

Lagrangian formulation of gyrokinetic theory with a single expansion parameter

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- 1 Motivation
- 2 Preliminaries: phase-space Lagrangian formalism and gyrokinetic ordering
- 3 Gyrokinetic change of coordinates
- 4 Gyrokinetic Poisson equation
- 5 Conclusions and further work

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Setting of the talk: Time independent magnetic field. We only deal with the dynamics of ions.

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Notation: $L \sim |\nabla(\ln |\mathbf{A}|)|^{-1}$, v_t the thermal speed, ρ_i a typical ion gyroradius, Ω a typical gyrofrequency, $\varphi(\mathbf{r}, t)$ the electrostatic potential.

- $\epsilon = \rho_i/L \ll 1$.
- $Ze\varphi/T_i \sim \epsilon$.
- $\mathcal{F}[\varphi(\cdot, t)](\mathbf{k})$ is localized in the region
 - $|\mathbf{k}_\perp|\rho_i \sim 1$, i.e. $|\mathbf{k}_\perp| \sim \epsilon^{-1}/L$,
 - $|\mathbf{k}_\parallel| \sim L^{-1}$.
- $\mathcal{F}[\varphi(\mathbf{r}, \cdot)](\omega)$ localized in the region $\omega/\Omega \sim \epsilon$.

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- $\mathcal{F}[\varphi(\mathbf{r}, \cdot)](\omega)$ localized in the region $\omega/\Omega \sim \epsilon$.

However, modern derivations of the gyrokinetic equations are carried out by performing two **independent** expansions:

- (I) Guiding-center expansion in $\epsilon = \rho_i/L$.
- (II) Gyrokinetic expansion in powers of a parameter, ϵ_φ , giving the size of the electrostatic fluctuations.

We will implement the gyrokinetic ordering, defined only by ϵ , in the phase-space Lagrangian formalism and give a fully explicit expression for the ϵ^2 term of the gyrokinetic Hamiltonian, $\overline{H}^{(2)}$, in general geometry.

- Calculations are much more complicated.
- Guiding-center and gyrokinetic dynamics are **inextricably tied together** and new terms arise that are absent in the customary derivations.
 - Similar remark made by A. M. Dimits in *Phys. Plasmas* (2010).

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Phase-space Lagrangian formalism

Take a dynamical system defined by a Hamiltonian $H(\mathbf{q}, \mathbf{p}, t)$ and the canonical Poisson bracket $\{q^i, p_j\} = \delta_j^i$. Then, Hamilton equations can be obtained by making stationary the functional

$$S[\mathbf{q}, \mathbf{p}] = \int_{t_1}^{t_2} \mathcal{L}(\mathbf{q}(t), \mathbf{p}(t), \dot{\mathbf{q}}(t), \dot{\mathbf{p}}(t), t) dt,$$

where

$$\mathcal{L}(\mathbf{q}, \mathbf{p}, \dot{\mathbf{q}}, \dot{\mathbf{p}}, t) = p_i \dot{q}^i - H(\mathbf{q}, \mathbf{p}, t).$$

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This formulation combines

- scalar invariance and variational principle from Lagrangian mechanics,
- flexibility of coordinate transformations in phase-space from Hamiltonian mechanics.

Lagrangian of a charged particle in an electromagnetic field

The Lagrangian in Gaussian units is given by

$$\mathcal{L}(\mathbf{r}, \mathbf{v}, \dot{\mathbf{r}}, \dot{\mathbf{v}}, t) = \left((Ze/c)\mathbf{A}(\mathbf{r}) + M\mathbf{v} \right) \cdot \dot{\mathbf{r}} - H(\mathbf{r}, \mathbf{v}, t),$$

with the Hamiltonian

$$H(\mathbf{r}, \mathbf{v}, t) = Mv^2/2 + Ze\varphi(\mathbf{r}, t).$$

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In dimensionless variables

$$\check{\mathbf{r}} = \frac{\mathbf{r}}{L}, \check{\mathbf{v}} = \frac{\mathbf{v}}{v_t}, \check{t} = \frac{v_t t}{L}, \check{\varphi} = \frac{Ze\varphi}{\epsilon M v_t^2}, \check{\mathbf{A}} = \frac{\mathbf{A}}{B_0 L}, \check{H} = \frac{H}{M v_t^2}, \check{\mathcal{L}} = \frac{\mathcal{L}}{M v_t^2},$$

$$\check{\mathcal{L}} = (\epsilon^{-1} \check{\mathbf{A}}(\check{\mathbf{r}}) + \check{\mathbf{v}}) \cdot \dot{\check{\mathbf{r}}} - (\check{v}^2/2 + \epsilon \check{\varphi}(\check{\mathbf{r}}, \check{t})).$$

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$$\check{\mathcal{L}} = (\epsilon^{-1} \check{\mathbf{A}}(\check{\mathbf{r}}) + \check{\mathbf{v}}) \cdot \dot{\check{\mathbf{r}}} - (\check{v}^2/2 + \epsilon \check{\varphi}(\check{\mathbf{r}}, \check{t})).$$

Our ordering is completed by assuming that

$$\nabla_{\mathbf{r}_\perp} \check{\varphi} \sim O(\epsilon^{-1}). \quad \text{Formally, } \check{\varphi}(\check{\mathbf{r}}_\perp/\epsilon, \check{\mathbf{r}}_\parallel, \check{t}).$$

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General strategy of the change of coordinates

$$\mathcal{L}(\mathbf{q}, \mathbf{p}, \dot{\mathbf{q}}, \dot{\mathbf{p}}, t) = p_i \dot{q}^i - H(\mathbf{q}, \mathbf{p}, t).$$

We obtain equivalent Lagrangians via:

- Change of coordinates: $\mathbf{q}(\mathbf{Q}, \mathbf{P}, t)$, $\mathbf{p}(\mathbf{Q}, \mathbf{P}, t)$.
- Addition of total time derivatives: $\mathcal{L} \mapsto \mathcal{L} + dS/dt$.

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Applying both transformations:

$$\begin{aligned} \mathcal{L}'(\mathbf{Q}, \mathbf{P}, \dot{\mathbf{Q}}, \dot{\mathbf{P}}, t) &= \left(p_i(\mathbf{Q}, \mathbf{P}, t) \frac{\partial q^i}{\partial Q^j} + \frac{\partial S}{\partial Q^j} \right) \dot{Q}^j + \left(p_i(\mathbf{Q}, \mathbf{P}, t) \frac{\partial q^i}{\partial P^j} + \frac{\partial S}{\partial P^j} \right) \dot{P}^j \\ &\quad - \left(H(\mathbf{q}(\mathbf{Q}, \mathbf{P}, t), \mathbf{p}(\mathbf{Q}, \mathbf{P}, t), t) - p_i(\mathbf{Q}, \mathbf{P}, t) \frac{\partial q^i}{\partial t} - \frac{\partial S}{\partial t} \right). \end{aligned}$$

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Roughly speaking, in gyrokinetics $(\mathbf{Q}, \mathbf{P}) \equiv \{\mathbf{R}, u, \mu, \theta\}$ and one wants to answer the following question: *Do new coordinates $\{\mathbf{R}, u, \mu, \theta\}$ and a function S exist such that \mathcal{L}' is independent of θ and μ is a constant of motion?*

General strategy of the change of coordinates

Just for technical convenience, we will write the gyrokinetic transformation T_ϵ as the composition of two transformations,

$$T_\epsilon = T_{1\epsilon} T_{2\epsilon}.$$

- $T_{1\epsilon}$ is a prescribed, preparatory change of variables adapted to the magnetic field structure.
- $T_{2\epsilon}$ is determined order by order by imposing the gyrophase-invariance of the Lagrangian and the conservation of μ .

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Remark: We will not use a geometrical formulation (no mention of pull-backs and differential forms in the talk, except this one!). For those familiar with this machinery, however, it will be obvious how to translate every step of the computation into that language.

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Remark: In order to determine $\overline{H}^{(2)}$, computation of the term proportional to $\dot{\theta}$ in the Lagrangian is needed at order ϵ^3 . We will use the notation $\mathcal{L} = \dots + O(\epsilon^3, \epsilon^4)$.

Transformed Lagrangian under $T_{1\epsilon}$

$\hat{\mathbf{e}}_1(\mathbf{r})$, $\hat{\mathbf{e}}_2(\mathbf{r})$, orthonormal vector fields such that $\hat{\mathbf{e}}_1(\mathbf{r}) \times \hat{\mathbf{e}}_2(\mathbf{r}) = \hat{\mathbf{b}}(\mathbf{r})$.

Then, $(\mathbf{r}, \mathbf{v}) = T_{1\epsilon}(\mathbf{R}_g, v_{\parallel g}, \mu_g, \theta_g)$ is defined by

$$(\mathbf{r}, \mathbf{v}) = \left(\mathbf{R}_g + \epsilon \boldsymbol{\rho}(\mathbf{R}_g, \mu_g, \theta_g), \quad v_{\parallel g} \hat{\mathbf{b}}(\mathbf{R}_g) + \boldsymbol{\rho}(\mathbf{R}_g, \mu_g, \theta_g) \times \mathbf{B}(\mathbf{R}_g) \right).$$

$$\boldsymbol{\rho}(\mathbf{R}_g, \mu_g, \theta_g) := -\sqrt{2\mu_g/B(\mathbf{R}_g)} [\sin \theta_g \hat{\mathbf{e}}_1(\mathbf{R}_g) - \cos \theta_g \hat{\mathbf{e}}_2(\mathbf{R}_g)].$$

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Now, we compute the transformed Lagrangian at the required order and add the total time derivative

$$-\frac{d}{dt} \left\{ \mathbf{A}_g \cdot \boldsymbol{\rho}_g + \frac{\epsilon}{2} \boldsymbol{\rho}_g \boldsymbol{\rho}_g : \nabla_{\mathbf{R}_g} \mathbf{A}_g + \frac{\epsilon^2}{6} \boldsymbol{\rho}_g \boldsymbol{\rho}_g : \nabla_{\mathbf{R}_g} \nabla_{\mathbf{R}_g} \mathbf{A}_g \cdot \boldsymbol{\rho}_g \right. \\ \left. - \epsilon^3 \int_0^{\theta_g} \left[\tilde{\mathbf{A}}_3 \cdot (\boldsymbol{\rho}_g \times \hat{\mathbf{b}}) - \langle \tilde{\mathbf{A}}_3 \cdot (\boldsymbol{\rho}_g \times \hat{\mathbf{b}}) \rangle \right] d\theta'_g \right\}$$

Transformed Lagrangian under $T_{1\epsilon}$

$$\begin{aligned} \mathcal{L} \circ T_{1\epsilon}(\mathbf{R}_g, v_{||g}, \mu_g, \theta_g) &= \left(\frac{1}{\epsilon} \mathbf{A}_g + v_{||g} \hat{\mathbf{b}}_g + \epsilon \mathbf{\Gamma}_{\mathbf{R}_g}^{(1)} + \epsilon^2 \mathbf{\Gamma}_{\mathbf{R}_g}^{(2)} \right) \cdot \dot{\mathbf{R}}_g \\ &+ \epsilon \left(-\mu_g + \epsilon \Gamma_{\theta_g}^{(1)} + \epsilon^2 \Gamma_{\theta_g}^{(2)} \right) \dot{\theta}_g - H^{(0)} - \epsilon H^{(1)} + O(\epsilon^3, \epsilon^4), \end{aligned}$$

where

$$\mathbf{\Gamma}_{\mathbf{R}_g}^{(1)} = \mu_g \nabla_{\mathbf{R}_g} \hat{\mathbf{e}}_{2g} \cdot \hat{\mathbf{e}}_{1g} - v_{||g} \nabla_{\mathbf{R}_g} \hat{\mathbf{b}}_g \cdot \boldsymbol{\rho}_g - \frac{1}{2} (\boldsymbol{\rho}_g \cdot \nabla_{\mathbf{R}_g} B_g) \boldsymbol{\rho}_g \times \hat{\mathbf{b}}_g + \frac{B_g}{2} \boldsymbol{\rho}_g \cdot \nabla_{\mathbf{R}_g} \hat{\mathbf{b}}_g \times \boldsymbol{\rho}_g,$$

$$\mathbf{\Gamma}_{\mathbf{R}_g}^{(2)} = \frac{1}{6} \boldsymbol{\rho}_g \boldsymbol{\rho}_g : \nabla_{\mathbf{R}_g} \nabla_{\mathbf{R}_g} \mathbf{B}_g \times \boldsymbol{\rho}_g - \frac{2\mu_g}{3B_g} (\boldsymbol{\rho}_g \cdot \nabla_{\mathbf{R}_g} B_g) \nabla_{\mathbf{R}_g} \hat{\mathbf{e}}_{2g} \cdot \hat{\mathbf{e}}_{1g}$$

$$- \frac{B_g}{3} [\boldsymbol{\rho}_g \cdot \nabla_{\mathbf{R}_g} \hat{\mathbf{b}}_g \cdot (\boldsymbol{\rho}_g \times \hat{\mathbf{b}}_g)] \nabla_{\mathbf{R}_g} \hat{\mathbf{b}}_g \cdot \boldsymbol{\rho}_g,$$

$$\Gamma_{\theta_g}^{(1)} = \frac{2\mu_g}{3B_g} \boldsymbol{\rho}_g \cdot \nabla_{\mathbf{R}_g} B_g, \quad \Gamma_{\theta_g}^{(2)} = \frac{\mu_g^2}{4B_g^2} (\mathbf{I} - \hat{\mathbf{b}}_g \hat{\mathbf{b}}_g) : \nabla_{\mathbf{R}_g} \nabla_{\mathbf{R}_g} \mathbf{B}_g \cdot \hat{\mathbf{b}}_g,$$

$$H^{(0)} = \frac{1}{2} v_{||g}^2 + \mu_g B_g, \quad H^{(1)} = \langle \phi \rangle (\mathbf{R}_{g\perp} / \epsilon, R_{g||}, \mu_g, t) + \tilde{\phi} (\mathbf{R}_{g\perp} / \epsilon, R_{g||}, \mu_g, \theta_g, t).$$

$$\phi (\mathbf{R}_{g\perp} / \epsilon, R_{g||}, \mu_g, \theta_g, t) := \varphi (\mathbf{R}_{g\perp} / \epsilon + \boldsymbol{\rho} (\mathbf{R}_g, \mu_g, \theta_g), R_{g||}, t).$$

Transformed Lagrangian under $T_{1\epsilon}T_{2\epsilon}$

The operator $(\mathbf{R}_g, v_{||g}, \mu_g, \theta_g) = T_{2\epsilon}(\mathbf{R}, u, \mu, \theta)$ can be written as a Lie-transform

$$(\mathbf{R}_g, v_{||g}, \mu_g, \theta_g) = \exp \sum_{n=1}^{\infty} \epsilon^n \left(\mathbf{R}_n \cdot \nabla_{\mathbf{R}} + u_n \partial_u + \mu_n \partial_\mu + \theta_n \partial_\theta \right) (\mathbf{R}, u, \mu, \theta),$$

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but the following variables are more useful for us ($\mathbf{R}_1 \equiv 0$):

$$\tilde{\mathbf{R}}_3 = \mathbf{R}_3 + \frac{1}{2} \mathbf{R}_2 \cdot \nabla_{(\mathbf{R}_\perp/\epsilon)} \mathbf{R}_2 + \frac{u_1}{2} \frac{\partial \mathbf{R}_2}{\partial u} + \frac{\mu_1}{2} \frac{\partial \mathbf{R}_2}{\partial \mu} + \frac{\theta_1}{2} \frac{\partial \mathbf{R}_2}{\partial \theta}.$$

$$\tilde{u}_2 = u_2 + \frac{u_1}{2} \frac{\partial u_1}{\partial u} + \frac{\mu_1}{2} \frac{\partial u_1}{\partial \mu} + \frac{\theta_1}{2} \frac{\partial u_1}{\partial \theta},$$

$$\tilde{\mu}_2 = \mu_2 + \frac{1}{2} \mathbf{R}_2 \cdot \nabla_{(\mathbf{R}_\perp/\epsilon)} \mu_1 + \frac{u_1}{2} \frac{\partial \mu_1}{\partial u} + \frac{\mu_1}{2} \frac{\partial \mu_1}{\partial \mu} + \frac{\theta_1}{2} \frac{\partial \mu_1}{\partial \theta},$$

$$\tilde{\theta}_2 = \theta_2 + \frac{1}{2} \mathbf{R}_2 \cdot \nabla_{(\mathbf{R}_\perp/\epsilon)} \theta_1 + \frac{u_1}{2} \frac{\partial \theta_1}{\partial u} + \frac{\mu_1}{2} \frac{\partial \theta_1}{\partial \mu} + \frac{\theta_1}{2} \frac{\partial \theta_1}{\partial \theta}$$

$$(\mathbf{R}_g, v_{||g}, \mu_g, \theta_g) =$$

$$(\mathbf{R} + \epsilon^2 \mathbf{R}_2 + \epsilon^3 \tilde{\mathbf{R}}_3, u + \epsilon u_1 + \epsilon^2 \tilde{u}_2, \mu + \epsilon \mu_1 + \epsilon^2 \tilde{\mu}_2, \theta + \epsilon \theta_1 + \epsilon^2 \tilde{\theta}_2) + \dots$$

$$\mathcal{L} \circ T_{1\epsilon} \circ T_{2\epsilon} =$$

$$\begin{aligned} & \left[\epsilon^{-1} \mathbf{A}_{\mathbf{R}} + u \hat{\mathbf{b}}_{\mathbf{R}} + \epsilon \left(\Gamma_{\mathbf{R}}^{(1)} + \mathbf{R}_2 \cdot \nabla_{\mathbf{R}} \mathbf{A}_{\mathbf{R}} + u_1 \hat{\mathbf{b}}_{\mathbf{R}} \right) + \epsilon^2 \left(\Gamma_{\mathbf{R}}^{(2)} + \tilde{\mathbf{R}}_3 \cdot \nabla_{\mathbf{R}} \mathbf{A}_{\mathbf{R}} \right. \right. \\ & \left. \left. + \tilde{u}_2 \hat{\mathbf{b}}_{\mathbf{R}} + u \mathbf{R}_2 \cdot \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} + u_1 \partial_u \Gamma_{\mathbf{R}}^{(1)} + \mu_1 \partial_\mu \Gamma_{\mathbf{R}}^{(1)} + \theta_1 \partial_\theta \Gamma_{\mathbf{R}}^{(1)} \right) \right] \cdot \dot{\mathbf{R}} + \epsilon \mathbf{A}_{\mathbf{R}} \cdot \dot{\mathbf{R}}_2 \\ & + \epsilon^2 \left[u \hat{\mathbf{b}}_{\mathbf{R}} \cdot \dot{\mathbf{R}}_2 + \mathbf{A}_{\mathbf{R}} \cdot \dot{\tilde{\mathbf{R}}}_3 \right] + \epsilon^3 \left[\mathbf{A}_{\mathbf{R}} \cdot \dot{\tilde{\mathbf{R}}}_4 + u \hat{\mathbf{b}}_{\mathbf{R}} \cdot \dot{\tilde{\mathbf{R}}}_3 \right. \\ & \left. + \left(\Gamma_{\mathbf{R}}^{(1)} + \mathbf{R}_2 \cdot \nabla_{\mathbf{R}} \mathbf{A}_{\mathbf{R}} + u_1 \hat{\mathbf{b}}_{\mathbf{R}} \right) \cdot \dot{\mathbf{R}}_2 \right] + \epsilon \left[-\mu + \epsilon (\Gamma_\theta^{(1)} - \mu_1) \right. \\ & \left. + \epsilon^2 \left(\Gamma_\theta^{(2)} - \tilde{\mu}_2 + \mu_1 \partial_\mu \Gamma_\theta^{(1)} + \theta_1 \partial_\theta \Gamma_\theta^{(1)} \right) \right] \dot{\theta} - \epsilon^2 \mu \dot{\theta}_1 \\ & + \epsilon^3 \left[(\Gamma_\theta^{(1)} - \mu_1) \dot{\theta}_1 - \mu \dot{\theta}_2 \right] - H_{\mathbf{R}}^{(0)} - \epsilon \left(H_{\mathbf{R}}^{(1)} + uu_1 + \mu_1 B_{\mathbf{R}} \right) \\ & - \epsilon^2 \left(u \tilde{u}_2 + \frac{u_1^2}{2} + \tilde{\mu}_2 B_{\mathbf{R}} + \mu \mathbf{R}_2 \cdot \nabla_{\mathbf{R}} B_{\mathbf{R}} + \mathbf{R}_2 \cdot \nabla_{(\mathbf{R}_\perp/\epsilon)} H_{\mathbf{R}}^{(1)} \right. \\ & \left. + \mu_1 \partial_\mu H_{\mathbf{R}}^{(1)} + \theta_1 \partial_\theta H_{\mathbf{R}}^{(1)} \right) + O(\epsilon^3, \epsilon^4). \end{aligned}$$

Allow the addition of a total derivative to the Lagrangian,

$$\frac{d}{dt} \left[\epsilon S^{(1)} + \epsilon^2 S^{(2)} + \epsilon^3 S^{(3)} \right] (\mathbf{R}_\perp / \epsilon, \mathbf{R}, u, \mu, \theta, t),$$

and impose that the Lagrangian be of the form

$$\mathcal{L} = \left[\frac{1}{\epsilon} \mathbf{A}_R + u \hat{\mathbf{b}}_R + \epsilon \bar{\Gamma}_R^{(1)} \right] \cdot \frac{d\mathbf{R}}{dt} - \epsilon \mu \frac{d\theta}{dt} - \bar{H}^{(0)} - \epsilon \bar{H}^{(1)} - \epsilon^2 \bar{H}^{(2)} + O(\epsilon^3, \epsilon^4),$$

where $\bar{H}^{(0)}, \bar{H}^{(1)}, \bar{H}^{(2)}$ are gyrophase independent and we choose

$$\bar{\Gamma}_R^{(1)} = \mu \nabla_R \hat{\mathbf{e}}_{2R} \cdot \hat{\mathbf{e}}_{1R} - \frac{\mu}{2} \hat{\mathbf{b}}_R \hat{\mathbf{b}}_R \cdot \nabla_R \times \hat{\mathbf{b}}_R.$$

These conditions allow to determine $\mathbf{R}_2, u_1, \mu_1, \theta_1, \tilde{\mathbf{R}}_{3\perp}, \tilde{u}_2, \tilde{\mu}_2, \bar{H}^{(0)}, \bar{H}^{(1)}, \bar{H}^{(2)}, S^{(1)}, S^{(2)}$ and $S^{(3)}$.

Explicit expressions for the change of coordinates

$$\mathbf{R}_{2\perp} = -\frac{u}{B_{\mathbf{R}}} \hat{\mathbf{b}}_{\mathbf{R}} \times \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \boldsymbol{\rho} - \frac{1}{2B_{\mathbf{R}}} \boldsymbol{\rho} \boldsymbol{\rho} \cdot \nabla_{\mathbf{R}} B_{\mathbf{R}} - \frac{1}{B_{\mathbf{R}}^2} \hat{\mathbf{b}}_{\mathbf{R}} \times \nabla_{(\mathbf{R}_{\perp}/\epsilon)} \tilde{\Phi},$$

$$R_{2\parallel} = -\frac{2u}{B_{\mathbf{R}}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla \hat{\mathbf{b}}_{\mathbf{R}} \cdot (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) - \frac{1}{8} \left[\boldsymbol{\rho} \boldsymbol{\rho} - (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) \right] : \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}$$

$$u_1 = u \hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla \hat{\mathbf{b}}_{\mathbf{R}} \cdot \boldsymbol{\rho} - \frac{B_{\mathbf{R}}}{4} \left[\boldsymbol{\rho}(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) + (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})\boldsymbol{\rho} \right] : \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}$$

$$\mu_1 = -\frac{\tilde{\phi}}{B_{\mathbf{R}}} - \frac{u^2}{B_{\mathbf{R}}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla \hat{\mathbf{b}}_{\mathbf{R}} \cdot \boldsymbol{\rho} + \frac{u}{4} \left[\boldsymbol{\rho}(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) + (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})\boldsymbol{\rho} \right] : \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}.$$

$$\theta_1 = \frac{1}{B_{\mathbf{R}}} \frac{\partial \tilde{\Phi}}{\partial \mu} + \frac{u^2}{2\mu B_{\mathbf{R}}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla \hat{\mathbf{b}}_{\mathbf{R}} \cdot (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) \\ + \frac{u}{8\mu} \left[\boldsymbol{\rho} \boldsymbol{\rho} - (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) \right] : \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} + \frac{1}{B_{\mathbf{R}}} (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) \cdot \nabla_{\mathbf{R}} B_{\mathbf{R}}.$$

$$\tilde{\Phi} := \int^{\theta} \tilde{\phi} d\theta'.$$

Explicit expressions for the gyrokinetic Hamiltonian

$$\bar{H}^{(0)} = \frac{1}{2}u^2 + \mu B_{\mathbf{R}},$$

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$$\begin{aligned} \bar{H}^{(2)} = & \frac{1}{2B_{\mathbf{R}}^2} \left\langle \nabla_{(\mathbf{R}_{\perp}/\epsilon)} \tilde{\Phi} \cdot \left(\hat{\mathbf{b}}_{\mathbf{R}} \times \nabla_{(\mathbf{R}_{\perp}/\epsilon)} \tilde{\phi} \right) \right\rangle - \frac{1}{2B_{\mathbf{R}}} \frac{\partial \langle \tilde{\phi}^2 \rangle}{\partial \mu} \\ & - \frac{u}{B_{\mathbf{R}}} \left\langle \left(\nabla_{(\mathbf{R}_{\perp}/\epsilon)} \tilde{\phi} \times \hat{\mathbf{b}}_{\mathbf{R}} \right) \cdot \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \boldsymbol{\rho} \right\rangle - \frac{\mu}{2B_{\mathbf{R}}^2} \nabla_{\mathbf{R}} B_{\mathbf{R}} \cdot \nabla_{(\mathbf{R}_{\perp}/\epsilon)} \langle \phi \rangle \\ & - \frac{1}{4B_{\mathbf{R}}} \left\langle \nabla_{(\mathbf{R}_{\perp}/\epsilon)} \tilde{\phi} \cdot \left[\boldsymbol{\rho} \boldsymbol{\rho} - (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) \right] \cdot \nabla_{\mathbf{R}} B_{\mathbf{R}} \right\rangle - \frac{1}{B_{\mathbf{R}}} \nabla_{\mathbf{R}} B_{\mathbf{R}} \cdot \langle \tilde{\phi} \boldsymbol{\rho} \rangle \\ & - \frac{u^2}{B_{\mathbf{R}}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} \cdot \frac{\partial}{\partial \mu} \langle \tilde{\phi} \boldsymbol{\rho} \rangle + \frac{u}{4} \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} : \frac{\partial}{\partial \mu} \left\langle \tilde{\phi} \left[\boldsymbol{\rho}(\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}}) + (\boldsymbol{\rho} \times \hat{\mathbf{b}}_{\mathbf{R}})\boldsymbol{\rho} \right] \right\rangle \\ & - \frac{3u^2\mu}{2B_{\mathbf{R}}^2} (\hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}) \cdot \nabla_{\mathbf{R}} B_{\mathbf{R}} + \frac{\mu^2}{4B_{\mathbf{R}}} (\mathbf{I} - \hat{\mathbf{b}}_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}) : \nabla_{\mathbf{R}} \nabla_{\mathbf{R}} B_{\mathbf{R}} \cdot \hat{\mathbf{b}}_{\mathbf{R}} \\ & - \frac{3\mu^2}{4B_{\mathbf{R}}^2} |\nabla_{\mathbf{R}\perp} B_{\mathbf{R}}|^2 + \frac{u^2\mu}{2B_{\mathbf{R}}} \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} : \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} + \left(\frac{\mu^2}{8} - \frac{\mu u^2}{4B_{\mathbf{R}}} \right) \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}} : (\nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}})^{\mathbf{T}} \\ & - \left(\frac{3\mu u^2}{8B_{\mathbf{R}}} + \frac{\mu^2}{16} \right) (\nabla_{\mathbf{R}} \cdot \hat{\mathbf{b}}_{\mathbf{R}})^2 + \left(\frac{3\mu u^2}{2B_{\mathbf{R}}} - \frac{u^4}{2B_{\mathbf{R}}^2} \right) |\hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla_{\mathbf{R}} \hat{\mathbf{b}}_{\mathbf{R}}|^2 \\ & + \left(\frac{\mu u^2}{8B_{\mathbf{R}}} - \frac{\mu^2}{16} \right) (\hat{\mathbf{b}}_{\mathbf{R}} \cdot \nabla_{\mathbf{R}} \times \hat{\mathbf{b}}_{\mathbf{R}})^2. \end{aligned}$$

- 1 Motivation
- 2 Preliminaries: phase-space Lagrangian formalism and gyrokinetic ordering
- 3 Gyrokinetic change of coordinates
- 4 Gyrokinetic Poisson equation
- 5 Conclusions and further work

Gyrokinetic Poisson equation

In dimensionless variables and denoting by λ_D the Debye length:

$$\nabla^2 \phi(\mathbf{r}, t) = -ZL^2/\lambda_D^2 \left(\int f(\mathbf{r}, \mathbf{v}, t) d^3\mathbf{v} - n_e(\mathbf{r}, t) \right).$$

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If $\lambda_D/\rho_i \ll 1$ and $\nabla_{\mathbf{R}} F \sim O(1)$, this equation gives

$$\int \mathcal{B}_{\parallel}^*(\mathbf{R}, u) [F(\mathbf{R}, u, \mu, t) - \epsilon u_{1\mathbf{R}} \partial_2 F(\mathbf{R}, u, \mu, t) - \epsilon \mu_{1\mathbf{R}} \partial_3 F(\mathbf{R}, u, \mu, t)] \\ \times \delta(\mathbf{R} + \epsilon \boldsymbol{\rho}(\mathbf{R}, \mu, \theta) - \mathbf{r}) d^3\mathbf{R} du d\mu d\theta - n_e(\mathbf{r}, t) + O(\epsilon^2) = 0,$$

where $\mathcal{B}_{\parallel}^*(\mathbf{R}, u) = \left(\mathbf{B}(\mathbf{R}) + \epsilon u \nabla \times \mathbf{b}(\mathbf{R}) \right) \cdot \mathbf{b}(\mathbf{R})$.

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A careful elimination of the delta function yields

$$0 = -n_e(\mathbf{r}, t) + \int \mathcal{B}_{\parallel}^*(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), u) \left[F(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), u, \mu, t) \right. \\ \left. - \epsilon u_{1\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta)} \partial_2 F(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), u, \mu, t) \right. \\ \left. - \epsilon \mu_{1\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta)} \partial_3 F(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), u, \mu, t) \right] \left(1 - \epsilon \nabla_{\mathbf{r}} \cdot \boldsymbol{\rho}(\mathbf{r}, \mu, \theta) \right) du d\mu d\theta + O(\epsilon^2).$$

Gyrokinetic Poisson equation

We can recover more familiar expressions assuming

$$F = F_0 + \epsilon F_1 + O(\epsilon^2), \quad \nabla_{\mathbf{R}} F \sim O(1),$$

Gyrokinetic Poisson equation

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Then,

$$\begin{aligned} & \int B_{\parallel}(\mathbf{r}) F_0(\mathbf{r}, u, \mu, t) \mathrm{d}u \mathrm{d}\mu \mathrm{d}\theta + \epsilon \int B_{\parallel}(\mathbf{r}) F_1(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), u, \mu, t) \mathrm{d}u \mathrm{d}\mu \mathrm{d}\theta \\ & + \frac{\epsilon}{B_{\mathbf{r}}} \int B_{\parallel}(\mathbf{r}) \tilde{\phi}(\mathbf{r} - \epsilon \boldsymbol{\rho}(\mathbf{r}, \mu, \theta), t) \partial_3 F_0(\mathbf{r}, u, \mu, t) \mathrm{d}u \mathrm{d}\mu \mathrm{d}\theta \\ & + \epsilon \mathbf{b}_{\mathbf{r}} \cdot (\nabla \times \mathbf{b}_{\mathbf{r}}) \int u F_0(\mathbf{r}, u, \mu, t) \mathrm{d}u \mathrm{d}\mu \mathrm{d}\theta - n_e(\mathbf{r}, t) + O(\epsilon^2) = 0. \end{aligned}$$

- The Jacobian coming from the delta function does not contribute.
- From the final expression we **deduce** that $\nabla_{\mathbf{R}_{\perp}} F_1 \sim \epsilon^{-1}$.

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Conclusions and further work

- We have formulated a gyrokinetic theory with a single expansion parameter, ϵ , in the Lagrangian formalism.
 - No splitting between guiding-center and gyrokinetic dynamics is possible.
- The ϵ^2 term of the gyrokinetic Hamiltonian in general geometry has been explicitly computed.
 - By setting $\varphi \equiv 0$ we get for free the guiding-center Lagrangian with zero electrostatic potential to an order higher than available in the literature, as far as we know.

Next steps:

- Work out the electromagnetic case.
- Include external flows.
- Implications?