

# Turbulence at High Plasma $\beta$

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Thanks to

F. Jenko, W.M. Nevins,  
T. Hauff, H. Doerk



Max-Planck-Institut  
für Plasmaphysik

*Gyrokinetics for ITER*

Vienna, Mar. 16, 2010

# Motivation and Experimental Results

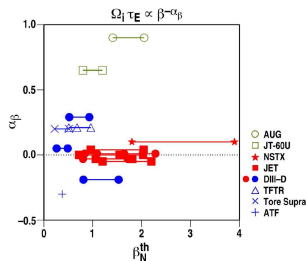
## Definition

$$\beta \equiv \beta_e = \frac{8\pi n_{e0} T_{\text{ref}}}{B_{\text{ref}}^2}$$

## Features in

- bootstrap fraction  $\propto \beta$
- fusion reaction rate  $\propto \beta^2$
- (kinetic) ballooning threshold
- magnetic fluctuations and transport

$\beta$  scalings of confinement time differ:



- JET '04: no  $\beta$  dependence
- ASDEX '07:  $\propto \beta^{-1}$
- JET '08:  $\propto \beta^{-1.4}$

**But:** fixing parameters difficult

# GENE and Previous Results

## The GENE code

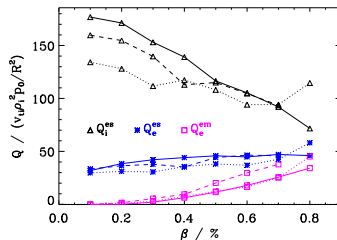
[www.ipp.mpg.de/~fsj/gene](http://www.ipp.mpg.de/~fsj/gene)

- nonlinear gyrokinetic equations
- multiple fully kinetic particle species
- collisions and electromagnetic effects
- linear Eigenvalue solver
- radially local and nonlocal modes
- open source

## Previous gyrokinetic simulations:

Jenko 2001, Parker 2004, Dannert 2004, Candy 2005, Pueschel 2008

## Good agreement:



## 1 $\beta$ Scans

- Cyclone Base Case
- Density Gradient Driven TEM Case
- Pure ITG Case

## 2 Magnetic Fluctuation Strength

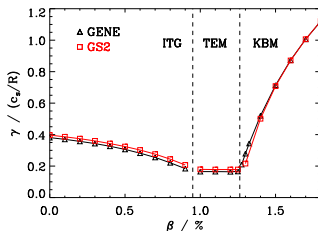
## 3 Magnetic Stochastization

## 4 Parallel Magnetic Fluctuations

# Growth Rates and KBM Onset

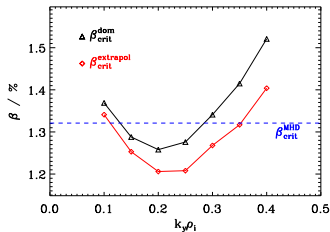
## Linear analysis

(here:  $k_y = 0.2$ ):  
depending on  $\beta$ , one gets  
dominant **ITG**, **TEM**, or **KBM**



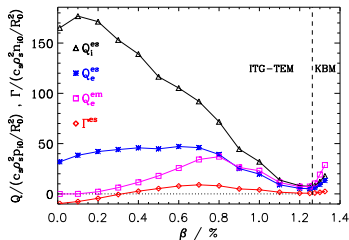
## KBM threshold:

destabilization at  $\beta_{crit}$

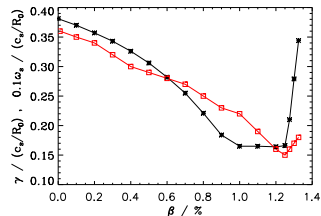
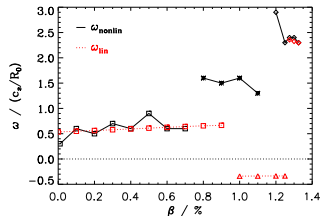


- minimal  $\beta_{crit}$  at nonlinear transport peak
- gradient dependence: only  $\omega_{Ti}$  significant,  
 $\beta_{crit}$  increases strongly for low values, exceeds  $\beta_{crit}^{MHD}$

# Nonlinear Transport Levels

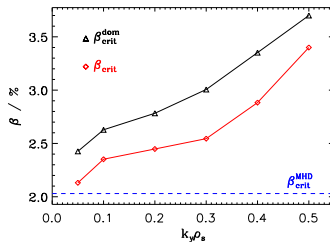
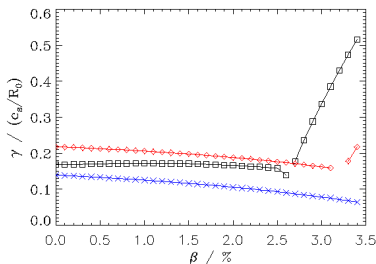


$\Rightarrow$  regimes not same as linear; high- $\beta$  **transport drop**:  
mode interaction (Merz 2008)  
and zonal flows



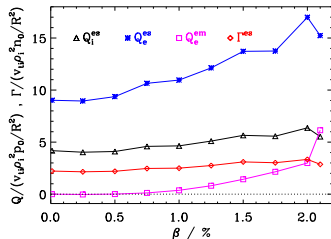
# Growth Rate and KBM Onset

Density gradient driven trapped electron mode:



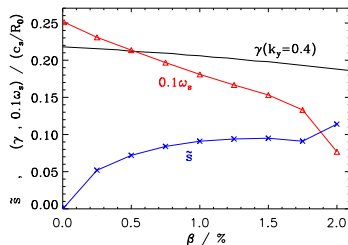
- KBMs appear first at lowest  $k_y$
- no KBMs appear in the MHD stability regime

# Transport and Zonal Flows



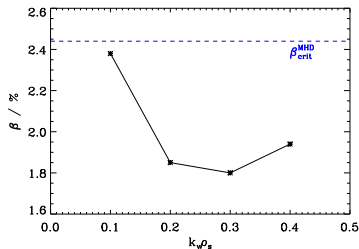
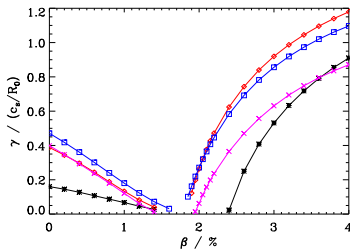
$\Rightarrow$  **different behavior:**  
 slight decrease linearly,  
 increase nonlinearly  
 KBM threshold: same as linear threshold

**Zonal flows** are  
 able to bridge the  
 linear-nonlinear gap



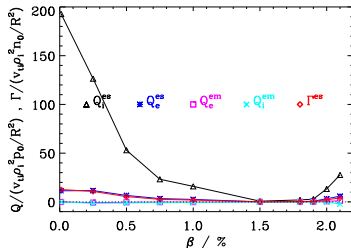
# Growth Rate and KBM Onset

Pure ITG scenario ( $\omega_{Te} = 0$ ),  
designed to have gap at intermediate  $\beta$   
between ITG and KBM



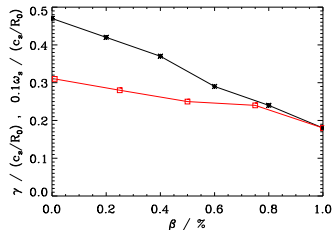
$\Rightarrow$  same as for Cyclone, except for missing TEM regime

# Transport and Zonal Flows



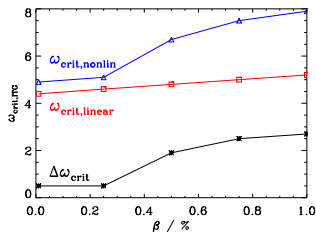
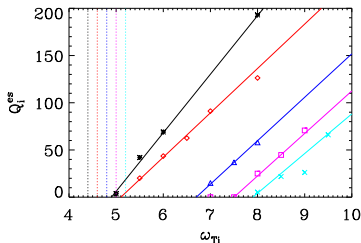
$\Rightarrow$  much **stronger decline**  
**nonlinearly** with  $\beta$ ,  
**KBM onset** at the same  $\beta_{crit}$

**Zonal flow impact on the transport curve:**



Sufficient to explain drop?

# Dimits Shift Study



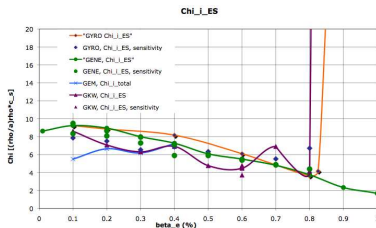
⇒ The **nonlinear upshift** of the critical gradient is **increased** for rising  $\beta$  values

Are zonal flows sufficient to explain this effect?  
(work in progress)

Note: one single runoff case was found

# Transport Runoff

CBC runoff ( $\beta \approx 0.85\%$ ): no saturation, transport at smallest  $k_y$   
Benchmark effort of gyrokinetic codes underway (Bill Nevins):



⇒ **good agreement** on runoff threshold

GENE findings: “stable” region around  $\beta \sim 1.2\%$

How to avoid: initial condition;  $\beta$  ramp-up

How not to avoid: larger box, higher resolution

## 1 $\beta$ Scans

- Cyclone Base Case
- Density Gradient Driven TEM Case
- Pure ITG Case

## 2 Magnetic Fluctuation Strength

## 3 Magnetic Stochastization

## 4 Parallel Magnetic Fluctuations

# Applications

The radial magnetic field fluctuation  $B_x$  influences, e.g.:

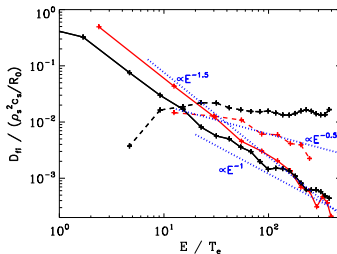
Model for electron heat transport along  
**perturbed field lines:**

Test particle transport model

$$\chi_e^{\text{em}} = q_0 R \left( \frac{T_e}{m_e} \right)^{1/2} \langle B_x^2 / B_{\text{ref}}^2 \rangle$$

Redistribution of NBI ions in turbulence (Hauff 2009):

$$D_{\text{em}} \propto B_x^2 E_{\text{beam}}^0$$



## Fluctuation Levels

Using  $\beta$  scans to fit  
straight lines to data:

$$\frac{B_x}{B_0} = C_x \frac{\beta}{\beta_{\text{crit}}^{\text{KBM}}} \frac{\rho_i}{R_0}$$

$\Rightarrow$  fluctuations strongest for  
**ITG** turbulence

### Radial Fluctuations

$C_{x,\text{CBC,ITG}}$	$\sim$	0.8
$C_{x,\text{ITG}}$	$\sim$	0.8
$C_{x,\text{CBC,TEM}}$	$\sim$	0.4
$C_{x,\text{TEM}}$	$\sim$	0.2
$C_{x,\text{MT-Doerk}}$	$\sim$	0.2

(**caveat:** CBC has  
no pure nonlinear TEM)

For the corresponding,  $C_y$ , slightly higher results are obtained

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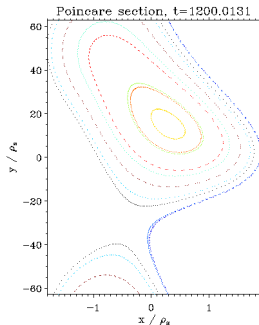
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# Field Line Integration

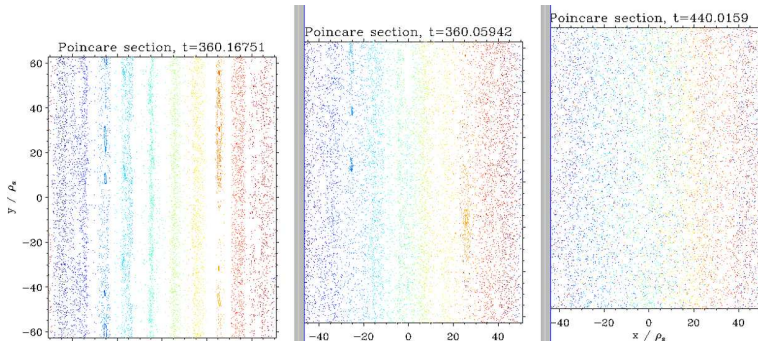
## Integration Scheme

- interpolate  $B_{x,y}(x, y, z)$
- seed lines at  $z = 0$
- obtain  $B_{x,y}(\text{lines})$
- calculate new position at  $\Delta z/2$ , obtain  $B_{x,y}$
- use new field at old position



- integrated with the GENE Diagnostics Tool
- optimized
- benchmarked (Bill Nevins)

## Flux Surface Destruction



$\Rightarrow$  flux surfaces **disintegrate** at moderate  $\beta$   
nonlinearly, islands are destroyed even at low  $\beta$   
stochastization **independent** of runoff, transport  
**Application:** (micro-)tearing, reconnection

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## $B_{\parallel}$ : Equations

If  $B_{\parallel}$  is not neglected, have coupled  $\Phi$ - $B_{\parallel}$  system  
( $A_{\parallel}$  remains decoupled):

$$\begin{aligned}\Phi &= \bar{C}_3 M_{00} - \bar{C}_2 M_{01} \\ B_{\parallel} &= \bar{C}_1 M_{01} - \bar{C}_2 M_{00} \\ M_{00} &\sim \sum_j \int J_0 g_j d^3 v \\ M_{01} &\sim \sum_j \int \mathcal{I}_1 g_j d^3 v\end{aligned}$$

Coupling in Vlasov equation via:  $\chi = \Phi - v_{Tj} v_{\parallel} A_{\parallel} + \mu T_j^{-1} q_j^{-1} B_{\parallel}$   
and via modified FLR corrections to the moments

**Independently** at high  $\beta$ : equilibrium  $j_{0\parallel}$  effects

## $B_{\parallel}$ : Transport and Impact

New  $B_{\parallel}$  component of the (electromagnetic) particle flux:

$$\Gamma_j = -n_{j0} \left( (\partial_y \Phi) \mathcal{M}_{00} - v_{Tj} (\partial_y A_{\parallel}) \mathcal{M}_{10} + T_{j0} q_j^{-1} B_0^{-1} (\partial_y B_{\parallel}) \mathcal{M}_{02} \right)$$

**Impact:** typically,  $\Delta\gamma \sim 10 - 20\%$  is observed  
near the KBM limit

Stellarators ( $\beta \sim 5\%$ ) and astrophysical applications  
(e.g.,  $\beta \gtrsim 1$ ) can have strong  $B_{\parallel}$  contributions

**Computational effort:**  $\sim 2\%$  more

**What is done in GS2?**

## Points for Discussion

- role of **zonal flows** at large  $\beta$  (Maxwell stress ...)
- **Dimits shift**: what causes the sudden change?
- magnetic **fluctuations**: universality, dependence on parameters
- magnetic **surfaces**: quantitative approaches, effect on turbulence
- **runoff** causes, prevention
- $B_{||}$ : impact, magnetic transport definitions