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## Introduction

- Experiments indicate tokamak turbulence is strongly affected by sheared flows Extensive theoretical work has been undertaken
  - flow shear can tear apart turbulent eddies, suppressing turbulence
  - not all shear flows are stabilising eg Parallel Velocity Gradient (PVG) instability
  - beyond critical perpendicular shear PVG linearly stabilised: transient perturbations remain [1] Strong toroidal flows of tokamaks have both perpendicular and parallel shear
  - What is the optimum shear flow for confinement do we want ever more toroidal flow shear?
- Investigate effect of flow shear on Ion Temperature Gradient (ITG) and PVG instabilities in sheared slab via dissipative fluid model: simple system allows clear interpretation

#### System Equations

- Analyse sheared slab, magnetic and flow fields:  $\boldsymbol{B} = B_0 \left( \hat{\mathbf{z}} + \frac{x}{t} \hat{\mathbf{y}} \right) = V_0 = V_0 \frac{x}{t} \hat{\mathbf{e}}_v$ ,  $V_0 \cdot \hat{\mathbf{x}} = 0$ Work in "twisting-shearing" representation <sup>[2]</sup>
- removes problems of shear and time dependence from background flow aligns coordinate lines with characteristics of plasma response: sound waves



- Take Boltzmann electron response; ions described by collisional fluid equations derived from avro-kinetic equation in doubly sheared coordinates with flow
  - normalise:  $(x', y') = \rho_s(x, y)$ ,  $z' = l_s z$ ,  $t' = (l_s / c_s) t$ ;  $c_s = \sqrt{(\gamma_e + \gamma_i)T/m}$ ,  $T_i = T_e$ ,  $\gamma_e = 1$ ,  $\gamma_i = 5/3$
  - usual orderings:  $\omega / \Omega \sim O(k_{\parallel}/k_{\perp}) \sim O(\delta f / F_0) \ll 1$
  - additional collisional orderings:  $\nu \gg \omega$ ,  $\omega^*$ ,  $v_{th}k_{\parallel}$ ,  $u_f k_{\parallel}$ ,  $\nu k^2 \rho^2$ ;  $k \rho \sim O(\sqrt{\omega/\nu}) << 1$
  - expand in \u03c8 / v: lowest order is perturbed Maxwellian closed equations for evolution of perturbed density  $\delta n$ , parallel velocity  $\delta V_{\parallel}$ , temperature  $\delta T$  from moment equations for particle,  $(n, V, T) = \frac{l_s}{\rho_s} \left( \frac{\delta n}{n_0}, \frac{\delta V_{\parallel}}{c_s}, \frac{\delta T}{T} \right)$ parallel momentum and energy conservation
- Restrict to linear case, take fields  $\propto \exp(iky)$ ,  $(v_k, \chi_k) = k^2(v_\perp, \chi_\perp)$ ,  $\nabla_\perp^2 = -k^2(1+z^2)$
- Dissipation: (ν<sub>⊥</sub>, χ<sub>⊥</sub>) ∝ (0.9, 1)ν<sub>ii</sub>
- Drift frequencies characterise driving x-gradients:  $(\omega_s^*, \omega_v^*, \omega_\tau^*) = \frac{3k}{8} \left( \frac{l_s}{l}, \frac{l_s}{l}, \frac{l_s}{l} \right); \quad \frac{1}{l} = \frac{1}{L} \frac{V_0}{c} \hat{\mathbf{e}}_v \cdot \hat{\mathbf{z}}$

$\left(\frac{\partial}{\partial t} + M \frac{\partial}{\partial z}\right) n$	+	$\frac{\partial V}{\partial z}$	=	i $\omega_n^*n$			$M = u_f / c_s$
$\left(\frac{\partial}{\partial t} + M \frac{\partial}{\partial z}\right) V$	+	$\frac{3}{8}\frac{\partial}{\partial z}(2n+T)$	=	i $\omega_v^* n$	-	$\nu_k\left(1+z^2\right)V$	$\omega_{_{\!\!\!\!\!\!\!\!\!\!\!\!\!\!}}^* \varpropto k \ M \ \cot \theta_{_{\!$
$\left(\frac{\partial}{\partial t} + M \frac{\partial}{\partial z}\right) T$	+	$\frac{2}{3}\frac{\partial V}{\partial z}$	=	i $\omega_T^* n$	-	$\chi_k\left(1+z^2\right)T$	$\theta_{i} \xrightarrow{\hat{\mathbf{e}}_{i}} \hat{\mathbf{z}}$

# Parallel Flow Instability

- · Formulate local dispersion relation neglecting dissipation
- look for solution  $\propto \exp\left(-i\omega t + ik_{\parallel}z\right)$ convective effect of perpendicular flow shear:  $\omega' = \omega - k_{\parallel}M$
- PVG mode  $\omega_n^* = \omega_T^* = 0$ :  $\omega^2 = k_{\parallel} (k_{\parallel} \omega_v^*)$
- instability for:  $\omega^* > k$
- effect of background parallel flow shear asymmetric
- Dissipative terms act to damp perturbations far along field lines: demonstrate for  $M = \chi = 0$ solution has form:  $V = s(z) \exp(i\omega_{x}^{*} z/2b) \exp(\gamma t)$

$$\frac{\partial^2 s}{\partial z^2} + \left[\frac{\omega_v^{*2}}{4b^2} - \frac{a}{b}(\gamma + v_k) - \frac{a}{b}v_k z^2\right] s = 0 \qquad \qquad a = \gamma - i\omega_n^*$$

$$b = 1 + \frac{i}{4\gamma} \left(\frac{3}{2}\omega_r^* - \omega_n^*\right)$$

$$= \text{ simple harmonic} \qquad s(z) = H \left[4\sqrt{\alpha_v v_k / b} + z\right] a^{-\sqrt{av_k / b}}(z)$$

- oscillator equation:  $s(z) = H_{n} \left[ \sqrt[4]{a v_{k} / b} z \right] e^{-\sqrt{a v_{k} / b}(z)}$ eigenvalue condition:  $\frac{a}{b}(\gamma + \nu_k) + (2n+1)\sqrt{\frac{a}{b}\nu_k} - \frac{\omega_v^{*2}}{4b^2} = 0$
- large k modes dominated by viscosity: FLR effects excluded

 Twisting mode localised where perpendicular gradients are small to minimise collisional dissipation







· Anticipate rapid convection sweeps growing instabilities to dissipative region: forces decay



entropy wave

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· Motivate by looking at characteristics of linear system

work in moving frame  $z = z_0 + M t$ , expand in large parameter M look for colution with time dependent growth rate  $o(1, c^2)$  dot

$$V(z_0,t) = s(z_0,t) \exp\left(i\omega_t^* z_0/2\right) \exp\left(\int \gamma(t')dt'\right) = \frac{\tau}{\gamma_0} = -t^2 + \sqrt{t^4 + 1} \qquad \frac{\tau}{\gamma_0} = -\omega_t^2/2$$

- instability grows exponentially for 
$$t \sim \frac{\sqrt{r_0}}{\sqrt{r_k M}} \propto \frac{1}{\sqrt{M}} \Rightarrow \text{amplification factor large} \sim \exp(c\sqrt{M})$$

- gives asymptotic form:
- Algebraic decay with time: amplified perturbation can linger

# Numerical Results

Impact of Convection

 Linear system solved numerically: 2<sup>nd</sup>-order upwind scheme, typical resolution ∆z = 0.1 growth rate maximised over k, consistency with orderings for  $\nu_{\perp} = 3.0$ ; here  $l_s / l_s = 0$ 





## Conclusions

- Investigated ITG and PVG stability in a sheared field with parallel and perpendicular flow shear Instabilities are twisting modes, convected along the field by perpendicular sheared flow
- Parallel flow shear drives instability, having complex interaction with ITG drive sensitivity to angle of flow favourable for instability suppression in spherical tokamaks
- Modes convected faster than the sound speed, M > 1, are swept downstream to be damped indicating suppression if flow shear convection faster than characteristic propagation speed of mode? Only transitory growth occurs for M > 1, but for large M growth can be substantial

mode decay slow; promotes subcritical turbulence?

This work was funded by the United Kingdom Engineering and Physical Sciences Research Council under grant EP/G003955 and the European Communities under the contract of Association between EURATOM and CCFE and EURATOM and IST. The views and opinions expressed herei do not necessarily reflect those of the European Commission.