Rotational shear in tokamak plasmas

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Fusion: how to spin it

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Overview

- Motivation
- Effect of rotational shear on turbulent transport
- Implications for local gradients (0D)
- Extension to radial profiles (1D)



Objective

- Identify mechanism(s) for achieving enhanced confinement
- Internal transport barriers observed with temperature gradients well above threshold
- Often accompanied by large E x B shear and low or negative magnetic shear
- Experimentally observed power threshold for formation

Connor et al. (2004)

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Model

GK equation with mean flow satisfying $\ \frac{\rho}{L} \ll M \ll 1$ but : $\nabla u \sim v_{th}/L$

$$\frac{dh}{dt} + \left(\mathbf{v}_{\parallel} + \mathbf{v}_{D} + \langle \mathbf{v}_{E} \rangle\right) \cdot \nabla h - \langle C[h] \rangle$$

$$= \frac{eF_{0}}{T} \frac{d\langle \varphi \rangle}{dt} - \langle \mathbf{v}_{E} \rangle \cdot \nabla \psi \left(\frac{dF_{0}}{d\psi} + \frac{mv_{\parallel}}{T} \frac{RB_{\phi}}{B} \frac{d\omega}{d\psi} F_{0}\right)$$

$$\frac{d}{dt} \equiv \frac{\partial}{\partial t} + R\omega(\psi)\hat{\mathbf{e}}_{\phi} \cdot \nabla$$
Local approximation:
$$\omega(\psi) \approx \omega(\psi_{0}) + (\psi - \psi_{0}) \frac{d\omega}{d\psi} \Big|_{\psi}$$

 $q d\psi v_{th}$

Linear stability

Cyclone base case: r/R = 0.18 q = 1.4 $\hat{s} = 0.8$



 \sim

- ITG drive at small shear
- ITG/PVG drive at moderate shear
- Stabilization at large shear
- Roughly linear dependence of critical flow shear on R/LT

Barnes et al., 2010

Transient growth



- Beyond critical shear value, transient linear growth
- Amplification of initial amplitude increases with shear
- Cf. Newton et al., 2010

Barnes et al., 2010





$$Q = \int d^3 v \, \frac{mv^2}{2} \left(\mathbf{v}_E \cdot \frac{\nabla \psi}{|\nabla \psi|} \right) \delta f$$
$$\Pi = \int d^3 v \, mR^2 \left(\mathbf{v} \cdot \nabla \phi \right) \left(\mathbf{v}_E \cdot \frac{\nabla \psi}{|\nabla \psi|} \right) \delta f$$

- Fluxes follow linear trends up to linear stabilization point
- Subcritical (linearly stable) turbulence beyond this point
- Optimal flow shear for confinement
- Possible hysteresis
- Maximum in momentum flux => possible bifurcation

Turbulent Prandtl number $\Pr = rac{ u_i}{\chi_i}$ $\Pi_i = -m_i v_{th} (qR_0/r) \nu_i \gamma_E$ $Q_i = -\chi_i dT_i/dr$



 Prandtl number tends to shearand R/LTindependent value of order unity (in both turbulence regimes)

Barnes et al., 2010

Stiffness



Barnes et al., 2010

- Complicated dependence on shear
- Generally, critical gradient shifts higher and stiffness increases at low shear
- Critical gradient shifts lower and stiffness decreases at high shear (when turbulence driven by shear instead of R/LT)

Zero magnetic shear



- Similar...sort of
- All turbulence subcritical

Highcock et al., 2010

Zero magnetic shear





- Similar...sort of
- All turbulence subcritical
- Very different critical flow shear values

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Two possible bifurcations

Adding neoclassical leads to a bifurcation

 It seems to contradict phenomenology

 Adding intrinsic rotation leads to another type of bifurcation

 Speculative

Balance w/o neoclassical



- Q = red lines • Π/Q = green lines $\frac{R}{L_T} = \frac{Pr_t}{\Pi/Q} \gamma_E$
- Critical gradient = • dashed line
- For given Π/Q and Q, only one solution No bifurcation!

Neoclassical energy flux



Neoclassical energy flux





• Neoclassical $\frac{R}{L_T} = \frac{Pr_n}{\Pi/Q} \gamma_E$

• Turbulent $\frac{R}{L_T} = \frac{Pr_t}{\Pi/Q} \gamma_E$

• Prandtl numbers $Pr_n < Pr_t$ Banana orbits give energy flux, not momentum



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Possible solutions



Possible solutions



Possible solutions







Bifurcations



- Consider inverse problem: for fixed fluxes, what are gradients?
- With inclusion of neoclassical fluxes, we see bifurcation to much larger flow shear and R/LT

Highcock et al., 2010



Smooth transitions to neoclassical





Slope increases for • Small Π/Q Favors neutral beam heating

• Small $\frac{d(R/L_T|_c)}{d\gamma_E}$

Favors small magnetic shear regions



Intersecting lines



Neoclassical bifurcation

- Bifurcation happens due to lower neoclassical Prandtl number
- Numerically tested
- Possible to obtain large jumps
- Favors neutral beams and low magnetic shear
- It is easier at lower power!

Intrinsic rotation terms Idea: expansion on poloidal gyroradius $\Pi = \Pi_0 + \alpha \rho_p \frac{\partial^2 T}{\partial r^2} + \beta \rho_p \frac{\partial^2 \omega}{\partial r^2} + \dots$ • For low flow, only temperature matters $\frac{\Pi_{t,n}}{Q_{t,n}} = Pr_{t,n}\frac{\gamma_E}{R/L_T} + \alpha_{t,n}\rho_p\frac{\partial^2 T}{\partial T}/\partial r^2$ - Generation of intrinsic rotation (Parra & Catto, PPCF 2010) • Assume $\frac{\partial^2 T/\partial r^2}{\partial T/\partial r} \sim \pm \frac{R}{L_T}$



 $\alpha_t \alpha_n > 0$



 $\alpha_t \alpha_n < 0$

Intrinsic rotation bifurcation



Total energy flux



Intrinsic rotation bifurcation

- There is a power threshold
- Very speculative
- Requires high energy input

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Why do radial profile analysis?

- Ultimately, we want to predict mean profiles
- Magnetic geometry varies radially. Want to know where barrier forms
- Can address Coriolis pinch, turbulent and viscous heating, temperature equilibration, etc
- Inverse problem more forgiving (stiffness phenomenon reversed)

Multiple scale problem

| Physics | Perpendicular spatial scale | Temporal scale |
|------------------------------|---|--------------------------------------|
| Turbulence from ETG modes | k_{\perp}^{-1} ~ 0.005 – 0.05 cm | ω_{*} ~ 0.5 - 5.0 MHz |
| Turbulence from ITG modes | k_{\perp}^{-1} ~ 0.3 - 3.0 cm | ω_{*} ~ 10 - 100 kHz |
| Transport barriers | Measurements suggest width ~ 1 - 10 cm | 100 ms or more in core? |
| Discharge evolution | Profile scales ~ 200 cm | Energy confinement time ~ 2 - 4 s |

simulation cost: $(L_{\parallel}/\Delta_{\parallel}) \times (L_{\perp}/\Delta_{\perp})^2 \times (L_v/\Delta_v)^2 \times (L_t/\Delta t) \sim 10^{21}$

Transport equations in GK

Moment equations for evolution of mean quantities:

$$\begin{aligned} \frac{\partial n_s}{\partial t} &= -\frac{1}{V'} \frac{\partial}{\partial \psi} \left(V' \left\langle \mathbf{\Gamma}_s \cdot \nabla \psi \right\rangle \right) + S_n \\ \frac{3}{2} \frac{\partial n_s T_s}{\partial t} &= -\frac{1}{V'} \frac{\partial}{\partial \psi} \left(V' \left\langle \mathbf{Q}_s \cdot \nabla \psi \right\rangle \right) \\ &+ T_s \left(\frac{\partial \ln n_s}{\partial \psi} - \frac{3}{2} \frac{\partial \ln T_s}{\partial \psi} \right) \left\langle \mathbf{\Gamma}_s \cdot \nabla \psi \right\rangle + \frac{\partial \ln T_s}{\partial \psi} \left\langle \mathbf{Q}_s \cdot \nabla \psi \right\rangle \\ &- \left\langle \int d^3 v \frac{h_s T_s}{F_{0s}} \left\langle C[h_s] \right\rangle_{\mathbf{R}} \right\rangle + n_s \nu_{\epsilon}^{su} \left(T_u - T_s \right) + S_p \\ \frac{\partial L}{\partial t} &= -\frac{1}{V'} \frac{\partial}{\partial \psi} \left(V' \sum_s \left\langle \pi_s \right\rangle \right) + S_L \end{aligned}$$

...depend on fluctuations

Sugama (1997), Abel (2010)

Multiscale grid

 Turbulent fluctuations calculated in small regions of fine space-time grid embedded in "coarse" grid (for mean quantities)



TRINITY schematic



Sampling profile with flux tubes



Sampling profile with flux tubes



Simulation volume reduced by factor of ~10s

TRINITY transport solver

- Transport equations are stiff, nonlinear PDEs. Implicit treatment via Newton's Method (multi-step BDF, adaptive time step) allows for time steps ~0.1 seconds (vs. turbulence sim time ~0.001 seconds)
- Challenge: requires computation of quantities like

$$\Gamma_j^{m+1} \approx \Gamma_j^m + \left(\mathbf{y}^{m+1} - \mathbf{y}^m\right) \frac{\partial \Gamma_j}{\partial \mathbf{y}} \bigg|_{\mathbf{y}^m} \qquad \mathbf{y} = \left[\{n_k\}, \{p_{i_k}\}, \{p_{e_k}\}\right]^T$$

- Local approximation: $\frac{\partial \Gamma_j}{\partial n_k} = \frac{\partial \Gamma_j}{\partial n_j} + \frac{\partial \Gamma_j}{\partial (R/L_n)_j} \frac{\partial (R/L_n)_j}{\partial n_k}$
- Simplifying assumption: normalized fluxes depend primarily on gradient scale lengths

TRINITY scaling

• Example calculation with 10 radial grid points:

- evolve density, toroidal angular momentum, and electron/ion pressures
- simultaneously calculate fluxes for equilibrium profile and for 4 separate profiles (one for each perturbed gradient scale length)
- total of 50 flux tube simulations running simultaneously
- ~2000-4000 processors per flux tube => scaling to over 100,000 processors with high efficiency

JET shot #42982



- ITER demo discharge
- H-mode D-T plasma, record fusion energy yield
- Miller local equilibrium model: q, shear, shaping
- B = 3.9 T on axis
- TRANSP fits to experimental data taken from ITER profile database

Evolving density profile



- 10 radial grid points
- Costs ~120k CPU hrs (<10 clock hrs)
- Dens and temp profiles agree within ~15% across device
- Energy off by 5%
- Incremental energy off by 15%
- Flow shear absent

Fluctuations



Conclusions

- Maximum temperature gradient for given heat flux. Occurs at finite flow shear.
- Turbulent Prandtl number is constant of order unity for moderate to large flow shear values.
- Stiffness modestly decreased for high flow shear (PVG driven turbulence). Main effect at low flow shear is upshift of critical temperature gradient
- Two possible bifurcation types in 0D model:
 - Neoclassical bifurcation (observed from GS2 simulations)
 - Intrinsic rotation bifurcation (demonstrates power threshold)
- Current work focuses on extension to self-consistent, 1D transport simulations

TRINITY transport solver

- Calculating flux derivative approximations:
 - at every radial grid point, simultaneously calculate $\Gamma_j[(R/L_n)_j^m]$ and $\Gamma_j[(R/L_n)_j^m+\delta]$ using 2 different flux tubes
 - use 2-point finite differences:

$$\frac{\partial \Gamma_j}{\partial (R/L_n)_j} \approx \frac{\Gamma_j[(R/L_n)_j^m] - \Gamma_j[(R/L_n)_j^m + \delta]}{\delta}$$

- possible because flux tubes independent (do not communicate during calculation)
- perfect parallelization (almost)