Sawtooth Period Prediction

Influence of neo-classical resistivity on scaling from JET to ITER.

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Content.

- Sawteeth: some general features.
- Modelling (a) Direct non-linear 3D
- (b) $1^{1}/_{2}$ D transport code
- Role of η ; neo-classical versus Spitzer
- Toy model numerical solutions for the Ramp.
- Some questions & conclusions.

Sawteeth in Ohmic Tokamaks were like this: 1

ViennaFigs.nb



Sawteeth in Large Tokamaks with auxiliary heating:



During flat top, only q evolves. What duration in ITER?

Sawtooth Period Prediction

- $\tau_{\rm s}$ ~ 100sec. Is too demanding for global kinetic simulation.
- Must resort to modelling by
- (a) employing dissipative fluid eqs. (as in Park-Monticello, Nucl. Fus. (1990) and Halpern et al. PPCF (2011))
 - OR
- (b) breaking cycle into 3 separate phases: Trigger → Collapse → Ramp, (as in Porcelli-Boucher-Rosenbluth, 1996)

Role for gyro-kinetics in Trigger & Collapse.

(a) Global, long duration, fluid simulations

- For recent toroidal simulations with 2-fluid effects see F.D.Halpern et al. PPCF 53(1) 015011 (2011). XTOR-2F code.
- Realistic simulation: with diamagnetic stabilisation effects giving long quiescent ramp and fast crash.
 But Spitzer η.
- W.Park & D.Monticello (1990). Single fluid, but investigated effect of Neo-classical η in such simulations.

(b) $1^{1/2}D$ Neoclassical η drives discharges towards 1/1 Instability.

- $\eta(\text{Neo}) \ge 4 \eta(\text{Spitzer})$ for r ~ a
- $\eta(\text{Neo}) \sim \eta(\text{Spitzer}) \text{ for } r \rightarrow 0$
- η (Neo) causes stronger peaking of J(r)
- And therefore lower q(0).

Steady state q(r)

1

Neo-QDiffusion.nb



3D Sawtooth simulation of Park-Monticello

- Following a Sawtooth collapse, neoclassical resistivity controls the subsequent q(r,t) evolution during quiescent ramp.
- W. Park & D.A. Monticello Nucl.Fus. (1990), 2413 --- MH3D simulation. Concluded:-
- Collapse like Kadomtsev reconnection;
- And large drop in q₀ during subsequent ramp, δq ~ 0.2, caused by neoclassical correction to η near r~0.

Ramp evolution of q(r,t) with $\eta \approx \eta_{Sp} (1 - \sqrt{(r/R)})^{-2}$

 Neoclassical correction is weak at r~0, but generates singularity in the diffusion equation as r→0.

$$\frac{\partial B_{\theta}}{\partial t} = -(\nabla \times \mathbf{E})_{\theta},$$

$$= \frac{\partial}{\partial r} \left[\frac{\eta}{\mu_0 r} \frac{\partial}{\partial r} (r B_{\theta}) \right]$$
(1)

or equivalently:

$$\frac{\partial}{\partial t} \left(\frac{1}{q}\right) = \frac{1}{r} \frac{\partial}{\partial r} \left[\frac{\eta}{\mu_0 r} \frac{\partial}{\partial r} \left(\frac{r^2}{q}\right)\right],
\frac{\partial q}{\partial \tau} = -4q^2 \frac{\partial}{\partial x} \hat{\eta}(x) \frac{\partial}{\partial x} \left(\frac{x}{q}\right),
\tau = \frac{t}{\tau_{\eta}}, \quad \tau_{\eta} = \frac{\mu_0 a^2}{\eta(0)}, \quad x = \frac{r^2}{a^2}$$
(2)

Regularity is restored by axial "collisional boundary layer"

 Transition out of Banana regime. Hirshman, Hawryluk & Birge (1977).

$$1 - \left(\frac{r}{R}\right)^{1/2}] \rightarrow \left[1 - \left(\frac{a}{R}\right)^{1/2} \frac{r^2}{r^{3/2} + \nu_*}\right], \qquad (1)$$

$$\nu_* = \frac{\nu_e Rq}{\left(\frac{a}{R}\right)^{3/2} V_{Te}} \qquad (2)$$

Applied on axis, this predicts

$$\frac{\partial q_0}{\partial t} \sim -\frac{8(\frac{a}{R})^{1/2}}{\nu_* \tau_\eta},\tag{1}$$

Implies "fast" linear decrease of $q_0=1-t/\tau_0$

Neoclassical Evolution of q(r)

- If the local "cuspy" structure of J(r) around magnetic axis is destroyed in the collapse (as in the Kadomtsev model), then fast downward evolution of q₀ ensues.
- Timescale is

•
$$\tau_0 \sim (\tau_\eta v_* / 8\sqrt{\epsilon}) \propto (a^{3/2} R^{3/2} N / T^{1/2})$$

<< $\tau_\eta \sim a^2 T^{3/2}$

Toy cylinder model: Neoclassical evolution of $q_0(t)$



Large drop after 0.1% of resistive diffusion time. - - - Initial analytic: ——— Numerical evolution.

Porcelli-Boucher-Rosenbluth (1996)

- 1 ¹/² D modelling, predicted Sawtooth period in ITER of ~ 150secs. for Kadomtsev model;
- Consistent with $\tau_s \propto T^{1.5}~a^2$ scaling from JET.
- Park-Monticello (after comparison of many devices) suggest:

 $\tau_{\rm s}$ = 9 T^{1.5} R² ms. = 0.0016 τ_{η}

- \Rightarrow Same scaling.
- But is this scaling valid at very high T_e ?

Example: equilibrium q evolution after a Kadomtsev reconnection. 40ms. to 4sec.



ITER q from 2.5s. to 250s.



Does v_* matter?

- Key trigger parameter may be the shear at q=1, s₁. (P-B-R, 1996, C-H-Z, 2012)
- $s_1(t)$ is influenced by 2 conflicting scalings.
 - $\tau_n \propto T^{1.5} a^2$ global resistive diffusion of q.

 $\tau_0 \propto a^3$ N / T^{1/2} from axial influence of v_*

Conclusions: (1) Trigger

- Trigger Collapse Ramp Modelling.
- Progress on analytic trigger criteria (Connor, Hastie & Zocco, PPCF, 54, p035003, 2012)
- Shows importance of $\beta/s^2,~\eta_e~\text{and}~\eta_i~$ as well as $\Delta'_{1/1}$
- Provides framework to compare with linear gyro-kinetic codes.
- Toroidal Δ' code for 1/1 mode, with α particles, is required. Work in progress.

Conclusions: (2) Collapse

- Can Gyro-kinetics simulate collapse for ~100 μs. or more?
- Is local axial structure of J(r) destroyed during turbulent collapse?
- Which collapse model is most realistic?
- Can fluid codes like XTOR-2F generate a new collapse model for use in 1¹/₂ D codes?
- ECE imaging is also providing invaluable insights.

Recent ECE data: Hyeon Park

Compares with Sawtooth Collapse models:

Kadomtsev reconnection; Kadomtsev Quasi-Interchange; Wesson Ideal Ballooning; Bussac&Pellat, Cowley&Wilson " No model is completely correct." "No model is totally wrong." H.Park(2012)

Conclusions: (3) Ramp simulation

- If axial structure of J(r) is destroyed during collapse, is q(r,t) evolution influenced by v_{*} as predicted in the Toy model?
- What timescale for, $s_1(t)$, shear at r_1 ?
- Do other pitch-angle scattering mechanisms broaden the axial "layer" in high T_e tokamaks?

Stochasticity during Collapse

- From Halpern, Leblond, Lutjens & Luciani. PPCF, 53(1), 015011, (2011)
- The following Poincare plots show initial growth of a 1/1 Kadomtsev island, followed by secondary instability and spread of stochasticity. Could explain 1980 observations of precursor oscillation and crash on TFR.
- Foundation for a new collapse model?

Cyclic regimes in $S - \omega_*$ phase space

Internal kink mode cyclic regimes Diamagnetic thresholds Discussion

Simulation model and setup

Magnetic field cross sections



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