

A CMB Millikan Experiment with Axion Strings

Remote Seminar
University of Oxford
May 6, 2020

Prateek Agrawal

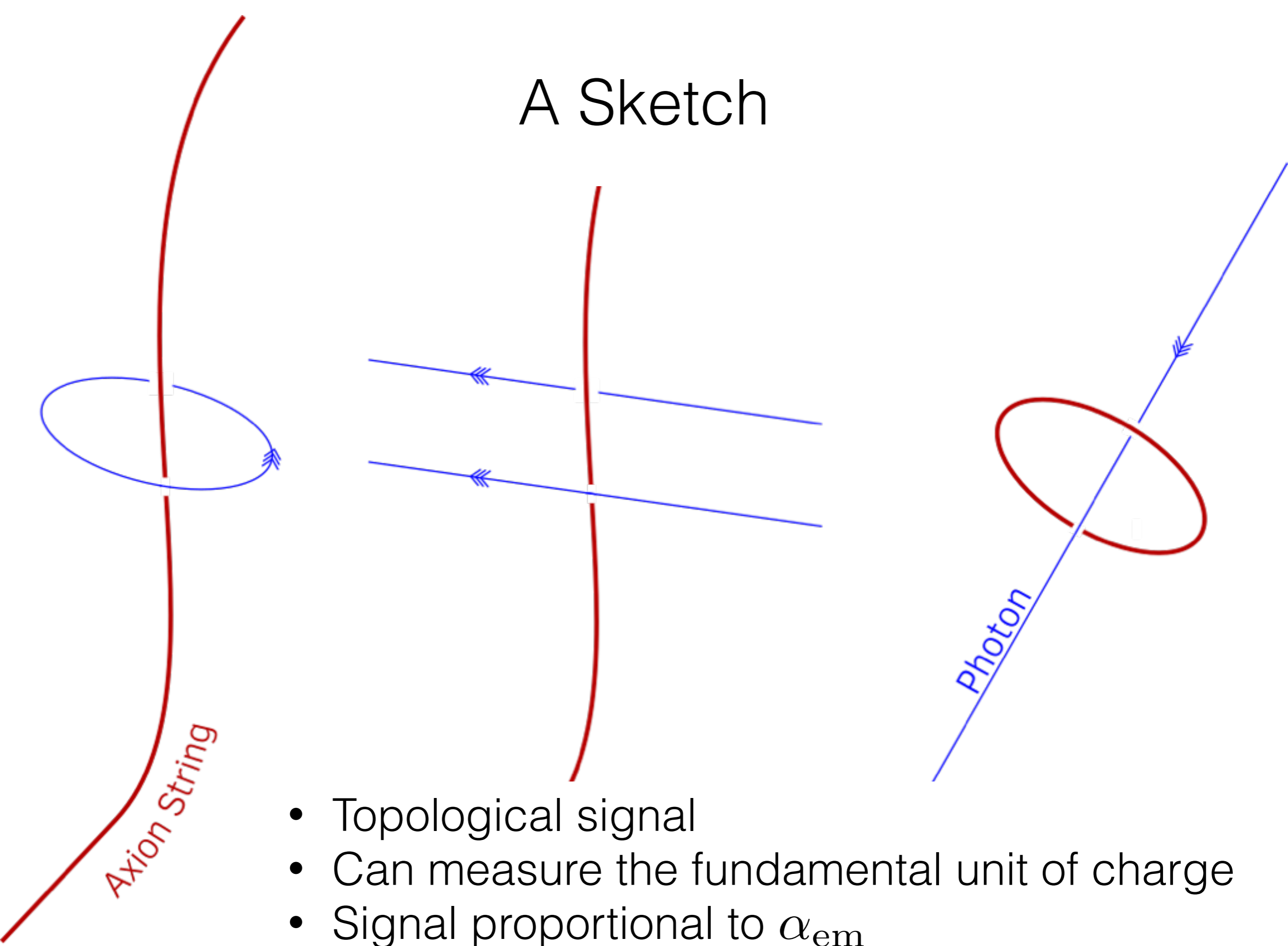


HARVARD
UNIVERSITY

[1912.02823]

PA, Anson Hook, Junwu Huang

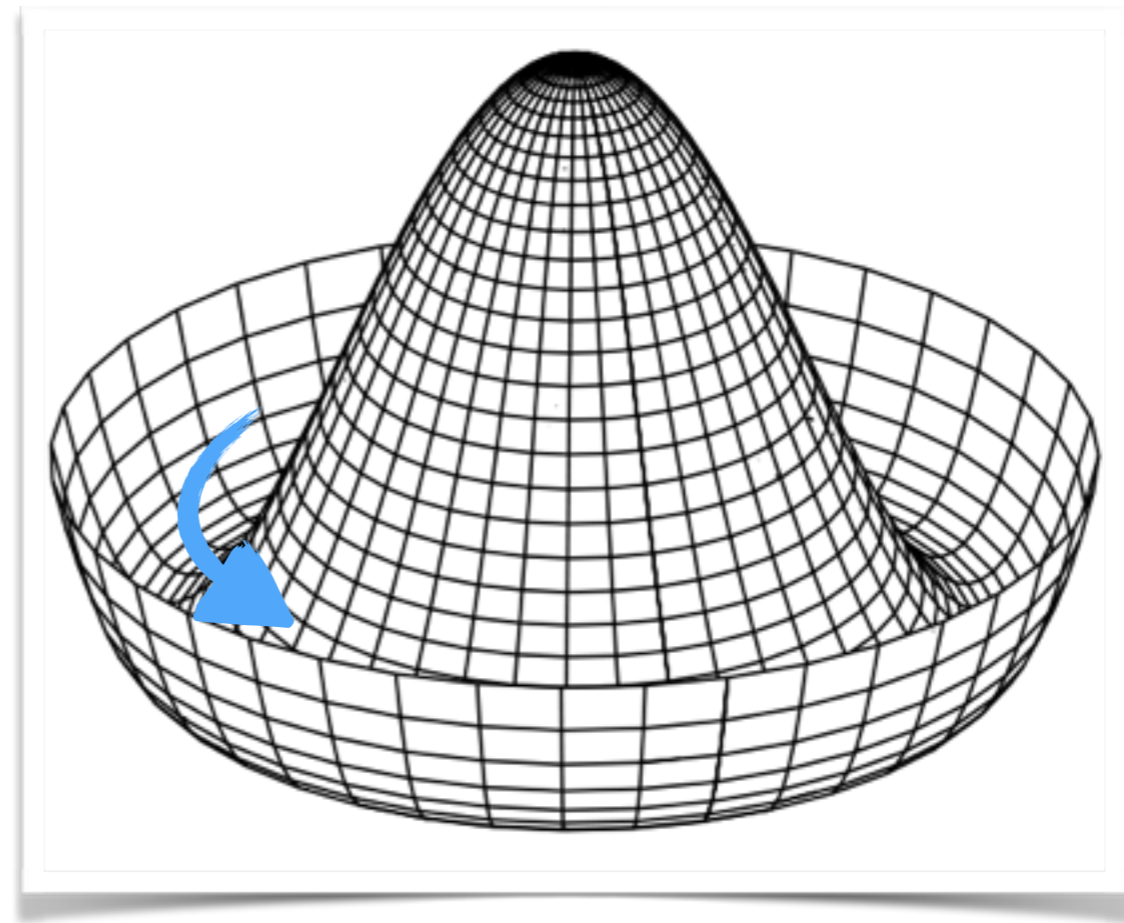
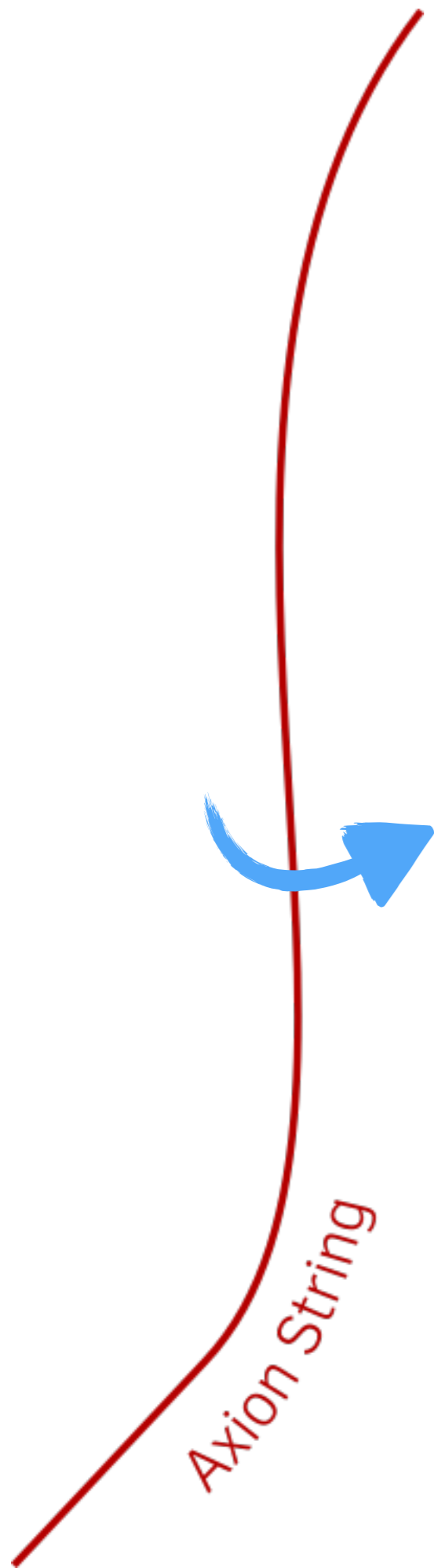
A Sketch



- Topological signal
- Can measure the fundamental unit of charge
- Signal proportional to α_{em}

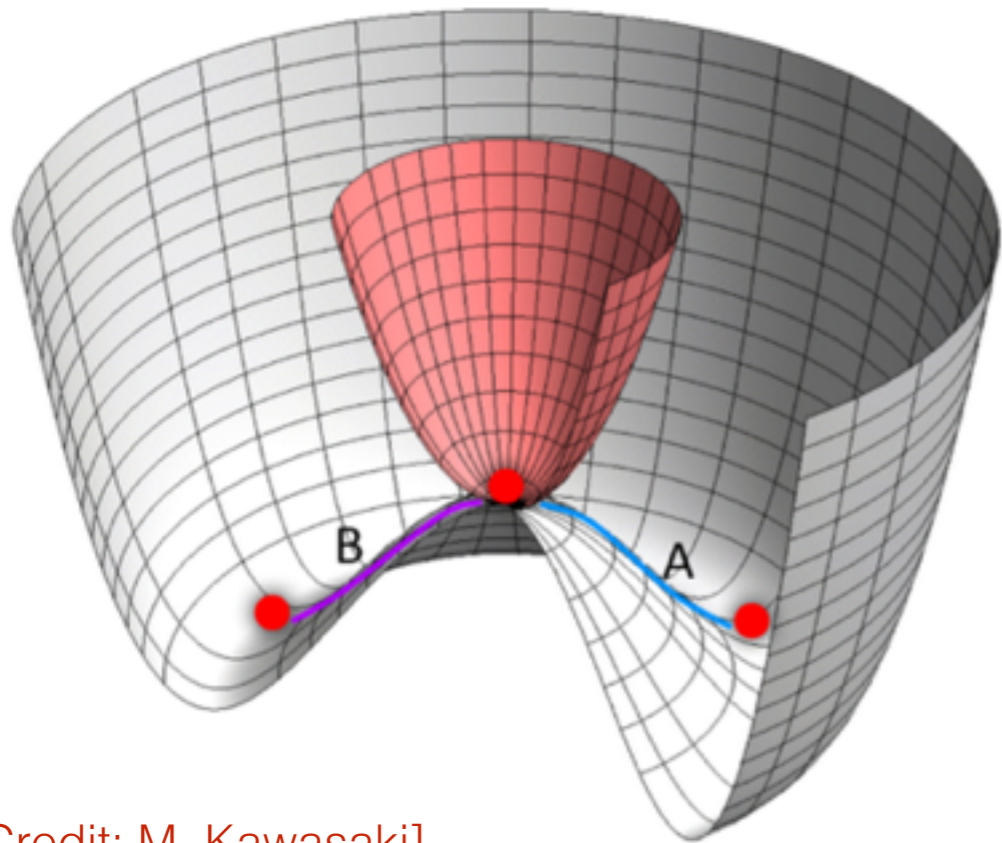
Axion Strings

Spontaneously broken Peccei-Quinn symmetry



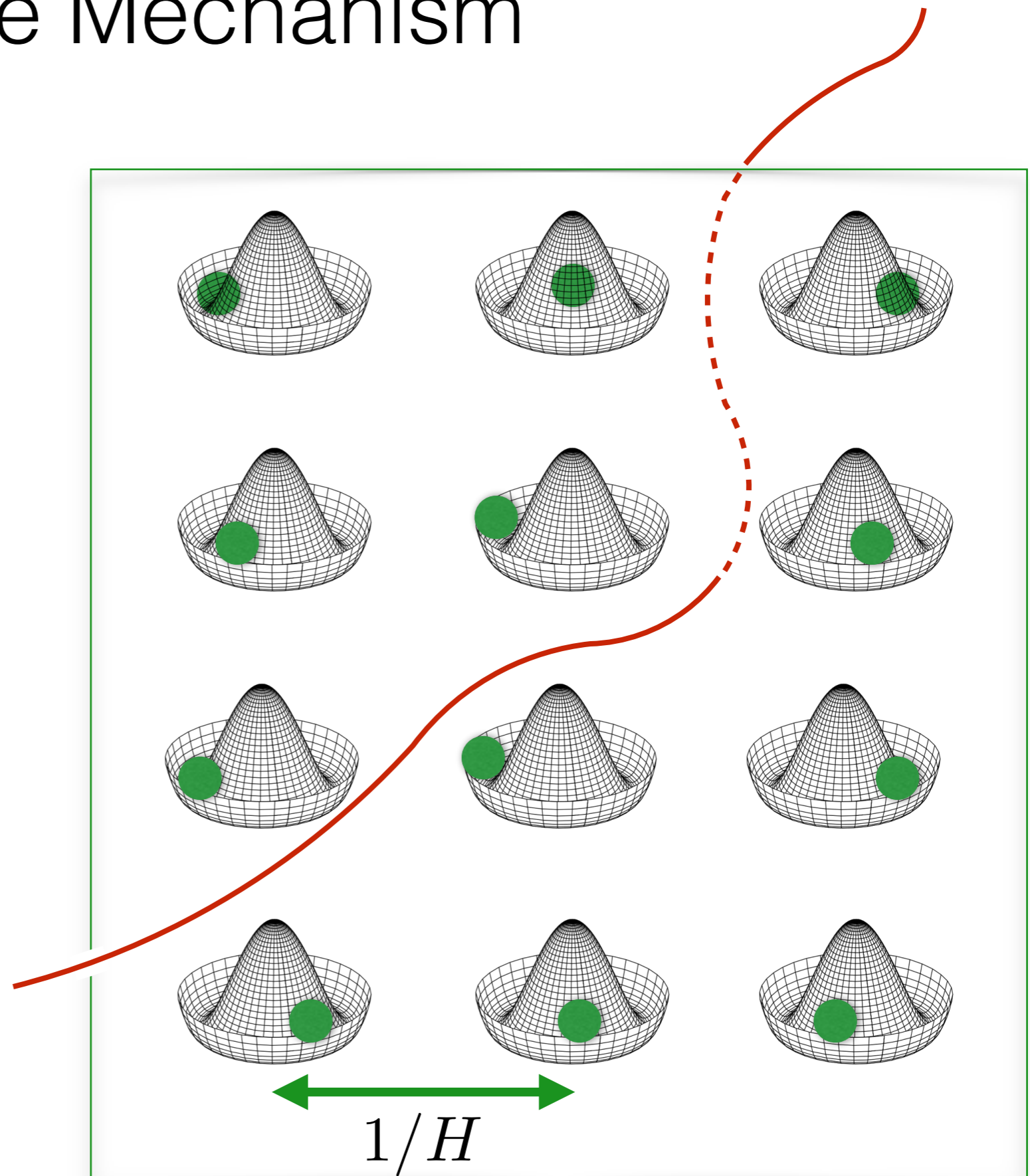
$$a \rightarrow a + 2\pi f$$

Kibble Mechanism

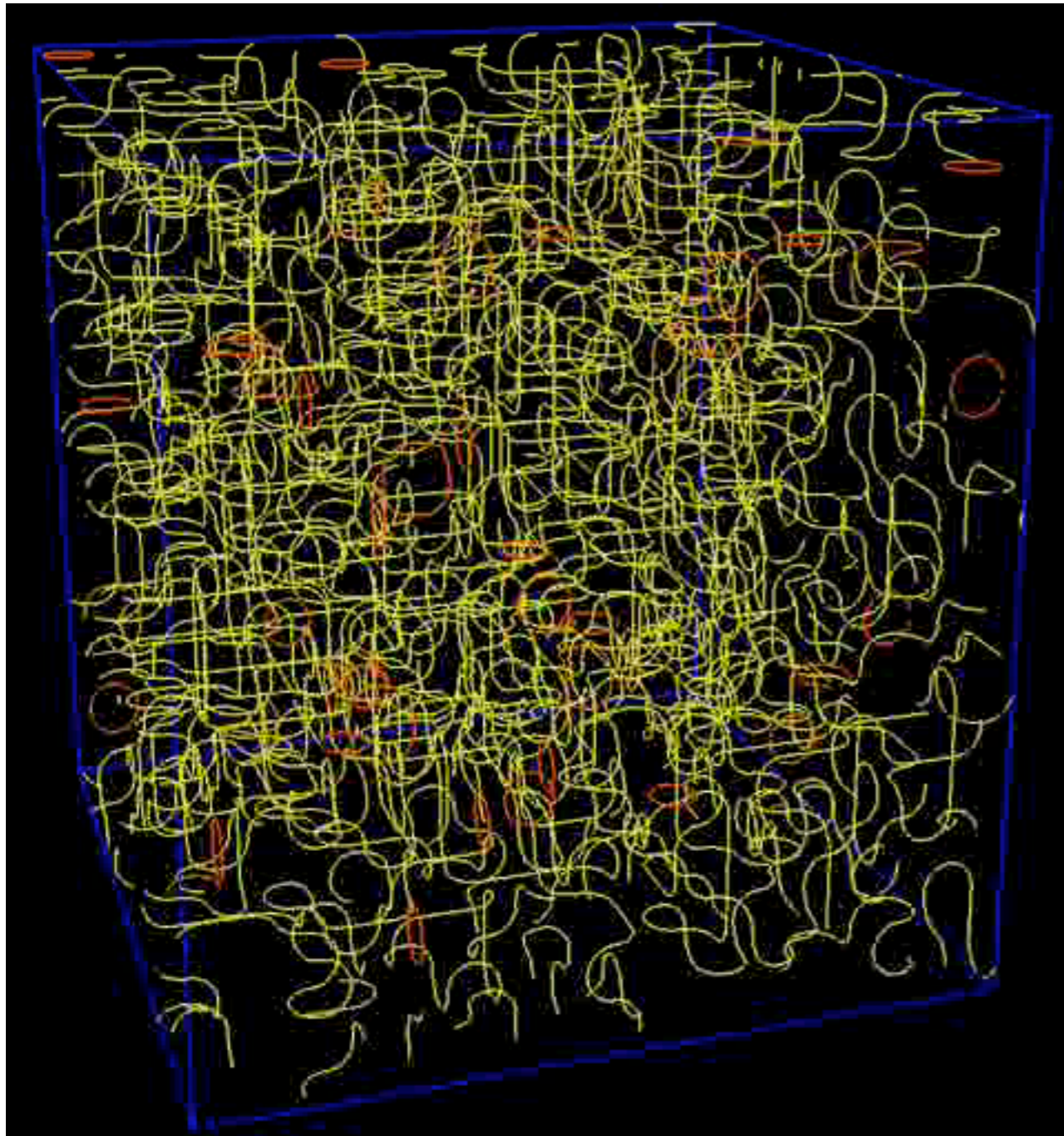


[Credit: M. Kawasaki]

Phase transition to the broken state in the early universe



The String Network



String interactions are complicated,
understood by numerical simulations

String energy density follows a scaling law

$$\rho_{\text{strings}} \simeq \xi \mu H^2$$

$$10^3 > \xi > 1$$

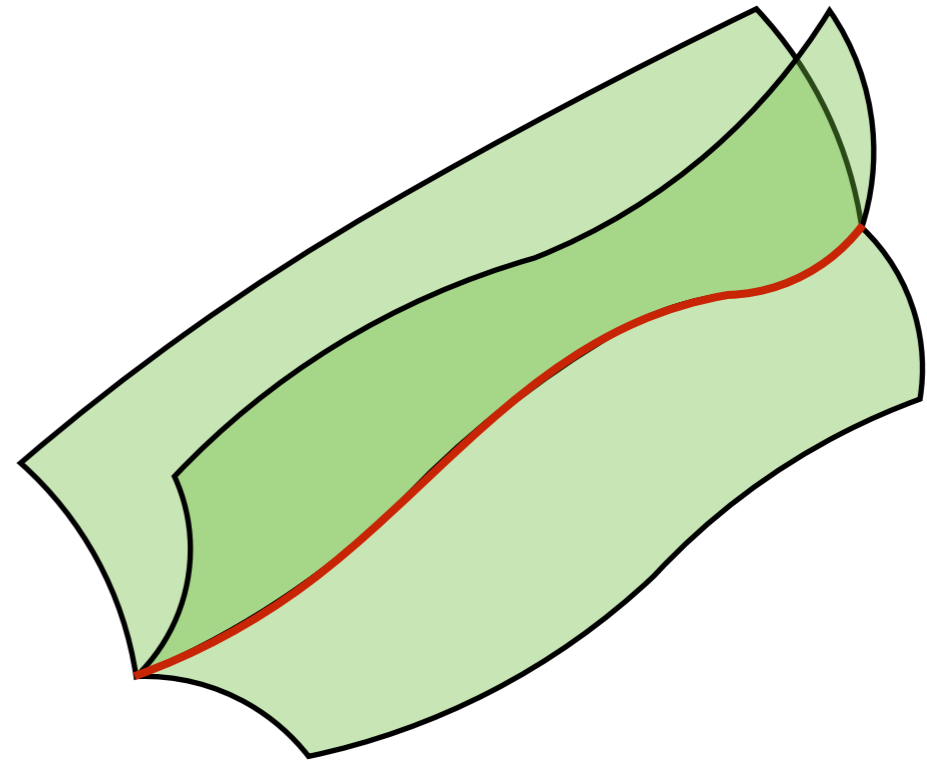
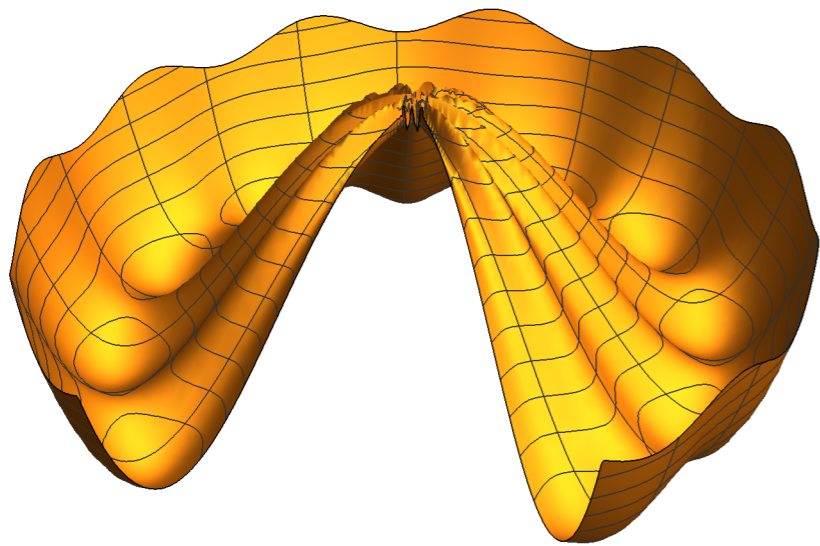
Equivalent to ξ strings per Hubble volume

Network is dominated by infinitely long
strings with structure at scale $1/H$

For massless axions:
Once formed, there are always a
few strings per Hubble

[C. Martins & E. P. Shellard]

Axion mass and domain walls

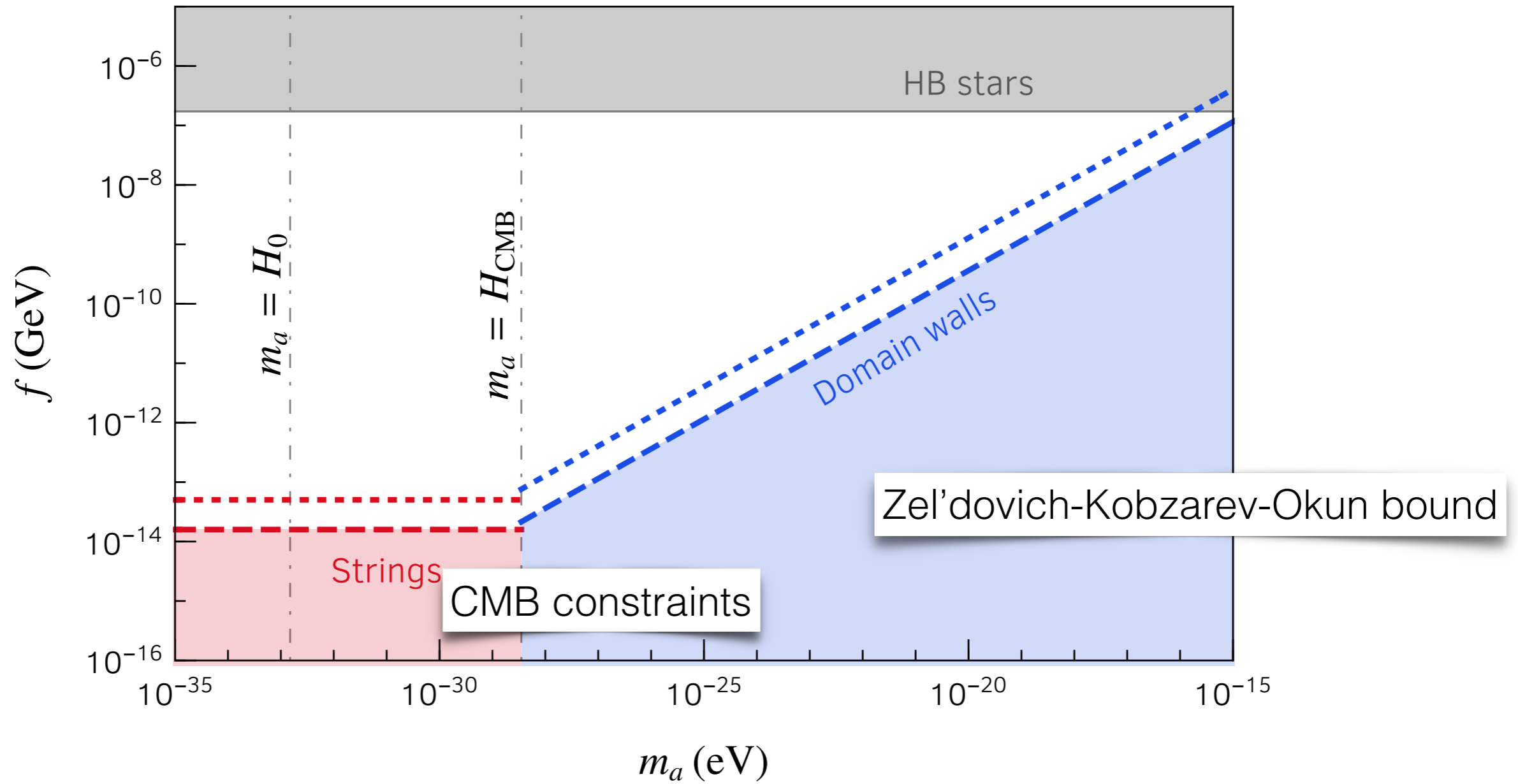


When $H < m_a$, domain walls ending on strings form

$N_{\text{DW}} = 1$ String network disappears soon after

$N_{\text{DW}} > 1$ String/domain wall network survives

Hyperlight axions



Not QCD axion, not dark matter

The String Axiverse

Hyperlight axions are ubiquitous in string compactifications

$$\mathcal{L} = \frac{A\alpha_{\text{em}}}{4\pi f} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

Axions are light, protected by an approximate shift symmetry

$$a \rightarrow a + c$$

Axions get a mass from instantons, can be exponentially suppressed

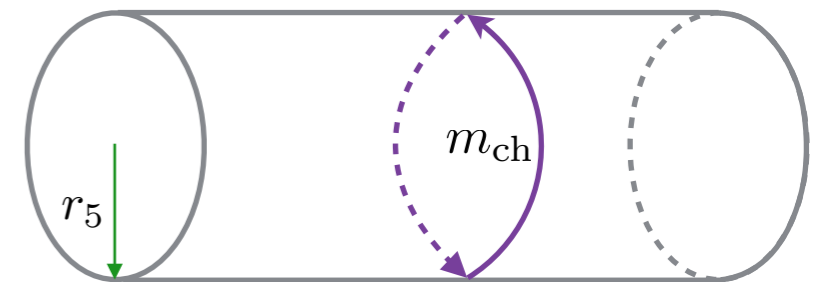
Toy example: Gauge theory in a theory with one extra dimension

$$A_M \equiv (A_\mu, A_5)$$

Only contribution to potential from charged particles around the circle

$$V(A_5) \sim \exp(-m_{\text{ch}} r_5) \cos(A_5 r_5)$$

$$V(a) \sim \exp(-M_{\text{pl}}/f) \cos(a/f)$$



“hundreds of axions, some of them massless”

Photons in Axion String Background

$$\mathcal{L} = \frac{\mathcal{A}\alpha_{\text{em}}}{4\pi f} a F_{\mu\nu} \tilde{F}^{\mu\nu} \propto a \vec{E} \cdot \vec{B}$$

Solve plane waves in axion electrodynamics

$$A_{\pm}(\eta, z) = A_{\pm}(0, 0) e^{i(kz - \omega\eta)} e^{\pm i\Delta\Phi(\eta, z)}$$

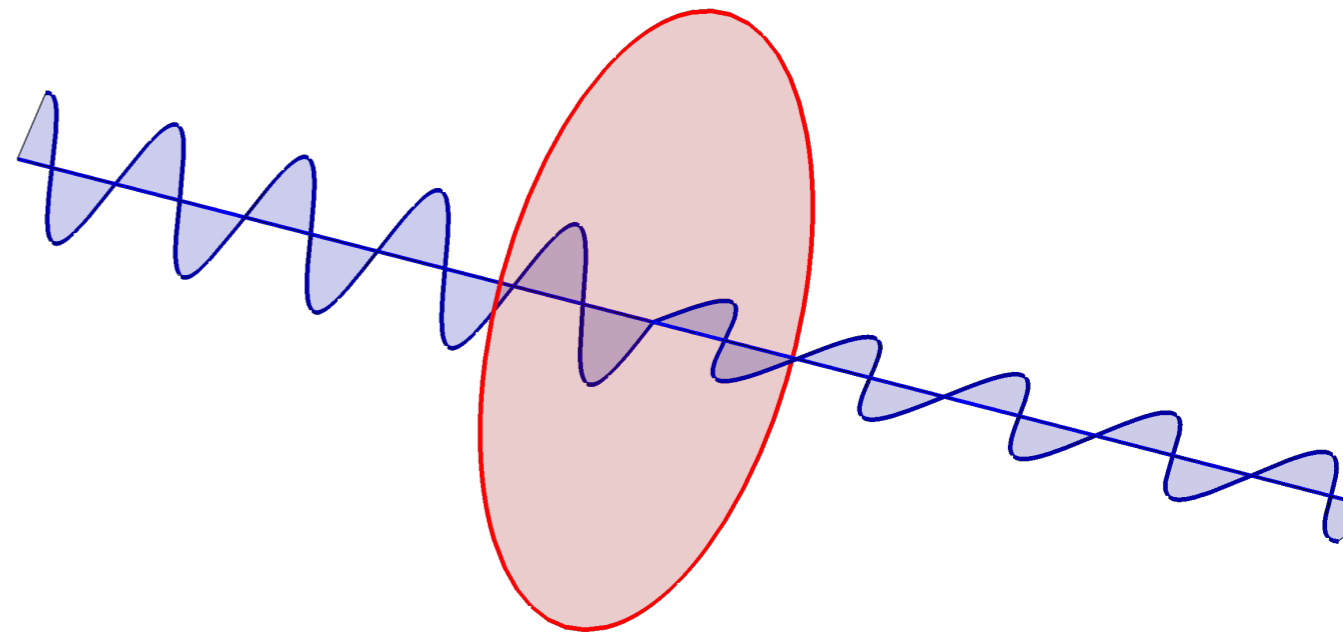
$$\Delta\Phi(\eta, z) = \frac{\mathcal{A}\alpha_{\text{em}}}{2\pi f} (a(\eta, z) - a(0, 0))$$

Rotation of linear polarization: axion birefringence

Aharonov-Bohm like effect for trajectory around a string $\Delta a = 2\pi f$

$$\Delta\Phi = \mathcal{A}\alpha_{\text{em}}$$

Access to measuring \mathcal{A} directly!



Axions and charge quantization

In the SM, all gauge invariant states (leptons, hadrons) carry integer electric charge

$$\mathcal{L} = \frac{\mathcal{A}\alpha_{\text{em}}}{4\pi f} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

The axion - photon coupling is quantized in units of fundamental EM charge

$$\mathcal{A} \in \frac{Q_{\text{fund}}}{Q_e} \times \mathbb{Z}$$

Usually, this is only true up to mass mixing effects for particles.

E.g. for the QCD axion, in the mass basis

$$2\mathcal{A} = \frac{E}{N} - 1.92 \sim \frac{E}{N} - \frac{m_a^2 f_a^2}{m_\pi^2 f_\pi^2}$$

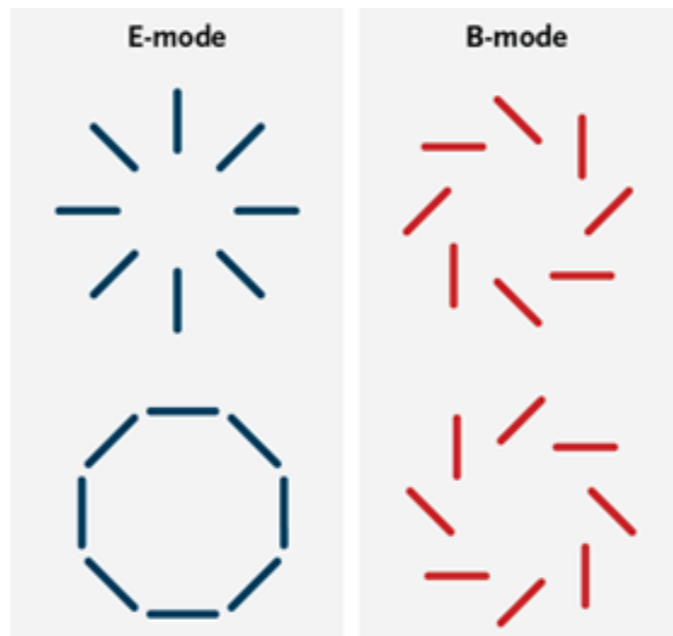
However, around axion strings, both the axion and the pion shift, so

$$\Delta\Phi = \mathcal{A}\alpha_{\text{em}} \quad \text{with} \quad \mathcal{A} \in \frac{Q_{\text{fund}}}{Q_e} \times \mathbb{Z}$$

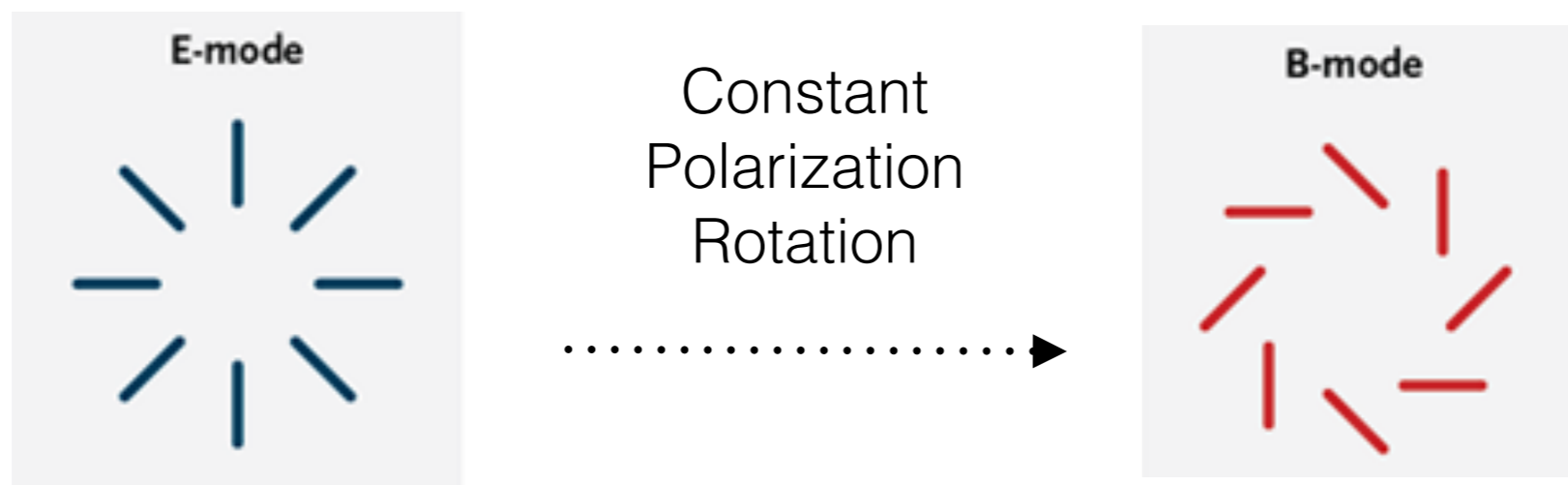
Measuring \mathcal{A} can test the fundamental unit of electric charge

CMB Observables

CMB polarization can be decomposed in curl-free (E-mode) and divergence-free (B-mode)



Correlated B-modes generated from E-modes




Cosmic Birefringence

For angle dependent rotation $\Phi(\hat{n})$, B-modes are convolution of Φ_{LM} and E-modes

$$B_{lm} = 2 \sum_{LM} \sum_{l'm'} \Phi_{LM} E_{l'm'} \Xi_{lm'l'm'}^{LM} H_{ll'}^L$$

Functions of Clebsch-Gordan coefficients



Estimator for Φ_{LM} from E- and B-mode maps

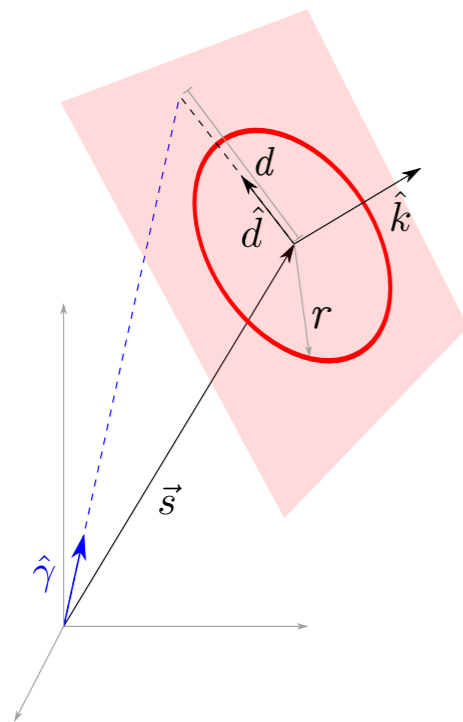
$$[\hat{\Phi}_{LM}^{E^i B^j}]_{ll'} = \frac{2\pi}{(2l+1)(2l'+1)C_l^{EE} H_{ll'}^L} \sum_{mm'} B_{lm}^i E_{l'm'}^{j*} \Xi_{lm'l'm'}^{LM}$$

Can be used to estimate the variance of the estimator from noise and background sources

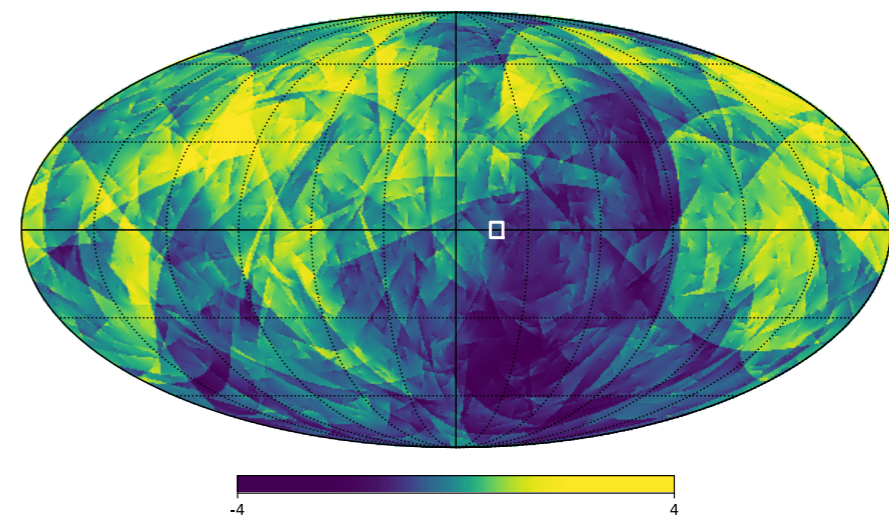
Theory predictions

Study the polarization rotation in two simplified settings

1. Semi-analytical approach



2. Simple numerical simulation



Future direction: Set up a string simulation for hyperlight axions combined with a CMB simulation

1. Semi-analytical approach

Model String network by

- Circular loops of comoving radius $1/aH$
- Total number of strings follow scaling $\rho_{\text{strings}} \simeq \xi \mu H^2$
- Spatially uniform, random orientation

Further assume that photons passing through the loop pick up rotation $\mathcal{A}\alpha_{\text{em}}$, and 0 otherwise

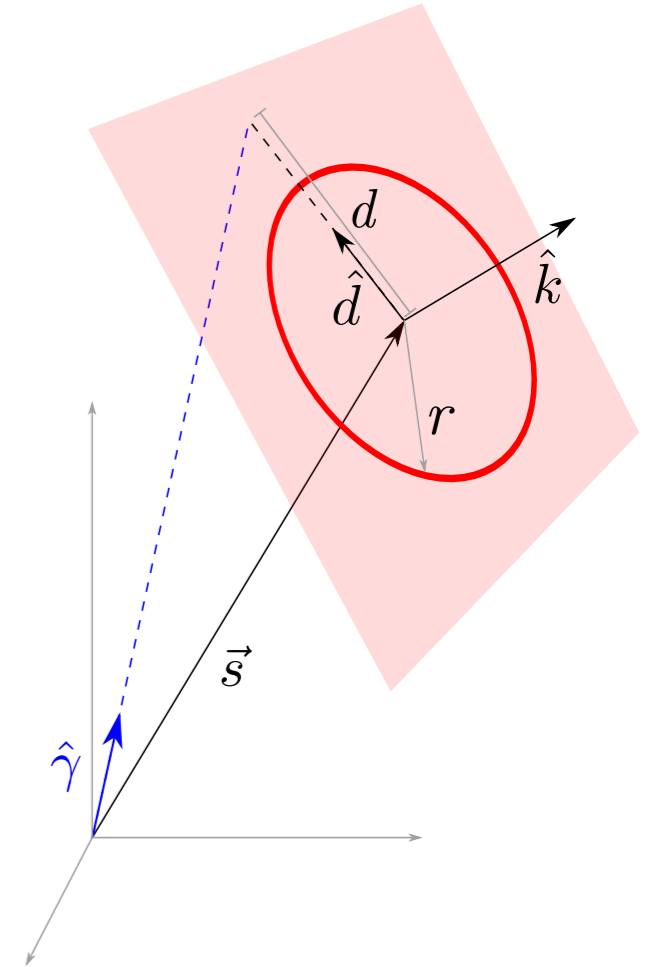
- An ok assumption for loops at smaller angular scales

Two-point function for the polarization rotation

$$\langle \Phi(\hat{\gamma}) \Phi(\hat{\gamma}') \rangle = (\mathcal{A}\alpha_{\text{em}})^2 \int d\eta \int d^2 \hat{s} \int d^2 \hat{k} (\eta_0 - \eta)^2 f(\eta) \\ \times \Theta \left(\frac{\eta}{2} - d(\hat{s}, \hat{\gamma}, \hat{k}, \eta) \right) \Theta \left(\frac{\eta}{2} - d(\hat{s}, \hat{\gamma}', \hat{k}, \eta) \right)$$

$$\langle \Phi(\hat{\gamma}) \Phi(\hat{\gamma}') \rangle = (\mathcal{A}\alpha_{\text{em}})^2 \int d[\text{string}] P([\text{string}]) \text{Pass}(\hat{\gamma}) \text{Pass}(\hat{\gamma}')$$

Variance: $\langle \Phi(0)^2 \rangle \simeq (\mathcal{A}\alpha_{\text{em}})^2 \xi \log \left(\frac{\eta_0}{\eta_{\text{CMB}}} \right)$



2. Simple Simulation

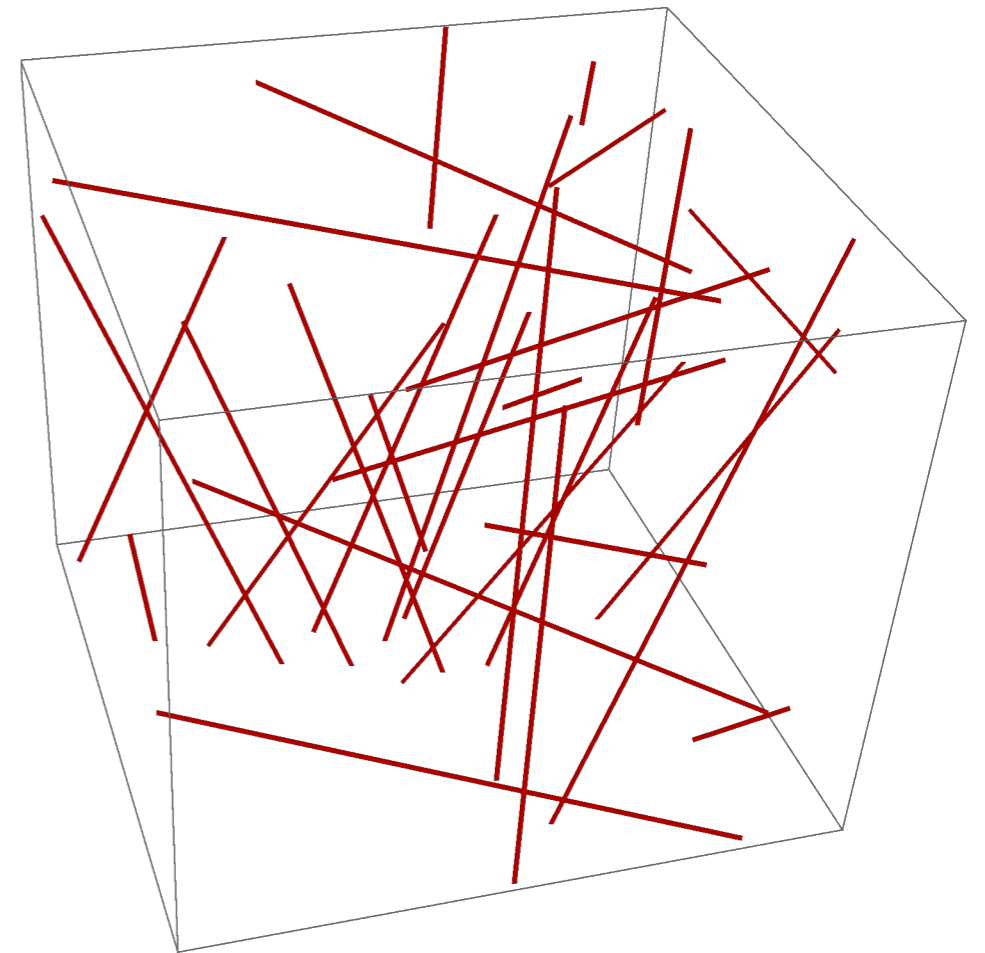
Model String network by

- Infinitely long, straight strings
- Total number of strings follow scaling $\rho_{\text{strings}} \simeq \xi \mu H^2$
- Spatially uniform, random orientation

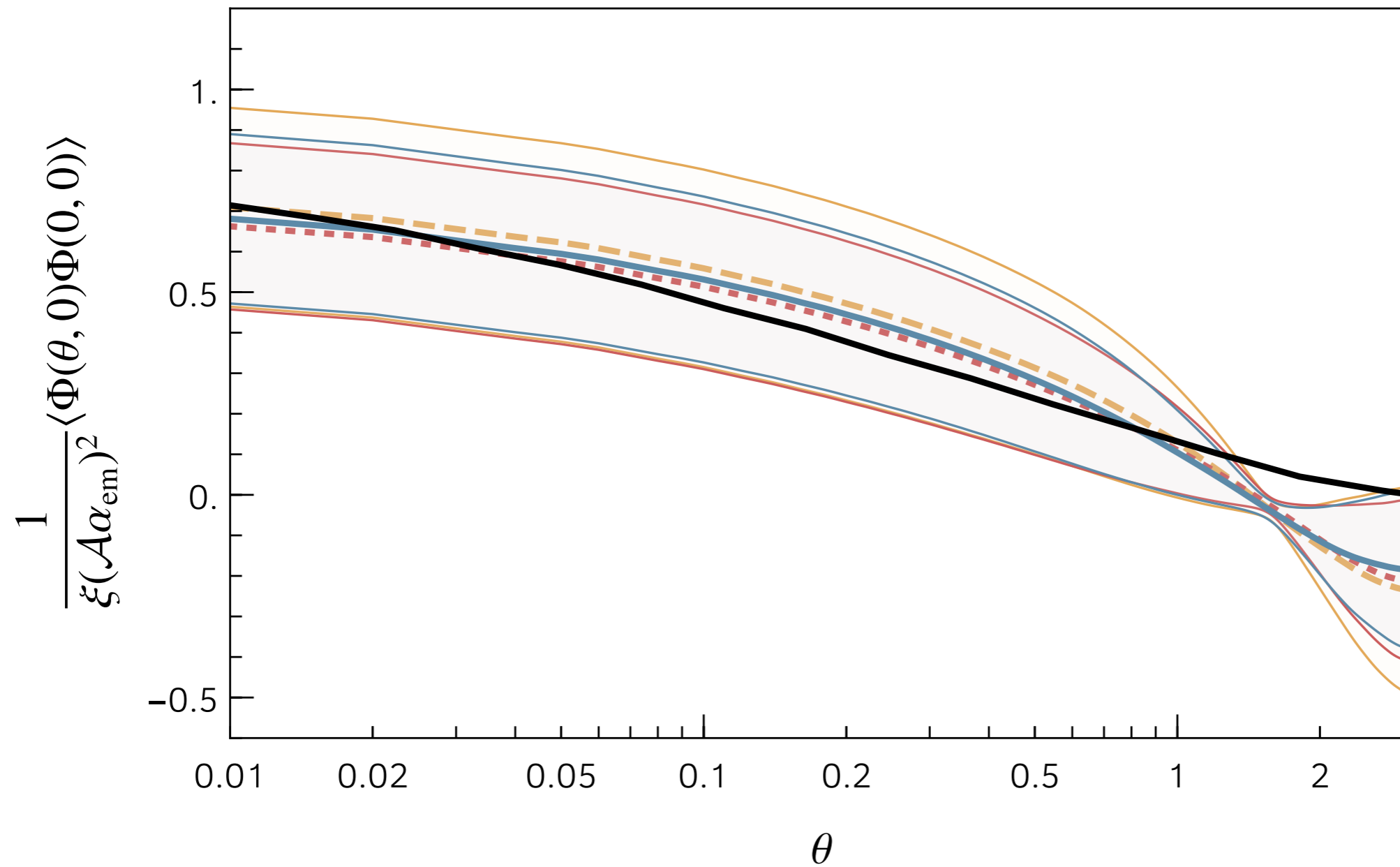
Strings are removed randomly to maintain scaling

Pass photons through this network, adding up their polarization rotations along trajectory

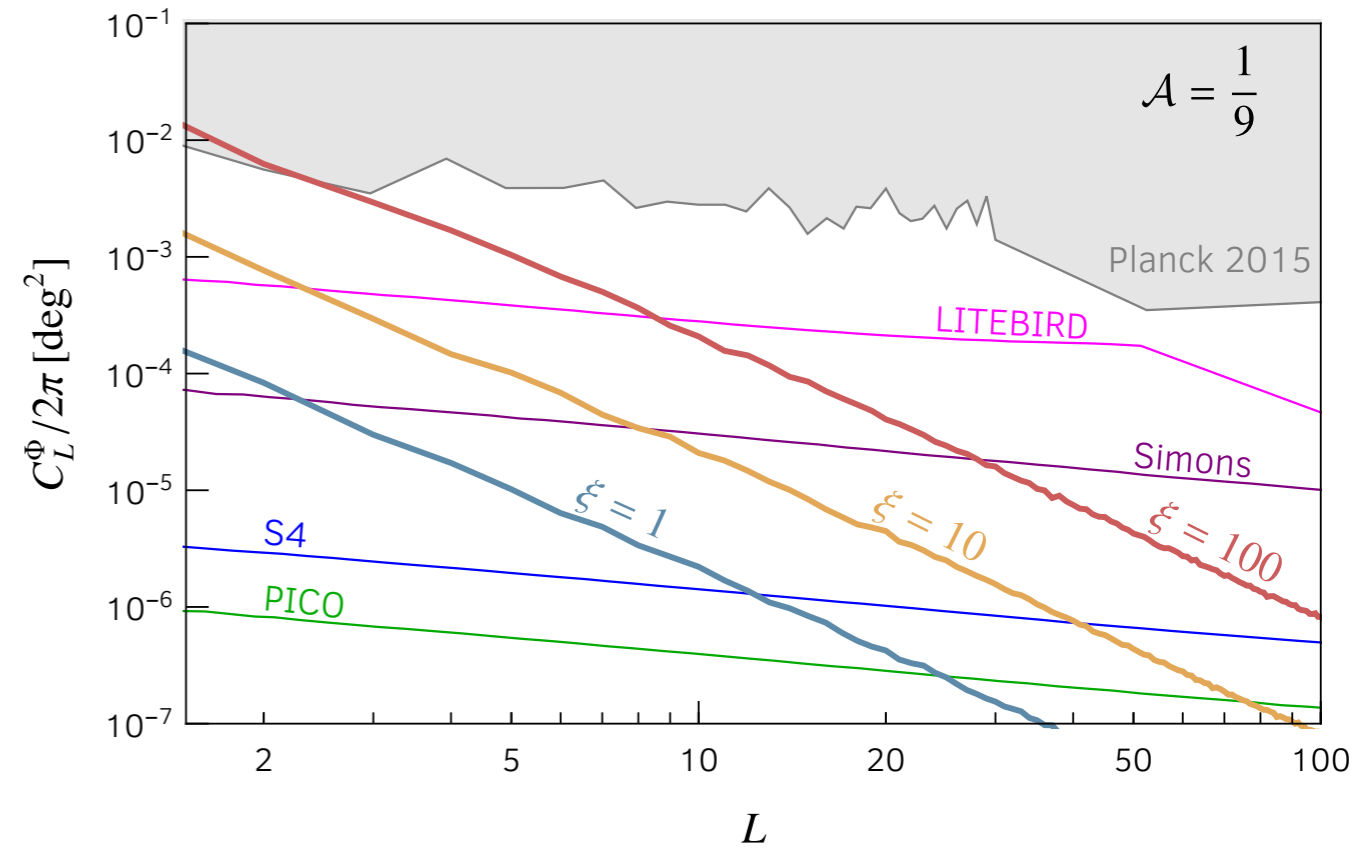
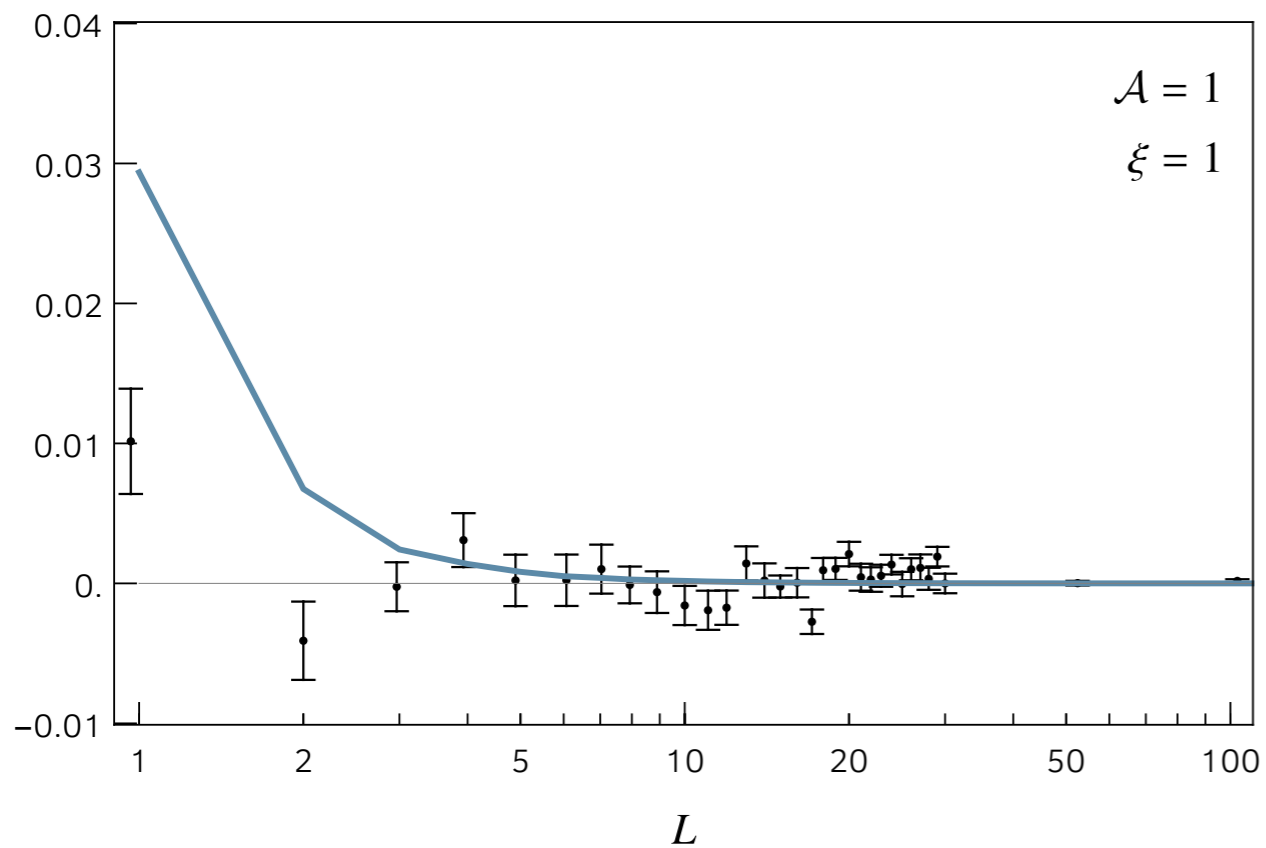
Captures larger angular scale correlations well



Two-point function



Angular power spectra



Data points

Planck 2015

Contreras, Boubel, Scott

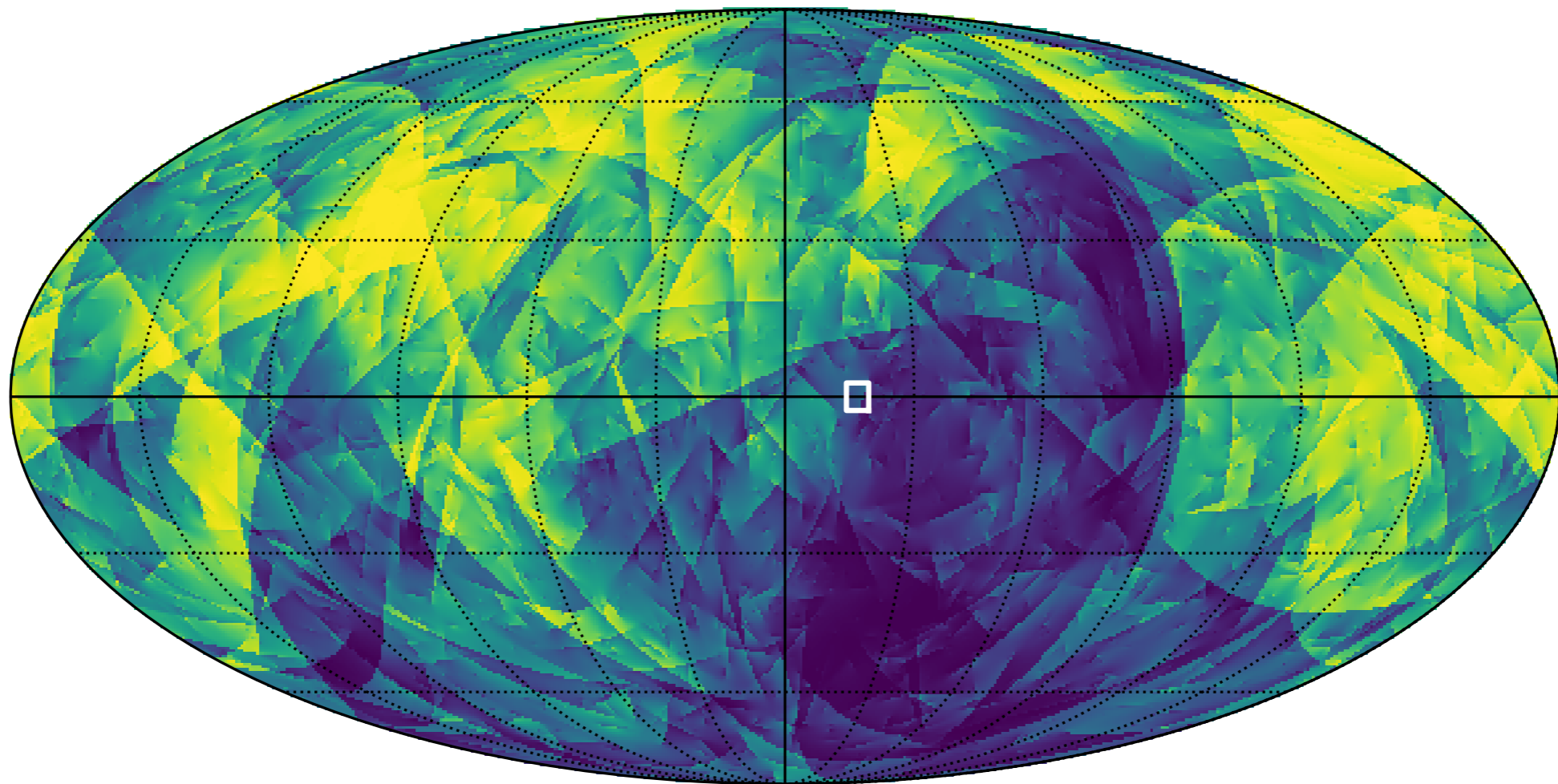
[arXiv:1705.06387]

Forecasts

Pogosian et al

[arXiv:1904.07855]

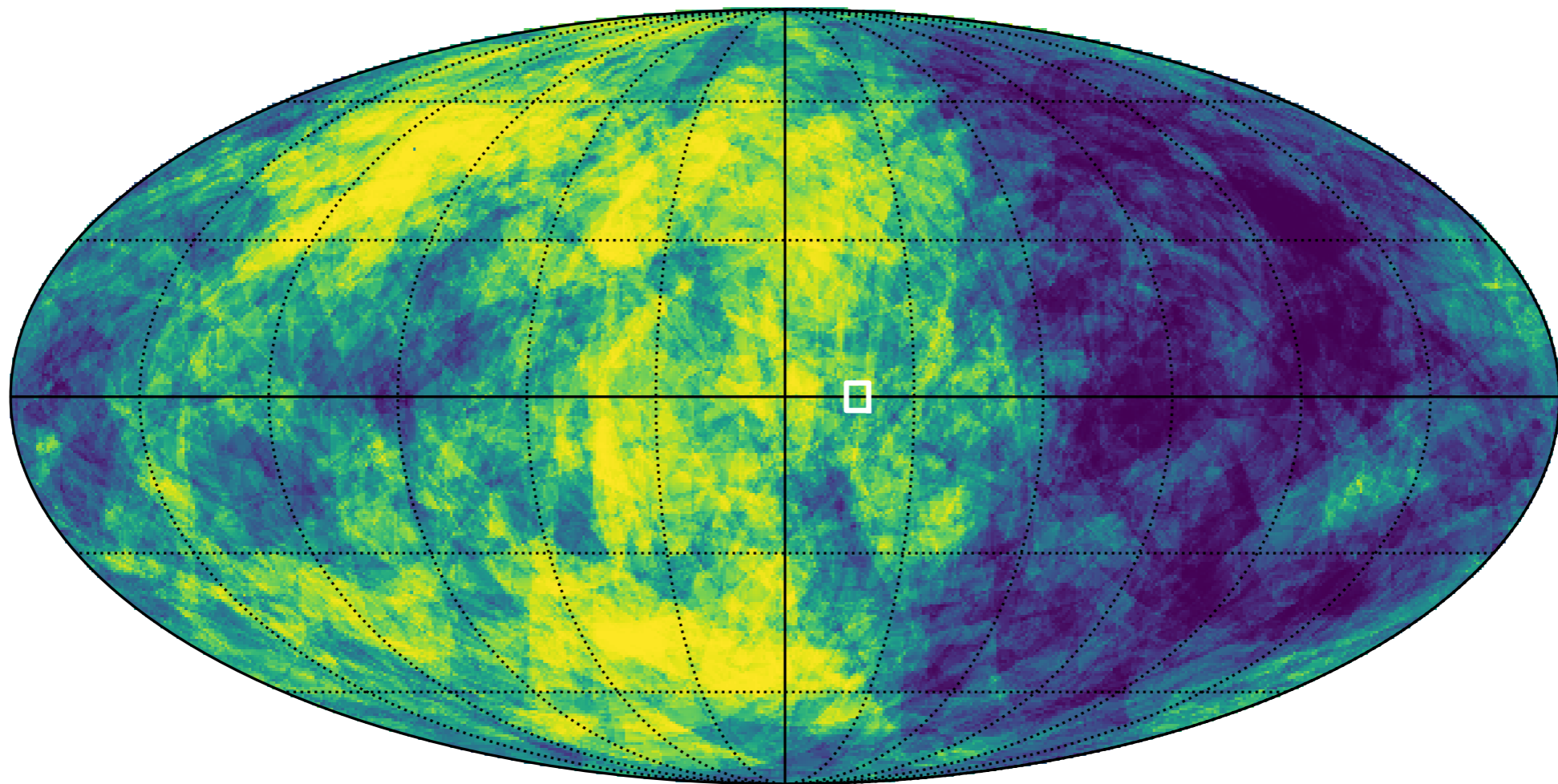
Sky maps



$$\frac{\Delta\Phi}{\sqrt{\xi}\mathcal{A}\alpha_{\text{em}}}$$

$$\xi = 1$$

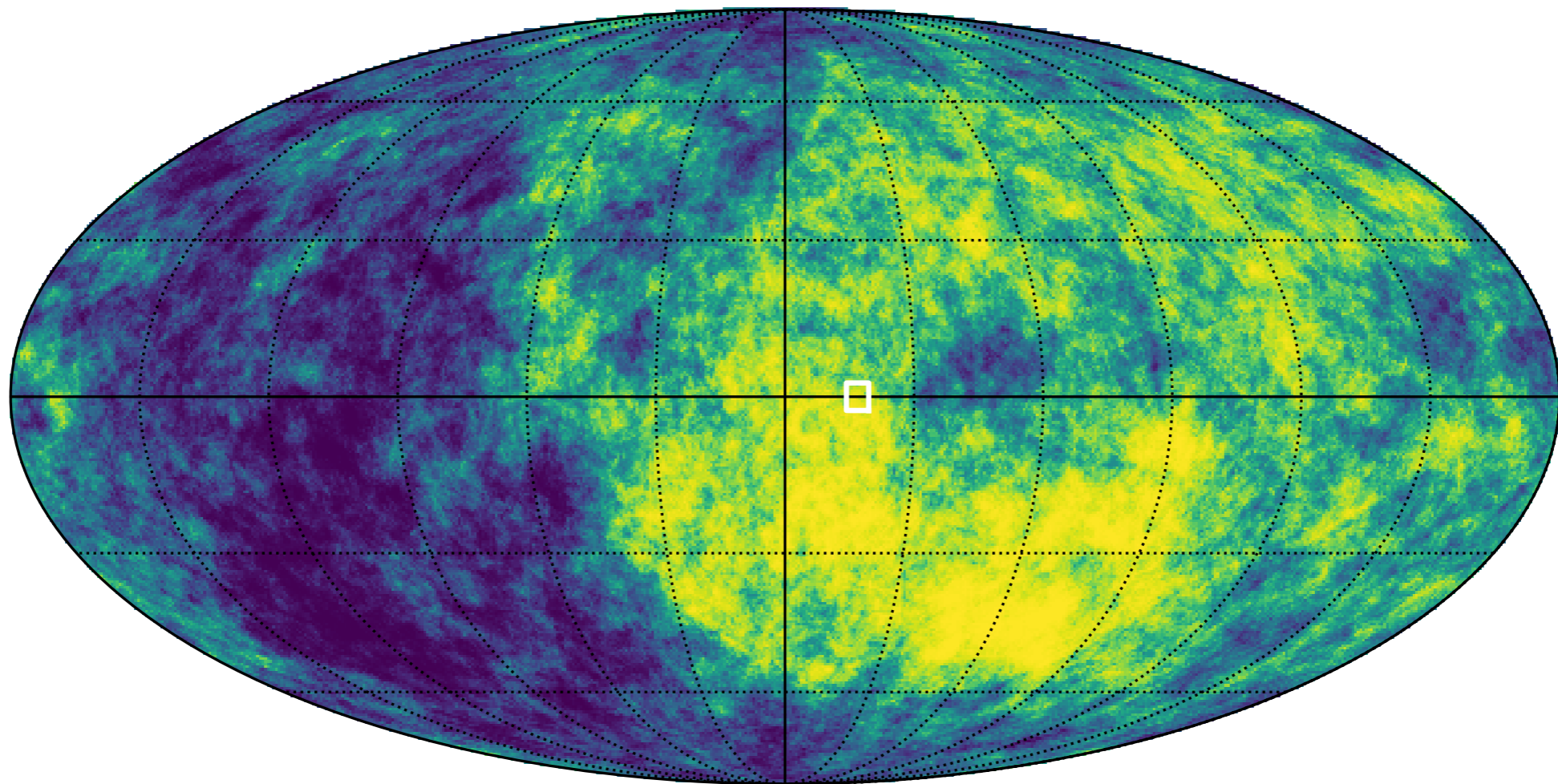
Sky maps



$$\frac{\Delta\Phi}{\sqrt{\xi}\mathcal{A}\alpha_{\text{em}}}$$

$$\xi = 10$$

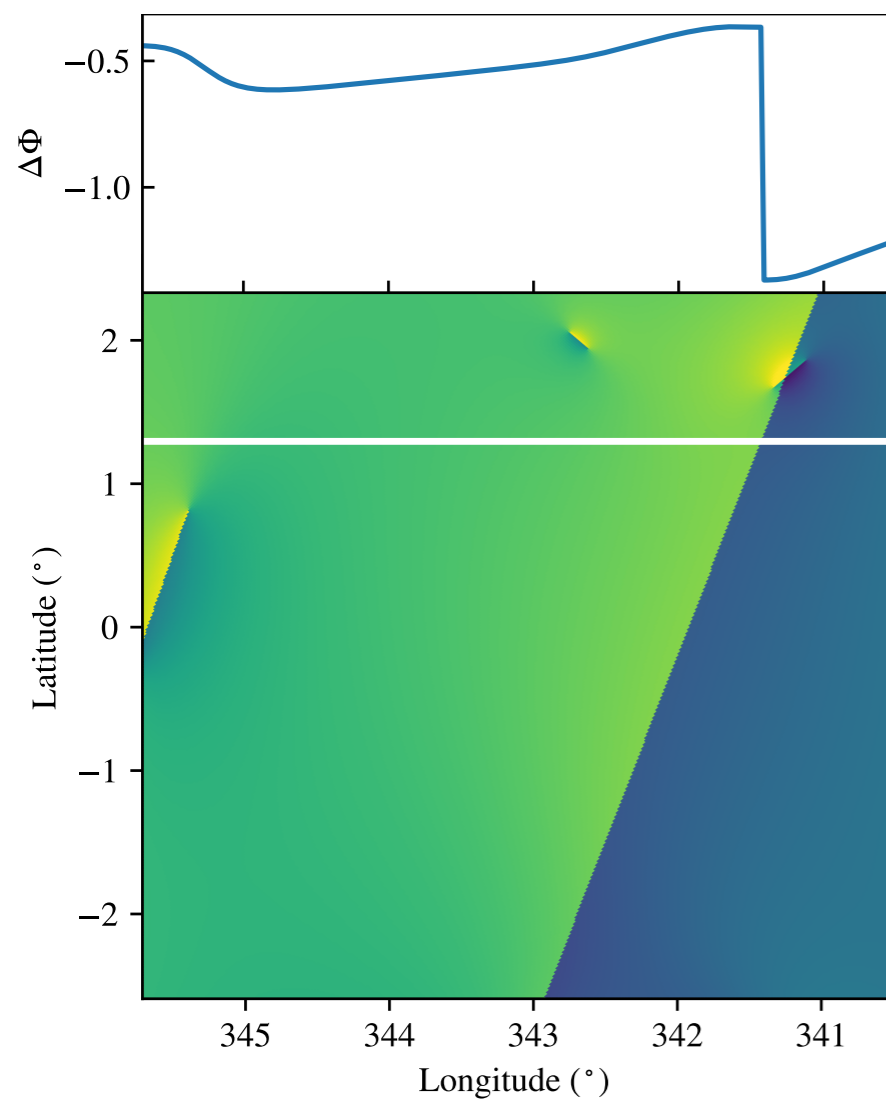
Sky maps



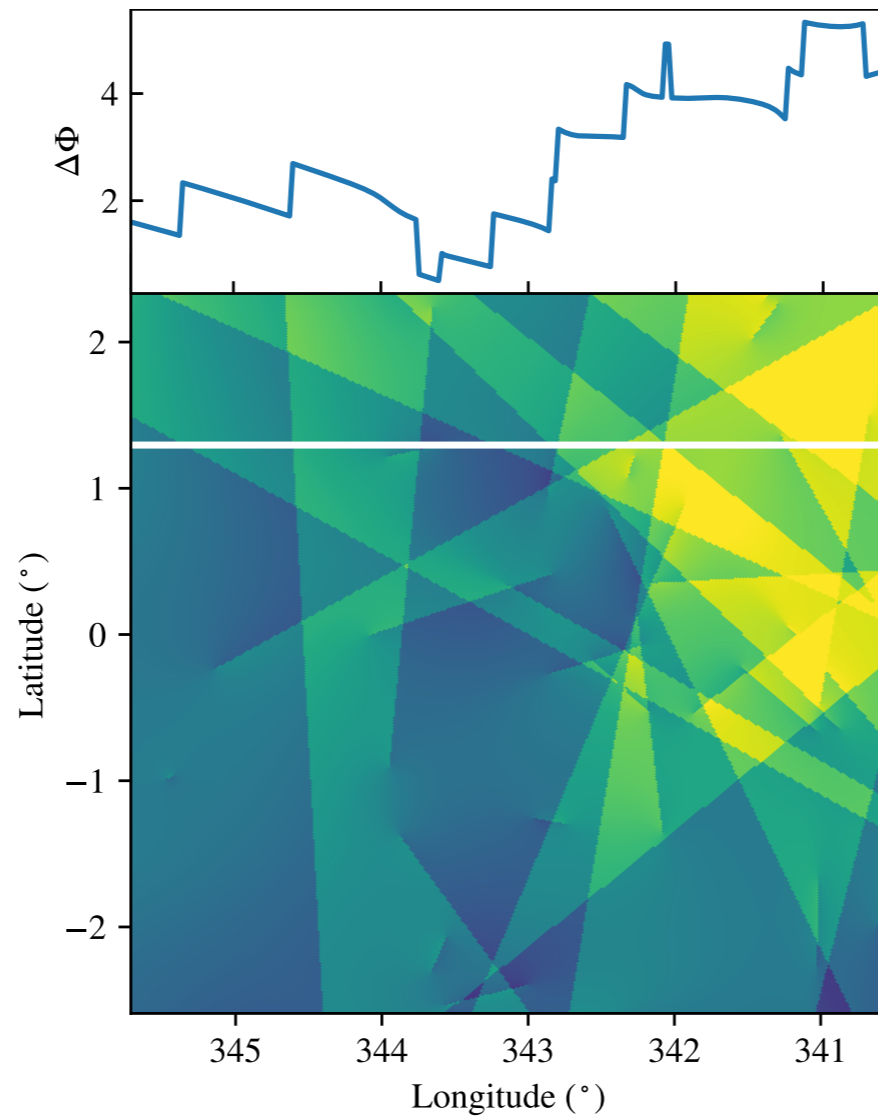
$$\frac{\Delta\Phi}{\sqrt{\xi} \mathcal{A} \alpha_{\text{em}}}$$

$$\xi = 100$$

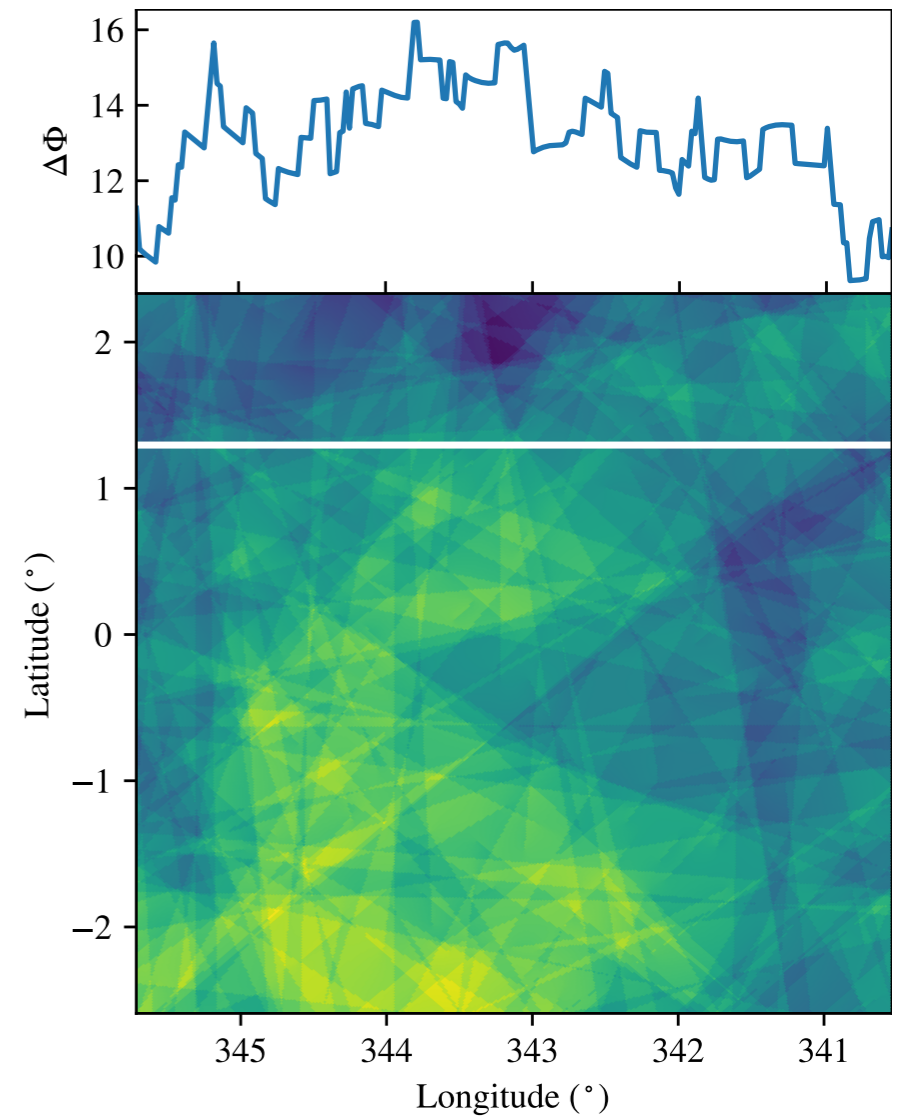
Edge Detection



$$\xi = 1$$

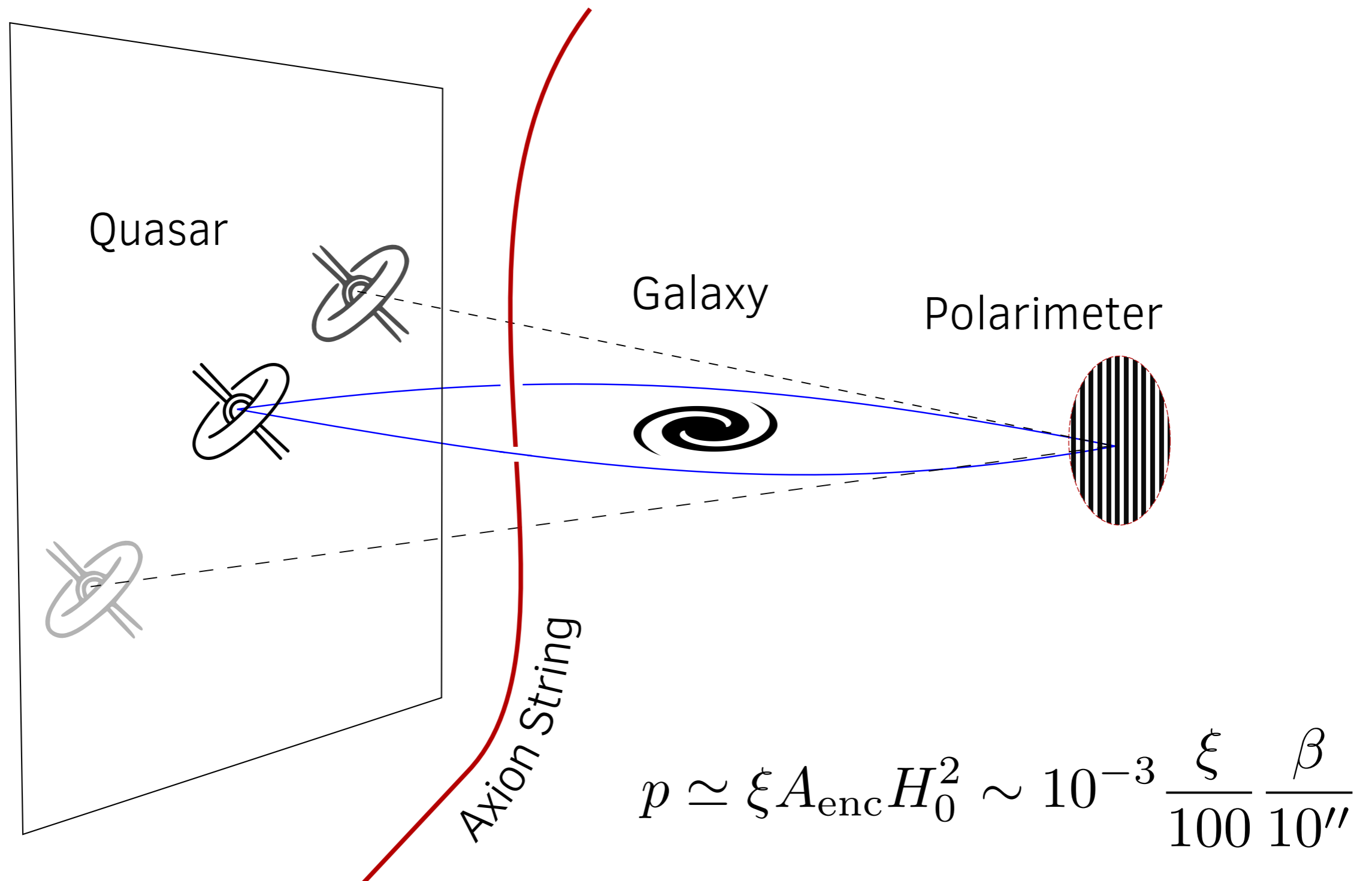


$$\xi = 10$$

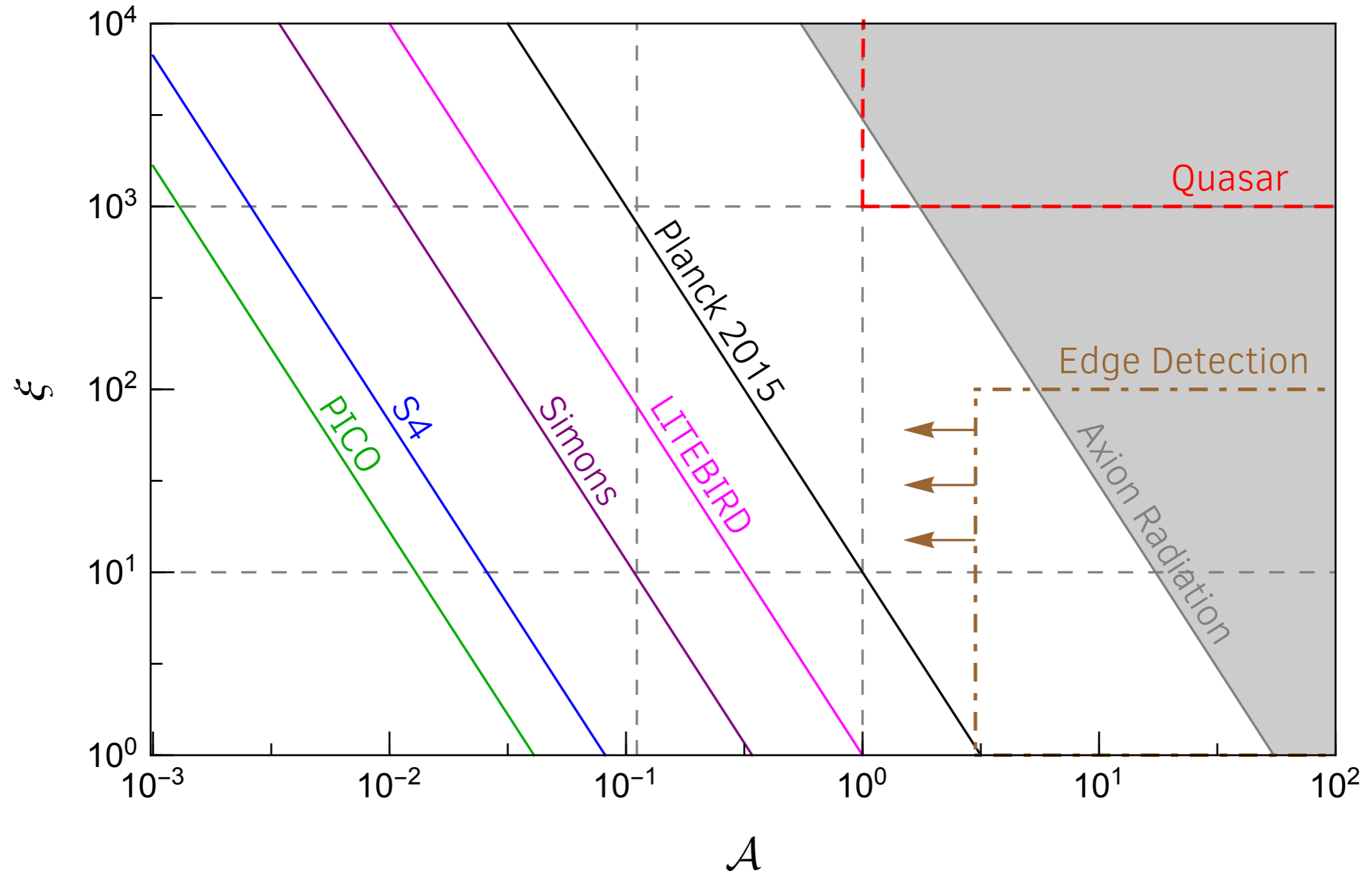


$$\xi = 100$$

Lensed Quasar systems



Reach Estimates



Thank You!