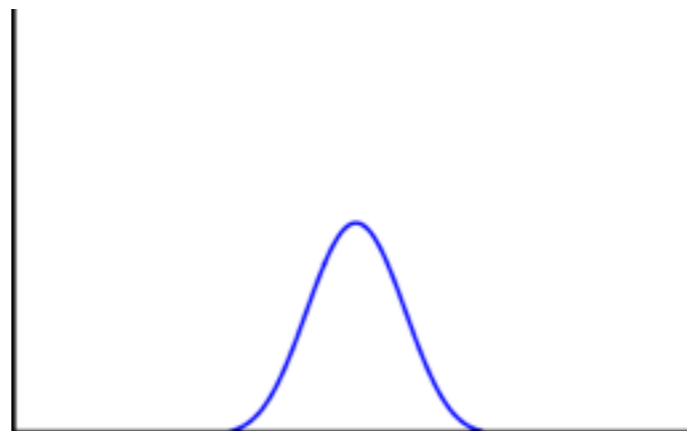


Turbulent strings in AdS/CFT

Takaaki Ishii
(University of Crete)

arXiv:1504.02190 with Keiju Murata



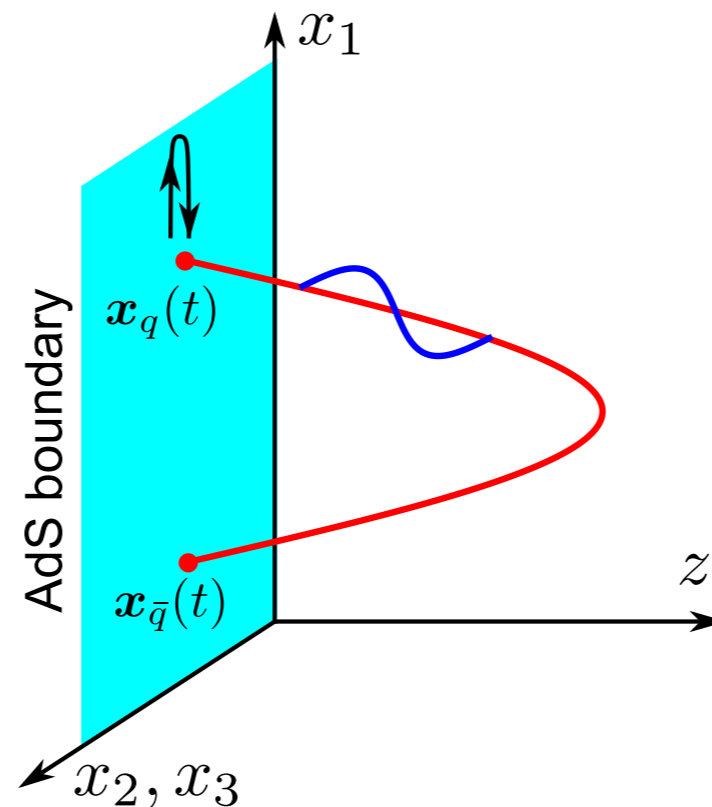
12 May 2015@Oxford

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What I will do

Perturb holographic quark-antiquark potential



We solve nonlinear time evolution

Motivation

I think time-dependent dynamics in gauge/gravity duality is interesting

- AdS thermalization

Relation to real QGP? New BH dynamics?

- AdS turbulent instability

What is essential? AdS? Einstein? Nonlinearity?

- c.f.) Dynamical meson melting

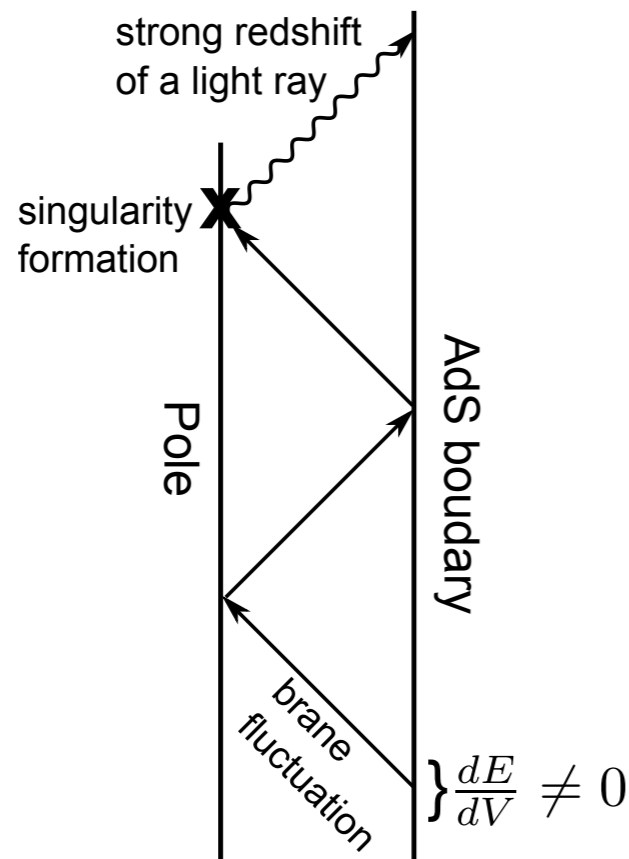
Time evolution in thermalizing D3/D7

[TI-Kinoshita-Murata-Tanahashi]

Turbulent instability in D3/D7

Singularity formation after some wave reflections

[Hashimoto-Kinoshita-Oka-Murata]



- Electric field quench: $0 \rightarrow E$
- “Meson turbulence”
- Probably due to nonlinearity in DBI

Considering F1 would be simpler

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Time-like holographic Wilson loop

[Maldacena, Rey-Yee]

$$\text{AdS}_5 \times \text{S}^5 \quad ds^2 = \frac{\ell^2}{z^2} (-dt^2 + dz^2 + d\mathbf{x}^2) + \ell^2 d\Omega_5^2$$

Static gauge: $(\tau, \sigma) = (t, z)$

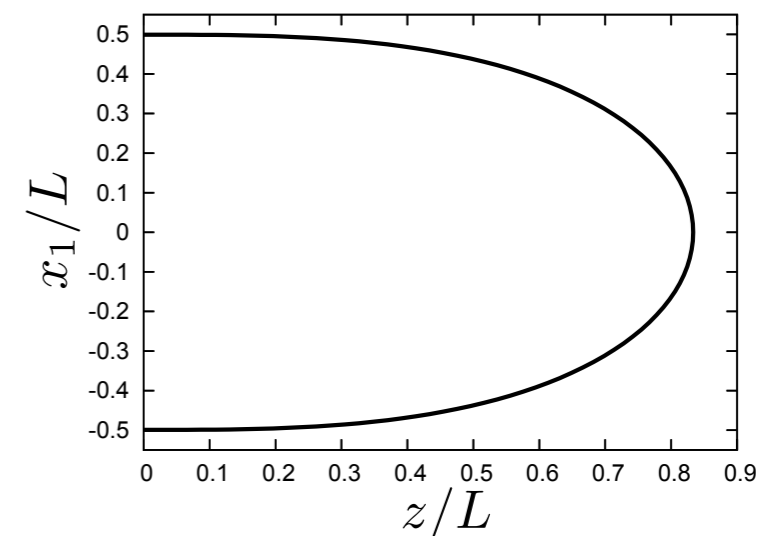
Target space embedding: $x_1 = X_1(z)$

Solution for separation L

$$\begin{aligned} X_1(z) &= \pm z_0 \int_{z/z_0}^1 dw \frac{w^2}{\sqrt{1-w^4}} \\ &= \pm z_0 \left[\Gamma_0 + F(z/z_0; i) - E(z/z_0; i) \right] \end{aligned}$$

\swarrow
 z_0 : string tip
 $\Gamma_0 = 0.599$

$$\frac{L}{2} = z_0 \Gamma_0$$



A convenient parametrization

Inverse function of $F(z;k)$ is $\text{sn}(x;k)$

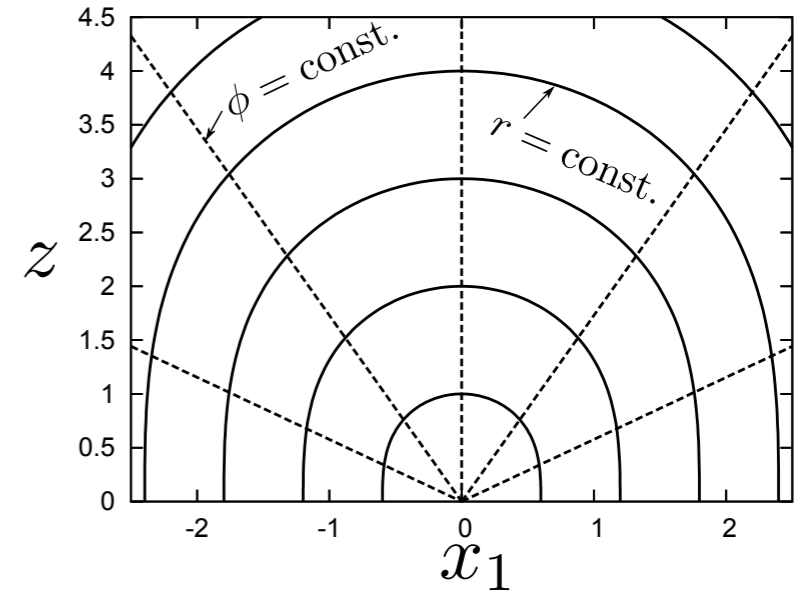
$$F(\text{sn}(x; k); k) = x$$

Polar-like coordinates (r, ϕ)
where the static solution is $r=z_0$

$$z = r f(\phi) = r \text{sn}(\phi; i)$$

$$x_1 = r g(\phi) = r \begin{cases} \phi - E(\text{sn}(\phi; i); i) + \Gamma_0 & (\phi \leq \beta_0/2) \\ \phi + E(\text{sn}(\phi; i); i) - \Gamma_0 - \beta_0 & (\phi > \beta_0/2) \end{cases}$$

A nice identity $f'(\phi)^2 + g'(\phi)^2 = 1$



Linearized perturbations

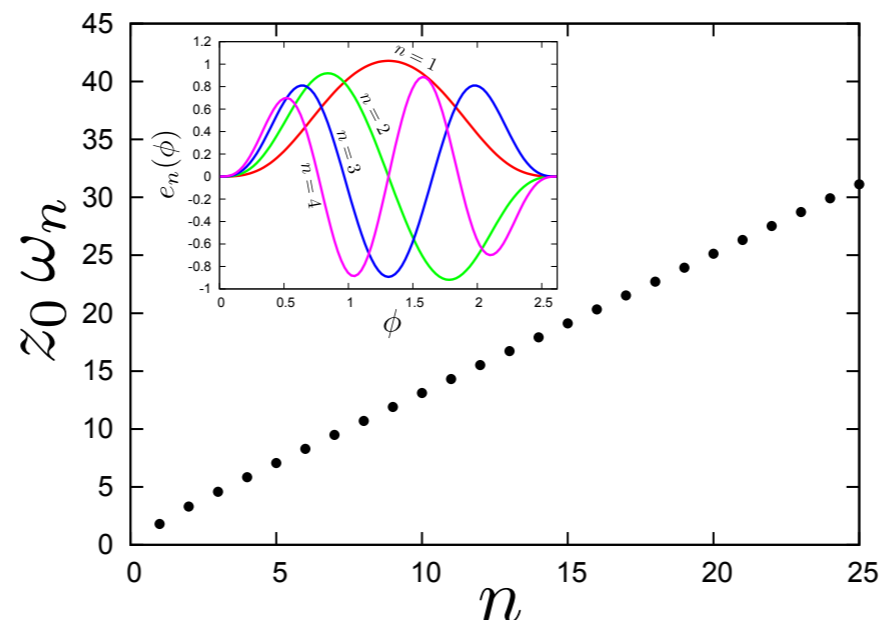
Longitudinal fluctuations around $r=z_0$

$$r = z_0[1 + \chi_1(t, \phi)]$$

[Callan-Guijosa, Klebanov-Maldacena-Thorn]

Linearized EoM for eigenvalues/functions

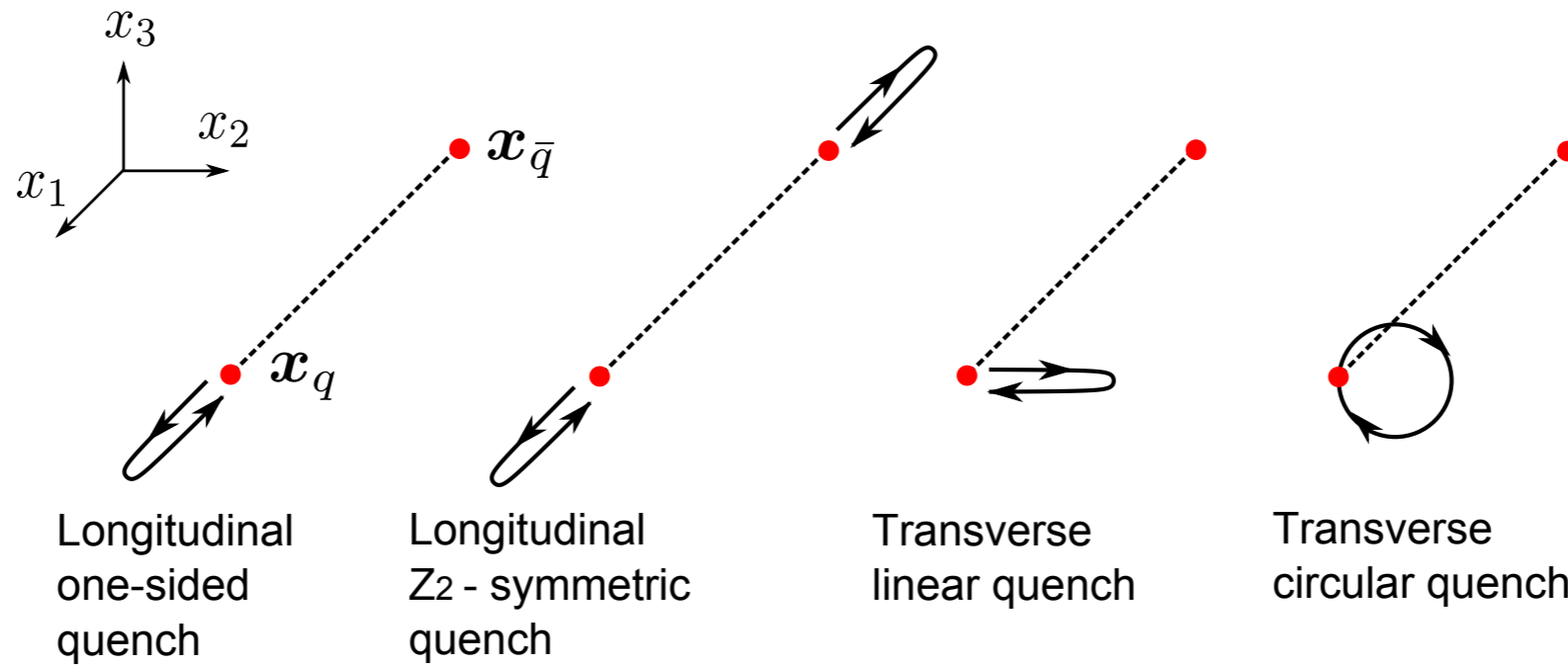
$$(\partial_t^2 + \mathcal{H})\chi_1 = 0 \quad \mathcal{H} \equiv -\frac{1}{z_0^2 h} \partial_\phi h \partial_\phi \quad h \equiv ((g/f)' f)^2$$



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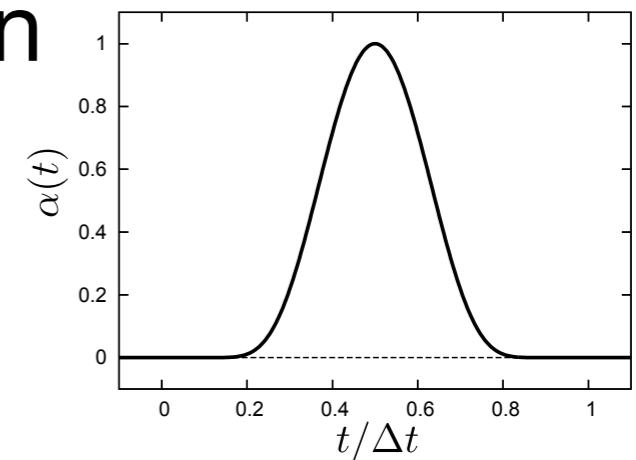
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Perturb the string endpoints



Quench profile: a compact C^∞ function

$$\alpha(t) = \exp \left[2 \left(\frac{\Delta t}{t - \Delta t} - \frac{\Delta t}{t} + 4 \right) \right] \quad (0 < t < \Delta t)$$



Worksheet double null coordinates

Induced metric $ds_{F_1}^2 = -2\gamma_{uv}dudv$

Worksheet: u, v

Target space: $T(u, v), Z(u, v), X_{1,2,3}(u, v)$

$$\gamma_{uv} = \frac{\ell^2}{Z^2} (-T_{,u}T_{,v} + Z_{,u}Z_{,v} + \mathbf{X}_{,u} \cdot \mathbf{X}_{,v})$$

Equations of motion

$$T_{,uv} = \frac{1}{Z} (T_{,u}Z_{,v} + Z_{,u}T_{,v})$$

$$Z_{,uv} = \frac{1}{Z} (T_{,u}T_{,v} + Z_{,u}Z_{,v} - \mathbf{X}_{,u} \cdot \mathbf{X}_{,v})$$

$$\mathbf{X}_{,uv} = \frac{1}{Z} (\mathbf{X}_{,u}Z_{,v} + Z_{,u}\mathbf{X}_{,v})$$

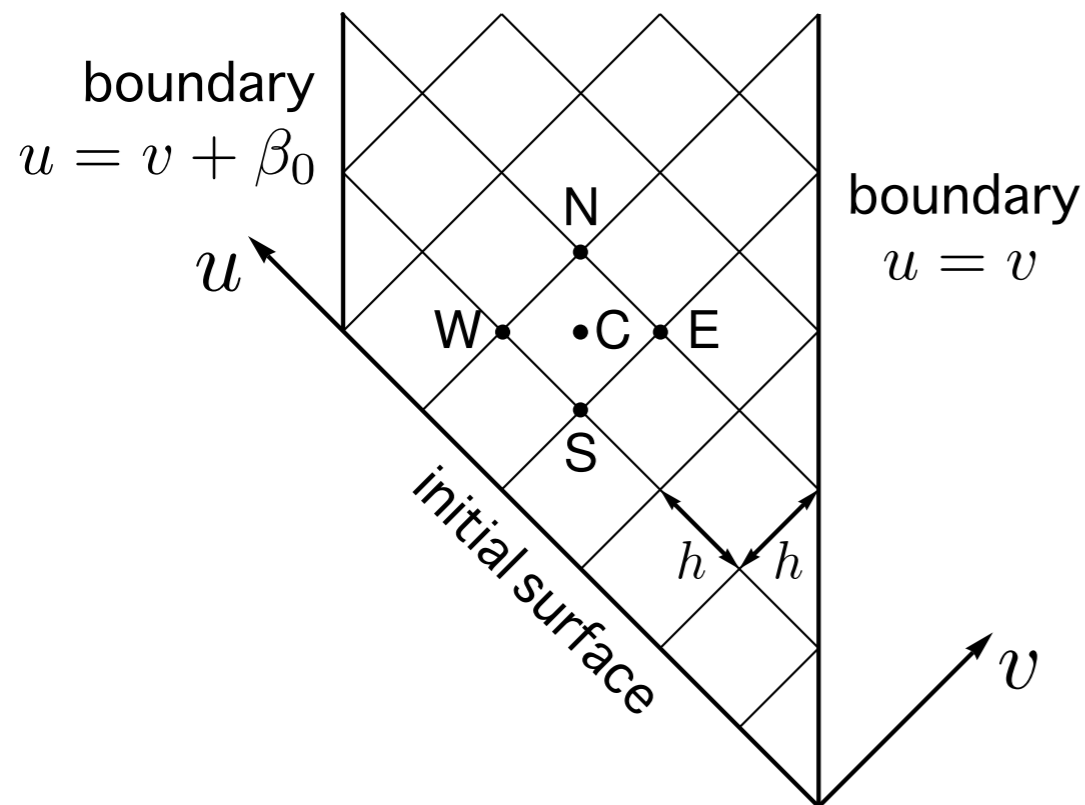
Constraints

$$\gamma_{uu} = \frac{\ell^2}{Z^2} (-T_{,u}^2 + Z_{,u}^2 + \mathbf{X}_{,u}^2) = 0$$

$$\gamma_{vv} = \frac{\ell^2}{Z^2} (-T_{,v}^2 + Z_{,v}^2 + \mathbf{X}_{,v}^2) = 0$$

Discretization

To solve EoMs, we use $O(h^2)$ central finite differential



$$\Psi_{,uv}|_C = \frac{\Psi_N - \Psi_E - \Psi_W + \Psi_S}{h^2}$$

$$\Psi_{,u}|_C = \frac{\Psi_N - \Psi_E + \Psi_W - \Psi_S}{2h}$$

$$\Psi_{,v}|_C = \frac{\Psi_N + \Psi_E - \Psi_W - \Psi_S}{2h}$$

$$\Psi|_C = \frac{\Psi_E + \Psi_W}{2}$$

Compute N by using EWS data

Initial data

Initial data satisfies the constraint

$$\gamma_{uu} = \frac{\ell^2}{Z^2} (-T_{,u}^2 + Z_{,u}^2 + \mathbf{X}_{,u}^2) = 0$$

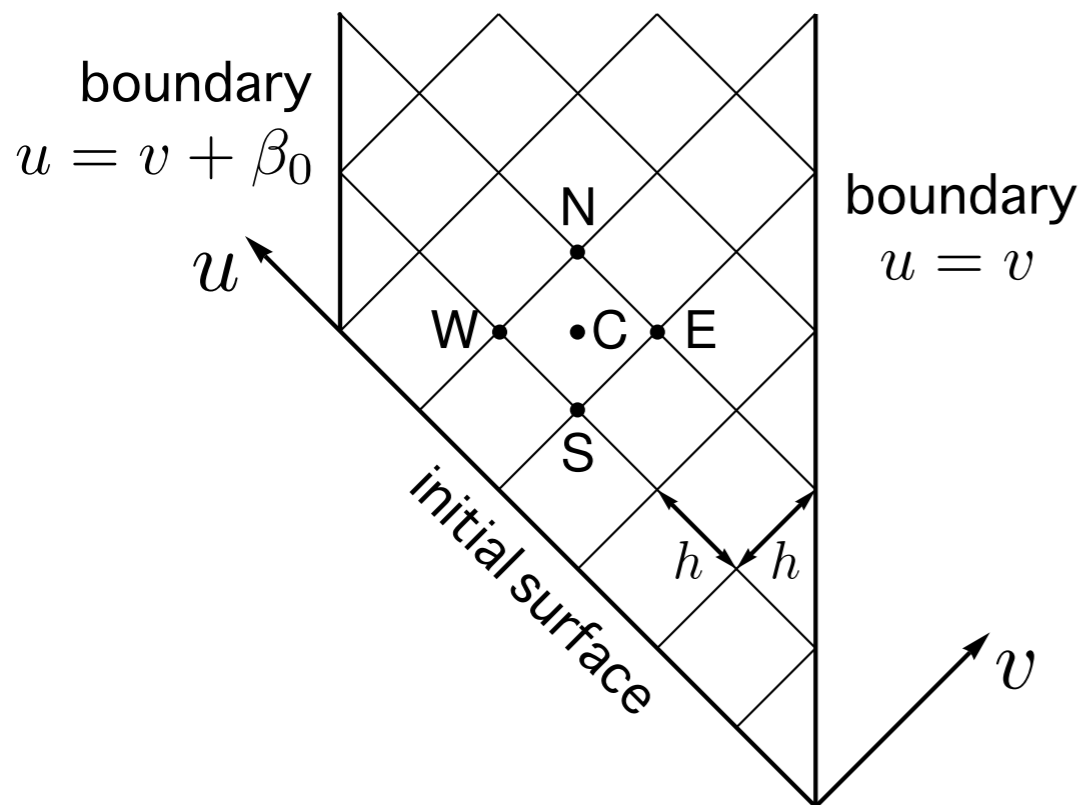
Solution (gauge: $\phi=u$ when $v=0$)

$$T(u, 0) = z_0 u$$

$$Z(u, 0) = z_0 f(u)$$

$$X_1(u, 0) = z_0 g(u)$$

where we used $f'(\phi)^2 + g'(\phi)^2 = 1$

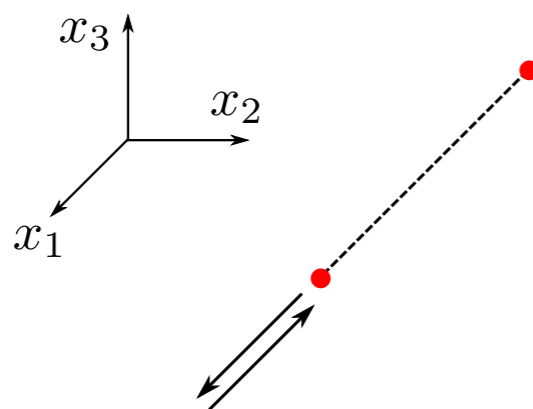
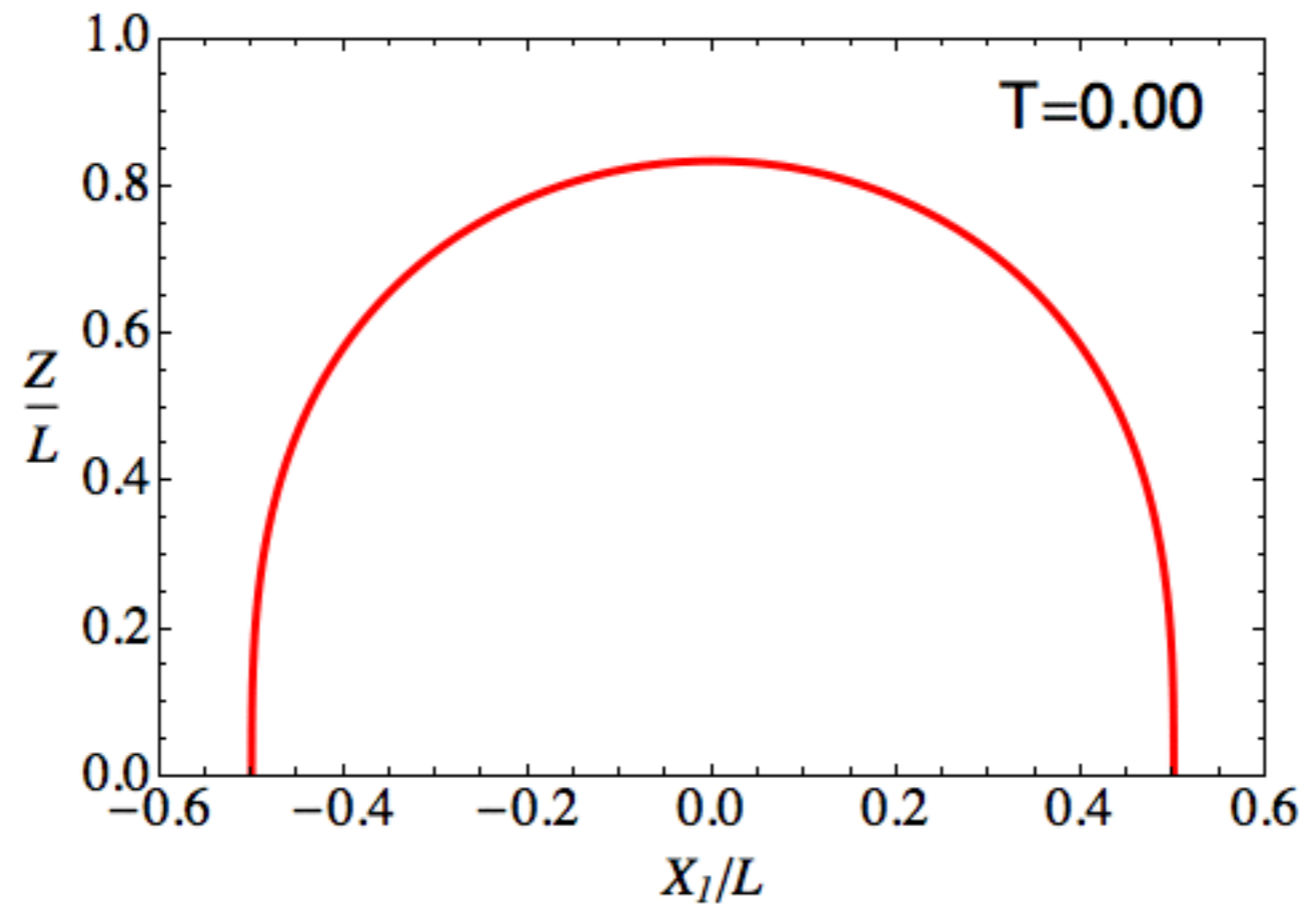


Boundary quench is then added at $0 < T_{\text{bdry}} < \Delta t$

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Longitudinal one-sided quench

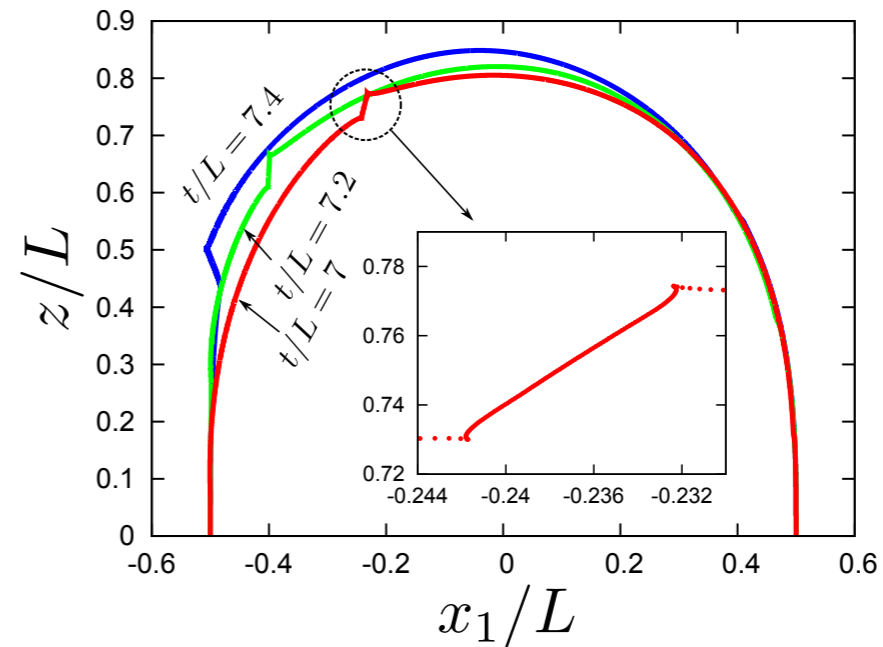
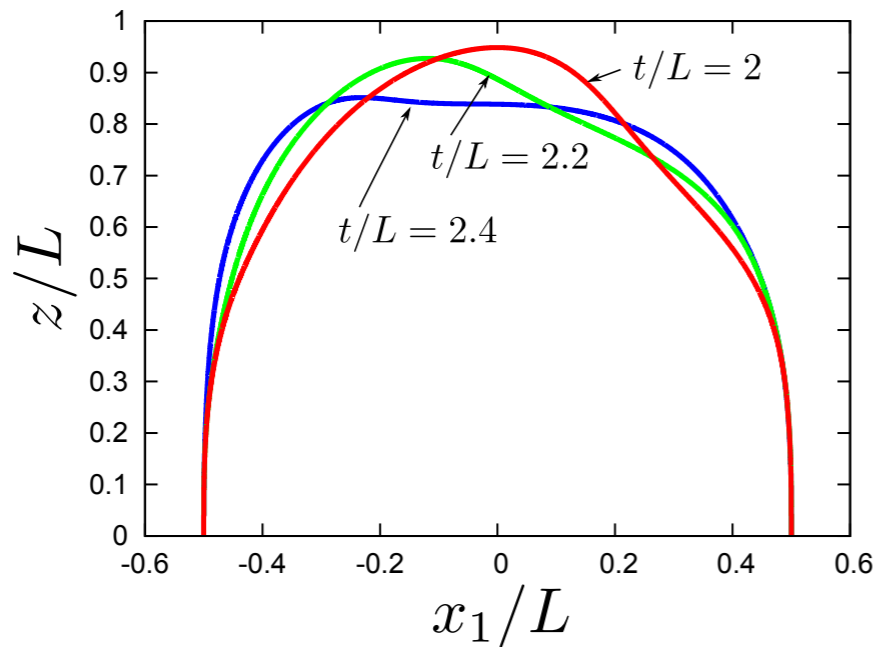


$$\varepsilon=0.03, \Delta t/L=2$$

Amplitude: $\varepsilon=\Delta x/L$

Duration: $\Delta t/L$

Cusp formation



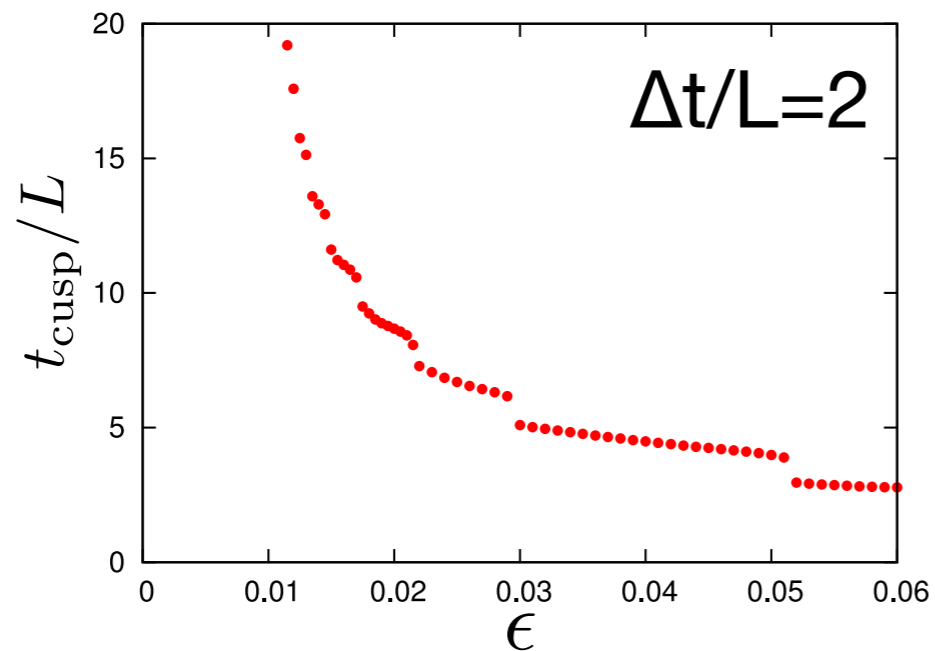
- Cusps are seen in target space (x,z) -coordinates
- Fields on worldsheet (u,v) -coordinates are regular
- Cusps are **created in a pair** (around $t/L \sim 5$)

Analysis 1: Cusp detection

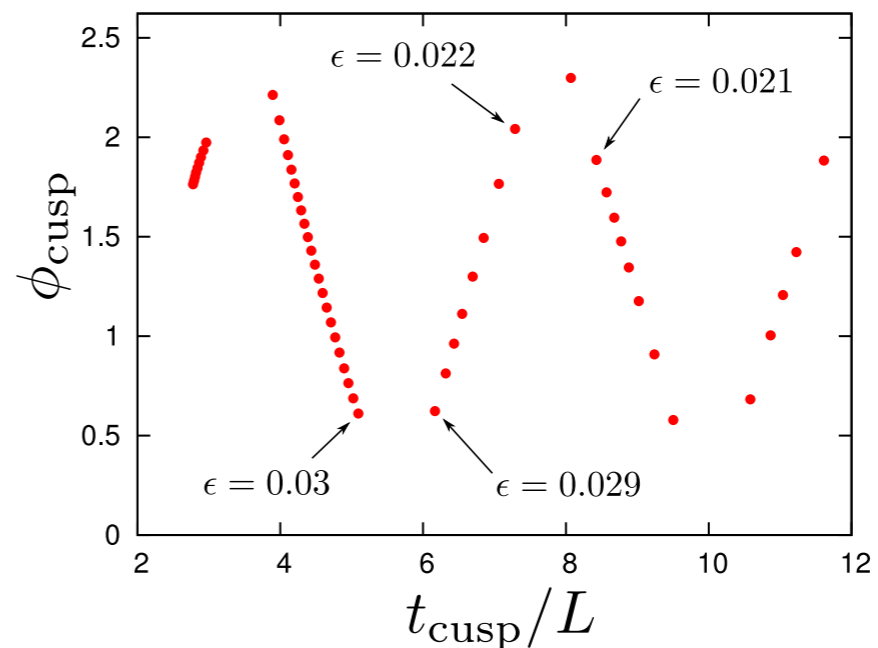
The conditions satisfied at a cusp:

$$J_z \equiv T_{,u}Z_{,v} - T_{,v}Z_{,u} = 0$$

$$J_i \equiv T_{,u}X_{i,v} - T_{,v}X_{i,u} = 0$$



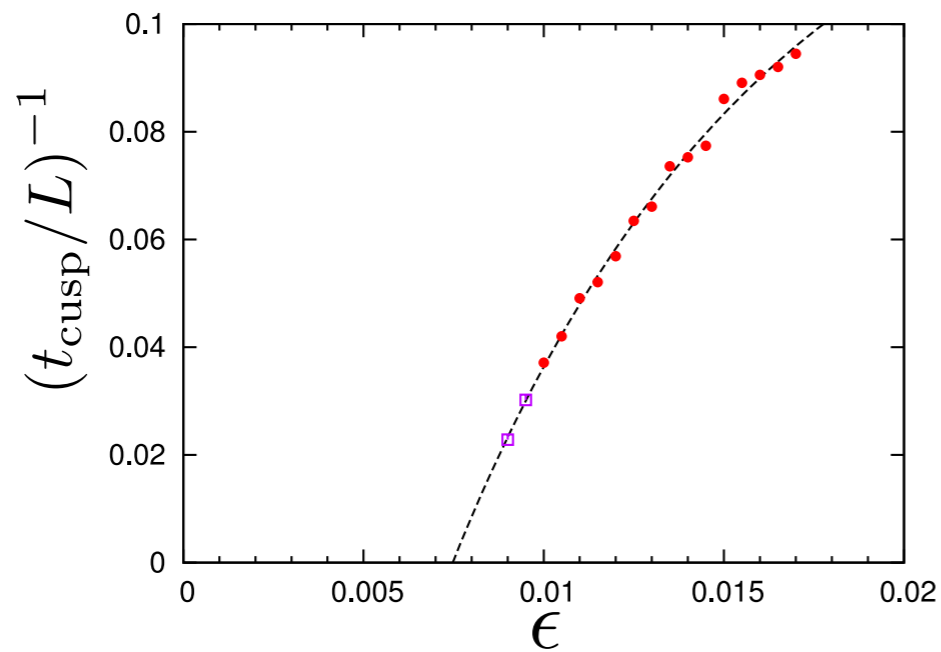
Cusp formation time
when ϵ is changed



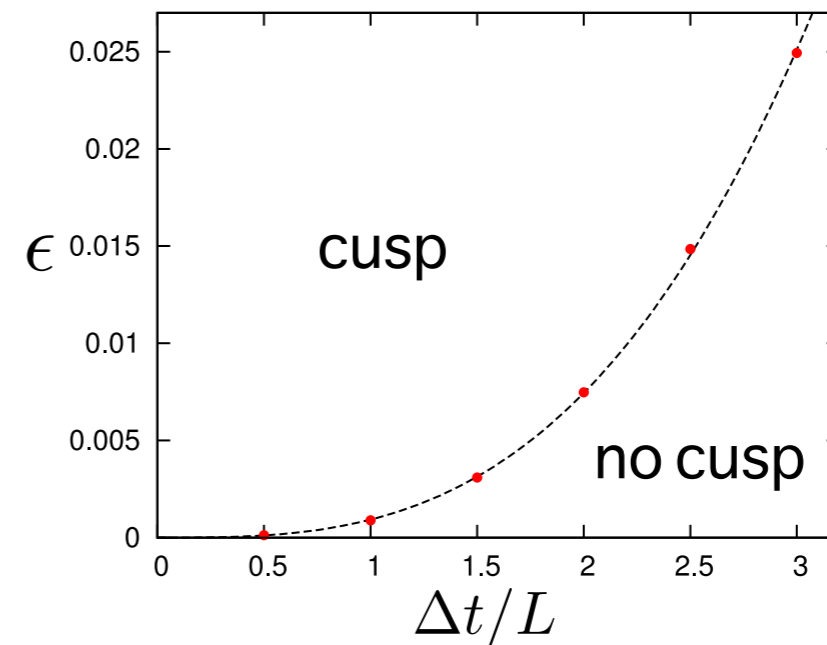
Corresponding
formation points

Critical amplitude

There is a minimal amplitude for cusp formation



An extrapolation to $t_{\text{cusp}} \sim \infty$:
 $\epsilon_{\text{crit}} \sim 0.075$ for $\Delta t/L = 2$



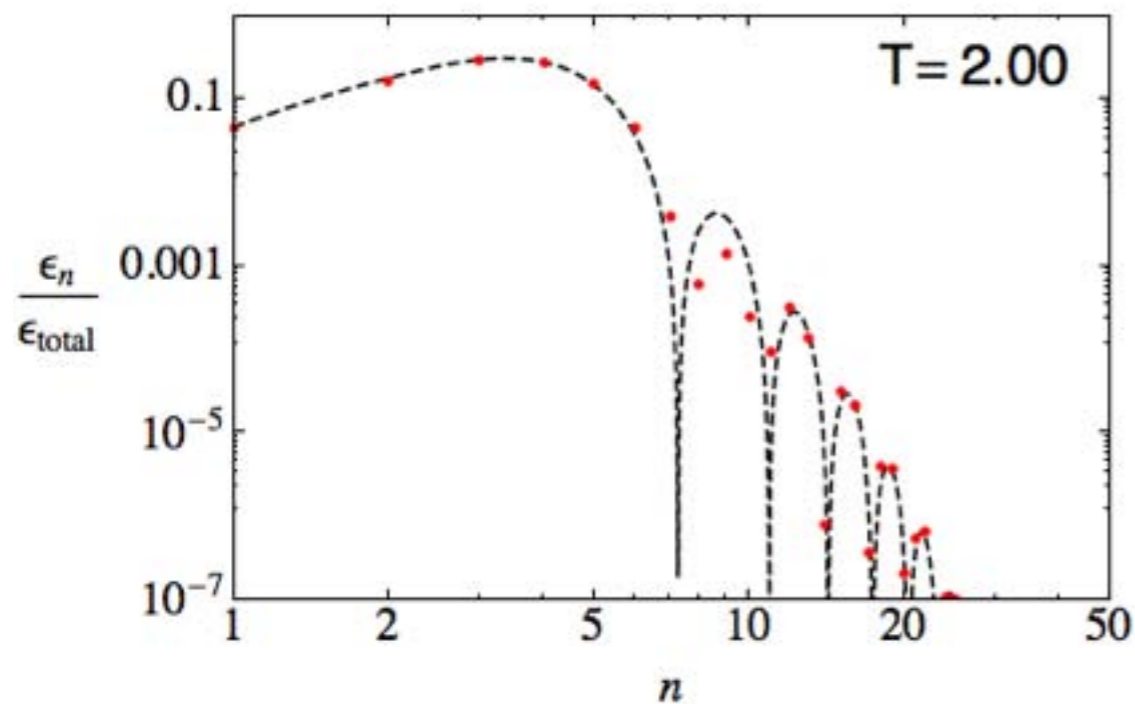
Scaling (in small $\Delta t/L$)
 $\epsilon_{\text{crit}} \sim (\Delta t/L)^3$

Analysis 2: Energy spectrum

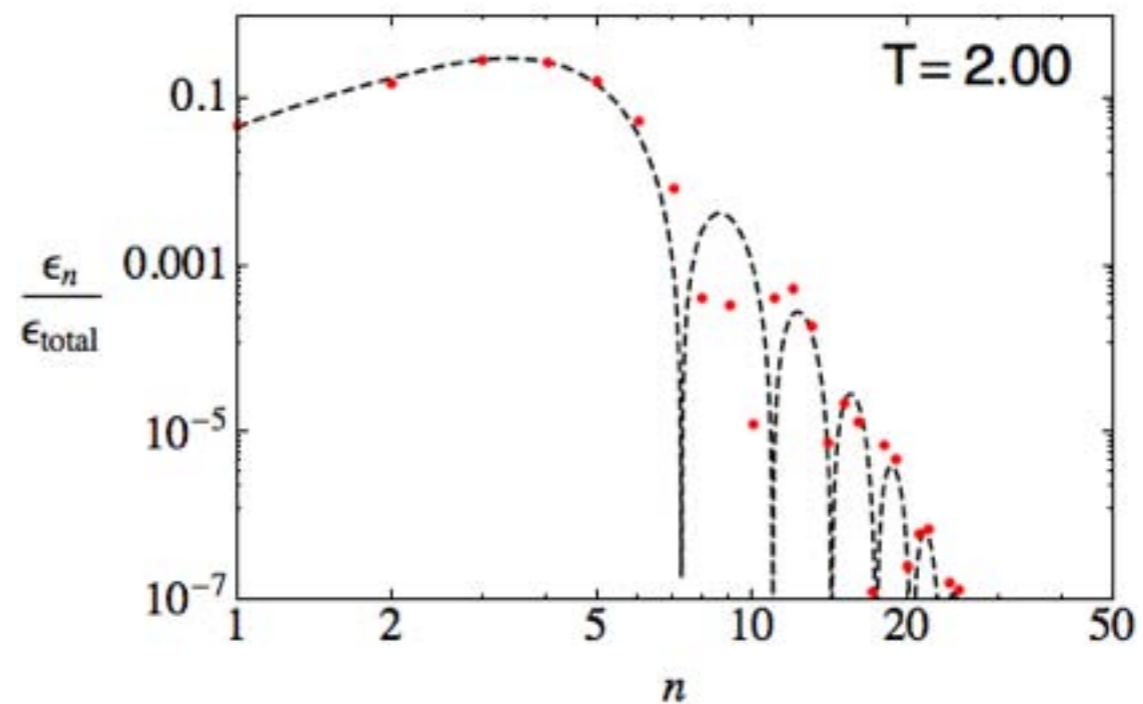
Decompose nonlinear solutions in linear eigenmodes e_n

$$\chi_1 = \sum_{n=1}^{\infty} c_n(t) e_n(\phi) \quad \varepsilon_n(t) = \frac{\sqrt{\lambda} z_0}{4\pi} (\dot{c}_n^2 + \omega_n^2 c_n^2)$$

Log-log plot



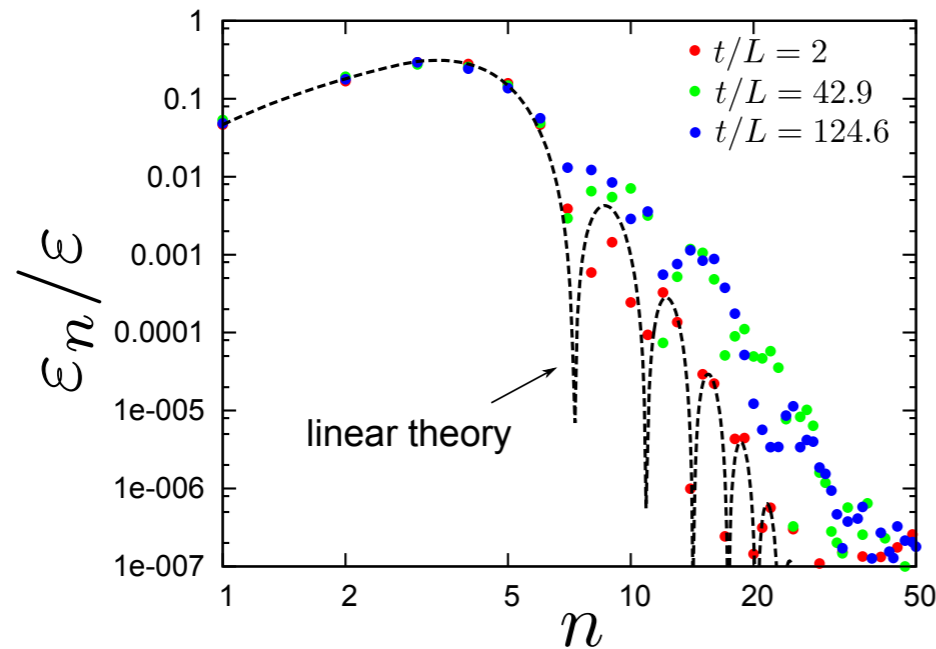
$\varepsilon=0.005$, $\Delta t/L=2$ (no cusp)



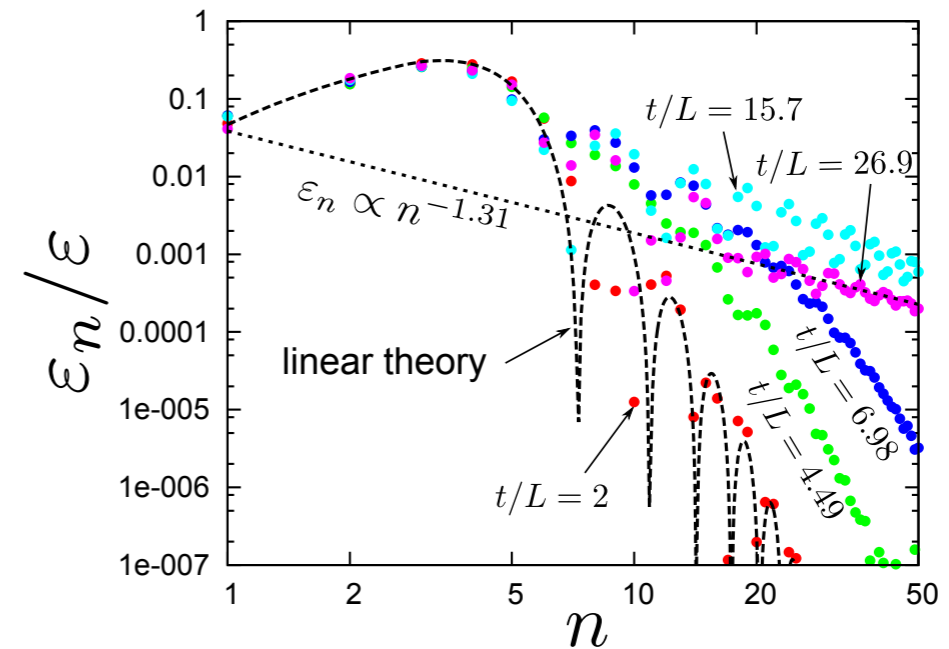
$\varepsilon=0.01$ (cusps $T \sim 27$)

***Dashed lines: linearized computations

Energy cascade



$\varepsilon=0.005$, $\Delta t/L=2$ (no cusp)



$\varepsilon=0.01$ (cusps $T \sim 27$)

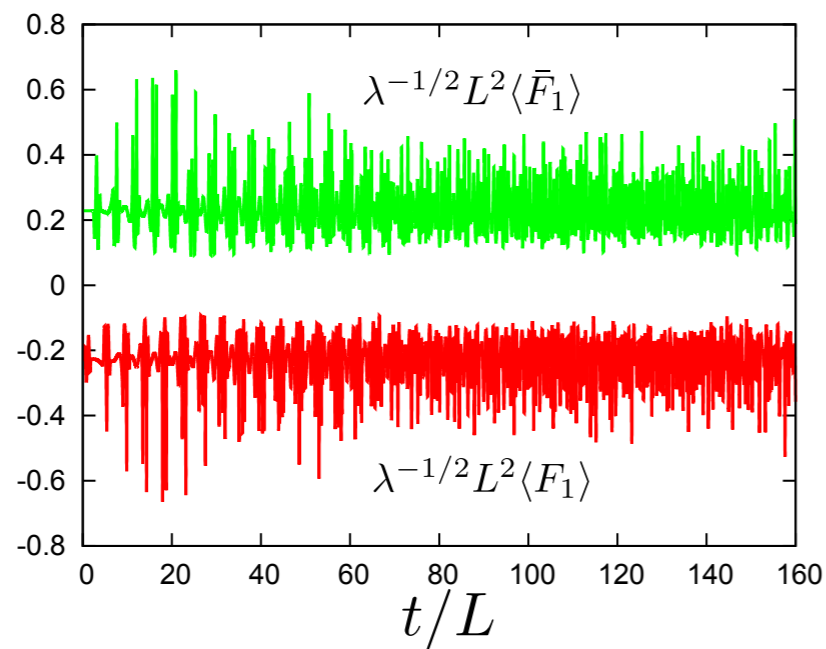
Cusp formation: **direct energy cascade** → power law

No cusp: no clear power law

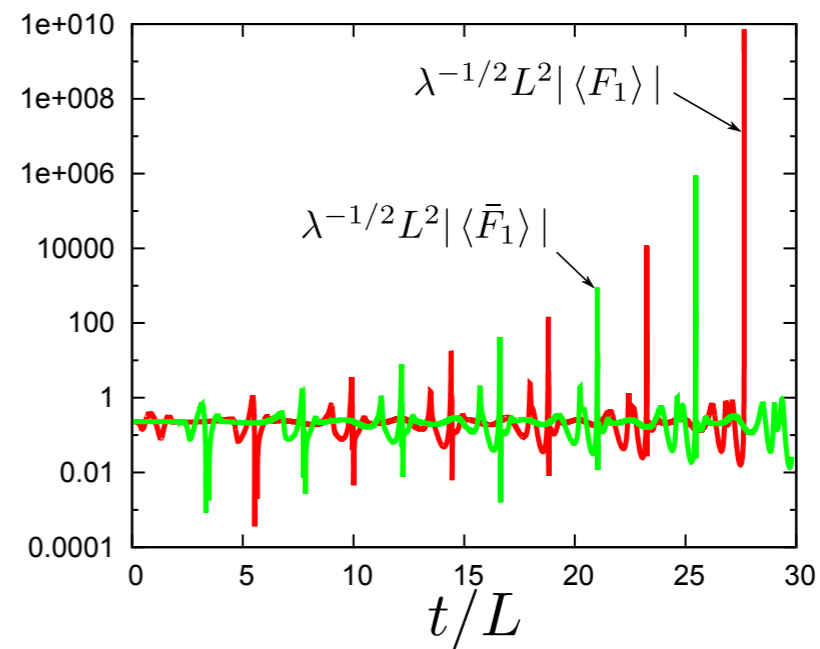
Analysis 3: Forces on the endpoints

$$\langle \mathbf{F}(t) \rangle = \frac{\delta S_{\text{on-shell}}}{\delta \mathbf{x}_q}$$

Force diverges when a cusp reaches the boundary



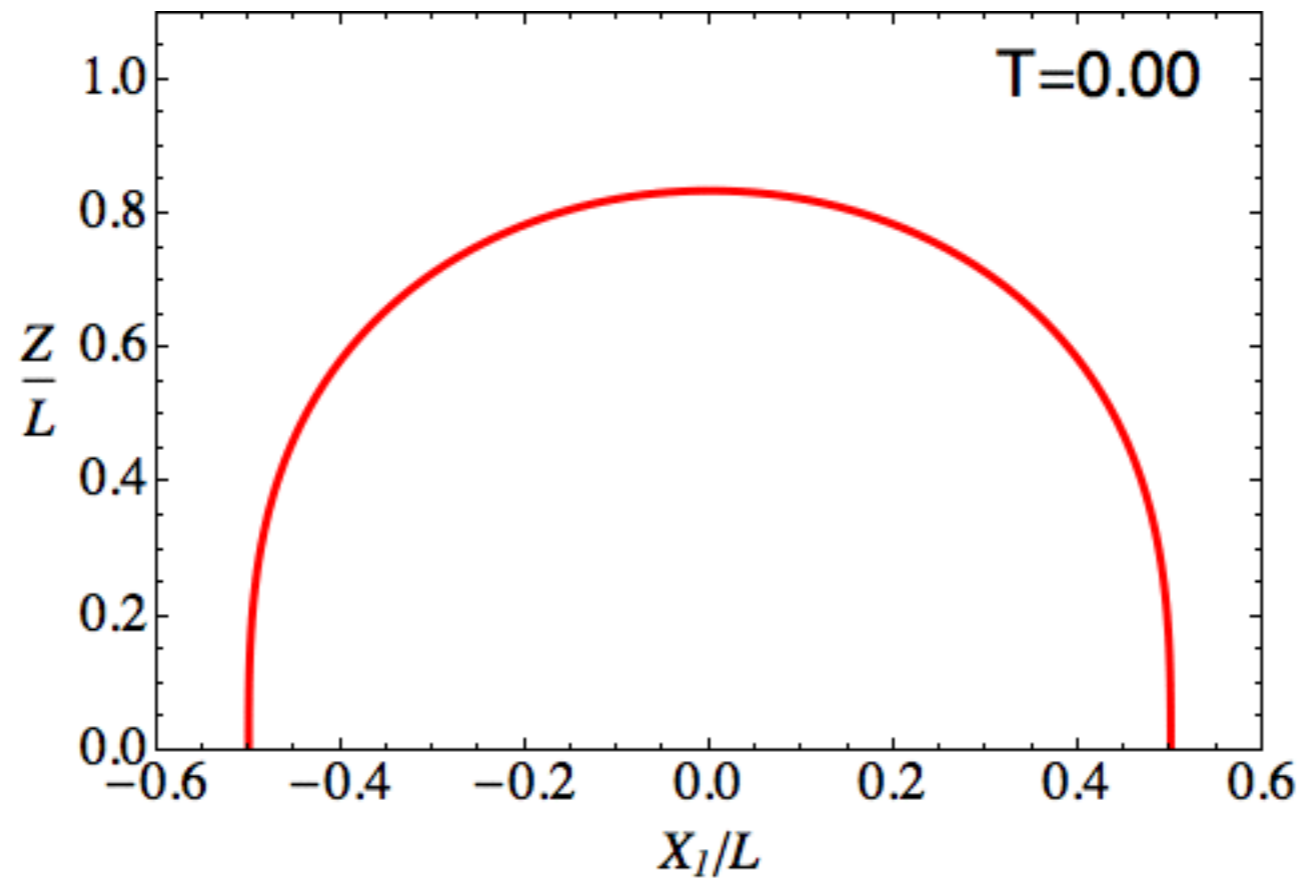
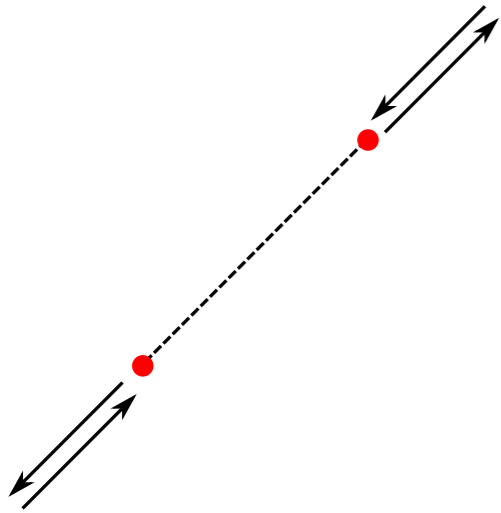
$\varepsilon=0.005, \Delta t/L=2$ (no cusp)



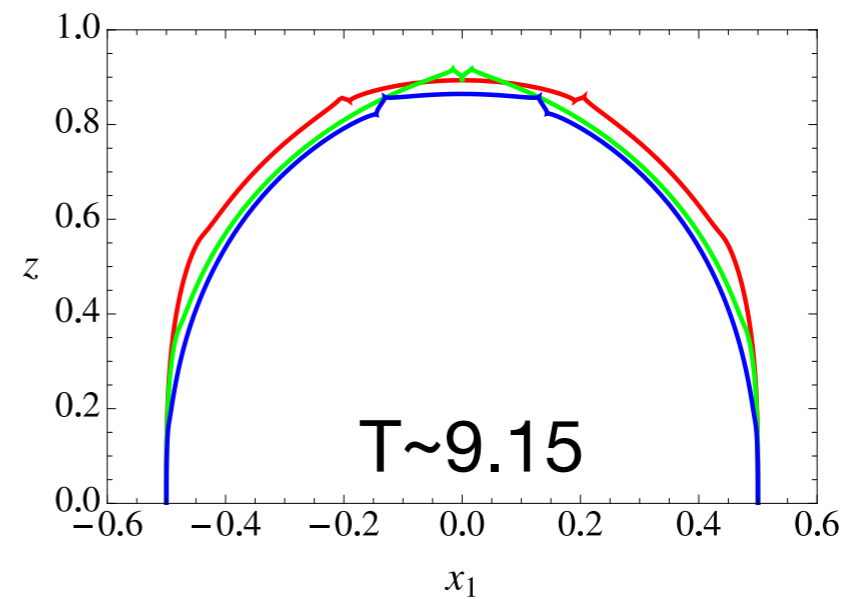
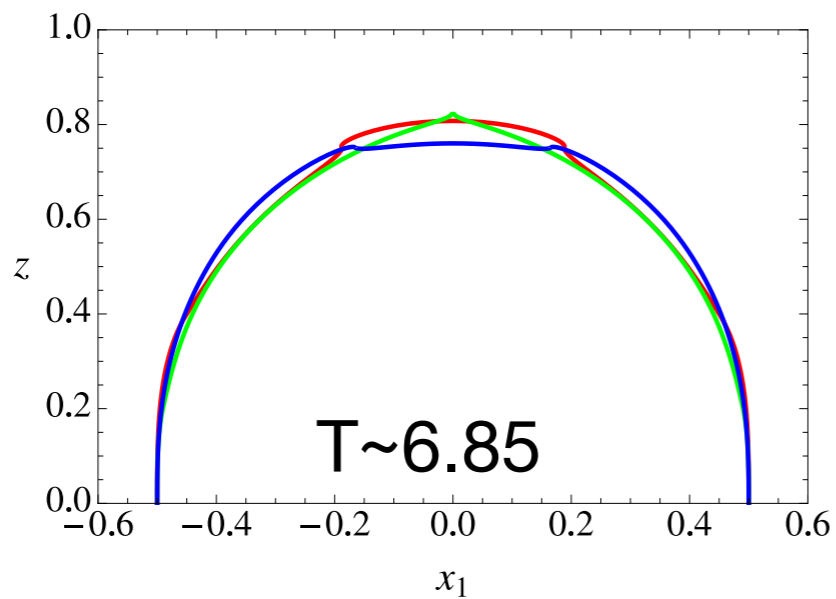
$\varepsilon=0.01$ (cusps $T \sim 27$)

***Red: $x=L/2$, green: $x=-L/2$

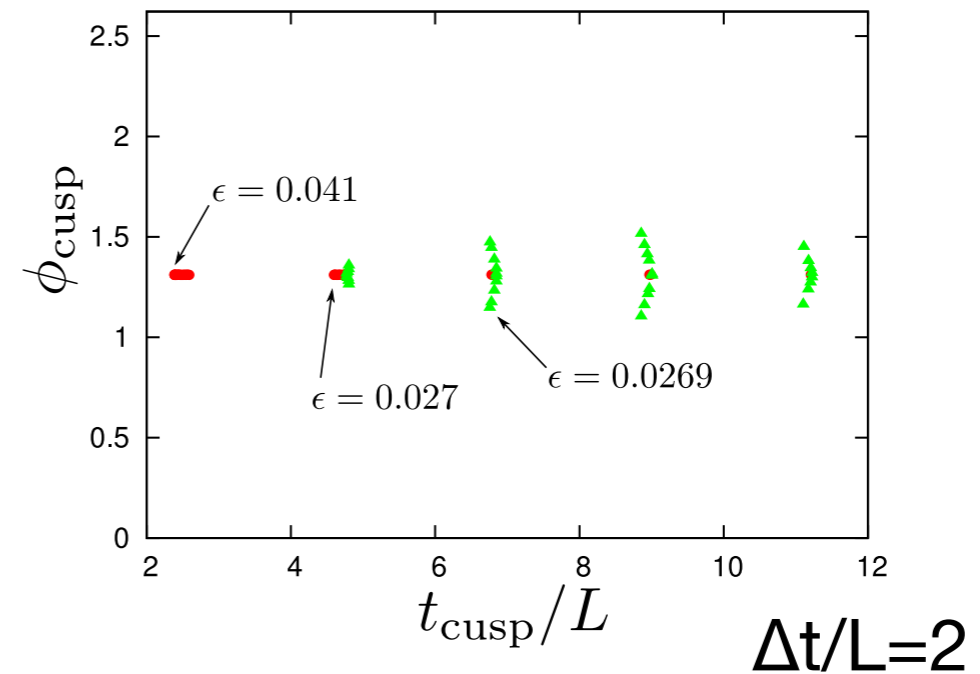
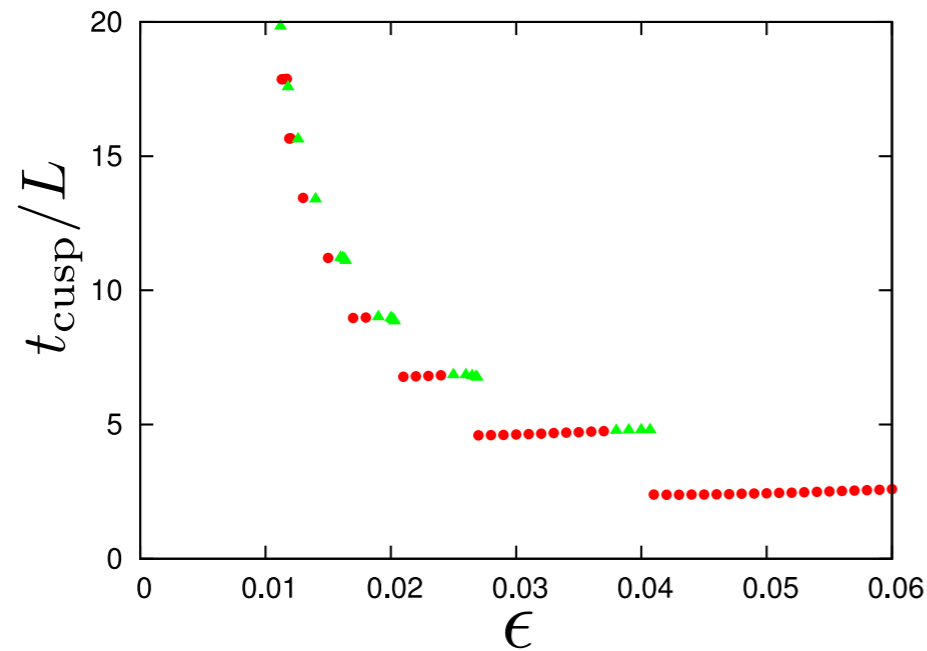
Z₂-symmetric quench



$\varepsilon=0.025$
 $\Delta t/L=2$

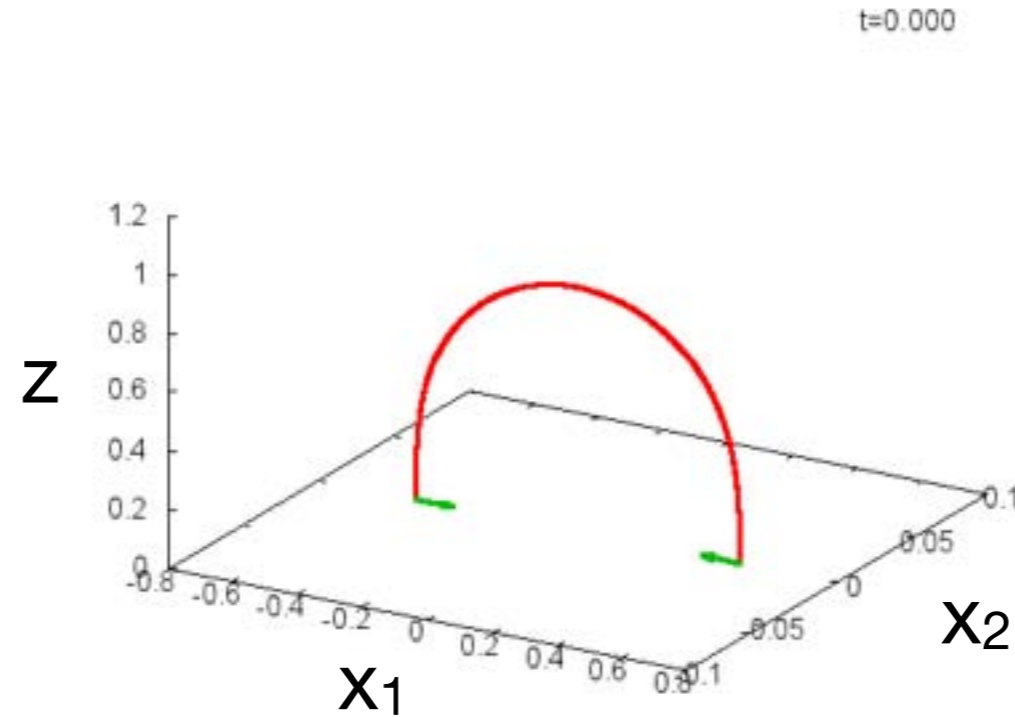
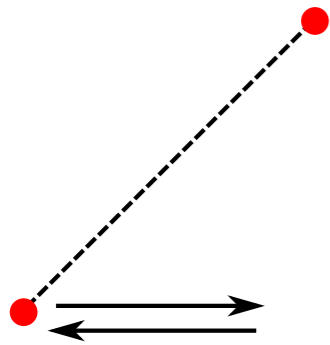


Z_2 -symmetric quench



- Formation times are discretized by wave collisions
- First cusp formations by such collisions (red ●).
The cusps are pair-created and annihilated.
- Traveling cusps can be formed first (green ▲)

Transverse linear quench

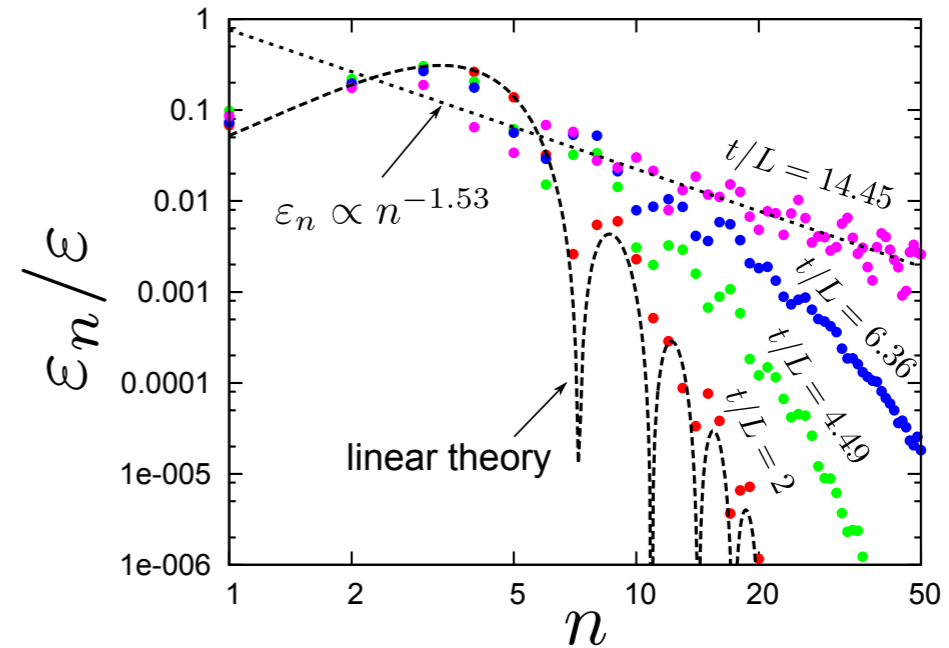
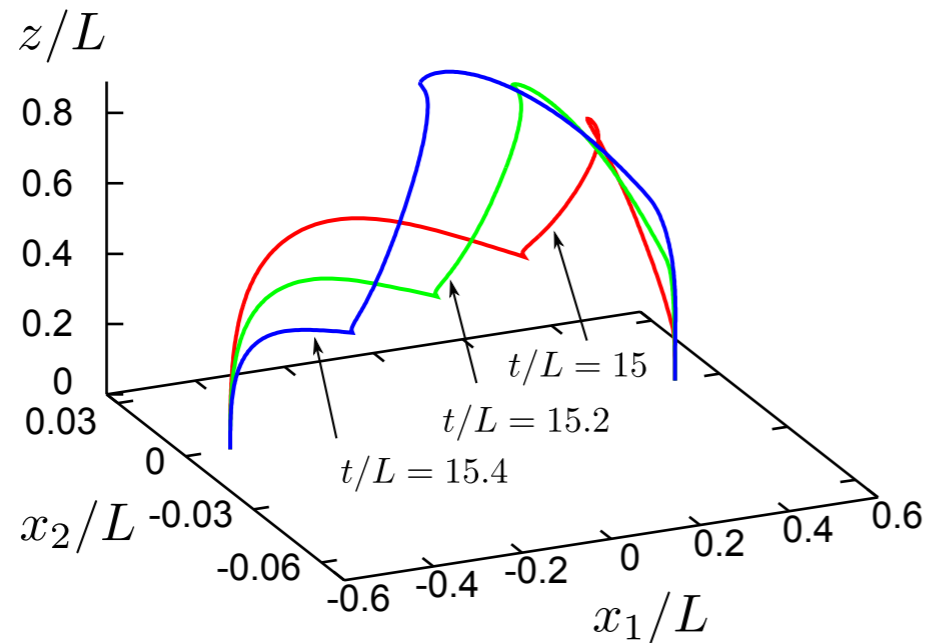


$$\varepsilon=0.03, \Delta t/L=2$$

***Green arrows: forces

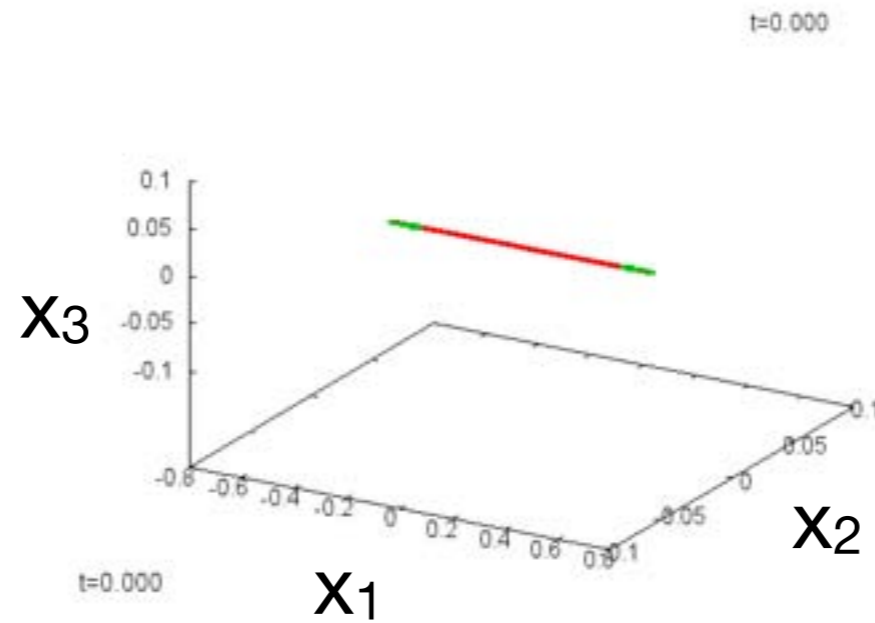
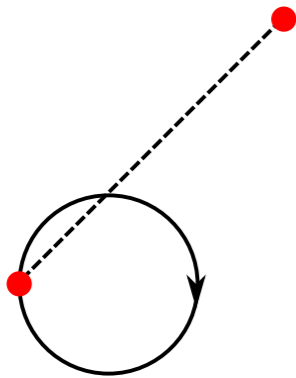
String oscillates in 1+3 dim (t, z, x_1, x_2)

Transverse linear quench

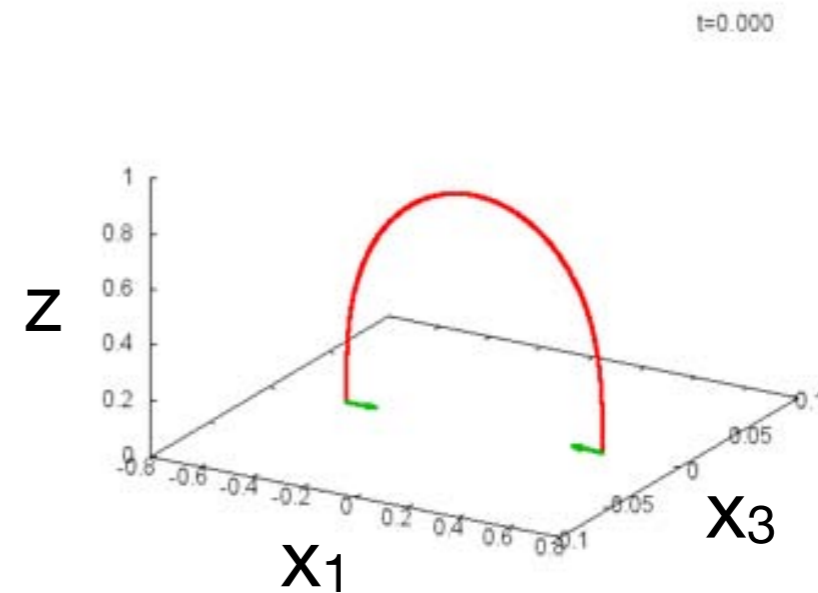
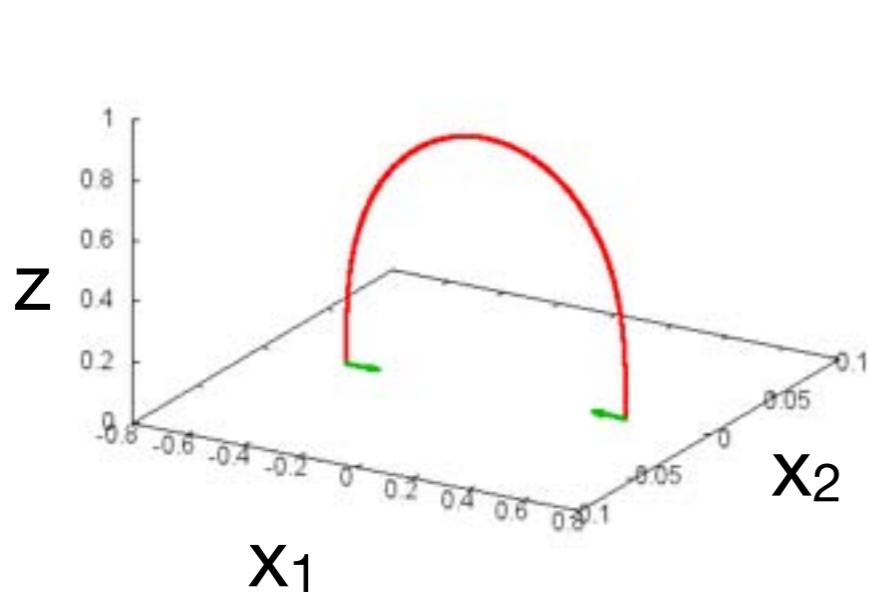


- Cusps are formed at $T \sim 14.45$
- The energy spectrum keeps a power law

Transverse circular quench

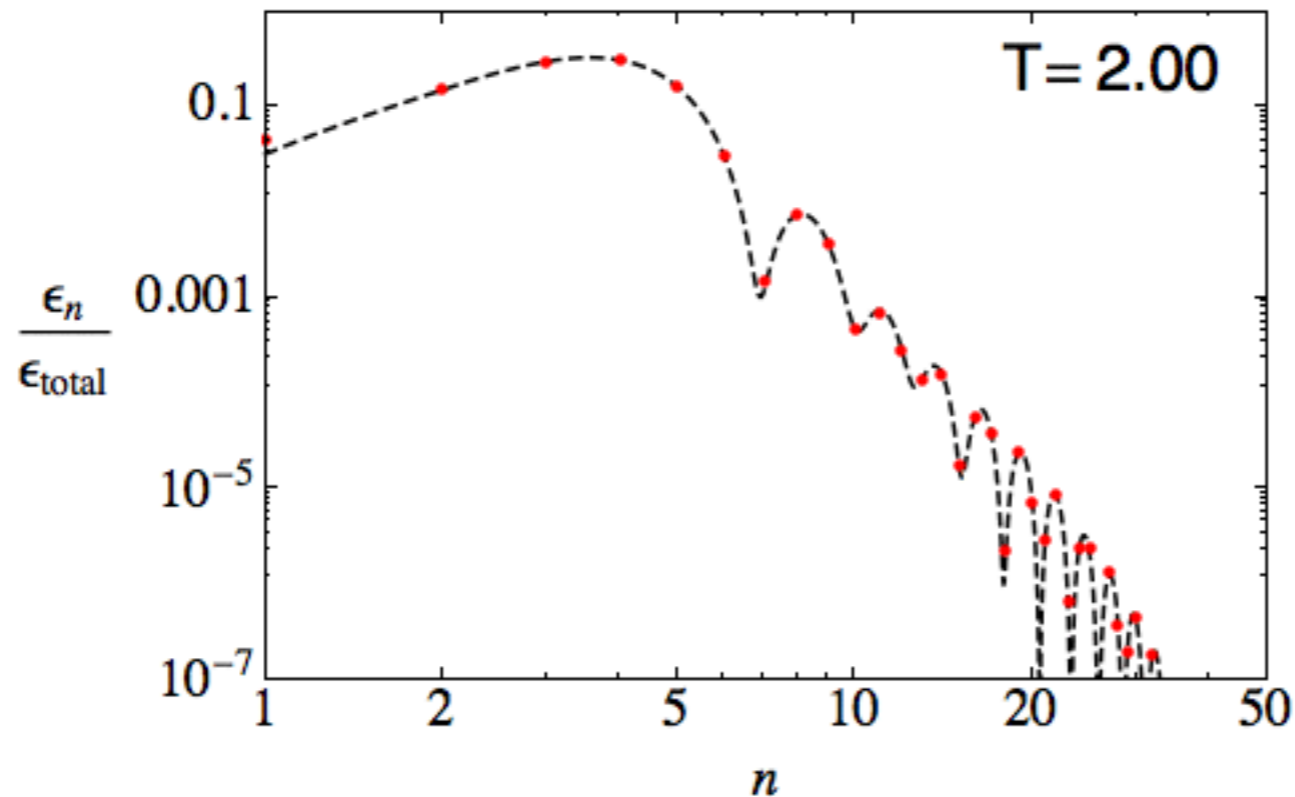


$$\varepsilon=0.02, \Delta t/L=2$$



String oscillates in all 1+4 dim (t, z, x_1, x_2, x_3)

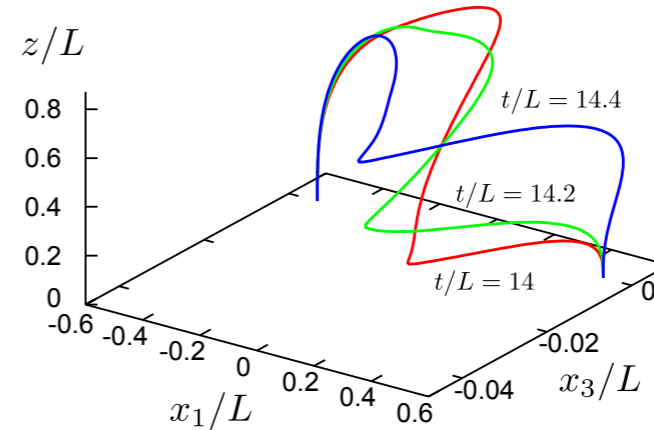
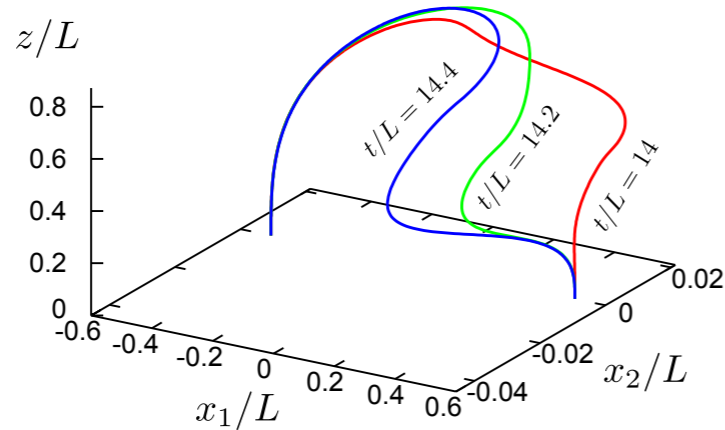
Energy spectrum (Log-log plot)



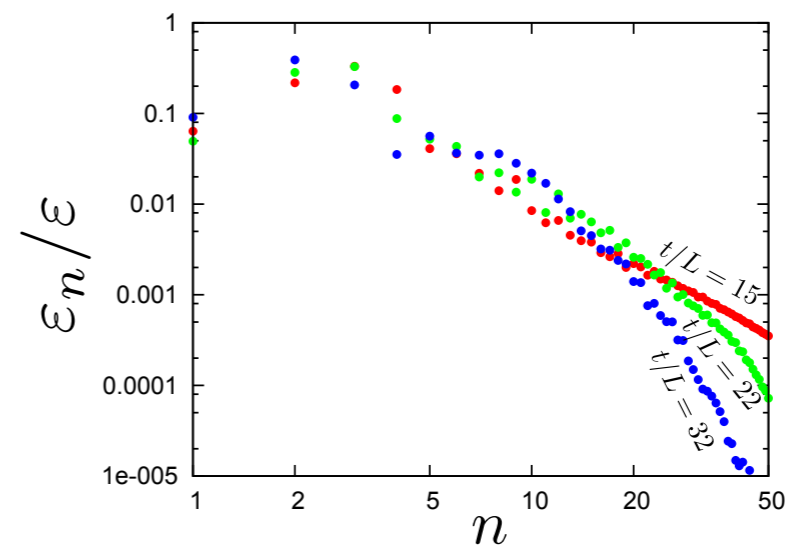
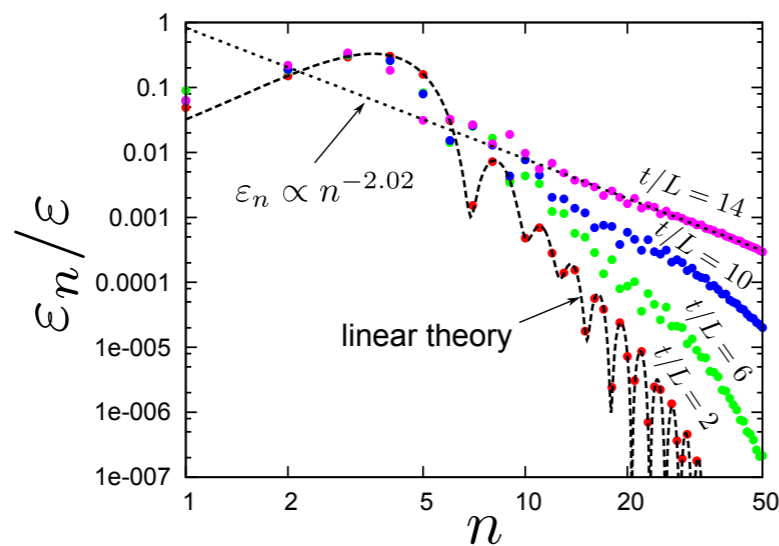
No cusp: no sustaining power law

c.f.) Probability of cusp formation is zero if $\text{dim} > 4$

Transverse circular quench



Cuspy, but not real cusps



Direct cascade → **inverse cascade**

Summary

We computed nonlinear dynamics of the quark-antiquark fundamental string in AdS

- Cusps and turbulent behavior in $\leq 1+3$ dim
- No cusp and direct/inverse cascades in $1+4$ dim
c.f.) Cosmic strings in flat space

Discussion

Gravitational backreaction may be necessary

- Curvature diverges at the cusps
- AdS gravitational wave bursts?
- Boundary interpretation: gluon bursts?

Future works

- Large amplitude/finite temperature
- Non-conformal backgrounds
- Application to drag force



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