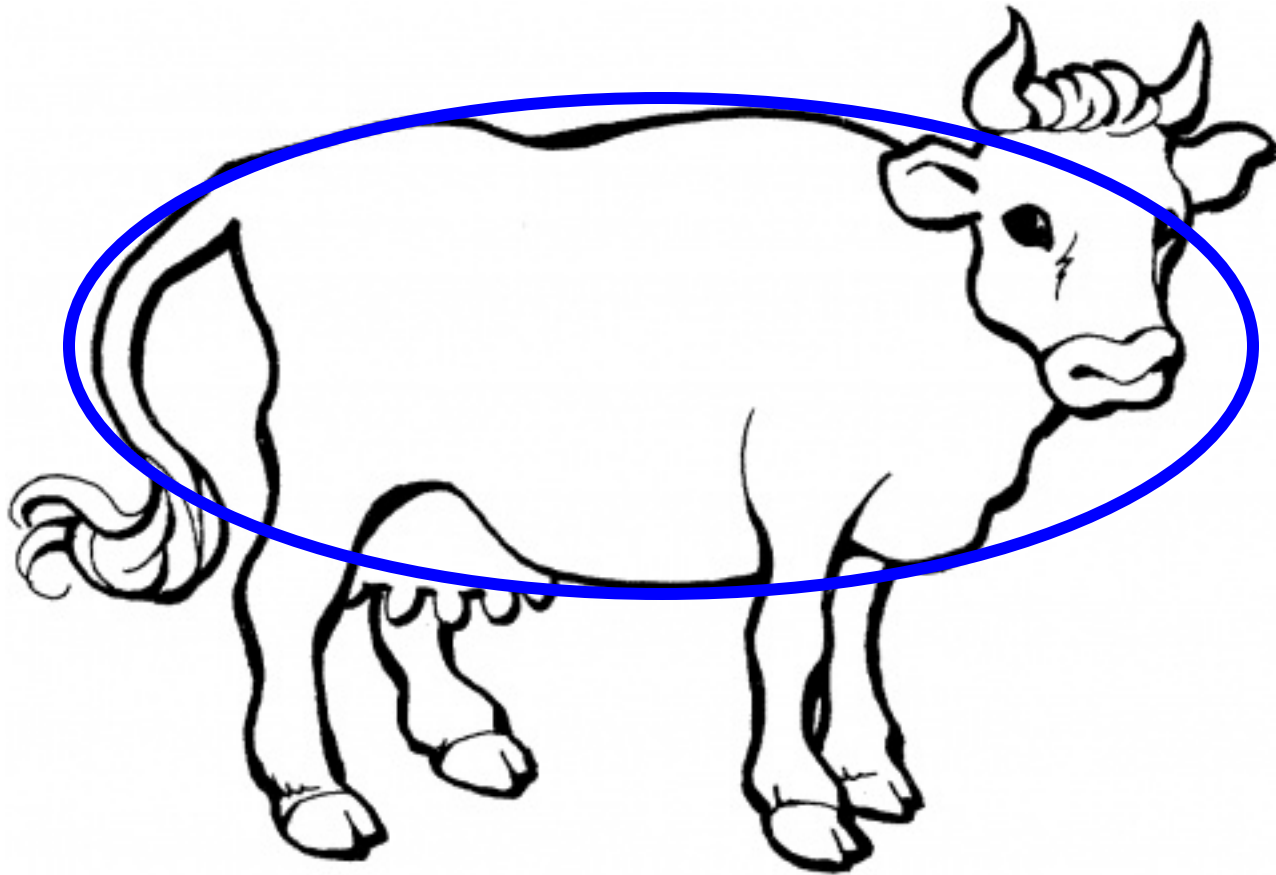
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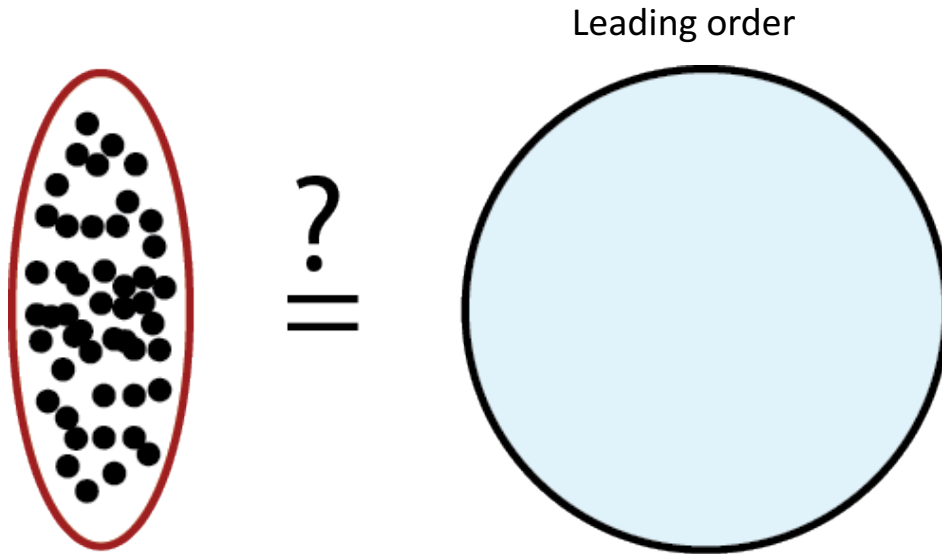
# Cows are not spheres



# Cows are more like ellipsoids!

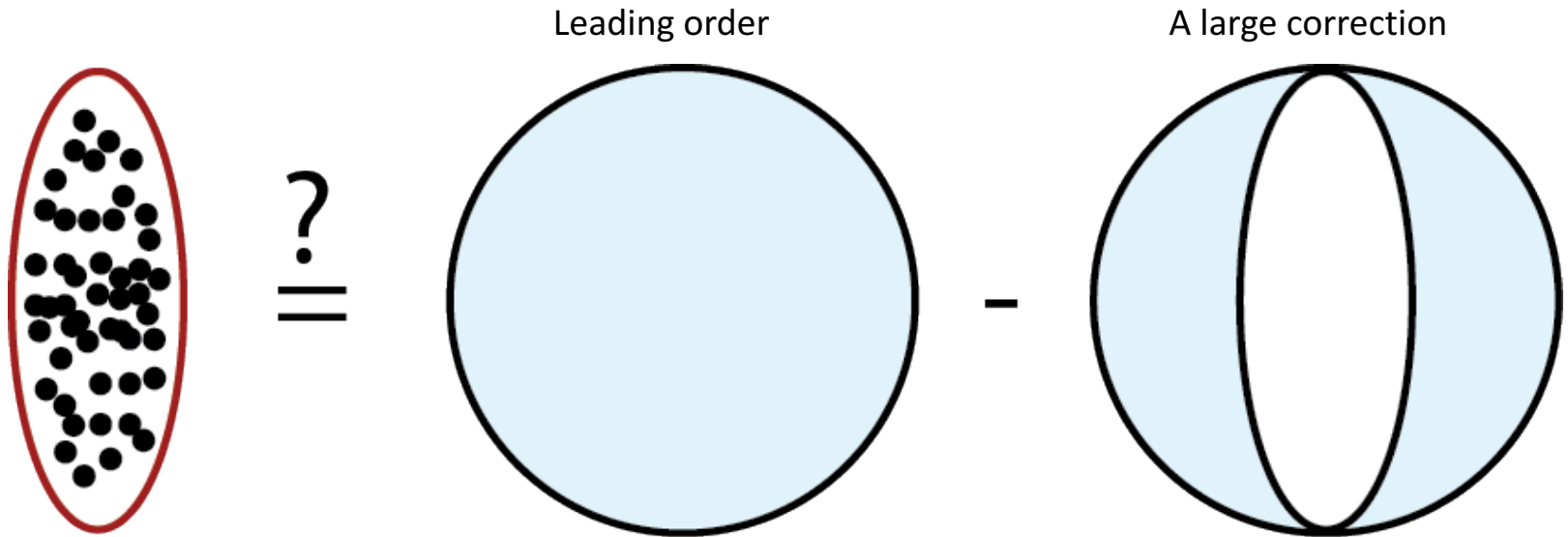


# Non-spherical cows



Viscous hydro says that we should approximate our particle momentum-space distribution to first order by a sphere in momentum space. However, if the system is highly anisotropic in momentum space, this will result in large corrections...

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# Spheroidal expansion method

## Viscous Hydrodynamics Expansion

$$f(\tau, \mathbf{x}, \mathbf{p}) = \underbrace{f_{\text{eq}}(\mathbf{p}, T(\tau, \mathbf{x}))}_{\text{Isotropic in momentum space}} + \delta f$$

See e.g.

- M. Martinez and MS, 1007.0889
- W. Florkowski and R. Ryblewski, 1007.0130
- D. Bazow, U. Heinz, and MS, 1311.6720
- D. Bazow, U. Heinz, and M. Martinez, 1503.07443
- E. Molnar, H. Niemi, and D. Rischke, 1602.00573; 1606.09019

## Anisotropic Hydrodynamics (aHydro) Expansion

$$f(\tau, \mathbf{x}, \mathbf{p}) = f_{\text{aniso}}(\mathbf{p}, \underbrace{\Lambda(\tau, \mathbf{x})}_{T_{\perp}}, \underbrace{\xi(\tau, \mathbf{x})}_{\text{anisotropy}}) + \delta \tilde{f}$$

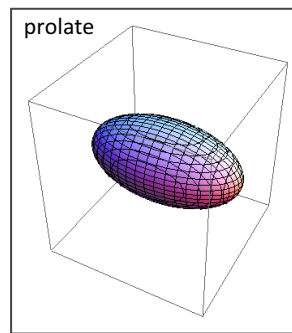
Treat this term perturbatively  
→ “NLO aHydro”

→ “Romatschke-Strickland” form in LRF

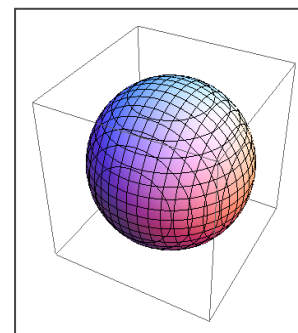
$$f_{\text{aniso}}^{LRF} = f_{\text{iso}} \left( \frac{\sqrt{\mathbf{p}^2 + \xi(\mathbf{x}, \tau) p_z^2}}{\Lambda(\mathbf{x}, \tau)} \right)$$

$$\xi = \frac{\langle p_T^2 \rangle}{2\langle p_L^2 \rangle} - 1$$

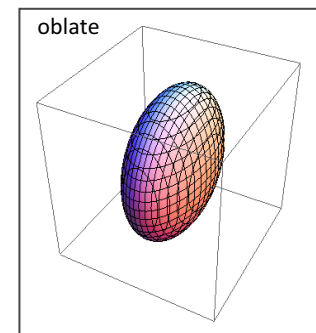
moo



$$-1 < \xi < 0$$



$$\xi = 0$$



$$\xi > 0$$



# Why spheroidal form at LO?

- What is special about this form at leading order?

$$f_{\text{aniso}}^{LRF} = f_{\text{iso}} \left( \frac{\sqrt{\mathbf{p}^2 + \xi(\mathbf{x}, \tau) p_z^2}}{\Lambda(\mathbf{x}, \tau)} \right)$$

- Gives the ideal hydro limit when  $\xi=0$  ( $\Lambda \rightarrow T$ )
- For longitudinal (0+1d) free streaming, the LRF distribution function is of spheroidal form; limit emerges automatically in conformal 0+1d aHydro

$$\xi_{\text{FS}}(\tau) = (1 + \xi_0) \left( \frac{\tau}{\tau_0} \right)^2 - 1$$

- Since  $f_{\text{iso}} \geq 0$ , the one-particle distribution function and pressures are  $\geq 0$  (not guaranteed in standard 2<sup>nd</sup>-order viscous hydro)
- Reduces to 2<sup>nd</sup>-order viscous hydrodynamics in limit of small anisotropies

M. Martinez and MS, 1007.0889

$$\frac{\Pi}{\mathcal{E}_{\text{eq}}} = \frac{8}{45} \xi + \mathcal{O}(\xi^2)$$

**For general (3+1d) proof of equivalence to second-order viscous hydrodynamics using generalized RS form in the near-equilibrium limit see Tinti 1411.7268.**



# The growing anisotropic hydrodynamics family

- There are two approaches being actively followed in the literature to address this problem
  - A. Linearize around a spheroidal distribution function and treat the perturbations using standard kinetic vHydro methods [“vaHydro”]  
Bazow, Heinz, Martinez, Molnar, Niemi, Rischke, MS
  - B. Introduce a generalized anisotropy tensor which replaces the entire viscous stress tensor at LO and then linearize around that instead  
Tinti, Ryblewski, Martinez, Nopoush, Alqahtani, Bluhm, Florkowski, Schaefer, MS
- Each of these methods has its own advantages.
- In what I will show today, I will use the generalized method (B) at leading order.

# Generalized aHydro formalism

In generalized aHydro, one assumes that the distribution function is of the form

$$f(x, p) = f_{\text{eq}} \left( \frac{\sqrt{p^\mu \Xi_{\mu\nu}(x) p^\nu}}{\lambda(x)}, \frac{\mu(x)}{\lambda(x)} \right) + \delta\tilde{f}(x, p)$$

$$\Xi^{\mu\nu} = \underbrace{u^\mu u^\nu}_{\text{LRF four velocity}} + \underbrace{\xi^{\mu\nu}}_{\text{Traceless symmetric anisotropy tensor}} - \underbrace{\Delta^{\mu\nu}}_{\text{Transverse projector}} \underbrace{\Phi}_{\text{"Bulk"}}$$

$$\begin{aligned} u^\mu u_\mu &= 1 \\ \xi^\mu{}_\mu &= 0 \\ \Delta^\mu{}_\mu &= 3 \\ u_\mu \xi^{\mu\nu} &= u_\mu \Delta^{\mu\nu} = 0 \end{aligned}$$

- 3 degrees of freedom in  $u^\mu$
  - 5 degrees of freedom in  $\xi^{\mu\nu}$
  - 1 degree of freedom in  $\Phi$
  - 1 degree of freedom in  $\lambda$
  - 1 degree of freedom in  $\mu$
- 11 DOFs

See e.g.

- M. Martinez, R. Ryblewski, and MS, 1204.1473
- L. Tinti and W. Florkowski, 1312.6614
- M. Nopoush, R. Ryblewski, and MS, 1405.1355

# Equations of Motion

- Herein the EOM are obtained from moments of the Boltzmann equation in the relaxation time approximation (RTA) including temperature -dependent quasiparticle mass

$$p^\mu \partial_\mu f + \frac{1}{2} \partial_i m^2 \partial_{(p)}^i f = -\mathcal{C}[f] \quad \mathcal{C}[f] = \frac{p^\mu u_\mu}{\tau_{\text{eq}}} (f - f_{\text{eq}})$$

- **It is relatively straightforward to use other collisional kernels (forthcoming)**
- 1 equation from the 0<sup>th</sup> moment [number (non-conservation)]
- 4 equations from the 1<sup>st</sup> moment [energy-momentum conservation]
- 6 equations from the 2<sup>nd</sup> moment [dissipative dynamics]
- We must also specify the relation between the equilibrium (isotropic) pressure and energy density (EoS). More on this later.

$$D_u n + n \theta_u = \frac{1}{\tau_{\text{eq}}} (n_{\text{eq}} - n)$$

$$\partial_\mu T^{\mu\nu} = 0$$

$$\partial_\mu \mathcal{I}^{\mu\nu\lambda} = \frac{1}{\tau_{\text{eq}}} (u_\mu \mathcal{I}_{\text{eq}}^{\mu\nu\lambda} - u_\mu \mathcal{I}^{\mu\nu\lambda})$$

$$\mathcal{I}^{\mu\nu\lambda} \equiv \int dP p^\mu p^\nu p^\lambda f(x, p).$$

# Is it really better?

aHydro reproduces exact solutions to the Boltzmann equation in a variety of expanding backgrounds better than standard viscous hydrodynamics.

## 0+1d Exact Solution

- Simple model: Boost-invariant transversally homogeneous Boltzmann equation in relaxation time approximation (RTA)
- Many results in this model, so we can compare with the literature
- Can be used to test different approximation schemes

Boltzmann EQ.  $p^\mu \partial_\mu f(x, p) = C[f(x, p)]$

RTA  $C[f] = \frac{p_\mu u^\mu}{\tau_{eq}} \left[ f_{eq}(p_\mu u^\mu, T(x)) - f(x, p) \right]$

Solution for the energy density (massless particle case)

$$\tilde{\mathcal{E}}(\tau) = D(\tau, \tau_0) \frac{\mathcal{R}(\xi_{FS}(\tau))}{\mathcal{R}(\xi_0)} + \int_{\tau_0}^{\tau} \frac{dr'}{\tau_{eq}(\tau')} D(\tau, \tau') \tilde{\mathcal{E}}(\tau') \mathcal{R}\left(\frac{\tau}{\tau'} - 1\right)$$

Time-dependent relaxation time	$\tau_{eq}(\tau) = \frac{5\eta}{T(\tau)}$	Damping Function	$D(\tau_2, \tau_1) = \exp\left[-\int_{\tau_1}^{\tau_2} dr \tau_{eq}^{-1}(\tau)\right]$
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See talk by R. Ryblewski for more details

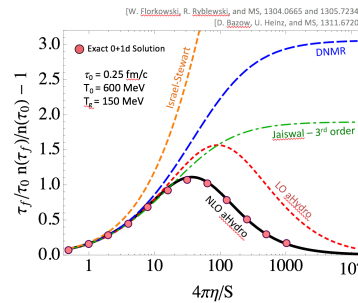
Massless Particles  
W. Florkowski, R. Ryblewski, and MS, 1304.0665 and 1305.7234

Massive Particles  
W. Florkowski, E. Makomura, R. Ryblewski, and MS, 1402.7348

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81

## Conformal 0+1d aHydro results

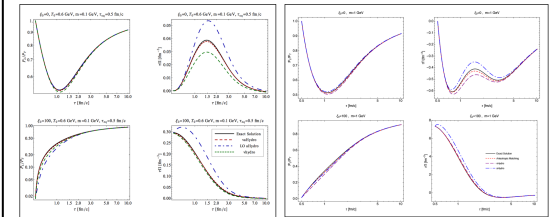


- Number (entropy) production vanishes in two limits: ideal hydrodynamic and free streaming limits
- In the conformal model which we are testing with, number density is proportional to entropy density

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30

## Non-conformal 0+1d aHydro results

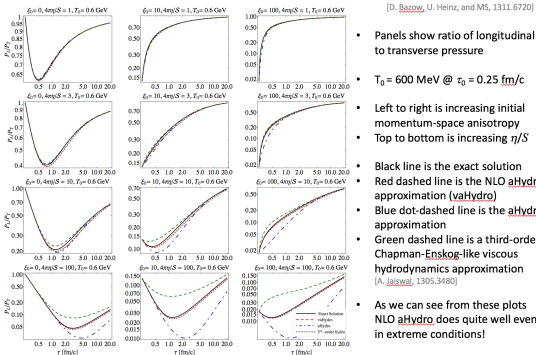


- Also works well in the non-conformal case
- Results on the left are from Bazow, Heinz, and Martinez [1503.07443]
- Results on the right are from Tinti [1506.07164]

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34

## Pressure Ratio Comparisons

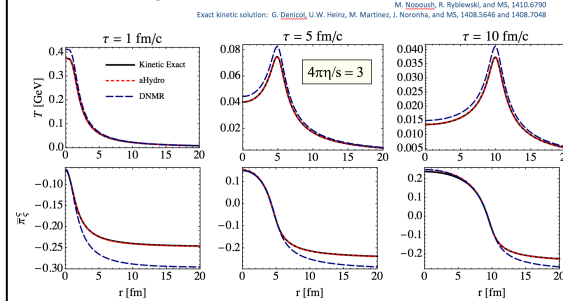


- Panels show ratio of longitudinal to transverse pressure
- $T_0 = 600 \text{ MeV} @ \tau_0 = 0.25 \text{ fm/c}$
- Left to right is increasing initial momentum-space anisotropy
- Top to bottom is increasing  $\eta/S$
- Black line is the exact solution
- Red dashed line is the NLO aHydro approximation (aHydro)
- Blue dot-dashed line is the aHydro approximation
- Green dashed line is a third-order Chapman-Enskog-like viscous hydrodynamics approximation [A. Jaiswal, 1305.3480]
- As we can see from these plots NLO aHydro does quite well even in extreme conditions!

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89

## 1+1d aHydro solution for Gubser Flow

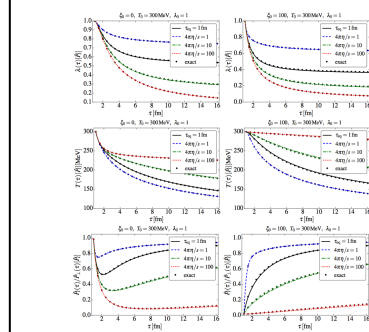


Once again, aHydro solution can be shown to reproduce the free streaming limit analytically. [M. Nopoush, R. Ryblewski, and MS, 1410.6790]

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33

## Conformal 0+1d aHydro results



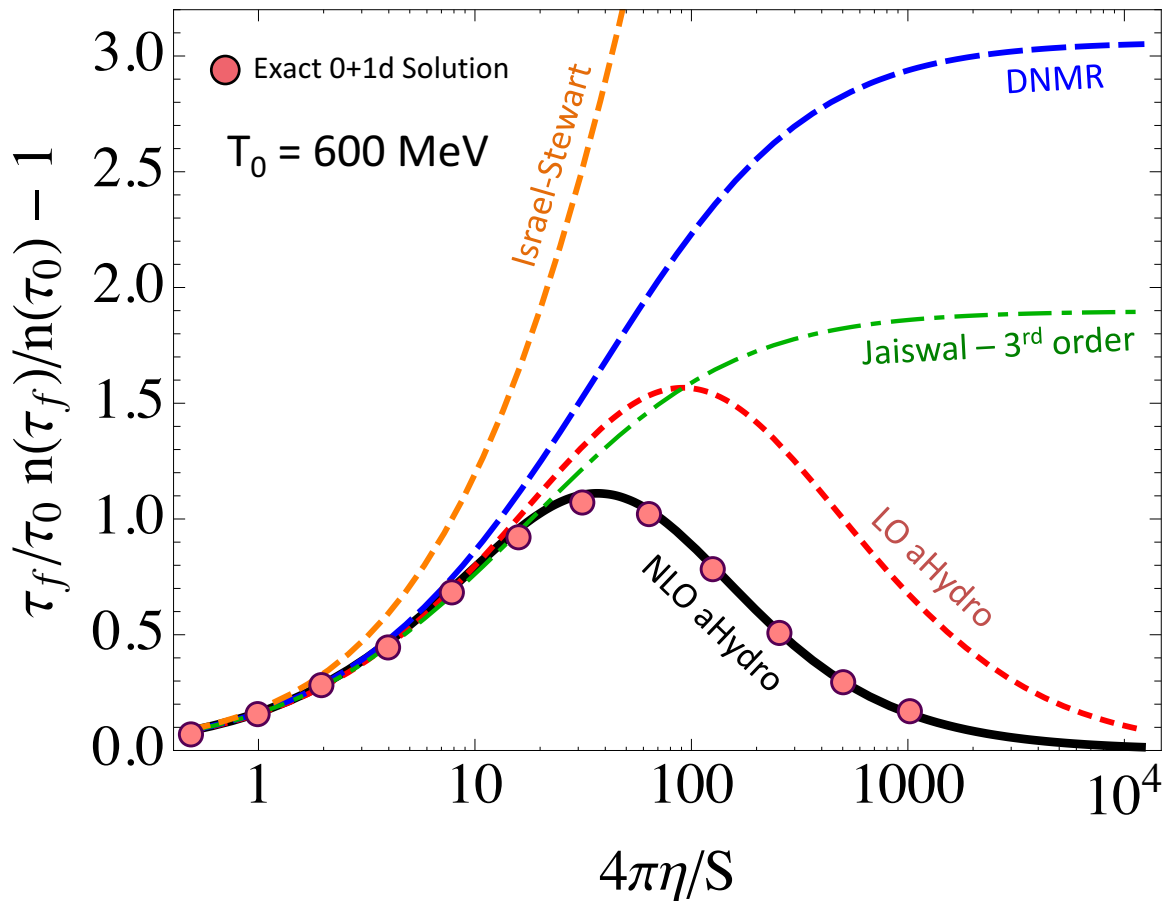
- Since our earlier papers, others have shown how to make things even better by a judicious choice of moments.
- Results on the left are from the recent paper of Molnar, Rischke, and Niemi [1606.09019]

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32

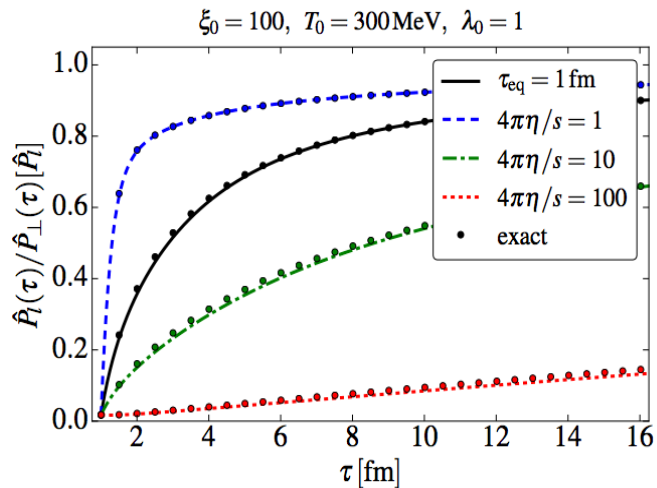
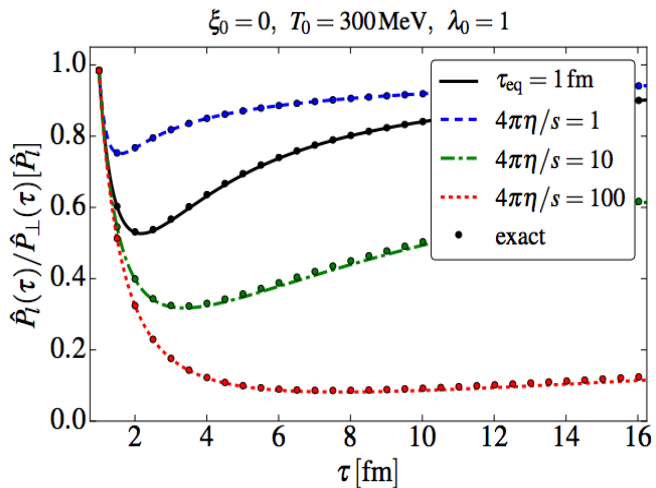
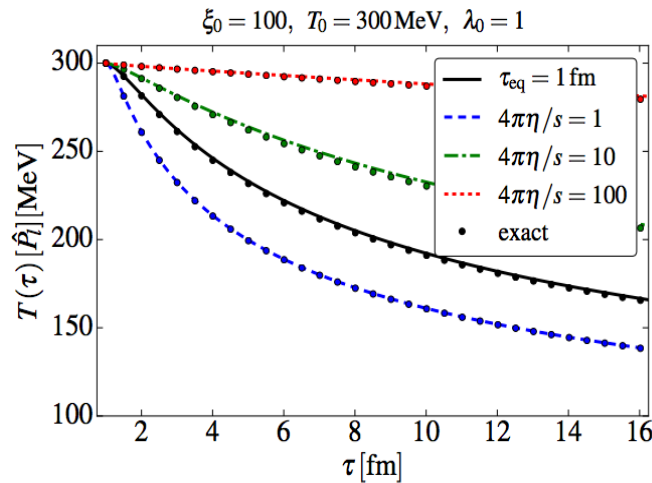
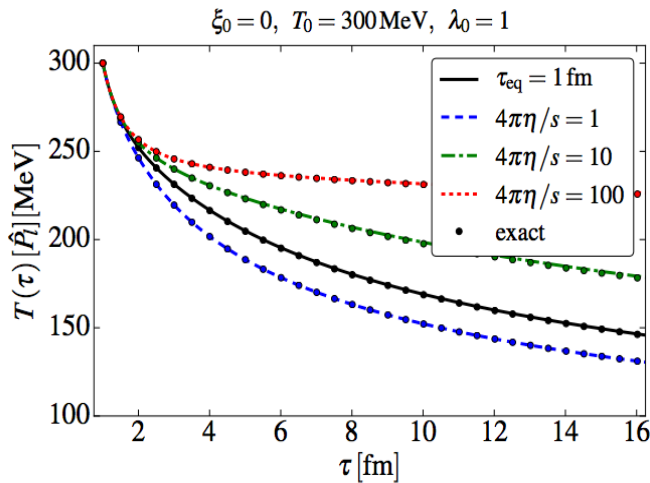
# Ex. 1: Dissipative particle production

[D. Bazow, U. Heinz, and MS, 1311.6720]



- Number (entropy) production vanishes in two limits: ideal hydrodynamic and free streaming limits
- In the conformal model which we are testing with, number density is proportional to entropy density

# Ex. 2: Conformal 0+1d aHydro results



- aHydro results (lines) on the left are from the recent paper of Molnar, Rischke, and Niemi [1606.09019]
- Exact solution is shown by dots [W. Florkowski, R. Ryblewski, and MS, 1304.0665 and 1305.7234]

# Ex 3: Gubser Flow

S. Gubser, 1006.0006

S. Gubser and Y.Yarom, 1012.1314

Gubser flow is a cylindrically-symmetric and boost-invariant flow that possesses a high degree of symmetry when mapped to Weyl-rescaled deSitter space

$SO(3)_q$	$\times$	$SO(1, 1)$	$\times$	$Z_2$
rotational symmetry around beam axis + conformal symmetry		boost invariance		reflection symmetry around the collision plane

The parameter  $q$  above is an arbitrary energy scale that sets the radial extent of the system at a given proper time.

## Polar Milne components

$$\tilde{u}^\tau = \cosh(\theta_\perp)$$

$$\tilde{u}^r = \sinh(\theta_\perp)$$

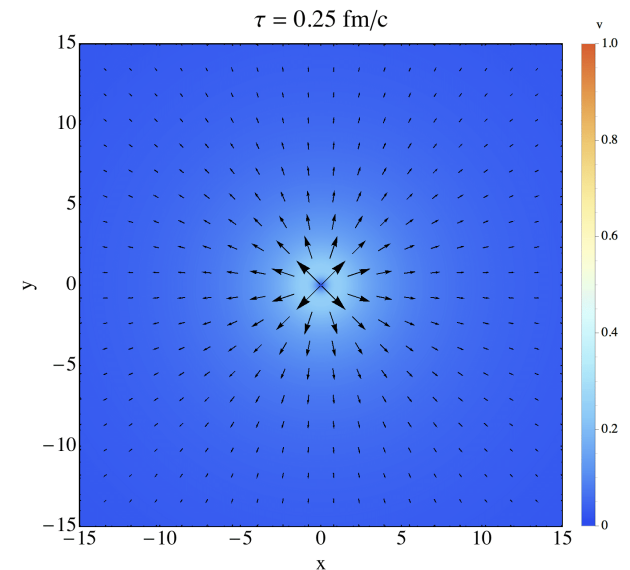
$$\tilde{u}^\phi = 0$$

$$\tilde{u}^s = 0$$

## Transverse rapidity

$$\theta_\perp = \tanh^{-1} \left( \frac{2q^2\tau r}{1 + q^2\tau^2 + q^2r^2} \right)$$

This flow is quite strong: The de Sitter space velocity gradients grow exponentially  $e^{|\rho|}$



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S. Gubser, 1006.0006

S. Gubser and Y. Yarom, 1012.1314

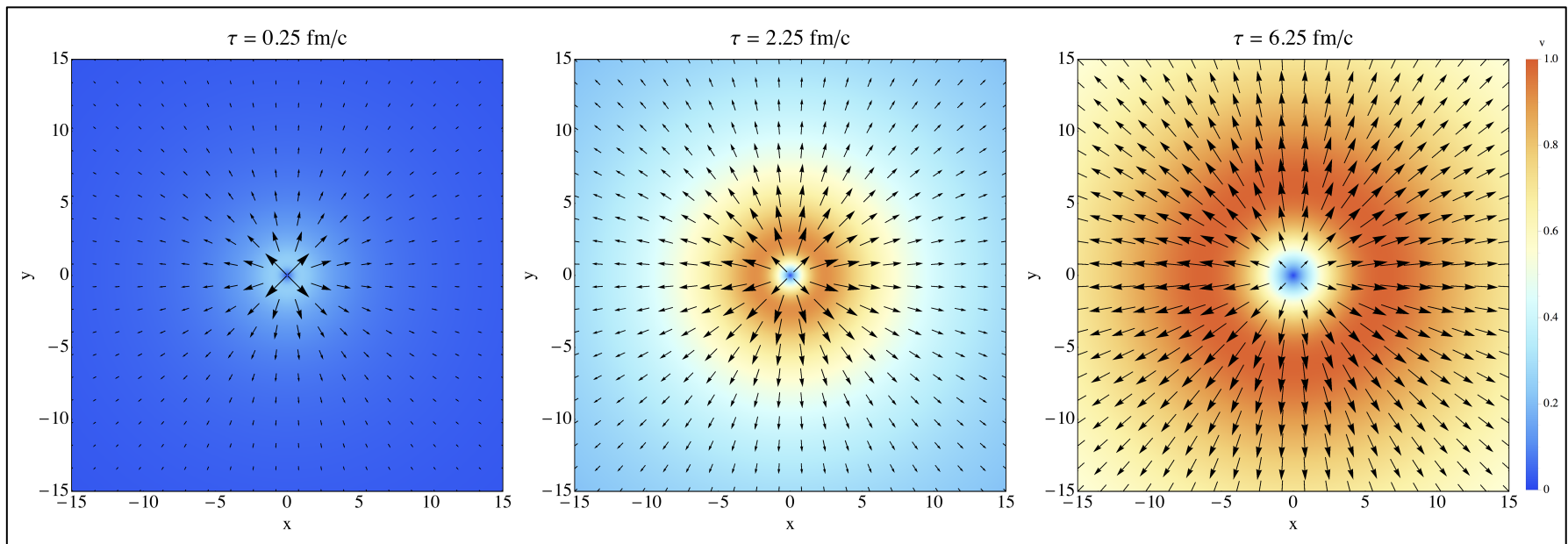
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rotational symmetry  
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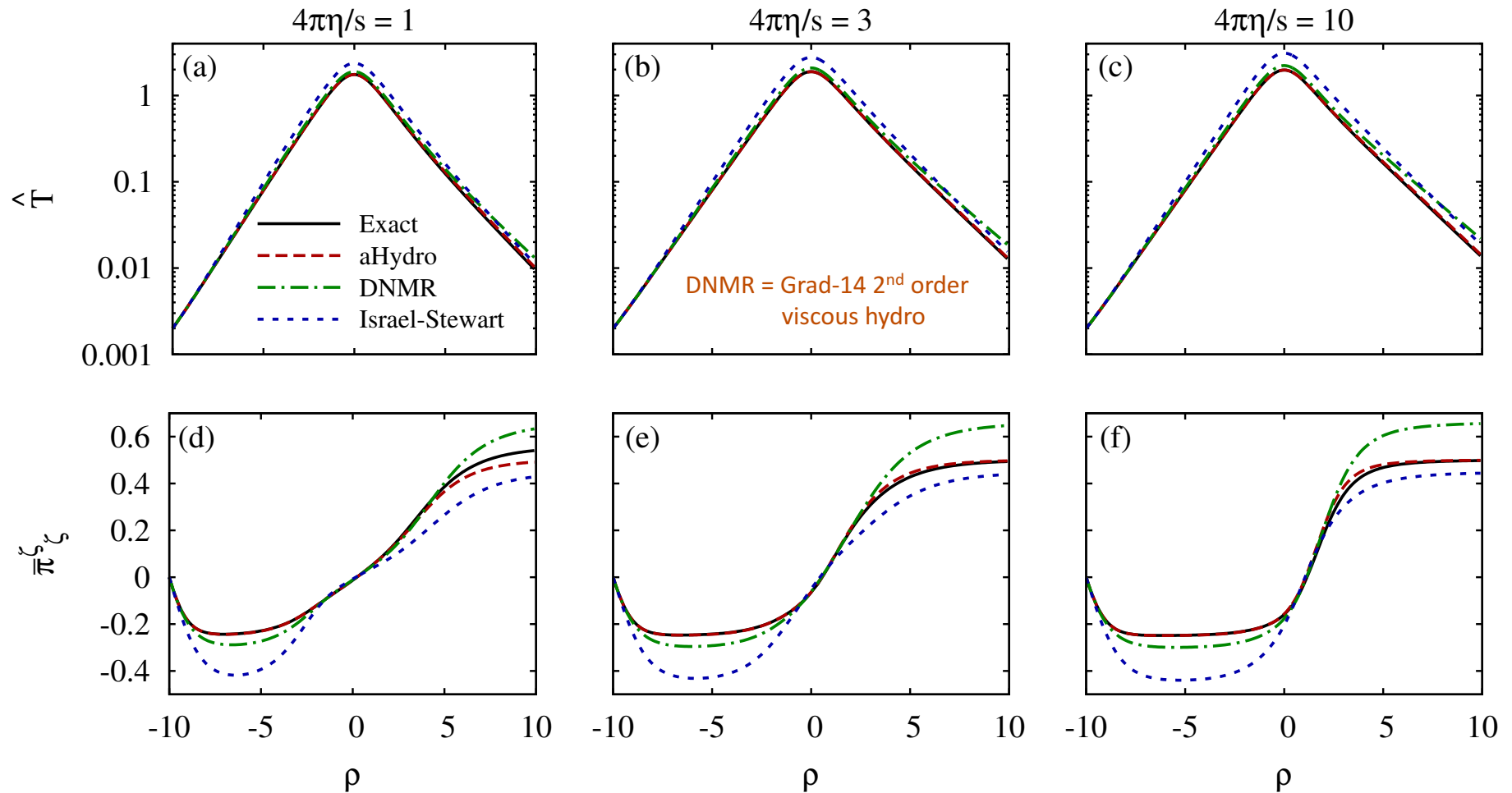




# Ex 3: LO aHydro for Gubser flow

M. Nopoush, R. Ryblewski, and MS, 1410.6790

Exact Solution: G. Denicol, U.W. Heinz, M. Martinez, J. Noronha, and MS, 1408.5646 and 1408.7048

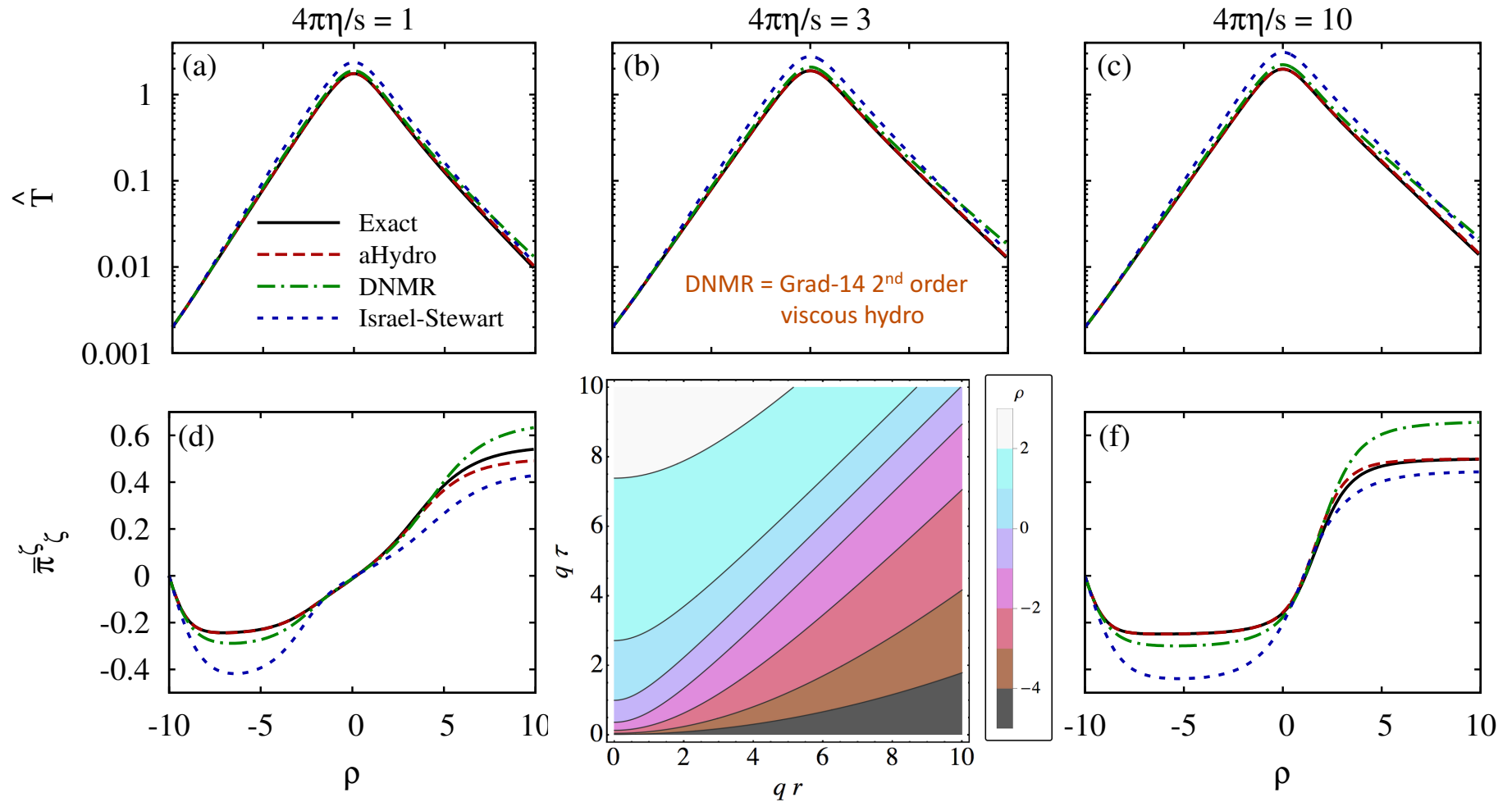


Isotropic initial conditions

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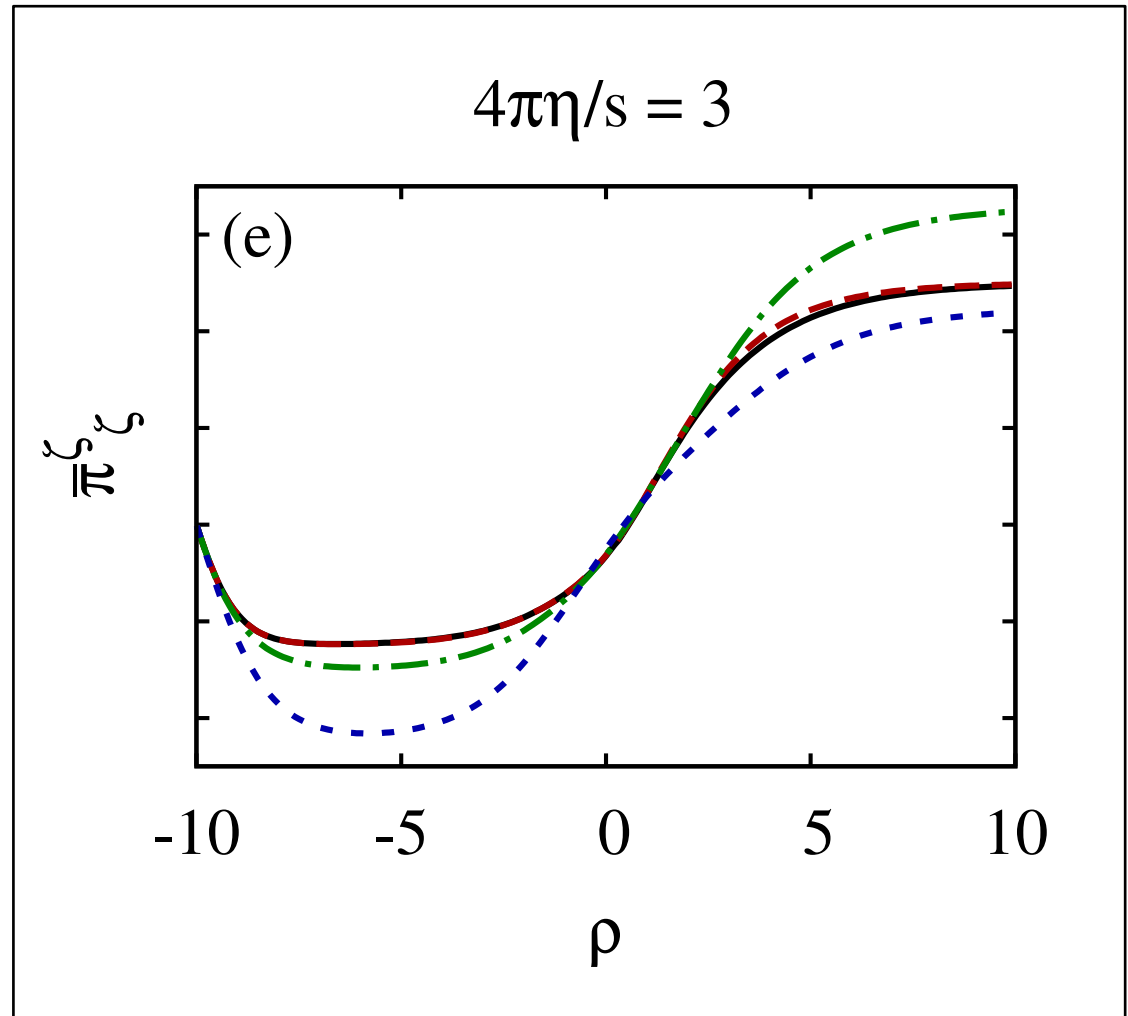
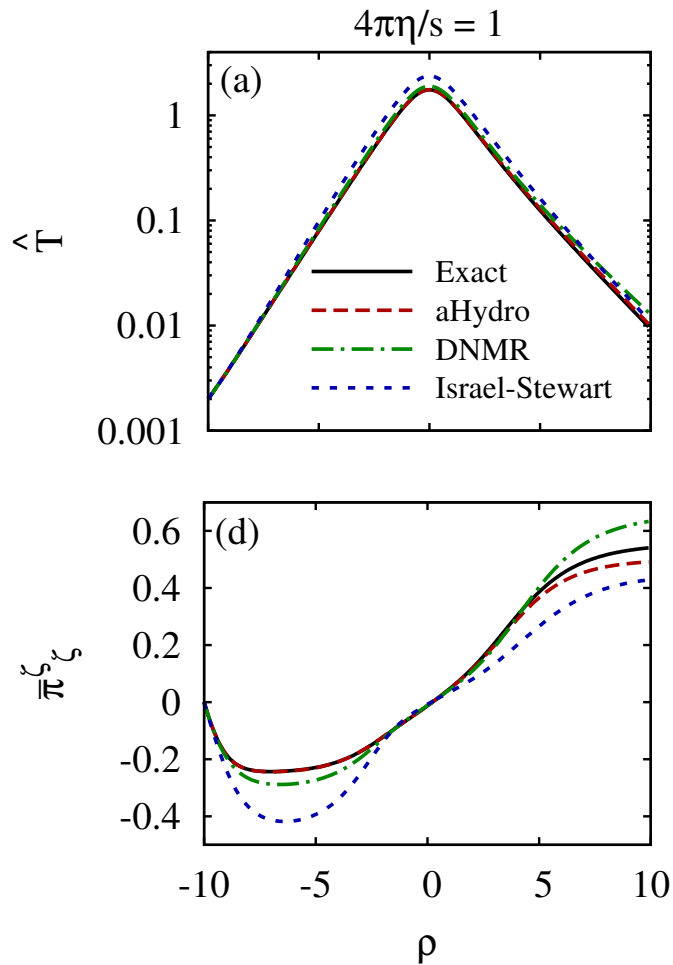


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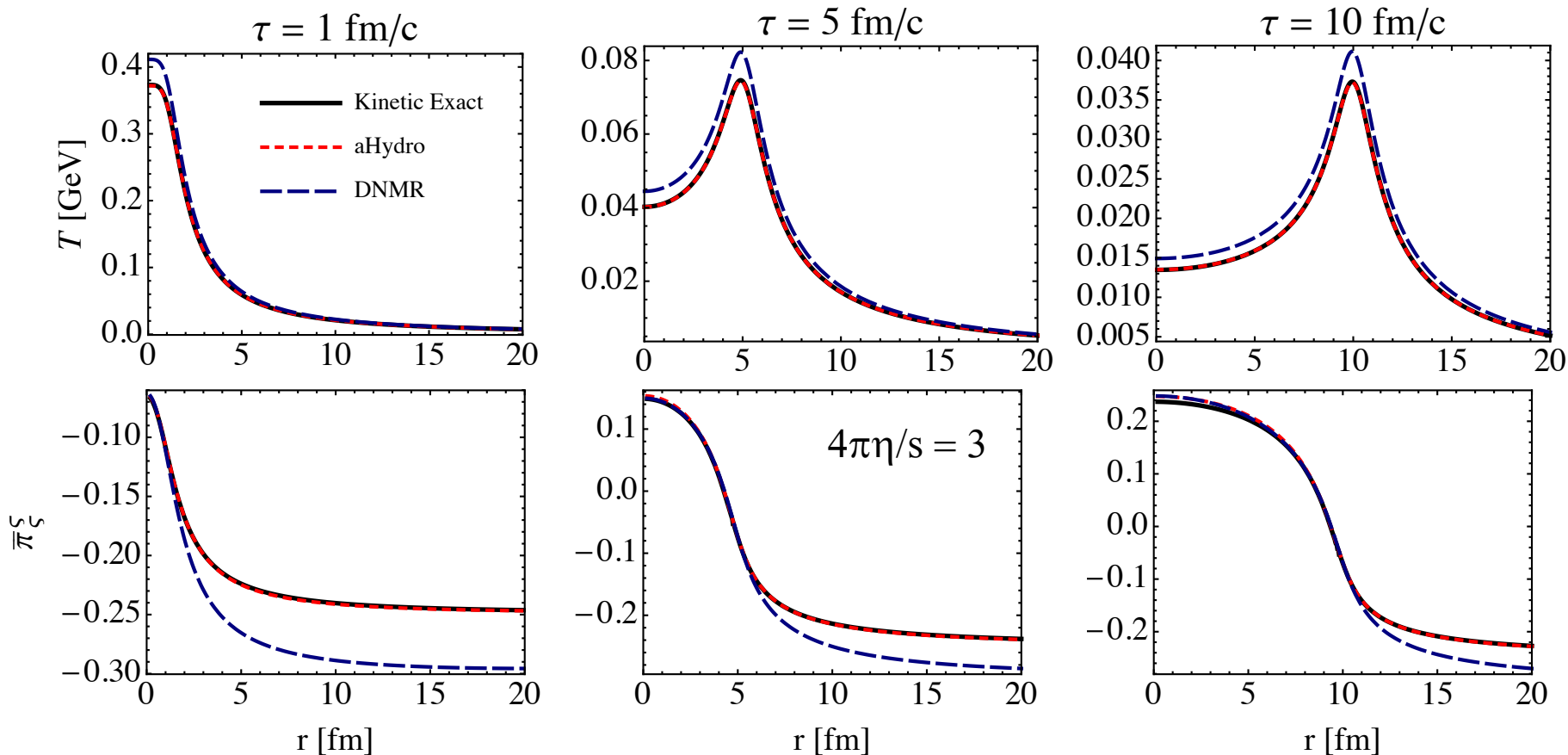


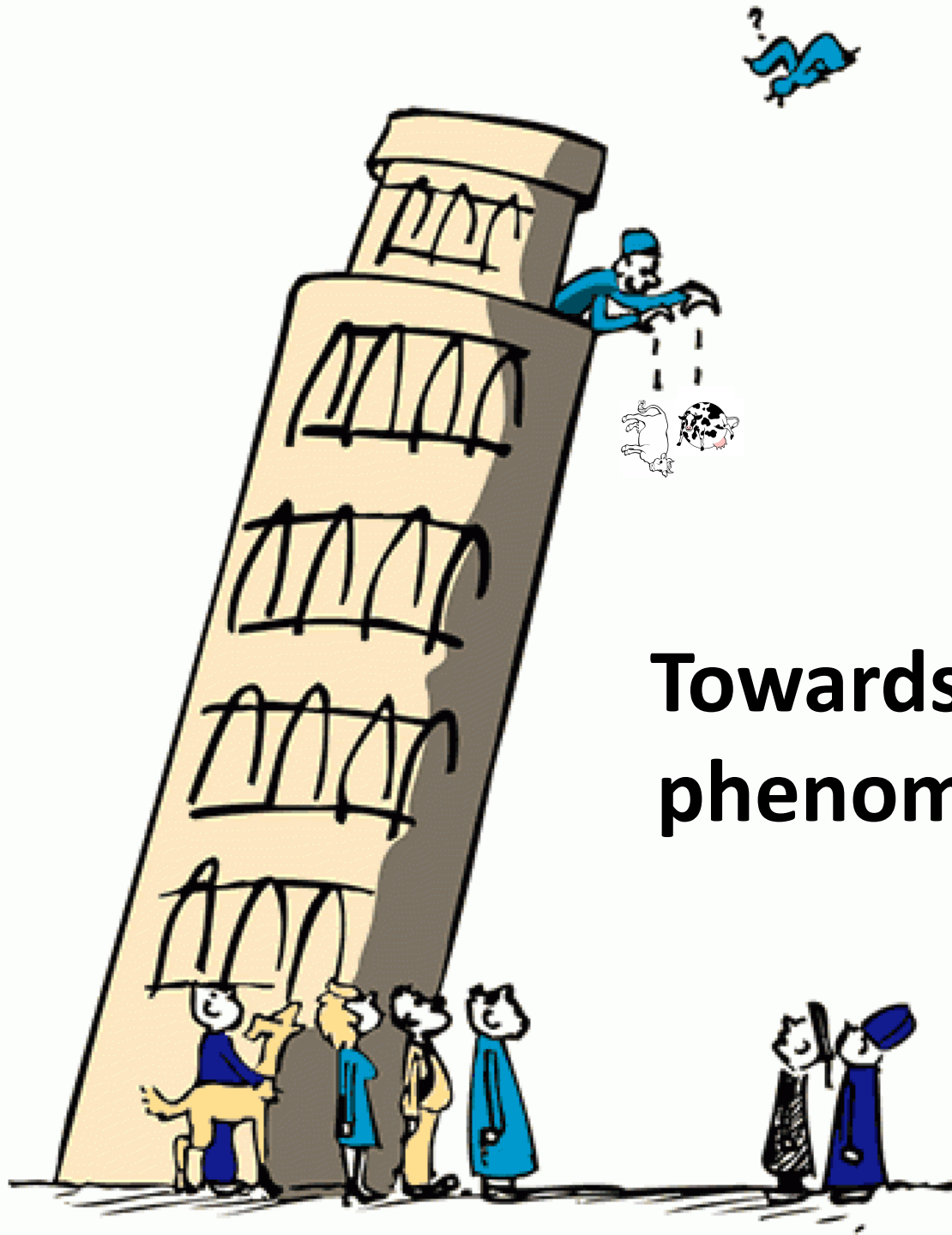
At NLO even better agreement with exact solution; see M. Martinez, M. McNelis, and U. Heinz, 1703.10955

# Ex 3: aHydro for Gubser flow

M. Nopoush, R. Ryblewski, and MS, 1410.6790

Exact Solution: G. Denicol, U.W. Heinz, M. Martinez, J. Noronha, and MS, 1408.5646; 1408.7048





# Towards realistic phenomenology

# 3+1d aHydro Equations of Motion

- Assuming an ellipsoidal form for the anisotropy tensor (ignoring off-diagonal components for now), one has seven degrees of freedom  $\xi_x, \xi_y, \xi_z, u_x, u_y, u_z$ , and  $\lambda$  which are all fields of space and time.
- Ignore  $\delta\tilde{f}$  for now

$$\begin{aligned}
 D_u \mathcal{E} + \mathcal{E} \theta_u + \mathcal{P}_x u_\mu D_x X^\mu + \mathcal{P}_y u_\mu D_y Y^\mu + \mathcal{P}_z u_\mu D_z Z^\mu &= 0, \\
 D_x \mathcal{P}_x + \mathcal{P}_x \theta_x - \mathcal{E} X_\mu D_u u^\mu - \mathcal{P}_y X_\mu D_y Y^\mu - \mathcal{P}_z X_\mu D_z Z^\mu &= 0, \\
 D_y \mathcal{P}_y + \mathcal{P}_y \theta_y - \mathcal{E} Y_\mu D_u u^\mu - \mathcal{P}_x Y_\mu D_x X^\mu - \mathcal{P}_z Y_\mu D_z Z^\mu &= 0, \\
 D_z \mathcal{P}_z + \mathcal{P}_z \theta_z - \mathcal{E} Z_\mu D_u u^\mu - \mathcal{P}_x Z_\mu D_x X^\mu - \mathcal{P}_y Z_\mu D_y Y^\mu &= 0.
 \end{aligned}$$

First Moment

$$\mathcal{I}^{\mu\nu\lambda} \equiv \int dP p^\mu p^\nu p^\lambda f(x, p).$$

$$\begin{aligned}
 \mathcal{I}_i &= \alpha \alpha_i^2 \mathcal{I}_{\text{eq}}(\lambda, m), \\
 \mathcal{I}_{\text{eq}}(\lambda, m) &= 4\pi \tilde{N} \lambda^5 \hat{m}^3 K_3(\hat{m}),
 \end{aligned}$$

$$\begin{aligned}
 D_u \mathcal{I}_x + \mathcal{I}_x (\theta_u + 2u_\mu D_x X^\mu) &= \frac{1}{\tau_{\text{eq}}} (\mathcal{I}_{\text{eq}} - \mathcal{I}_x), \\
 D_u \mathcal{I}_y + \mathcal{I}_y (\theta_u + 2u_\mu D_y Y^\mu) &= \frac{1}{\tau_{\text{eq}}} (\mathcal{I}_{\text{eq}} - \mathcal{I}_y), \\
 D_u \mathcal{I}_z + \mathcal{I}_z (\theta_u + 2u_\mu D_z Z^\mu) &= \frac{1}{\tau_{\text{eq}}} (\mathcal{I}_{\text{eq}} - \mathcal{I}_z).
 \end{aligned}$$

Second Moment

# Implementing the equation of state

M. Alqahtani, M. Nopoush, and MS, 1509.02913; 1605.02101

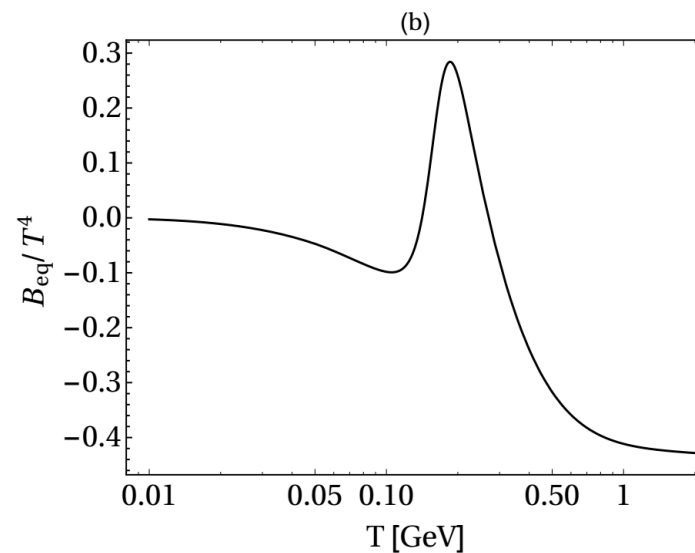
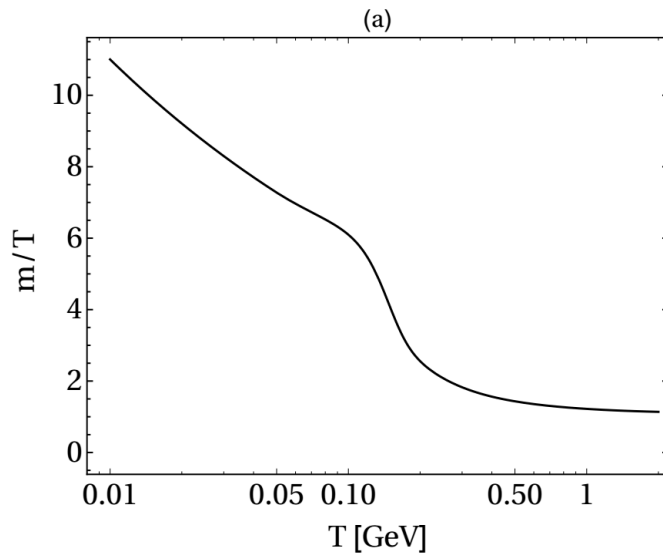
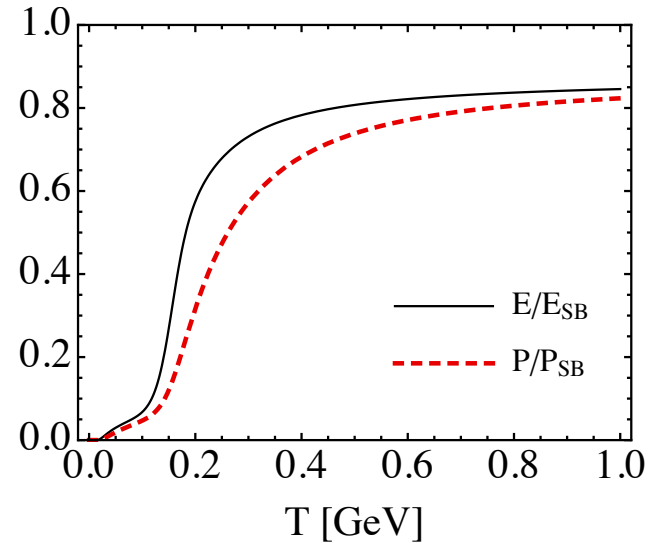
M. Alqahtani, M. Nopoush, R. Ryblewski, and MS, 1703.05808; 1705.10191

## Quasiparticle Method

$$T^{\mu\nu} = T_{\text{kinetic}}^{\mu\nu} + Bg^{\mu\nu}$$

$$p^\mu \partial_\mu f + \frac{1}{2} \partial_i m^2 \partial_{(p)}^i f = -\mathcal{C}[f]$$

$$\partial_\mu B = -\frac{1}{2} \partial_\mu m^2 \int dP f(x, p)$$



# Implementing the equation of state

M. Alqahtani, M. Nopoush, and MS, 1509.02913; 1605.02101  
M. Alqahtani, M. Nopoush, R. Ryblewski, and MS, 1703.05808; 1705.10191

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## Shear viscosity

Fix relaxation time as a function of the energy density by requiring fixed shear viscosity to entry density ratio.

$$\frac{\eta}{\tau_{\text{eq}}} = \frac{1}{T} I_{3,2}(\hat{m}_{\text{eq}})$$

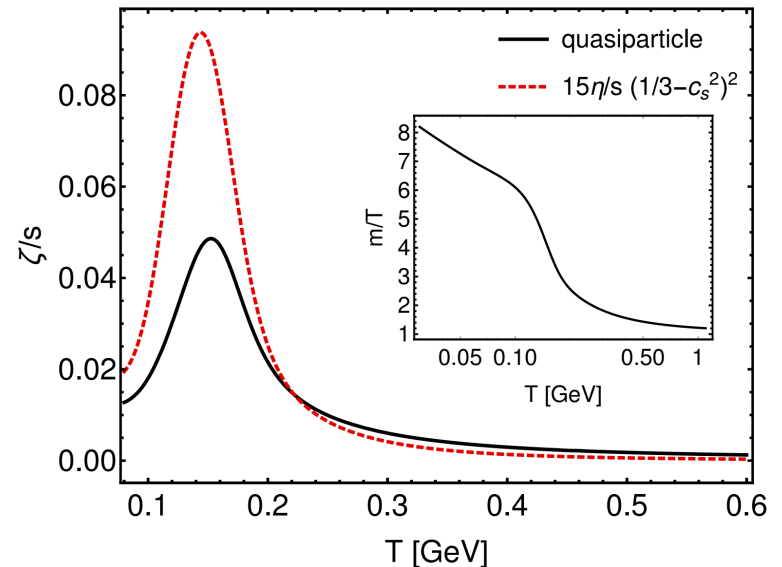
## Bulk viscosity

$$\frac{\zeta}{\tau_{\text{eq}}} = \frac{5}{3T} I_{3,2} - c_s^2 (\mathcal{E} + \mathcal{P}) + T \hat{m}^3 \frac{dm}{dT} I_{1,1}$$

$$I_{3,2}(x) = \frac{N_{\text{dof}} T^5 x^5}{30\pi^2} \left[ \frac{1}{16} (K_5(x) - 7K_3(x) + 22K_1(x)) - K_{i,1}(x) \right],$$

$$K_{i,1}(x) = \frac{\pi}{2} \left[ 1 - xK_0(x)\mathcal{S}_{-1}(x) - xK_1(x)\mathcal{S}_0(x) \right],$$

$$I_{1,1} = \frac{g m^3}{6\pi^2} \left[ \frac{1}{4} (K_3 - 5K_1) + K_{i,1} \right]$$





# Implementing the equation of state

M. Alqahtani, M. Nopoush, and MS, 1509.02913; 1605.02101

M. Alqahtani, M. Nopoush, R. Ryblewski, and MS, 1703.05808; 1705.10191

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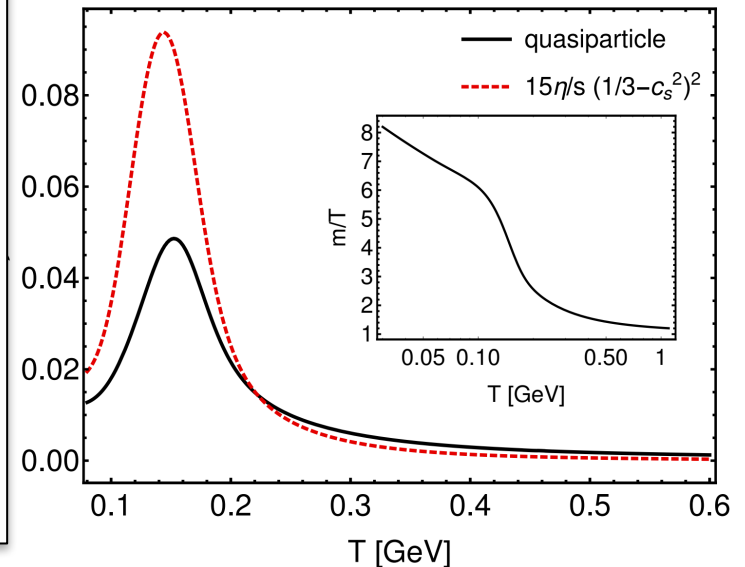
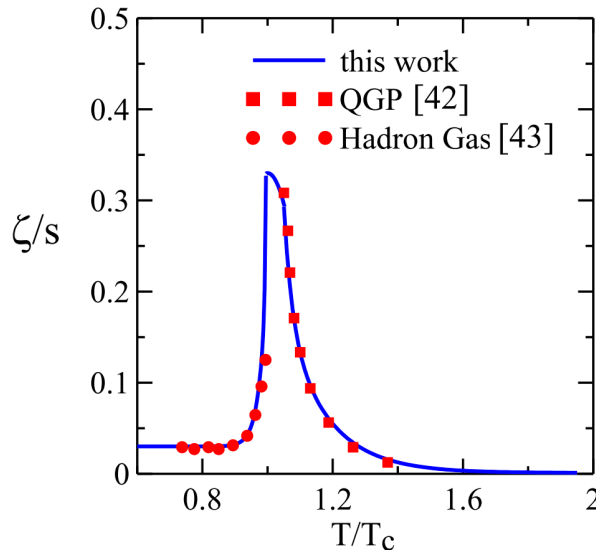
$$\frac{\zeta}{\tau_{\text{eq}}} = \frac{5}{3T} I_{3,2}$$

$$I_{3,2}(x) = \frac{N_{\text{dof}}}{30}$$

$$K_{i,1}(x) = \frac{\pi}{2} \left[ 1 - \frac{m^2}{x^2} \right]$$

$$I_{1,1} = \frac{gm^2}{6\pi^2}$$

Ryu et al, PRL 115 (2015) no.13, 132301



# Implementing the equation of state

M. Alqahtani, M. Nopoush, and MS, 1509.02913; 1605.02101

M. Alqahtani, M. Nopoush, R. Ryblewski, and MS, 1703.05808; 1705.10191

## Quasiparticle Method

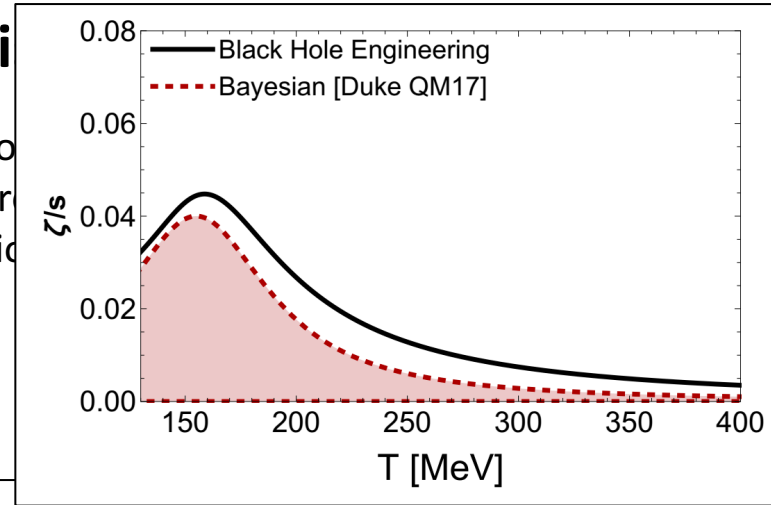
$$T^{\mu\nu} = T_{\text{kinetic}}^{\mu\nu} + Bg^{\mu\nu}$$

$$p^\mu \partial_\mu f + \frac{1}{2} \partial_i m^2 \partial_{(p)}^i f = -\mathcal{C}[f]$$

$$\partial_\mu B = -\frac{1}{2} \partial_\mu m^2 \int dP f(x, p)$$

## Shear viscosity

Fix relaxation  
density by r  
density ratio



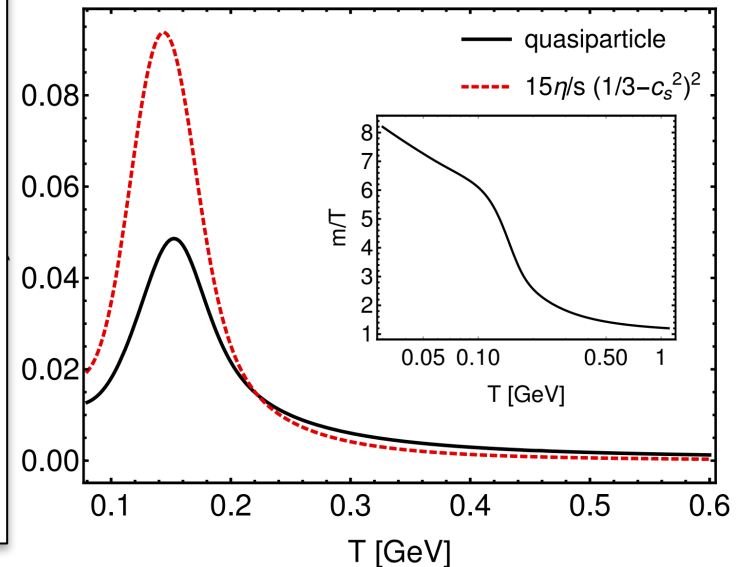
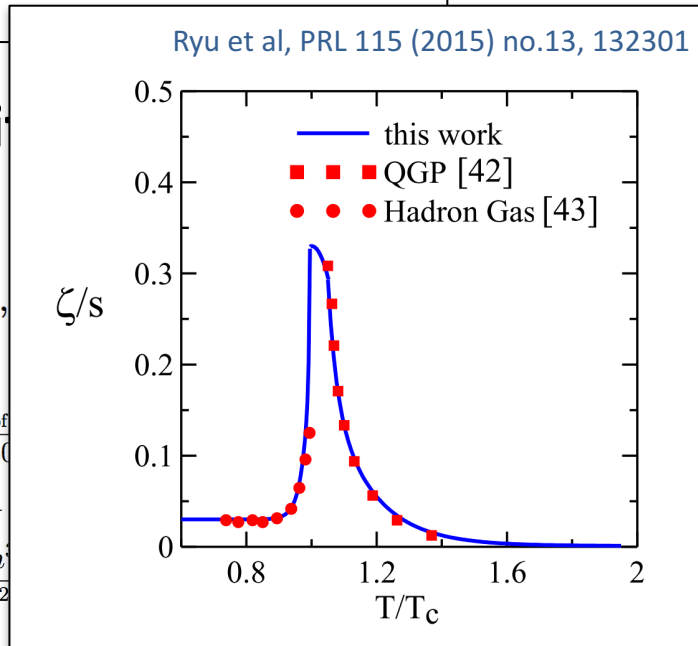
## Bulk viscosity

$$\frac{\zeta}{\tau_{\text{eq}}} = \frac{5}{3T} I_3,$$

$$I_{3,2}(x) = \frac{N_{\text{dof}}}{30}$$

$$K_{i,1}(x) = \frac{\pi}{2} \left[ 1 - \frac{1}{x} \right]$$

$$I_{1,1} = \frac{gm}{6\pi^2}$$



# Anisotropic Cooper-Frye Freezeout

M. Alqahtani, M. Nopoush, and MS, 1605.02101

M. Alqahtani, M. Nopoush, R. Ryblewski, and MS, 1703.05808; 1705.10191

- Use same generalized-RS form for “anisotropic freeze-out” at LO
- Form includes both shear and bulk corrections to the distribution function
- Use energy density (scalar) to determine the freeze-out hypersurface  $\Sigma \rightarrow$  e.g.  $T_{\text{eff,FO}} = 130$  MeV

$$f(x, p) = f_{\text{iso}} \left( \frac{1}{\lambda} \sqrt{p_\mu \Xi^{\mu\nu} p_\nu} \right)$$

$$\Xi^{\mu\nu} = \underbrace{u^\mu u^\nu}_{\text{isotropic}} + \underbrace{\xi^{\mu\nu}}_{\text{anisotropy tensor}} - \underbrace{\Phi \Delta^{\mu\nu}}_{\text{bulk correction}}$$

$$\xi_{\text{LRF}}^{\mu\nu} \equiv \text{diag}(0, \xi_x, \xi_y, \xi_z)$$

$$\xi^\mu{}_\mu = 0 \quad u_\mu \xi^\mu{}_\nu = 0$$

$$\left( p^0 \frac{dN}{dp^3} \right)_i = \frac{\mathcal{N}_i}{(2\pi)^3} \int f_i(x, p) p^\mu d\Sigma_\mu,$$

**NOTE:** Usual 2<sup>nd</sup>-order viscous hydro form

$$f(p, x) = f_{\text{eq}} \left[ 1 + (1 - a f_{\text{eq}}) \frac{p_\mu p_\nu \Pi^{\mu\nu}}{2(\epsilon + P)T^2} \right]$$

$$f_{\text{eq}} = 1 / [\exp(p \cdot u / T) + a] \quad a = -1, +1, \text{ or } 0$$

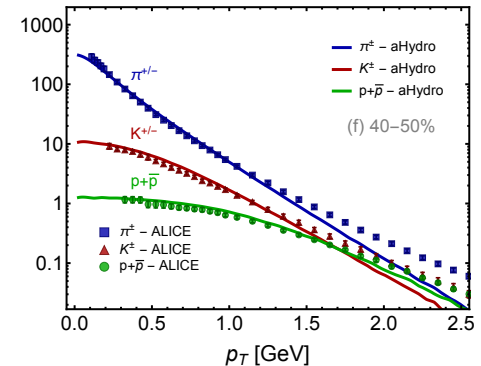
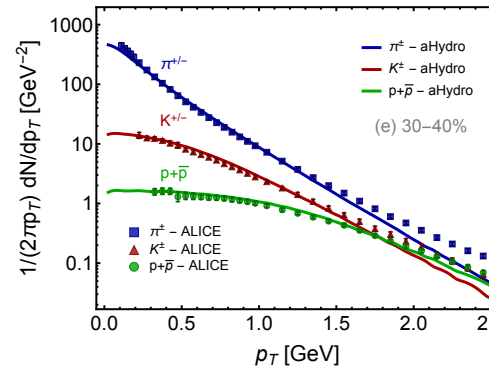
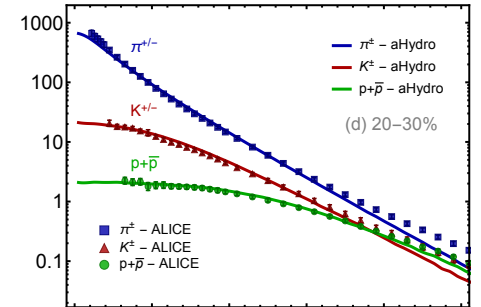
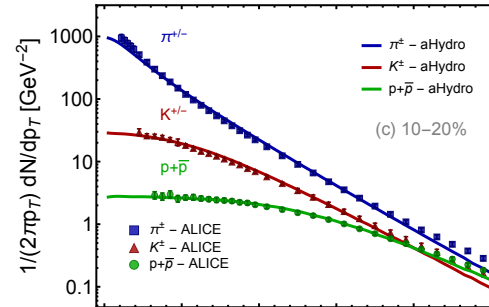
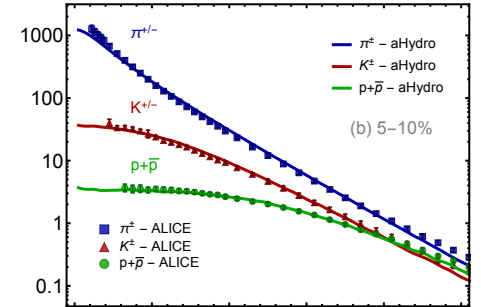
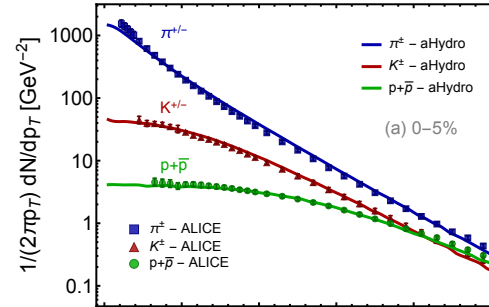
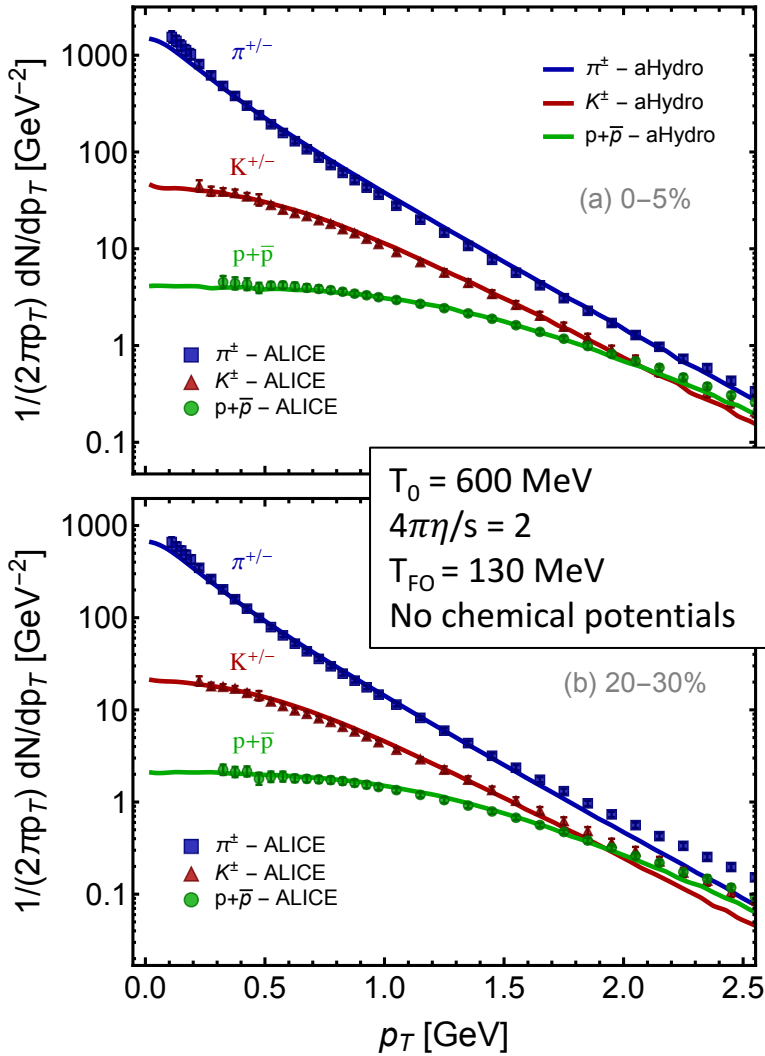
- This form suffers from the problem that the distribution function can be negative in some regions of phase space  $\rightarrow$  unphysical
- **Problem becomes worse when including the bulk viscous correction.**

# The phenomenological setup

- Keep it simple at first → smooth Glauber initial conditions
- Mixture of wounded nucleon and binary collision profiles with a binary mixing fraction of 0.15 (empirically suggested from prior viscous hydro studies)
- In the rapidity direction, we use a rapidity profile with a “tilted” central plateau and Gaussian “wings”
- We take the system to be initially isotropic in momentum space
- We then run the code and extract the freeze-out hypersurface
- The primordial particle production is then Monte-Carlo sampled using the Therminator 2 [Chojnacki, Kisiel, Florkowski, and Broniowski, arXiv:1102.0273]
- Therminator also takes care of all resonance feed downs
- All data shown are from the **ALICE collaboration**

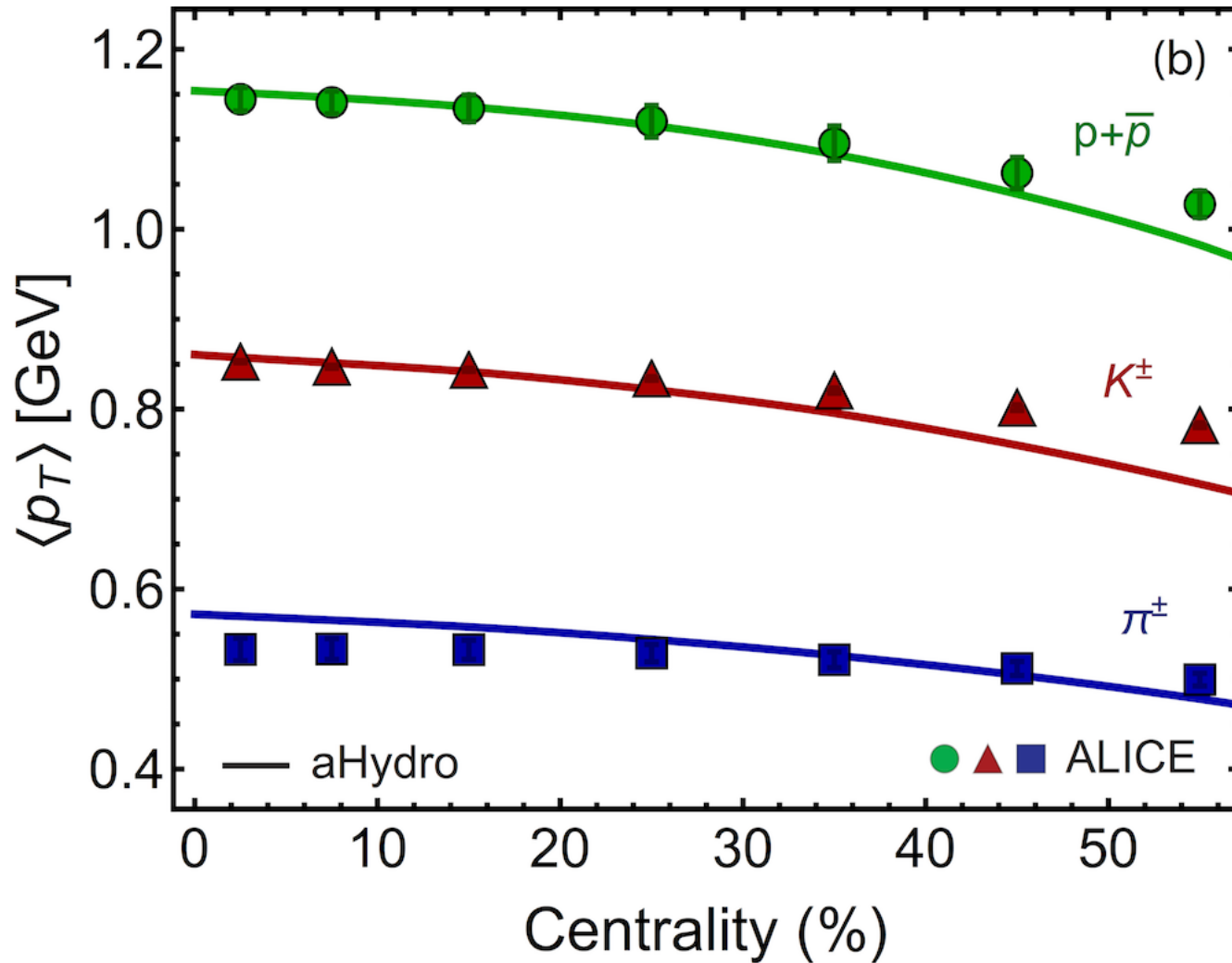
# Identified particle spectra

Alqahtani, Nopoush, Ryblewski, MS, 1703.05808; 1705.10191



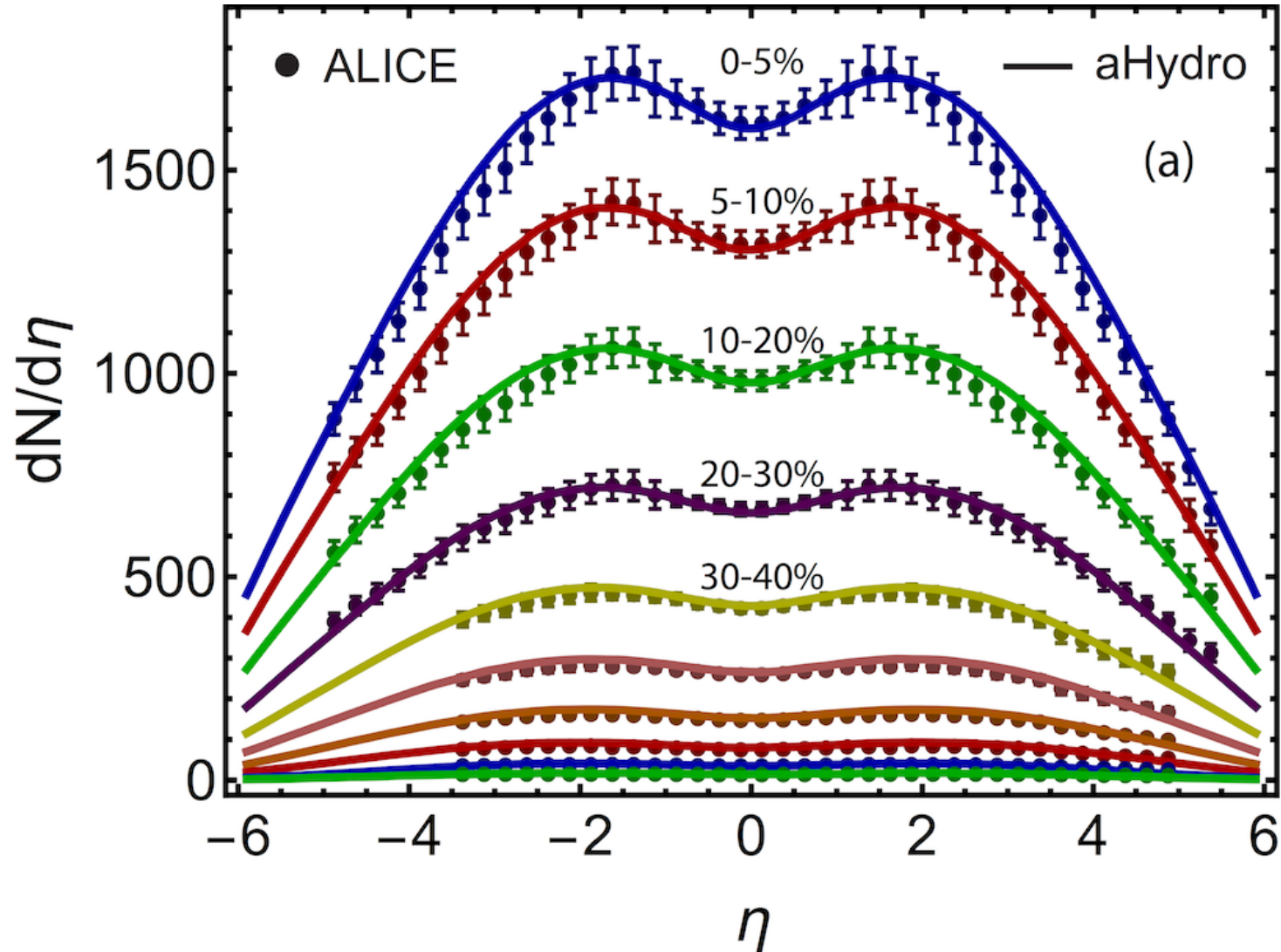
# Identified particle average $p_T$

Alqahtani, Nopoush, Ryblewski, MS, 1703.05808; 1705.10191



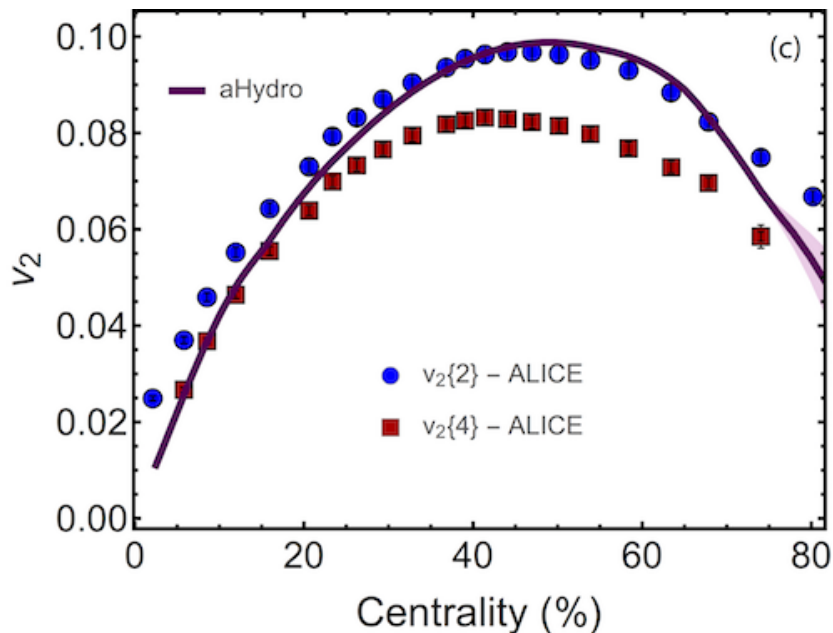
# Charged particle multiplicities

Alqahtani, Nopoush, Ryblewski, MS, 1703.05808; 1705.10191

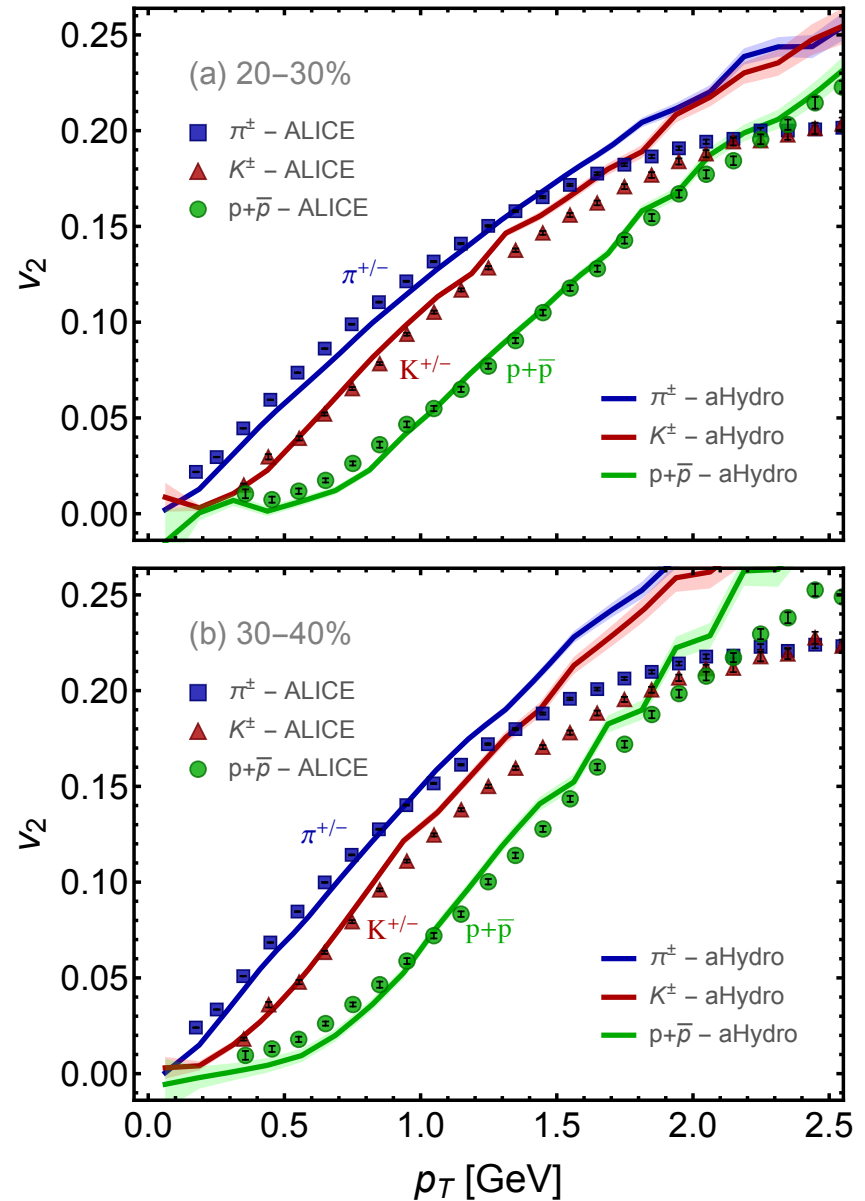


# Elliptic flow

- Quite good description of elliptic flow as well
- Problems for central collisions but this is to be expected since we have not included fluctuating initial conditions yet

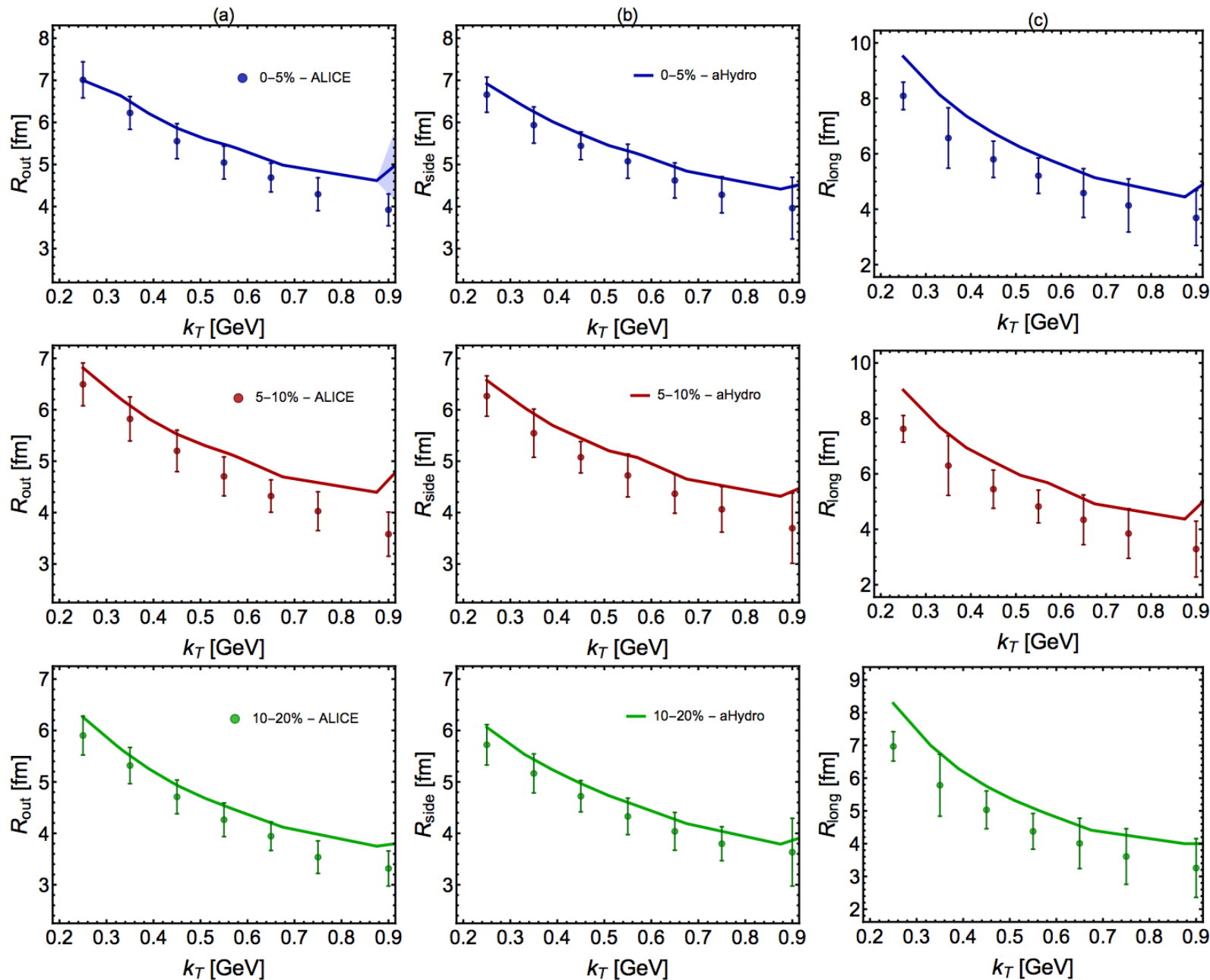


Alqahtani, Nopoush, Ryblewski, MS, 1703.05808; 1705.10191

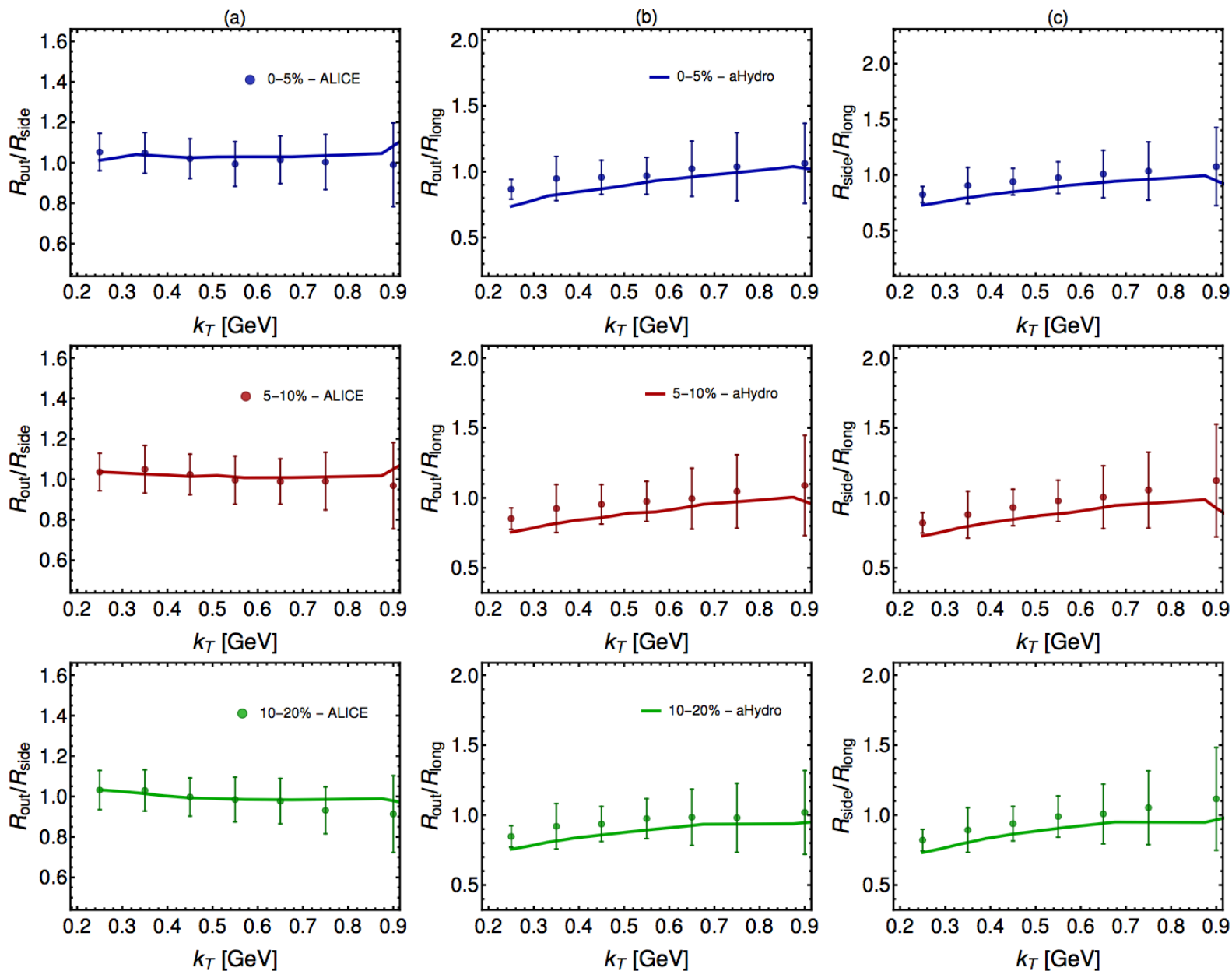


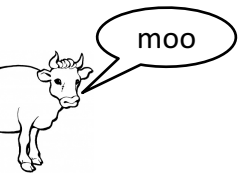


# HBT Radii



# HBT Radii Ratios



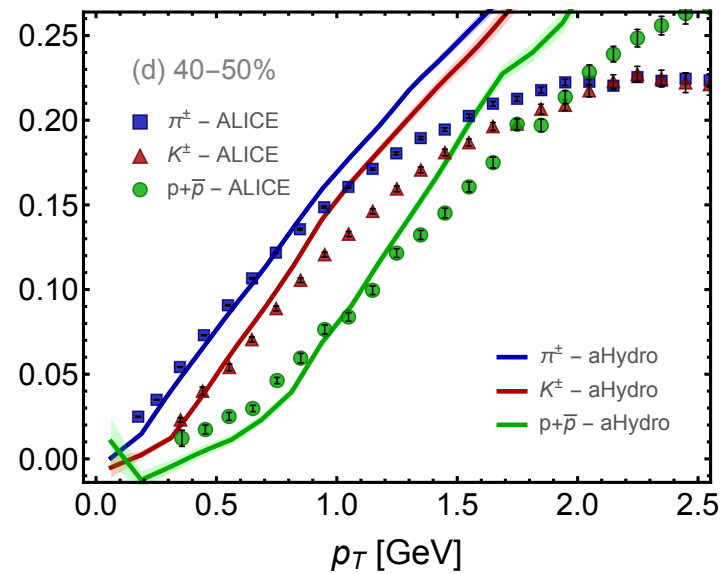
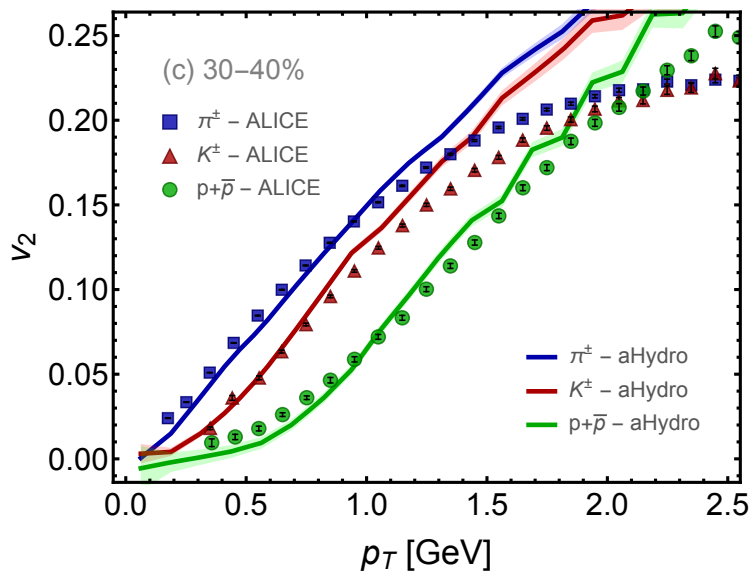
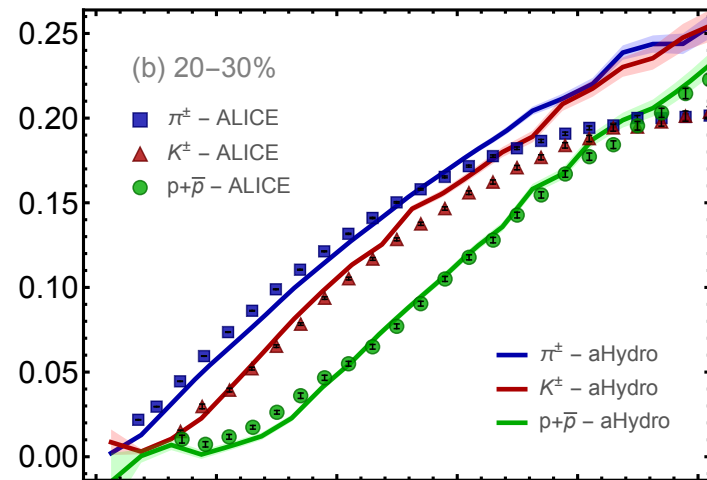
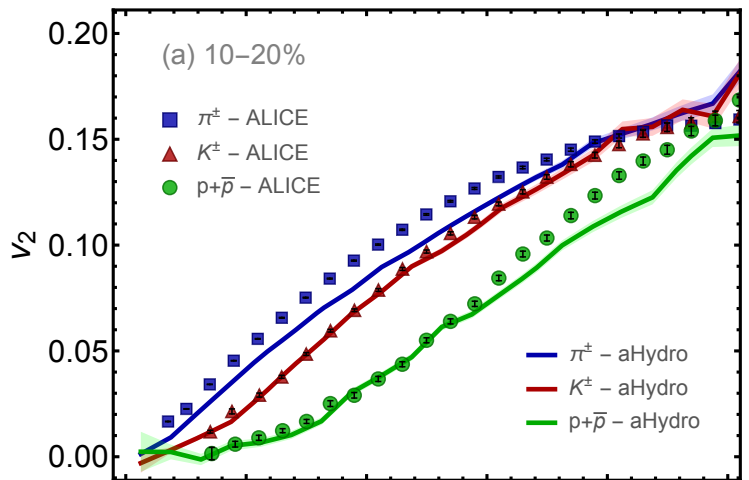


# Conclusions and Outlook

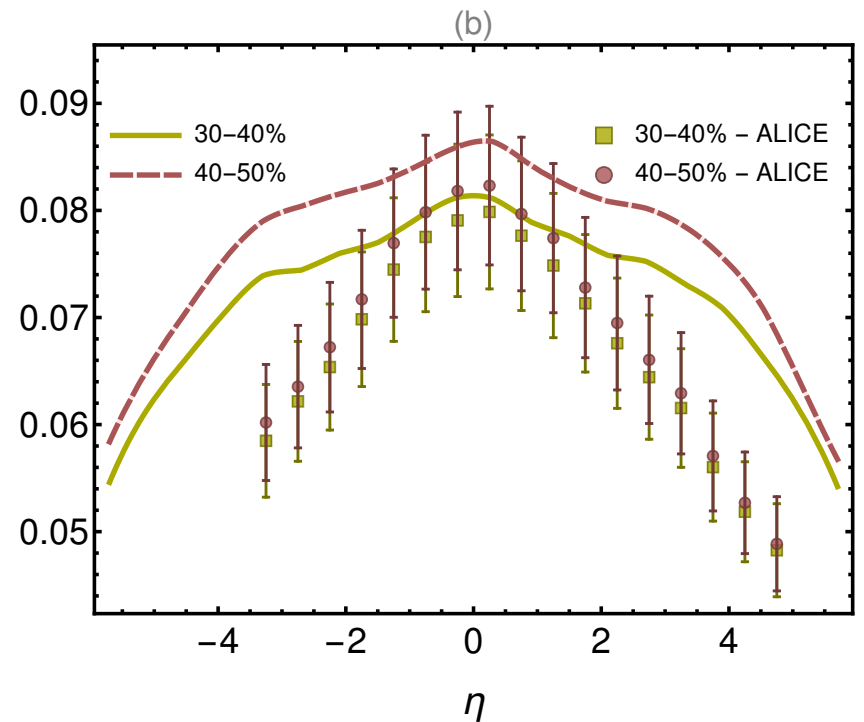
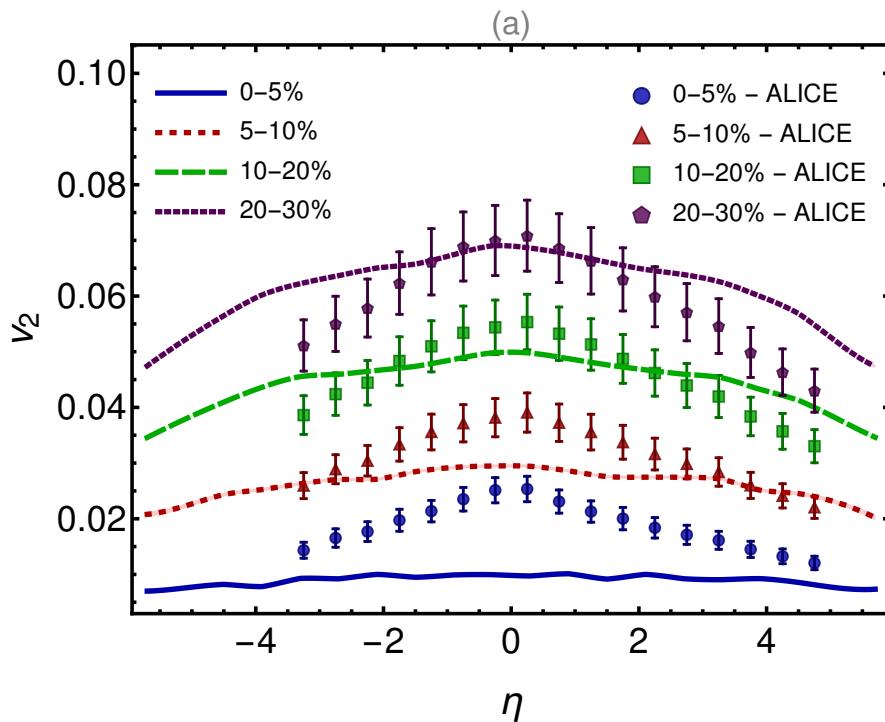
- Anisotropic hydrodynamics builds upon prior advances in relativistic hydrodynamics in an attempt to create an even more quantitatively reliable model of QGP evolution.
- It incorporates some “facts of life” specific to the conditions generated in relativistic heavy ion collisions and, in doing so, **optimizes the dissipative hydrodynamics approach for HIC.**
- We now have a running 3+1d “ellipsoidal” aHydro code with realistic EoS, anisotropic freeze-out, and fluctuating initial conditions.
- Our preliminary fits to experimental data using smooth Glauber initial conditions look quite nice.
- **Future:** off-diagonal anisotropies, turn on the fluctuating initial conditions, lower-energies/finite  $\mu_B$ , small systems...

# Backup slides

# More figures #1



# More figures #2

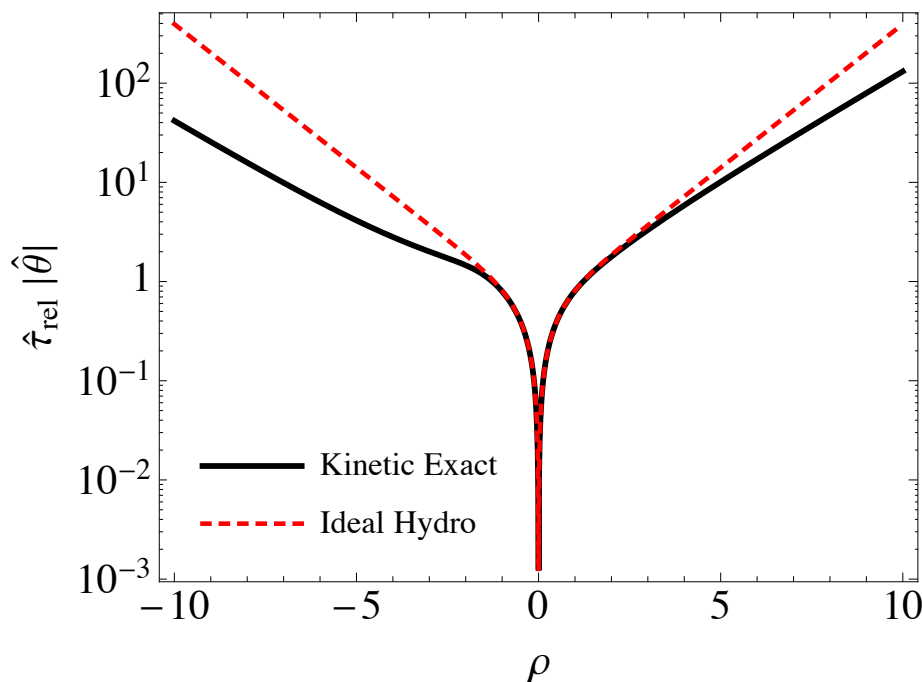


# Gubser flow is extreme!

Knudsen number in de Sitter coordinates

$$\text{Kn} = \hat{\tau}_{\text{micro}} / \hat{\tau}_{\text{macro}} = \hat{\tau}_{\text{rel}} |\hat{\theta}| \equiv \underbrace{\hat{\tau}_{\text{rel}}}_{c/\hat{T}} \underbrace{|\hat{\nabla} \cdot \hat{u}|}_{2 \tanh(\rho)}$$

$$4\pi\eta/s = 1 \quad \rho_0 = 0 \quad \hat{E}(\rho_0) = 1$$

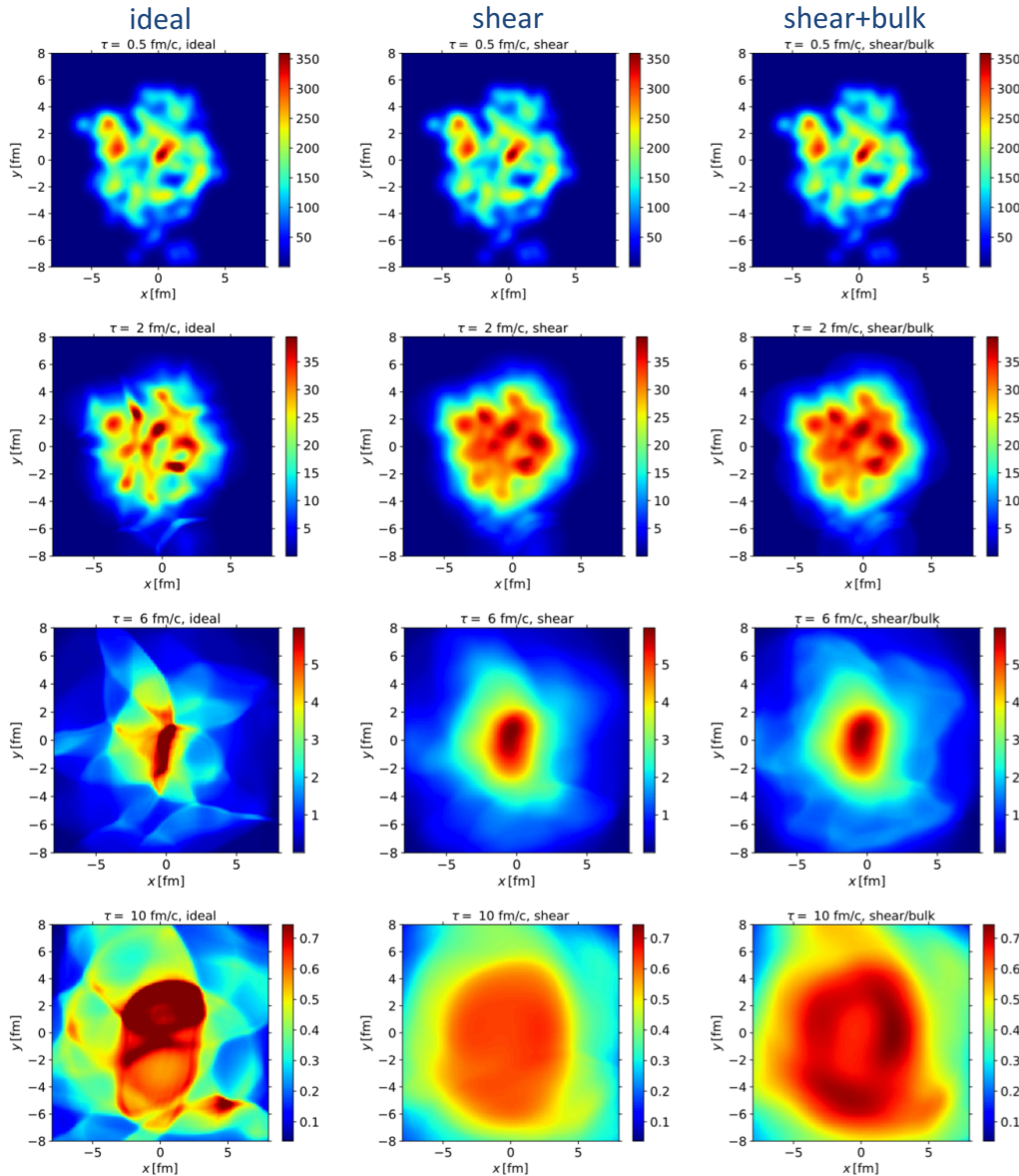


- Exponentially large gradients at early and late de Sitter times!
- Large gradients  $\rightarrow$  potential problems with hydro description

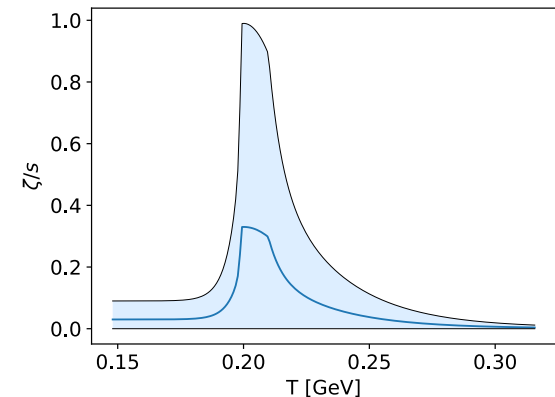
Ideal Solution: S. Gubser, 1006.0006; S. Gubser and Y. Yarom, 1012.1314

Exact Solution: G. Denicol, U.W. Heinz, M. Martinez, J. Noronha, and MS, 1408.5646; 1408.7048

# Some pretty pictures from 3d viscous hydro



- Left panels show output from the Ohio State/Kent State GPU-based viscous hydro code [Bazow, Heinz, and MS, 1608.06577]
- Solves the non-conformal DNMR (Denicol, Niemi, Molnar, Rischke) equations with a realistic EoS
- Parameterized  $\zeta/s$  (plot below)
- $\eta/s = 0.2$
- $T_0 = 600 \text{ MeV @ } t_0 = 0.5 \text{ fm/c}$





# Technicalities - A numerical challenge

- One of the most daunting challenges faced by the quasiparticle approach is that one has to evaluate a bunch of “ $\mathcal{H}$ ” functions, e.g.

$$\mathcal{E} = \mathcal{H}_3(\boldsymbol{\alpha}, \hat{m}) \lambda^4 + B$$

$$\mathcal{H}_3(\boldsymbol{\alpha}, \hat{m}) = \tilde{N}\alpha \int d^3\hat{p} \mathcal{R}(\boldsymbol{\alpha}, \hat{m}) f_{\text{eq}}\left(\sqrt{\hat{p}^2 + \hat{m}^2}\right)$$

$$\mathcal{R}(\boldsymbol{\alpha}, \hat{m}) = \sqrt{\alpha_x^2 p_x^2 + \alpha_y^2 p_y^2 + \alpha_z^2 p_z^2 + m^2}$$

- We evaluate these efficiently by expanding the integrand around the diagonal in anisotropy space up to 12<sup>th</sup> order.
- We do this around two points (1,1,1) and (2,2,2) and switch between these two expansions smoothly.
- With this method we were able to accelerate the evaluation of H functions by a factor of  $10^5$  while achieving  $< 0.1\%$  accuracy.