β ~ I, magnetized plasmas in the laboratory

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Summary/Outline

- Large Plasma Device (LAPD) at UCLA: Upgraded plasma source: LaB6 cathode provides up to 100x increase in plasma pressure + warm ions; with lowered B, $\beta \sim I$
- Modification of pressure-gradient-driven turbulence and transport with increasing $\boldsymbol{\beta}$
 - Magnetic fluctuations increase, with parallel magnetic fluctuations dominant ($2x \ \tilde{B}_{\perp}$ at highest β); Observations consistent with Gradient Driven Coupling (GDC) instability: favorable comparisons with GENE simulations
- Opportunities to study processes relevant to space/astrophysical plasmas in LAPD and new device ETPD
 - Linear and nonlinear physics of Alfvén waves, MHD turbulence in $\beta \sim 1$ plasmas
 - Pressure anisotropy: can mirror and firehose instabilities be studied in the laboratory?

The LArge Plasma Device (LAPD) at UCLA



- Solenoidal magnetic field, cathode discharge plasma (BaO and LaB₆)
- BaO Cathode: $n \sim 10^{12} \text{ cm}^{-3}$, $T_e \sim 5-10 \text{ eV}$, $T_i \approx 1 \text{ eV}$
- LaB₆ Cathode: $n \sim 5 \times 10^{13} \text{ cm}^{-3}$, $T_e \sim 10-15 \text{ eV}$, $T_i \sim 6-10 \text{ eV}$
- B up to 2.5kG (with control of axial field profile)
- Large plasma size, 18m long, D~60cm (BaO) (1kG: ~300 ρ_i, ~100 ρ_s);
 D ~ 20cm (LaB₆ prototype)
- High repetition rate: I Hz
- US DOE/NSF user facility for basic plasma science: the Basic Plasma Science Facility or BaPSF (international users are welcome!)

LAPD BaO Plasma source



Examples of recent research using BaPSF/LAPD



 Decay instabilities of large amplitude shear Alfvén waves [Dorfman & Carter, PRL 116, 195002 (2016)]



 Excitation of chirping whistler waves by energetic electrons [Van Compernolle, et al., PRL 114, 245002 (2015)]



 Laser-driven magnetized collisionless shocks [Bondarenko, et al., Nat. Phys. (2017)]

LAPD prototype LaB₆ Cathode





- LaB6 cathode (operates at 1800C) 20cm square prototype cathode installed (larger version in development)
- Much better emissivity leads to 50x higher electron density, up to a factor of ~2 higher electron temperature, factor of ~10 higher ion temperature
- With reduced magnetic field, high β (order unity) achievable with magnetized ions (marginally so at $\beta \sim 1$)

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Enormous Toroidal Plasma Device at UCLA





- Former Electric Tokamak, (5m major radius, 1m minor radius)
- Operating now with LaB₆ cathode discharge into toroidal+vertical field
- Produces ~120m long, magnetized, high beta plasma (up to ~5x10¹³ cm⁻³, Te, Ti ~ 15-30eV, B~200G, β ~ 1).

High beta, hot ion plasmas in ETPD



- T_e ~ T_i ~ 20 eV measured (passive spectroscopy of He II 4686 line).
- With B~250G, plasma beta of order unity is achieved

Possible studies in ETPD



- Alfvén waves, damping at β~I (underway, data above), many (~I00)
 Alfvén parallel wavelengths in device; Wave-wave interactions,
 driven Alfvénic cascade at β~I
- Gradient-driven/interchange turbulence at high β
- Mirror/firehose: Drive anisotropy, higher beta through expansion (drive plasma into low field region)
- Reconnection, Shock physics

Why does β matter for pressure-gradient-driven turbulence and transport?

- Modifications to drift instabilities at finite β
 - Drift-Alfvén wave coupling at low β (mass ratio): leads to magnetic fluctuations (δA_{II} or δB_{\perp}) (seen in low β LAPD discharges)
 - ITG stabilization with increasing β
- Electromagnetic transport, e.g. magnetic flutter-transport arises with B_{\perp} fluctuations (Rechester, 1978)
- New instabilities may develop at finite β
 - Kinetic ballooning, etc.
 - Gradient-driven drift-coupling mode (GDC)
 - Couples to collisionless tearing modes and drives faster magnetic reconnection rates (Pueschel, 2015)

LAPD: β up to 15% produced with magnetic field scan



- Magnetic field varied from 1kG to 175G, 0.1% $\lesssim \beta \lesssim 15\%$ in the core
- ρ_i varies from ~3mm to ~2.5cm; marginal magnetization at highest β ($\rho_i/L_n \sim 4$; FLR effects likely important)

Fluctuation profiles: δn , δB_{\parallel} peak on gradient; δB_{\perp} core localized



Fluctuation profiles: δn , δB_{II} peak on gradient; δB_{\perp} core localized



 δB_{\perp} profile consistent with dominance of low-m modes

Magnetic fluctuations increase with β ; surprise is that δB_{I} dominates



Magnetic fluctuations increase with β , density fluctuations somewhat reduced



Strong parallel magnetic field fluctuations seen as β is increased



• Surprisingly, B₁ fluctuations dominate above $\beta \sim 1\%$

• Evidence for emergence of new instability?

Strong parallel magnetic field fluctuations seen as β is increased



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- Evidence for emergence of new instability?

Cross-correlation measurements reveal low-m structure and anti-correlation between $\tilde{n}_{_{\rm e}}$ and $\tilde{B}_{^{\rm I}}$



- Most fluctuation power in azimuthal modes m=1 to m=3
- Cross-correlation between density (Isat) and magnetic field fluctuations: two fields are ~π out of phase

Cross-correlation measurements reveal low-m structure and anti-correlation between $\tilde{n}_{_{\rm e}}$ and $\tilde{B}_{^{\|}}$



Electrostatic particle flux decreases with increasing β



 Peak ES particle flux decreases with β, primarily due to decreased density and electric field fluctuation amplitudes

Electromagnetic transport?



- Initial magnetic flux estimate from $\Gamma \sim \tilde{n} \ \tilde{v}_{\nabla B} \propto \tilde{n} \ \tilde{B}_{\scriptscriptstyle \|}$
- Decrease at higher β due to ion FLR? More work needed here (e.g. global particle balance, measurement of magnetic flutter transport)

Observations are consistent with newly* predicted instability, the Gradient Drift Coupling (GDC) instability

- GDC recently discovered in the context of magnetic reconnection in high-β current sheets [Pueschel, et al. PoP 2015]
 - Linear and nonlinear calculations show that it should be active in these LAPD experiments
 - Not likely to be excited in tokamaks due to magnetic shear, but expected in space/astro plasmas
- Similar to universal instability/interchange instability but relies on B_{I} perturbations that arise with density fluctuations in finite beta plasmas and the associated ∇B drifts in the perturbed fields
 - Predicts density, parallel magnetic fluctuations have π phase shift

Observations are consistent with newly predicted instability, the Gradient Drift Coupling (GDC) instability



GDC linear growth rate from GENE, using experimental parameters: growing, and trend reminiscent of trends in saturated amplitudes





Solar wind campaign: physics of $\beta \sim I$, warm ion plasmas

- Kinetic instabilities, waves and turbulence at high plasma beta ($v_A \sim v_{th,i}$) with warm ions
- Warm ions provide opportunity to study ion kinetic effects in waves and instabilities: e.g. ion FLR effects on Alfvén wave propagation; ion cyclotron absorption; modification to nonlinear Alfven wave interactions; MHD turbulence
- With lower field, plasma beta can be increased substantially to study, e.g., modifications to Alfvén wave dispersion and damping (e.g. ion Landau/ Barnes damping). Can temperature anisotropy driven instabilities (mirror and firehose) be observed in these plasmas?

Campaign Leader: Greg Howes (U. Iowa)



Previous studies of nonlinear Alfvén waves in LAPD

- Series of experiments exploring three-wave interactions and decay instabilities. Motivations include studying Alfvénic turbulence in the lab
- Collision of two antenna-launched shear Alfvén waves:
 - Two co-propagating AWs produce a quasimode [Carter, et al., PRL, 96, 155001 (2006)]
 - Two co-propagating KAWs drive drift waves, lead to control/ suppression of unstable modes (in favor of driven stable mode) [Auerbach, et al., PRL, 105, 135005 (2010)]
 - Two counter-propagating AWs, one long wavelength (k_I ≈ 0), produce daughter AW (building block of MHD turbulent cascade) [Howes, et al., PRL, 109, 255001 (2012)]
 - Two counter-propagating AWs nonlinearly excite an ion acoustic wave [Dorfman & Carter, PRL, 110, 195001 (2013)]
- Parametric instability of single large-amplitude shear wave [Dorfman & Carter, PRL, 116, 195002 (2016)]

Large amplitude Alfvén wave generation



- Antennas can generate AWs with $\delta B/B \sim 1\%$ (~10G or 1mT); large amplitude from several points of view:

 - Wave beta is of order unity $\beta_w = \frac{2\mu_o p}{\langle \delta B^2 \rangle} \approx 1$ Wave Poynting flux ~ 200 kW/m², same as discharge heating power density
 - From GS theory: stronger nonlinearity for anisotropic waves; here $k_{\parallel}/k_{\perp} \sim \delta B/B$

MHD-cascade relevant collisions: AW+AW \rightarrow AW

- Initial attempts in LAPD (Carter, Boldyrev, et al.): no strong evidence for daughter wave production/cascade (instead see beat waves, heating, harmonic generation, etc). Used local interaction, trying to look for perp. cascade.
- New idea (Howes): have one of the two interacting (pump) waves be $k_{\parallel} \approx 0$, theoretical prediction for stronger NL interaction in this case B_{0}



 UCLA Loop antenna (large amplitude) versus U. Iowa ASW antenna (small amplitude but precise k_⊥ control)

First laboratory observation of daughter AW production: consistent with weak turbulence theory



Howes et al., PRL 109, 255001 (2012)

• Perpendicular wavenumber spectrum consistent with threewave matching $(k_1 + k_2 = k_3)$

How do we achieve an MHD turbulent cascade in the lab?

- Have studied interesting AW interactions, but no success yet in LAPD in getting multiple collisions/development of inertial range
 - Two issues: (1) LAPD is long, but not long enough (not enough space for multiple collisions to occur/need to reflect waves, get a second pass) (2) Damping of pump wave is significant (we are near dissipation scale with pump)
 - Possible remedies: (1) new LAPD regime with higher beta, higher density (effective LAPD length gets longer, c/wpi smaller; lowered damping?) (2) Use ETPD to get much longer plasma length?
- Another approach: cascade driven by instabilities? One possibility: EM pressure gradient driven instabilities to act as stirring for cascade?

Goal: Can we excite temperature-anisotropy-driven modes in the lab: mirror and firehose instabilities



- Ion temperature anisotropy in the solar wind: limits explained by action of mirror and firehose modes
- Could play important role in solar wind thermodynamics; also thought to be important in other astrophysical plasmas, e.g. accretion disks

Mirror/Firehose in the laboratory?

- MHD threshold for firehose: $~~eta_{\parallel}-eta_{\perp}\geq 2$
- Resonant mirror/firehose can be triggered below the MHD threshold, but not far below; need high beta [Hellinger]
- Need to operate at low field to get high beta, but need magnetized ions. Also, typical scale size of firehose instability is c/ω_{pi}, need large enough density to include this scale in the experiment
- Need low collisionality! Growth time for firehose can be ~ f_{ci}, need collisionality at least lower than this tough when factoring in requirement for high density.

Firehose in LAPD/ETPD?

 $\beta_{\parallel} - \beta_{\perp} \ge 2 \qquad \text{Need } \mathsf{Mv}_{\mathsf{A}}^2 \text{ to be low enough}$ $k_B T_{\parallel} - k_B T_{\perp} \ge M v_A^2 \qquad \text{(I0 eV)?}$

 At LAPD/ETPD densities, this requires very low field: n=10¹³ /cc, B=50G, Mv_A² ~ 12 eV (He)

Helium	Hydrogen
$T_i \sim 20 \mathrm{eV}$	$T_i \sim 20 \mathrm{eV}$
$\nu_{ii} \sim 13 {\rm kHz}$	$ u_{ii} \sim 26 \mathrm{kHz}$
$f_{ci} \sim 20 \mathrm{kHz}$	$f_{ci} \sim 76 \mathrm{kHz}$
$ ho_i \sim 18 { m cm}$	$ ho_i \sim 9 { m cm}$

 Difficult, but may be possible to get to these conditions; need a large plasma to have magnetized ions (need ETPD?)

One idea: use beam populations to drive anisotropy



- 25 keV, IOA beam of H or He ions can be produced using existing beam
 - Beam β large enough to reach firehose threshold at low B (~100G) [following Chen, et al., Astrophys. J. Lett. 825 L26 (2016)]
 - Excitation of shear & fast Alfvén waves via Doppler-shifted IC resonance observed in low β plasmas

Ion Cyclotron Resonance Heating for generation of energetic ions



- High power (~200 kW) RF driver and fast wave antenna available.
- Initial experiments: good coupling (~30G wave amplitude), some evidence of perpendicular ion heating via fundamental minority resonance (H in He plasma). Can we generate an energetic tail, excite mirror instability?

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