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# Holographic superconductors and spatial modulation

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work with A.Donos and J.P.Gauntlett, 1310.5741[hep-th]

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# Introduction: AdS/CMT

Use the gauge/gravity correspondence as a tool to investigate the dynamics of strongly coupled CFTs at finite temperature and charge density and/or placed in an external magnetic field.

- Attempt to understand universal features of strongly coupled condensed matter systems found in the vicinity of 'quantum critical points'.
- Explore black hole physics: construct and study novel charged black hole solutions that asymptote to AdS.

# Introduction: Top-Down Vs Bottom-Up

#### Top-Down Approach:

- Consider theories obtained by consistent truncations of the D=10,11 supergravities.
- Difficult to obtain CFTs of interest; involved calculations.
- CFT guaranteed to be well defined.

Bottom-Up Approach:

- Consider phenomenological gravity theories with only few additional d.o.f. that are dual to CFTs with the desirable features.
- No guarantee to have a string embedding.
- Simple calculations.

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#### Introduction: Questions

Specific questions that can be addressed:

- What type of phases are possible and what are the transport properties of each phase?
- What kind of ground states are possible? classification of IR geometries, investigation of new emergent IR scaling behaviours?
- How do these phases compete? What is the dual phase diagram?

# Holographic superconductors

High  $T_c$ -superconductors consitute one the most challenging problems in condensed matter physics. Remain mysterious due to their strongly coulped nature.

Minimal ingredients:

- finite temperature ightarrow black holes in the bulk
- finite charge density ightarrow U(1) gauge field
- an "order parameter" that spontaneously acquires an expectation value, e.g. a charged scalar for s-wave superconductors.

Superconducting instabilities: As the temperature is reduced, a new branch of black holes supported by non-vanishing hair emerges at some critical temperature. The U(1) symmetry is now spontaneously broken.

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#### Spatially modulated phases

In condensed matter, phases with *spontaneously* broken translational invariance are very common. Realised in various configurations, e.g. stripes, checkerboards and helices.

- Spin Density Wave
- Charge Density Wave
- Current density wave



The modulation is fixed by an order parameter with non-zero momentum and it's not related to the underlying lattice of the material.

Holographically, SM phases dual to BHs with broken translational invariance.

Introduction

Conclusions

Key point to get SM phases = mixing modes

- Consider a perturbations of a charged scalar in  $AdS_2 \times R^2$ . The e.o.m. becomes  $\Box_{Ads2}\phi M^2\phi = 0$ , where  $M^2 = m^2 ce^2 + k^2$ . Unstable region is centered around k = 0.
- Finding the temprature at which the instability sets in, one gets a "bell curve" with the max. at *k* = 0.
- Mixing of modes would introduce off-diagonal terms in the mass matrix that drive the most unstable mode off k = 0, shifting of the "bell curve" to k ≠ 0.



Examples, at finite charge density:

- in D=5[Nakamura,Ooguri,Park] and in D=4 [Donos,Gauntlett] with the mixing introduced by CS terms and axions respectively.
- PT does not necessarily need to be broken [Donos,Gauntlett]. Examples, in magnetic field:
  - in D=4,5 with mixing term  $\phi * F \land G$  [Donos,Gauntlett,CP]
  - in  $U(1)^3$  and  $U(1)^4$  sugra with metric components mixing as well [Donos,Gauntlett,CP]. Interesting interplay with susy solutions.
- $\rightarrow$  Spatially modulated phases are more the rule, than the exception.

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#### Plan for this talk

#### Question:

Is it possible to have spatially modulates superconducting states in holography?[Donos,Gauntlett]

- The FFLO state describes an s-wave superconductor in which the cooper pair has non-vanishing momentum.
- This was conjectured to exist in the '60s and possibly has been seen experimentally in heavy fermion materials and some organic superconductors, eg CeColn<sub>5</sub>.

#### The model of interest

Aim: Study p-wave superconductors in D=5. The order parameter can be either an SU(2) vector or a two-form. Both cases give similar results; we focus on the second case.

Consider a theory of gravity in D=5 coupled to a U(1) gauge field and a complex two-form

$$\mathcal{L} = (R+12) * 1 - \frac{1}{2} * F \wedge F - \frac{1}{2} * C \wedge \bar{C} - \frac{i}{2m}C \wedge \bar{H}$$

where F = dA and  $H = dC + ieA \wedge C$ .

• For particular values of (e, m), this theory can be obtained as a consistent truncation from D = 10, 11; corresponds to a subsector of the D = 5 Romans theory.

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- This theory admits a unit radius  $AdS_5$  vacuum solution with A = C = 0 which is dual to a d=4 CFT with a conserved U(1) current and a tensor operator with charge, e, and scaling dimension  $\Delta = 2 + m$ .
- Another solution of the theory is the electrically charged AdS-RN black hole.

$$ds^2 = -gdt^2 + g^{-1}dr^2 + r^2(dx_i^2), \qquad A = \mu(1 - \frac{r_+^2}{r_-^2})dt.$$

This corresponds to the high temperature, spatially homogeneous and isotropic phase of the dual CFTs when held at finite chemical potential  $\mu$ .

# Step 1: Instabilities in the NHL

Study perturbations around the near-horizon limit of the electric AdS-RN:  $AdS_2 \times R^{D-2}$  .

- Modes tend to be more unstable in this region.
- Use the  $AdS_d$  BF bound criterion to check for instabilities: if the bound is violated, the theory is unstable (converse not always true)

$$L^2 M^2 \geq -\frac{(d-1)^2}{4}$$

#### In our model:

In the NHL, peturbations of the two-form,  $\delta C$ , decouple; consider the ansatz [Donos,Gauntlett]

$$\delta C = \cdots + dx_1 \wedge (u_3 dx^3 + v_3 dx^2),$$

where 
$$u_3 = d_3 cos(kx_1), v_3 = d_3 sin(kx_1) \rightarrow p$$
-wave  
or  $u_3 = d_3 e^{ikx_1}, v_3 = id_3 e^{ikx_1} \rightarrow p$ +ip-wave.

• For both cases, if  $e^2 > \frac{m^2}{2}$ , the BF bound is violated for a certain k. Most unstable mode has  $k \neq 0$ ; when heated up, we expect the preferred branch to be modulated.

**p-wave:**  $\delta C = \cdots + d_3(r) dx_1 \wedge [\sin(kx_1) dx_2 + \cos(kx_1) dx_3]$ 

k = 0: order parameter pointing in  $-dx_2$ 3 translations and rotations in  $(x_1, x_3)$ .  $k \neq 0$ : dualise to see the helical structure, pitch is  $2\pi/k$ .  $x_2, x_3$  translations,  $x_1$  translation combined with a rotation,  $(x_2, x_3)$  rotation (Bianchi VII<sub>0</sub>)



#### (p+ip)-wave: $\delta C = \cdots + e^{-ikx_1}id_3(r)dx_1 \wedge (dx_2 - idx_3)$

k = 0: order parameter pointing in  $dx_2 + idx_3$ 3 translations and  $(x_2, x_3)$  rotations upto const gauge transf.  $k \neq 0$ : same as before, but  $x_1$  translations are compensated by a gauge transf.

No helical structure, no symmetry reduction.

# Step 2: Perturbations around AdS<sub>5</sub>-RN

Consider linearised perturbations around the full AdS-RN black hole. [Donos,Gauntlett]

- Specify the critical temperature at which the instability sets in.
- Allows to search for instabilities localised far from the horizon.

#### In our model:

- Two-form perturbations,  $\delta C$ , decouple again. Consider the same ansatz as before.
- All the action is included in the last term ~ d<sub>3</sub>(r): regular at the horizon and spontaneously breaking the U(1).

$$d_3 = d_{3+} + \mathcal{O}(r - r_+), \qquad d_3 = c_{d_3}r^{-|m|} + \cdots$$

• Obtain an one-parameter family of solutions as expected. Plot the critical temperatures  $T_c$  versus k for the existence of normalisable static perturbations of the two-form for fixed (m, e).



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- For fixed *m*, *T<sub>c</sub>* increases as *e* increases. For fixed *e*, *T<sub>c</sub>* decreases as *m* increases.
- Depending on (e, m) this plot may not cross the k = 0 axis.
- p- and (p+ip)-wave set in at the same temperature.

# Step 3: Backreacted solutions

Construct the backreacted BHs to get information about the thermodynamics: is the new branch of black holes preferred?

- In principle, one needs to solve PDEs to study spatial modulation.
- Here, we can get away with solving ODEs: the three-dimensional Euclidean group breaks down to Bianchi VII<sub>0</sub>.
- Use the left-invariant one-form of this Lie algebra when constructing the ansatz.

$$\begin{split} \omega_1 &= dx_1 ,\\ \omega_2 &= \cos(kx_1) \ dx_2 - \sin(kx_1) \ dx_3 ,\\ \omega_3 &= \sin(kx_1) \ dx_2 + \cos(kx_1) \ dx_3 . \end{split}$$

#### p-wave helical superconductors

Consider the following ansatz:

$$ds^{2} = -g f^{2} dt^{2} + g^{-1} dr^{2} + h^{2} \omega_{1}^{2} + r^{2} \left( e^{2\alpha} \omega_{2}^{2} + e^{-2\alpha} \omega_{3}^{2} \right) ,$$
  

$$A = a dt ,$$
  

$$C = (i c_{1} dt + c_{2} dr) \wedge \omega_{2} + c_{3} \omega_{1} \wedge \omega_{3} ,$$

- E.o.m. boil down to a set of ODEs which is solved subject to boundary conditions: regularity at the horizon and  $AdS_5$  asymptotics compatible with spontaneous symmetry breaking no sources.
- Obtain a two parameter family of solutions, labeled by (k, T), consistent with the bell curves of step 2.

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- All the solutions have a lower free energy than AdS-RN: the CFT undergoes a phase transition to a helical superconducting phase.
- The preferred ones lie along the red locus: as the temperature is lowered, the pitch is increasing. (e, m) = (2, 2)



- Interestingly, preferred solutions satisfy c<sub>h</sub> = 0. This is understood by varying the free energy with respect to k: k∂<sub>k</sub>w = 8c<sub>h</sub> = 0.
- Plot the condensate along the preferred branch. The phase transition is second order: near  $T_c$ , we have a mean field behaviour.



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- On the k = 0 branch, c<sub>h</sub> = c<sub>α</sub> = 0, signaling the existence of a reduction of the ansatz: h = re<sup>-α</sup>. Consequence of enhanced symmetry.
- Boundary stress tensor is inhomogeneous and exhibits anisotropic nature of p-wave superconductors: (also, traceless and conserved)

• For large enough *e*, the preferred locus exhibit inversion of the pitch of the helix - this point is completely regular. This phenomenon was seen experimentally in e.g. helimagnets.



Solutions with k = 0 are only relevant at the inversion point.
 Finite k solutions can not be ignored.

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#### Ground States

#### Scaling solutions:

The system admits the following scaling solutions in the IR:

- If k = 0, the IR fixed point is invariant under the anisotropic scaling: similar to [Taylor]  $r \to \lambda^{-1}r$ ,  $t \to \lambda^{z}t$ ,  $x_{1,3} \to \lambda^{1-\gamma}x_{1,3}$ ,  $x_{2} \to \lambda^{1+\gamma}x_{2}$ . The entropy density scales like  $S \sim T^{(3-\gamma)/z}$ .
- If k ≠ 0, there is a helical scaling symmetry: similar to [Kachru et al.]
  - $r \to \lambda^{-1}r, \quad t \to \lambda^{z}t, \quad x_{2,3} \to \lambda x_{2,3}, \quad x_1 \to x_1.$ The entropy density scales like  $S \sim T^{2/z}$ .

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- If  $k \neq 0$ , possible to have the same  $AdS_5$  in the IR as in the UV.
  - [Horowitz,Roberts]: s-wave superconductor. In the condensed phase, the gauge field is transformed to an irrelevant operator. -In our case, the condensate vanishes at T = 0 and we are left only with relevant modes. Puzzling!
  - [Sachdev et al.]: introducing a lattice. Exploid the k-dependence of the relevant modes.
  - k-dependence allows us to suppress relevant modes, e.g.

$$c_3(r) = c_3^0 \frac{e^{-kx/r}}{r^{1/2}} + \cdots$$

#### Black Holes at T=0:

The T=0 limit of our black holes approaches smooth domain walls that interpolate between  $AdS_5$  in the UV and the following IR:



#### Question:

Both, the  $AdS_5$  and the helical scaling, exist at  $k \neq 0$  but emerge at k < 0 and k > 0 respectively. Is this related to thermodynamics or is there a problem with constructing the domain walls?

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# (p+ip)-wave superconductors

A very similar story for the (p+ip)-case:

$$ds^{2} = -gf^{2}dt^{2} + g^{-1}dr^{2} + h^{2}(dx_{1} + Qdt)^{2} + r^{2}(dx_{2}^{2} + dx_{3}^{2}),$$
  

$$A = adt + wdx_{1},$$
  

$$C = (ic_{1}dt + c_{2}dr + ic_{3}dx_{1}) \wedge (\omega_{2} - i\omega_{3}),$$

- The ansatz is now stationary, but not static.
- By solving a set of seven coupled ODEs subject to boundary conditions, we obtain a 2-parameter family of (p+ip)-wave superconductors. Label solutions by (k,T) as before.

- These solutions have lower free energy than the AdS-RN as well; p-wave and (p+ip)-wave phases compete.
- Apart from the standard smarr-type formula, all the solutions satisfy two additional contrains on the UV data [Donos,Gauntlett].

$$4c_h - \frac{k}{e}c_w = 0$$
$$2c_Q + c_w\mu = 0$$

- The preferred locus of solutions satisfy  $c_w = 0 \Rightarrow c_h = c_Q = 0$ . Also understood by varying the action with respect to k.
- On the k = 0 branch,  $c_h = 0$ , but  $c_w, c_Q \neq 0$ : no symmetry enhancement.

• Boundary stress tensor is homogeneous and exhibits the isotropic nature of (p+ip)-wave superconductor:

• For large *e*, *k* becomes slightly negative as the temperature is lowered.

- Have not yet been able to pin down the precise behaviour of the solution when  $T \rightarrow 0$ ; scenarion of  $AdS_5$  in the IR.
- The entropy of the ground states goes to zero. For small values of the charge, there is an interesting crossover in its behaviour.



#### Competition between p and p+ip

Both instabilities set in at the same  $T_c$ . Which is preferred?



- p-wave is preferred for small e and (p+ip)-wave for large e.
- For intermediate values of *e*, there is a first order transition between the two at *T*<sub>\*</sub>; the black holes have the same free energy, but do not intersect on field space.
   c.f. [Gubser,Pufu]

#### Further remarks

- The discussion of homogeneous p-wave superconductors has been generalised to include spatial modulation.
- Can you find two forms with *e*,*m* in top down setting, perhaps with addition of extra fields, that are unstable?
- Spatially modulated phases are more fundamental than expected.
- The field is moving towards solving PDEs [Donos],[Withers], [Rozali et al.]

Introduction 000000 0 Main Part 0000000 00000000 Conclusions

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# Thank you!