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Nature and role of microtearing turbulence in conventional tokamaks

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N<sup>th</sup> Vienna meeting (N»1, thanks to Alex) March 19<sup>th</sup>-30<sup>th</sup>, 2012



# Outline of the talk

- Historical context
- Studies of the linear microtearing **instability** 
  - expect existence in conventional tokamaks?
  - what are the critical plasma parameters?
- Nonlinear dynamics: microtearing **turbulence** 
  - magnetic fluctuation amplitudes?
    - role of magnetic stochasticity?
- Microtearing modes in **finite**  $\beta$  **ITG/TEM** turbulence

# IPP

# A brief history of microtearing research

- 1973: Small magnetic perturbations can lead to stochastic fields and enhanced electron heat flux (  $\tilde{B}/B_0\sim 10^{-4}$  ) [Stix]
- 1975-77: Microtearing modes driven by  $\nabla T_e$  are a possible source [Hazeltine et al., Drake et al., Chen et al.]
- **1980s and 1990s:** Linear theory of microtearing modes [Catto, Connor, Cowley, Drake, Hassam, Hastie, ...]
- **Since 2000:** Linear gyrokinetic simulations; microtearing modes responsible for electron heat flux in **spherical tokamaks**? [Kotschenreuther et al., Redi et al., Applegate et al.]
- 2007/08: Microtearing modes also in **medium aspect ratio tokamaks**? [Applegate et al., Vermare et al., Told et al.]
- **2010/11: Nonlinear gyrokinetic microtearing** simulations point to relevance for NSTX [Guttenfelder] and AUG (as well as ITER) [this work]



# Scope of this work

Key questions:

- Properties of linear microtearing instabilities
- Electromagnetic **heat transport** due to microtearing turbulence
- Nonlinear saturation
   mechanism



### Strategy:

- Gyrokinetic simulations
- Influence of **plasma parameters** (temperature gradient, beta, collisions...)
- Comparison to **analytical models**

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# Studies of the linear microtearing instability

# Comparison of microinstabilities in ballooning space



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ΠD

# Global linear simulations of ASDEX Upgrade shot 26459



Microtearing instabilities expected to exist in AUG



# System size effects

 $\rho^*$  scan ( $\rho^*=\rho_i/a$ ):

- Even for  $\rho^*{\sim}1/100,$  the local results are fairly accurate
- $k_v \rho_i$ =0.12 is n=1 for  $\rho^*$ =1/50
- More toroidal mode numbers are microtearing unstable in larger tokamaks



Linear modes are well described in the local approximation

### IPP

# Growth rate spectra involving microtearing modes





# Influence of electron temperature gradient and $\beta_e$



a/L<sub>Te</sub> and  $\beta_{e}$  are critical plasma parameters

# **Collisional effects** ( $a/L_{Te}$ and $\beta$ kept constant)

- Including collisions is important
- Growth rate depends on collision frequency only **moderately** (in agreement, e.g., with Applegate `07)
- Experiments are in the semicollisional to collisionless regime

$$u_{
m c} \sim {
m n} \; {
m T}^{-3/2}$$
 but:



Microtearing modes can also be present in hot (core) plasmas



# Thin current layers at resonant flux surfaces

#### Model of Drake (1977):

 ${\bf k}_{\parallel}=0$  at rational flux surface  ${\bf x}={\bf 0}$ : Nonadiabatic electron response until  ${\bf k}_{\parallel}{\bf v}_{te}\approx\omega$  with  ${\bf k}_{\parallel}\approx {\bf k}_{{\bf y}}{\bf x}/{\bf L}_{{\bf s}}$  and  ${\bf L}_{{\bf s}}={\bf q}{\bf R}/\hat{{\bf s}}$ 



# Nonlinear dynamics: Microtearing turbulence



# Turbulent electron heat flux spectrum

#### Simulation setup:

- sufficiently large radial box
- high radial resolution

384x64x24x32x16 grid points in 5D (x,y,z,v<sub>||</sub>,  $\mu$ ) phase space

Very challenging simulations!

#### Simplifications:

- **ITG** drive switched **off** (avoid multimode drive)
- **Circular** magnetic geometry [Lapillonne et al. 2009] (allow for maximum flexibility)



Heat flux is dominated by magnetic component at low k<sub>v</sub>

# Magnetic fluctuation level in microtearing turbulence



#### $R/L_{Te}$ scaling

- "scatter": reduction of the nominal **resolution** by a factor of 2-3 in various phase-space dimensions
- $\bullet$  Other parameters like  $\beta$  are important

Drake's formula yields good estimate, but neglects  $\boldsymbol{\beta}$  effects etc.

Model by Drake (1980):  $ilde{\mathbf{B}}/\mathbf{B}\sim oldsymbol{
ho}_e/\mathbf{L}_{ ext{Te}}$ 





# The role of magnetic stochasticity



#### Diffusivity model:

 $\chi_e^{em} = v_{te} D_M$ 

 $D_M = L_C (\delta B/B_0)^2$  (quasilinear result) use  $L_C = qR$  as correlation length

 $\chi^{\mathbf{e}}_{\mathbf{em}} = 1.37 \mathbf{q} \mathbf{R} \mathbf{v}_{\mathbf{te}} (\mathbf{\tilde{B}_x} / \mathbf{B_0})^2$ 

- Effective threshold for onset of strong transport (A to B)
- In stochastic cases **(B)**, GENE results confirm test-particle models (e.g. Liewer 1985, collisionless case)

# Nonlinear saturation of microtearing turbulence



Free energy transfer to medium  $k_v$ , where it is dissipated

# Microtearing modes in finite $\beta$ ITG/TEM turbulence

Detailed account: D. Hatch (next week)

## IPP

# Motivation: Recent results on finite $\beta$ turbulence

- Electron magnetic heat transport can approach (or even surpass) electrostatic transport as β increases [Candy PoP `05, Pueschel PoP `08]
- Magnetic transport violates quasilinear theory  $\beta^2$ -scaling



Observation of near-ubiquitous **magnetic** 

**stochasticity** – even at low values of  $\beta$ 

(Nevins PRL '11, Wang PoP '11)

2/(v.o.<sup>2</sup>p<sub>0</sub>/R<sup>2</sup>) , Γ/(v.o.<sup>2</sup>n<sub>0</sub>/R<sup>2</sup>) 150 ∆Q<sup>ee</sup> ITG-TEM KBM × 0." 100 0 Qem 50 0.2 0.41.0 1.2 0.0 0.60.8 1.4% β Pueschel PoP '08 β=0.1% Nevins PRL`11 ...needs an explanation! 19

## IPP

# What is the reason for magnetic stochasticity?

- **Ballooning parity** modes: no reconnection/stochastic fields
- Tearing parity modes allow reconnection
- Analysis: **POD** of vector potential (optimal basis) :  $A_{\parallel k}(z,t) = \sum A_{\parallel k}^{(n)}(z)h_k^{(n)}(t)$
- First two modes (ballooning and tearing) plus residual modes





# Magnetic transport – superposition of ITG and tearing



- **`Dip'** in the magnetic **electron heat transport spectrum** (at k<sub>y</sub> of electrostatic transport peak) [Candy, PoP`05]
- Understood by attributing transport to ballooning and tearing part of B<sub>x</sub>
- Stochastic transport **explains**  $Q_e^{EM} \sim \beta^2$

# Origin of the magnetic electron heat transport



- GENE eigenmode solver finds all gyrokinetic eigenmodes
- Nonlinearly evolved distribution function projected onto 1000 orthogonalized linear eigenmodes (k<sub>v</sub>ρ<sub>i</sub>=0.2,k<sub>x</sub>ρ<sub>i</sub>=0)
- One damped eigenmode dominates, has properties of microtearing mode
- Analysis of nonlinear transfer: Excitation via coupling to zonal modes

Nonlinearly excited microtearing mode causes electromagnetic heat flux

# Summary & Outlook



# Summary & Outlook

- **Microtearing modes** are expected to exist in **ASDEX Upgrade** and presumably also in other medium-aspect-ratio tokamaks like **ITER**
- Nonlinear gyrokinetic simulations establish microtearing as a candidate to explain electron heat flux in outer core of tokamaks: χ<sup>e</sup><sub>em</sub>~1m<sup>2</sup>/s
- Magnetic stochasticity occurs when magnetic fluctuation amplitude exceeds an effective threshold (in analogy to Dimits shift)
- Nonlinearly excited (linearly stable) microtearing modes cause stochasticity and magnetic electron heat flux in ITG turbulence
- Underway: Microtearing + ITG turbulence (preliminary result: coexistence!)

**References:** 

- Doerk et al., PRL 2011
- Doerk et al., APS / PoP
- Hatch et al., submitted