

- Many-body rel. system. Can have $T \gg mc^2$
- $\alpha(t)$ vs $\Gamma(t)$

①

- freeze-out: out of equil. \rightarrow reaction stops

• In equil: $\mu_{\alpha_1} + \mu_{\alpha_2} + \dots = \mu_{\beta_1} + \mu_{\beta_2} + \dots$
 $\alpha_1 + \alpha_2 + \dots \rightarrow \beta_1 + \beta_2 + \dots$

$$\mu_\gamma = 0$$

$$\mu_{e^+} + \mu_{e^-} = 0 \Rightarrow \mu_{e^+} = -\mu_{e^-}$$

B, L_e, L_μ, L_τ, Q

$$dN_{i}^{F/B} = \frac{g_i}{(2\pi\hbar)^3} \frac{1}{e^{\frac{\epsilon - \mu}{kT}} \pm 1} d^3p d^3x$$

Consider ultrarel. particles with $\mu = 0$

Photons: (e.g. $\mu = 0$) $\epsilon = pc$ (Rel) $g = 2$
 Bosons

$$dN_B = \frac{2V}{(2\pi\hbar)^3} \frac{4\pi p^2 dp}{e^{\frac{pc}{kT}} - 1}$$

(2)

$$N_B = \frac{8\pi V}{(2\pi\hbar)^3} \int_0^\infty \frac{p^2 dp}{e^{\frac{pc}{kT}} - 1} \left(\frac{kT}{c}\right)^3$$

$$= \frac{8\pi V}{(2\pi\hbar)^3} \left(\frac{kT}{c}\right)^3 \int_0^\infty \frac{x^2 dx}{e^x - 1} = \frac{2(kT)^3 \zeta(3)}{\pi^2 \hbar^3 c^3}$$

||
2 ζ(3)

$$N_B = \frac{2 \zeta(3)}{\pi^2} \left(\frac{kT}{\hbar c}\right)^3 \sim T^3$$

Rel. fermions with g_i , $\epsilon = pc$

$$dN_f = \frac{g_i V}{(2\pi\hbar)^3} \frac{4\pi p^2 dp}{e^{\frac{\epsilon}{kT}} + 1}$$

|| $\frac{3}{2} \zeta(3)$

$$N_f = \frac{g_i V \cdot 4\pi}{(2\pi\hbar)^3} \int_0^\infty \frac{p^2 dp}{e^{pc/kT} + 1} = \frac{g_i V \cdot 4\pi}{(2\pi\hbar)^3} \left(\frac{kT}{c}\right)^3 \int_0^\infty \frac{x^2 dx}{e^x + 1}$$

(3)

$$N_f = \frac{g_i V}{4\pi^2} \left(\frac{kT}{hc} \right)^3 \zeta(3) \sim T^3$$

Moral: Number density

• rel. bosons: $N_B = \frac{g \zeta(3)}{\pi^2} \left(\frac{kT}{hc} \right)^3$

• rel fermions: $N_F = \frac{g \zeta(3)}{\pi^2} \left(\frac{kT}{hc} \right)^3 \cdot \frac{3}{4}$

Energy density: rel. fermi/bose gas

$$d\varepsilon_{B/F} = \frac{g_i}{(2\pi\hbar)^3} \frac{\varepsilon}{e^{\frac{\varepsilon-\mu}{kT}} \pm 1} d^3p d^3x$$

For $\varepsilon = pc \Rightarrow d^3p = \frac{4\pi \varepsilon^2 d\varepsilon}{c^3}$

let $\mu = 0$:

$$\varepsilon_{B/F} = \frac{g_i V}{(2\pi\hbar c)^3} \int_0^{\infty} \frac{\varepsilon^3 d\varepsilon}{e^{\varepsilon/kT} \pm 1}$$

$$\rho_{B/F} c^2 = \frac{\varepsilon_{B/F}}{V_B} = \frac{g_i \cancel{4\pi} (kT)^4}{(2\pi \hbar c)^3} \int_0^\infty \frac{x^3 dx}{e^x + 1} \quad (4)$$

$$\frac{\pi^4}{15} \quad (B)$$

$$\frac{7\pi^4}{120} \quad (F)$$

$$\rho_B c^2 = \frac{g_i \cancel{\pi^2} (kT)^4}{30 (\hbar c)^3} \quad (B)$$

$$\rho_F c^2 = \frac{g_i \pi^2 (kT)^4}{2 \cdot 120 (\hbar c)^3} = \frac{7}{8} g_i \frac{\pi^2 (kT)^4}{30 (\hbar c)^3} \quad (F)$$

g_i : spin d.o.f. number
 = 2 for photons;

~~$= 2SA$~~ for massive particles
 ~~$= 2SA$~~ (= 2 for electrons)
 ~~$= SA$~~ for neutrinos (weyl)

Moral: ultrarelat. gas ($\epsilon \approx pc$)

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$$pc^2 = g_{\text{eff}} \frac{\pi^2}{30} \frac{(kT)^4}{(\hbar c)^3}$$

$$g_{\text{eff}} = \sum_{\text{Bosons}} g_i + \frac{7}{8} \sum_{\text{Fermions}} g_i$$

$(mc^2 \ll kT) \qquad (mc^2 \ll kT)$

$$\text{EOS: } P_i = \frac{1}{3} \rho_i c^2 \Rightarrow (w = 1/3)$$

Non-rel. gas: $Mc^2 \gg kT$

$$n_i = g_i \left(\frac{m_i kT}{2\pi \hbar^2} \right)^{3/2} e^{-\frac{m_i c^2}{kT}}$$

$$p_i c^2 = M_i c^2 n_i$$

$$P_i = n_i kT \ll M_i c^2 n_i = p_i c^2 \Rightarrow P \approx 0$$

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$$CMB: kT_0 a_0 = kT_1 a_1$$

$$\frac{\lambda_{obs}}{\lambda_{emit}} = \frac{a(t_0)}{a(t_1)} \quad \lambda \nu = c$$

$$\frac{\nu_{emit}}{\nu_{obs}} = \frac{a_{obs}}{a_{emit}} \Rightarrow \nu_1 a_1 = \nu_2 a_2$$

$$\text{Planck: } e^{h\nu/kT} - 1 \rightarrow$$

$$e^{h\nu_1/kT_1} - 1 = e^{h\nu_2/kT_2} - 1$$

$$\text{form-invariant: } \boxed{T \cdot a = \text{const}}$$

$$1+z = \frac{1}{a}$$

Can we use thermal equilibrium?

(7)

Rate:

$$l = c^2$$

$$H^2 \sim G \rho$$

$$G \sim 1/M_P^2$$

$$\rho \sim T^4$$

$$\Rightarrow \frac{\dot{a}}{a} \sim \frac{T^2}{M_P} \sim \frac{1}{\text{time}} \sim \Gamma_{\text{exp}}$$

Rate of reactions: $\Gamma \sim \frac{1}{T^2}$

(chemical rate - different)

$$\sigma \sim \frac{\alpha_{\text{em}}^2}{T^2}$$

$$\alpha_{\text{em}} = e^2/4\pi\epsilon_0 \sim 1/137$$

$$n \sim T^3 \Rightarrow n\sigma \sim \alpha^2 T$$

$$\boxed{\Gamma_{\text{react}} \sim \alpha^2 T}$$

~ Equilibrium: $\Gamma_{\text{react}} \gg \Gamma_{\text{expansion}}$

$$\alpha^2 T \gg \frac{T^2}{M_P} \Rightarrow T \ll \alpha^2 M_P \sim 10^{14} \text{ GeV}$$

$$T \ll \alpha^2 M_p$$

$$\alpha^2 \sim 10^{-4} \quad M_p \sim 10^{19} \text{ GeV}$$

Non-rel. particles at equilibrium

Equil: $\mu_1 = \mu_2$ (chemical)

Suppose $m_1 \gtrsim m_2$ (but $m_1 \sim m_2$).

Then:

$$\frac{n_1}{n_2} = \frac{g_1}{g_2} \left(\frac{m_1}{m_2} \right)^{3/2} e^{\frac{(m_2 - m_1)c^2}{kT}} \rightarrow 0$$

as $T \rightarrow 0$

\Rightarrow heavier particles (e.g. neutrons vs p) disappear...

\Rightarrow not an equil.

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$$\frac{d \ln N_1}{d \ln a} = - \frac{\Gamma}{H} \left[1 - \left(\frac{N_1^{eq}}{N_1} \right)^2 \right]$$

$$N_1^{eq} = n_1^{eq} a^3$$

$$\frac{\Gamma}{H} \gg 1 \Rightarrow N_1 \sim N_1^{eq}$$

$$\frac{\Gamma}{H} \ll 1 \Rightarrow \frac{d \ln N_1}{d \ln a} \approx 0 \quad \text{Chemical freeze-out}$$

$N_1 \sim \text{const}$ (comoving), i.e.

$n_1 \sim N_1 / a^3$ with time.

at $\Gamma \sim H$

Abundances of elements i can be measured.