Cold holographic matter and a color group-breaking instability

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- \Rightarrow Take (super)Yang–Mills theories at strong coupling and large N
- \Rightarrow Add fundamental matter (quarks) at finite charge density
- \Rightarrow Then...

...the simplest^(*) ground states are described in the IR by a non-relativistic field theory with specific scaling properties...

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Table of Contents

Motivation

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A cartoon of the QCD phase diagram



A cartoon of the QCD phase diagram



Interest and use of holography

- ⇒ Why care? Color-superconductivity phases and transitions; neutron stars...
- $\Rightarrow\,$ Region of strong coupling suggests a holographic approach, but no QCD dual, we take $\mathcal{N}=4SYM$ as ballpark
- ⇒ Try to extract qualitative lessons of the effects of the chemical potential in strongly coupled systems
 - What is the equation of state?
 - How to observe color superconductivity?

Strings ho!

I describe results from top-down models, where we extremize type II SUGRA, DBI+WZ and NG actions.

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- ⇒ Take the geometry sourced by a set of Dp-branes on (flat) spacetime
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- ⇒ Take the geometry sourced by a set of Dp-branes on (flat) spacetime
- \Rightarrow Backreact Dq-branes with gauge field on the w.v. turned on
- \Rightarrow Then...

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- ⇒ Take the geometry sourced by a set of Dp-branes on (flat) spacetime
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...the simplest^(*) supergravity solutions at T = 0 are described in the IR by a Hyperscaling-Violating Lifshitz spacetime...

...and they seem to be unstable towards Higgs-branch phases

In this talk I will have the gauge/gravity duality in mind

(non-)AdS spacetime \iff (non-)conformal theory

 $\mathsf{Domain} \ \mathsf{wall} \ \mathsf{solution} \ \ \longleftrightarrow \ \ \mathsf{Renormalization} \ \mathsf{group} \ \mathsf{flow}$

Boundary $\leftrightarrow \to UV$ (high energies)

Origin \leftrightarrow IR (low energies)

System action

$$S = \frac{1}{2\kappa^2} \int e^{-2\phi} \left(R * 1 - \frac{1}{2}H \wedge *H - \frac{1}{2}d\phi \wedge *d\phi \right) - \frac{1}{2\kappa^2} \int \frac{1}{2}F_1 \wedge *F_1 + \frac{1}{2}F_3 \wedge *F_3 + \frac{1}{4}F_5 \wedge *F_5 - \frac{1}{2\kappa^2} \int \frac{1}{2}C_4 \wedge H \wedge F_3 - N_f T_7 \int d^8x \, e^{-\phi} \sqrt{-|\hat{G} + dA + \hat{B}|} + N_f T_7 \int e^{dA + \hat{B}} \hat{C}_q$$

Table of Contents

Motivation

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SYM theories from type II SUGRA



SU(N) SYM with 16 supercharges

$$ds^{2} = h^{-1/2} dx_{1,p}^{2} + h^{1/2} \left(dr^{2} + r^{2} d\Omega_{8-p}^{2} \right)$$
$$\int_{S^{8-p}} *F_{8-p} \sim N$$

SYM theories from type II SUGRA [Itzhaki et al '98]

 \Rightarrow For D3-branes one has $\mathcal{N}=4$ SYM, conformal

 \Rightarrow For D2-branes λ is dimensionful and there is a running



 \Rightarrow The holographic radius is related to the energy scale $r = E \ell_S^2$.

SYM theories from type II SUGRA



SYM with other gauge group and less supercharges

$$ds^{2} = h^{-1/2} dx_{1,p}^{2} + h^{1/2} \left(dr^{2} + r^{2} d\Sigma_{8-p}^{2} \right)$$
$$\int_{M_{8-p}} *F_{8-p} \sim N$$

SYM theories from type II SUGRA

To preserve $\mathcal{N}=1$ supersymmetry the internal manifold must admit one Killing spinor. This constrains the possible choices. From now on

	dim	base	cone
D3-branes	5	Sasaki-Einstein	Calabi-Yau
D2-branes	6	nearly Kähler	G ₂ -cone

Table of Contents

Motivation

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Fundamental matter [Karch and Katz '03]

	x^{μ}	r	$\theta^{1,2,3}$	$\theta^{4,\cdots,8-p}$
Dp	×	_	_	_
D(p+4)	×	×	×	_

 \Rightarrow The action as sum of two parts: type II SUGRA and

$$S_{flavor} = -T_{p+4} \int d^{p+1} x \, d^4 y \, e^{-\phi} \, \sqrt{\hat{G}} + T_{p+4} \int \hat{C}_{5+p}$$

Backreaction of the flavor branes

 \Rightarrow Consider the RR form the flavor brane sources

$$S_{IIB+D7} \supset \frac{1}{2} \int \mathrm{d}C_8 \wedge * \mathrm{d}C_8 + \int C_8 \wedge \underbrace{(\delta(f_1)\delta(f_2)\mathrm{d}f_1 \wedge \mathrm{d}f_2)}_{\Xi_2}$$

which implies the Bianchi identity for a sourced RR form

$$\mathrm{d}F_1 = -\Xi_2$$

 $\Rightarrow\,$ The number of flavor branes is given by Gauss law

$$\int F_1 \sim N_f$$

Backreaction with smearing (D3/D7) [Benini et al '06]

- \Rightarrow Two things I have shown in previous slides
 - 1. For D3-branes the compact manifold is a SE
 - 2. $S_{D7} = T_7 \int \left(-d^8 x \, e^{-\phi} \sqrt{-G} + C_8 \right) \wedge \Xi_2$ with Ξ_2 exact
- \Rightarrow SE manifolds can be expressed as U(1) fibrations over KE manifolds and are equiped with an SU(2)-structure

$$\mathrm{d}\eta_{\mathsf{K}\mathsf{E}} = 2J_{\mathsf{K}\mathsf{E}} \;, \qquad \mathsf{vol}(\mathsf{S}\mathsf{E}) = \frac{1}{2}J_{\mathsf{K}\mathsf{E}} \wedge J_{\mathsf{K}\mathsf{E}} \wedge \eta_{\mathsf{K}\mathsf{E}}$$

⇒ Idea: to identify $\Xi_2 \sim J_{KE}$ and use the SU(2)-structure to write a *consistent radial ansatz* for the IIB+sources action

$$F_1 \sim N_f \eta_{KE} \quad \Rightarrow \quad \mathrm{d}F_1 \sim N_f J_{KE}$$

A taste of the smeared solution (D3/D7) [Benini et al '06]

 \Rightarrow With a simple ansatz

$$\mathrm{d}s^2 = g_1(r)\,\mathrm{d}x_{1,3}^2 + g_2(r)\,\mathrm{d}r^2 + g_3(r)\,\mathrm{d}s_{KE}^2 + g_4(r)\,\eta_{KE}^2 \ ,$$

with dilaton and RR forms

$$F_5 \sim N (1+*) J_{KE} \wedge J_{KE} \wedge \eta_{KE} , \qquad F_1 \sim N_f \eta_{KE} ,$$

 \Rightarrow A SUSY solution exists

$$\phi' = N_f e^{\phi} \quad \Rightarrow \quad e^{\phi} = \frac{1}{N_f (r_{LP} - r)}$$

When is backreaction needed? (D3/D7)

- ⇒ When can we omit backreaction (probe approximation) and when is it necessary?
- \Rightarrow Compare energies (effect on metric)

$$\frac{|F_1|}{|F_5|} \sim \lambda \frac{N_f}{N}$$

and one concludes (wrongly) that for $\lambda \frac{N_f}{N} \ll 1$ probe approx. is enough at all scales.

Backreaction and smearing (D2/D6) [Faedo et al '15]

For the D2/D6 case a similar situation holds

- ⇒ Start with a NK (6d) manifold and $S_{D6} = T_6 \int (-d^7 x e^{-\phi} \sqrt{-G} + C_7) \wedge \Xi_3$ with Ξ_3 exact
- \Rightarrow There is a SU(3)-structure

$$\mathrm{d}\mathcal{J} = 3\mathrm{Im}\Omega \;, \quad \mathrm{d}\mathsf{Re}\Omega = 2\mathcal{J}\wedge\mathcal{J} \;, \quad \textit{vol}(\mathit{NK}) = \frac{1}{6}\mathcal{J}^3$$

 \Rightarrow In this case

$$F_2 \sim N_f \mathcal{J} \quad \Rightarrow \quad F_2 \sim N_f \operatorname{Im} \Omega \sim \Xi_3$$

When is backreaction needed? (D2/D6)

 \Rightarrow Backreaction matters in the IR

$$\frac{|F_2|}{|F_6|} \sim \frac{\lambda \frac{N_f}{N}}{E} \equiv \frac{E_{flavor}}{E}$$

and there is a change in the dynamics of the theory with a crossover at $E \sim E_{\it flavor}.$

 \Rightarrow UV boundary conditions

A taste of the smeared solution (D2/D6) [Faedo et al '15] The D2/D6 solution



Including charge in the setup

⇒ My motivation was to add a 'quark' density to these setups, but keeping vanishing temperature.

 \Rightarrow Dissolved strings in the flavor branes

U(1) global current $\stackrel{\text{dual}}{\longleftrightarrow}$ U(1) gauge field

 \Rightarrow In particular a charge density corresponds to $A = A_t(r) dt$

$$S_{D7} = -T_7 \int d^8 x \, e^{-\phi} \sqrt{-|G+F+B|} + T_7 \int C_8 - C_6 \wedge (F+B) + \cdots$$

When does the charge density matter [Chen et al '09] [Bigazzi et al '11]

 \Rightarrow Take the e.o.m. for the NS form (and set B = 0)

$$\mathrm{d}(e^{-2\phi}*H) = 0 = F_3 \wedge F_5 + F_1 \wedge *F_3 + (DBI) * \mathrm{d}t \wedge \mathrm{d}r$$

 \Rightarrow From the second term in RHS we deduce a component

$$F_3 \supset \mathcal{C}'(r) \mathrm{d}t \wedge \mathrm{d}r \wedge \eta_{KE} + \cdots$$

⇒ From the first term in RHS we deduce a constant (density) term

$$F_3 \supset N_q \,\mathrm{d} x^1 \wedge \mathrm{d} x^2 \wedge \mathrm{d} x^3$$

When does the charge density matter [Faedo et al '14]

⇒ Same game as before: When is the effect of charge comparable to the effect of color physics?

$$\frac{|F_3|}{|F_5|} = \left(\frac{\lambda^{2/3} (N_q/N^2)^{1/3}}{E}\right)^3 \equiv \left(\frac{E_{charge}}{E}\right)^3$$

so charge becomes important in IR for $E < E_{charge}$.

 \Rightarrow Similarly, we can show from $|F_1|/|F_3|$ that the charge dominates *always* in the IR.

When does the charge density matter [Faedo et al '14]

 \Rightarrow The same argument also works for the D2/D6 system

$$\frac{|F_2|}{|F_6|} = \left(\frac{\lambda^{1/2} (N_q/N^2)^{1/4}}{E}\right)^4 \equiv \left(\frac{E_{charge}}{E}\right)^4$$

so charge becomes important in IR for $E < E_{charge}$.

⇒ Similarly, we can show from $|F_2^{flavor}|/|F_2^{charge}|$ that the charge dominates *always* in the IR.

How does the charge density affect dynamics [Kumar '12] [Faedo et al '14]

- ⇒ To see charge effects we can take a limit in which we discard distracting flavor effects.
- ⇒ The IR turns out to be a non-relativistic theory (Lifshitz HV metric)

$$t \to \xi^z t \;, \qquad \vec{x} \to \xi \vec{x}$$

with hyperscaling-violation

$$\mathrm{d}s^2 \to \xi^{\frac{\theta}{d-1}} \mathrm{d}s^2 \quad \Rightarrow \quad F \sim T^{\frac{d+z-1-\theta}{z}}$$

$$\Rightarrow$$
 For $p = 3$ one gets $z = 7$, $\theta = 0$

$$\Rightarrow$$
 For $p = 2$ instead $z = 5$, $\theta = 1$

3d SYM theory with quark density [Faedo et al '15]





3d SYM theory with quark density [Faedo et al '15]



4d SYM theory with quark density work in progress

- \Rightarrow In 4d SYM there is only one *classical* scale
- \Rightarrow Presence of the Landau pole complicates numerical analysis
- ⇒ However the IR analysis showing the existence of a Lifshitz solution in the IR still holds

Table of Contents

Motivation

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Potential instabilities

If the IR we have proposed is realized one has to wonder about its stability

 \Rightarrow Thermodynamically it is stable

$$\lim_{T\to 0} S = 0 \text{ and } T \frac{\partial S}{\partial T} > 0$$

- ⇒ One possible instability would be towards striped (or other inhomogenous) phase
- \Rightarrow Or maybe there is a dynamic instability of some field that wants to condense. The BF bound in Lifshitz is

$$m^2 \ge -\frac{(p+z-\theta)^2}{4} = -25$$

An instability of the solution

 \Rightarrow In particular, take the U(1) BI vector field with one component in the internal directions

$$A_{\mu} \mathrm{d} x^{\mu} \supset \Psi(r) \eta_{KE}$$

 \Rightarrow In $\mathcal{N} = 4$ SYM the scalar is dual to

 ${\cal O}^{\prime} \sim {\cal Q}^{\dagger} \, \sigma^{\prime} \, {\cal Q}$

with mass squared on the BF bound

⇒ In Lifshitz spacetime this scalar has mass below the BF bound, so we expect it to condense if the Lifshitz region is large

An instability of the solution

⇒ Backreaction of the mode in the supergravity fields affects the Gauss law for the D3-branes

$$\int_{S^5} F_5 \sim N + \frac{1}{2} N_f \Psi(r)^2$$

the color branes are separated: U(N). Since color symmetry is broken we have superconductivity

 \Rightarrow In fact this is the only field in the smeared setup with this property: any color superconductor in our setup must have non vanishing Ψ

Table of Contents

Motivation

A quick review of strongly coupled SYM theories

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Looking for an instability

Conclusions

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- \Rightarrow We have identified a IR phase of cold YM theories with charge density, given in the gravity side by a HV-Lifshitz metric
- \Rightarrow We have worked out the numerical solution for 3d SYM, and this is work in progress for the 4d version.
- ⇒ The IR appears to be unstable towards condensation of $\mathcal{O}^{\prime} \sim Q^{\dagger} \sigma^{\prime} Q$.
- ⇒ Condensation of the dual scalar field gives rise to a color superconductor phase with order parameter given by $\langle O^I \rangle$.

Thank you