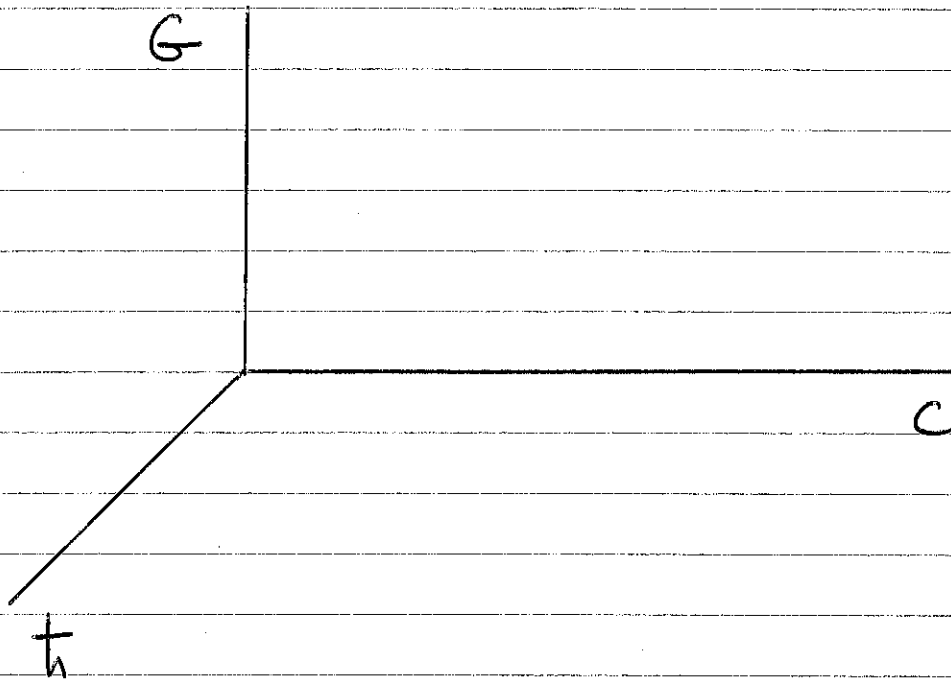


①

CG: Introduction to Quantum Field Theory

Goal: extrapolate QM to relativistic domain (i.e. beyond $v/c \ll 1$, $E/mc^2 \ll 1$)



Appropriate dimensionless parameters:

$$SR: v/c$$

$$QM: \lambda/l$$

$$\lambda = \frac{h}{mc}$$

$$QG: l_p/l$$

Reminder: Planck units

$$l_p = \sqrt{\frac{Gh}{c^3}} \sim 1.6 \cdot 10^{-33} \text{ cm}$$

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$$t_p = \sqrt{\frac{\hbar}{c^3}} \sim 0.5 \cdot 10^{-43} \text{ s}$$

$$m_p = \sqrt{\frac{\hbar c}{G}} \sim 2 \cdot 10^{-8} \text{ kg}$$

$$E_p \sim 1.2 \cdot 10^{19} \text{ GeV} \quad (E_{4HC} \sim 10^4 \text{ GeV})$$

Note: grav. coupling has dim. of L^2

$$[G] = L^2$$

Note: quantum grav. becomes important for $\lambda \sim l$ (size of grav. system,

e.g. $l \sim r_s = 2GM/c^2$)

$$\frac{2MG}{c^2} \sim \frac{\hbar}{Mc}$$

$$\Rightarrow M \sim m_p \quad (\text{Planck mass})$$

Moral: for $E \ll E_p$ can treat gravity as classical background. Moreover, to

a good approx., $g_{\mu\nu}(x) \approx \eta_{\mu\nu}$ (Mink)

\Rightarrow QFT in flat space-time.

4d Mink $ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2$

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Symmetries: 4 transl. ($t \rightarrow t+a$ etc)

Lorentz transf. $\left\{ \begin{array}{l} 3 \text{ rot.} \\ 3 \text{ boosts } (tx, ty, tz) \end{array} \right.$

\Rightarrow 10-parameter Poincaré group =
= transl. + Lorentz group

More precisely, $\mathbb{R}^{1,3} \rtimes O(1,3)$
(semi-direct product)

Moral: with gravity non-dynamical and flat, fundamental symmetry of the "arena" is Poincaré symmetry

Thus, we build a theory which has Poincaré-inv. action and covariant e.o.m.

Note: this is not God-given - validity of SR is an experim. fact; keep looking for violations of Lor-invar.

Where do we expect to find rel. corr. to QM?

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Spectroscopy of hydrogen atom

$$\text{QM: } E_{nl} = - \frac{\alpha^2}{2n^2} mc^2 = - \frac{me^4}{2n^2 (4\pi\epsilon_0)^2 \hbar^2}$$

($l=0, \dots, n-1$)

$$\text{where } \alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c} \approx 1/137$$

$$\alpha \approx 0.0072973525693 \quad (\text{PDG } 2018)$$

$$\alpha^{-1} \approx 137.035999084$$

Electron moves in atom with $v/c \sim 10^{-2}$

\Rightarrow expect rel. corrections to E_{nl}

of order $v^2/c^2 \sim 10^{-4} \sim O(\alpha^2)$,

$$\text{i.e. } E_{nl} = - \frac{\alpha^2}{2n^2} mc^2 + O(\alpha^4)$$

• Generalise Schrödinger eq to include rel. effects

Non-rel. particles / waves (de Broglie)

$$E = \bar{p}^2 / 2m \quad \text{disp. relation is}$$

reproduced from wave $\psi \sim e^{-i\omega t + i\bar{p}\bar{x}}$

with $E \rightarrow i\hbar \partial / \partial t$

$$p_i \rightarrow -i\hbar \frac{\partial}{\partial x_i}$$

$$\Rightarrow \text{eq. for } \psi \text{ is } i\hbar \frac{\partial \psi}{\partial t} = \hat{H} \psi,$$

$$\hat{H} = \hat{p}^2 / 2m.$$

This eq. is parabolic and is not Lor-covariant.

Parab. eqs (diffusion, Navier-Stokes)

\Rightarrow infinite prop. speed.

$$\text{E.g. } \partial_t \phi = D \nabla^2 \phi \quad \text{in } n \text{ dim:}$$

$$\Rightarrow \phi = \frac{1}{(4\pi Dt)^{n/2}} e^{-\bar{x}^2 / 4Dt}$$

With $\phi(t=0, \bar{x}) = \delta(\bar{x})$;

at any $t > 0$, the solution is

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non-zero at any $|\vec{x}|$, no matter how large \Rightarrow signal propagates with infinite speed.

Need hyperbolic eq, such as Maxwell eq:

$$\square A^\mu = \left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) A^\mu = 0$$

to respect SR.

• In SR, the disp. rel. for a free particle is $E^2/c^2 - \vec{p}^2 = m^2 c^2$

With $E \rightarrow i\hbar \partial_t$, $\vec{p}_i \rightarrow -i\hbar \partial/\partial x_i$ we get

$$\left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} + \frac{m^2 c^2}{\hbar^2} \right) \Phi(t, \vec{x}) = 0$$

$$\text{or } (\square + m^2) \phi = 0$$

This is Klein-Gordon-Fock eq (1926), originally derived by Schrödinger (1926)

Note: $(\square - m^2)\phi = 0$ in $-+++$ metric, where $\square = -\frac{1}{c^2} \frac{\partial^2}{\partial t^2} + \nabla^2$.

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KG eq. with external electromagnetic field $A^\mu = (\phi/c, \vec{A})$:

$$\left\{ \begin{array}{l} \vec{p} \rightarrow \vec{p} - \frac{e}{c} \vec{A} \\ E \rightarrow E - e\phi \end{array} \right.$$

(minimal coupling - this gives correct classical eq. of motion e.g. Lorentz force - see LL vol 2)

$$\left(i\hbar \frac{\partial}{\partial t} - e\phi \right)^2 \psi = \left(\frac{\hbar c}{i} \vec{\nabla} - e\vec{A} \right)^2 \psi + m^2 c^4 \psi$$

Hydrogen atom: $\left\{ \begin{array}{l} e\phi = -Ze^2/r \quad (Z=1) \\ \vec{A} = 0 \end{array} \right.$

Spectrum:

$$E_{nl} = mc^2 \left[1 - \frac{\alpha^2}{2n^2} - \frac{\alpha^4}{2n^4} \left(\frac{n}{l+1/2} - \frac{3}{4} \right) + \dots \right]$$

fine structure term disagrees with spectroscopy experim.