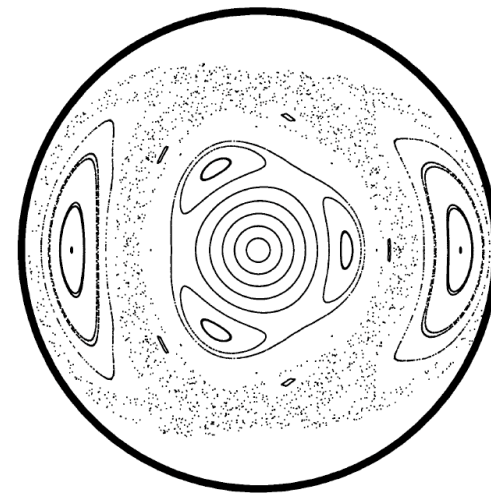


# Are the field lines in Tokamaks stochastic?

*Steve Cowley, UKAEA Culham and Imperial College.*



# Introduction.

- Is transport in tokamaks due to wandering field lines?
- Are field lines in tokamak drift-wave turbulence stochastic?
- I will extend the frozen-in theorem for ion scale turbulence (*ITG and TEM*).
- I will suggest that the field is only stochastic on electron scales.

# Transport Stochasticity and Turbulence -- $\tau_E$

*Mapping of field lines -- the need for magnetic surfaces*

SOME PROPERTIES OF ROTATIONAL TRANSFORMS

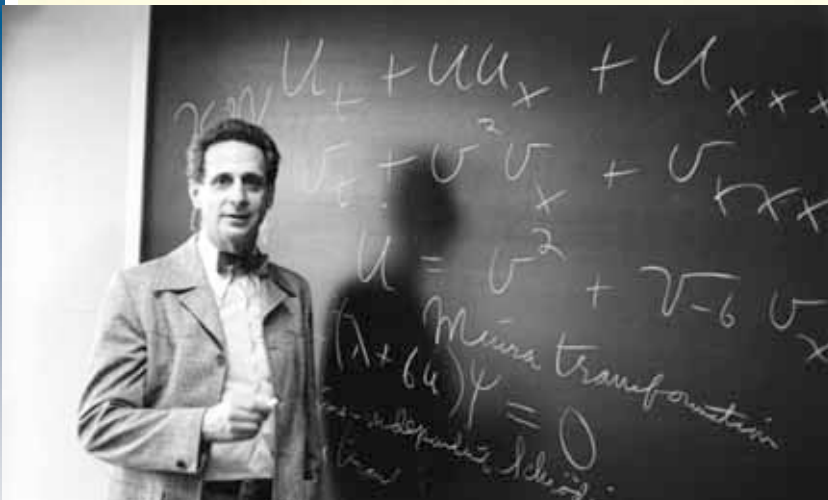
February 18, 1952

NYO-998

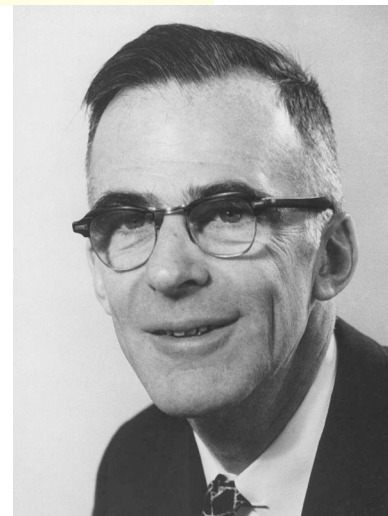
PM-S-5

$$\frac{1}{g} (X_{n+1} - X_n) = \sum_i g^i U^{(i)}(X_n);$$

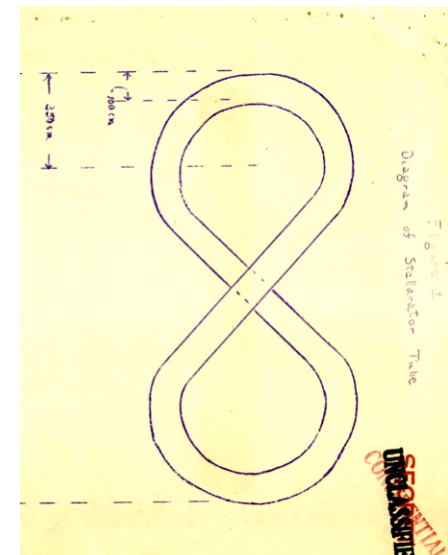
Martin D. Kruskal



Martin D. Kruskal



Lyman Spitzer



# Transport Stochasticity and Turbulence -- $\tau_E$



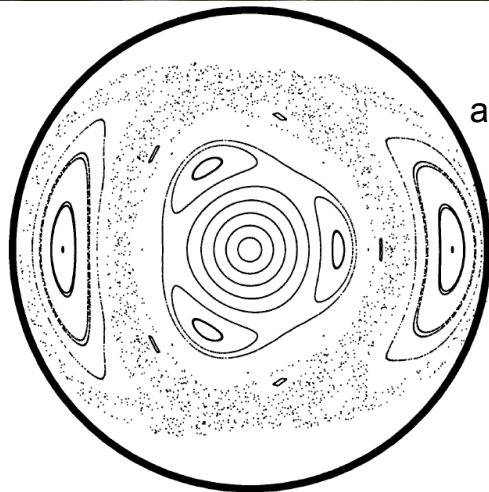
*Chirikov-Taylor map*

$$\begin{aligned} p_{n+1} &= p_n + K \sin(\theta_n) \\ \theta_{n+1} &= \theta_n + p_{n+1} \end{aligned}$$

*Field Line Diffusion  
Island overlap.*



N. Rosenbluth, R. Z. Sagdeev, J. B. Taylor,  
and G. M. Zaslavsky, Nucl. Fusion 6 , 297 (1966)



$$D_B = L_c \left( \frac{\delta B_r}{B} \right)^2$$



# Test particle transport -- $\tau_E$

*Test particles moving in prescribed stochastic fields.*



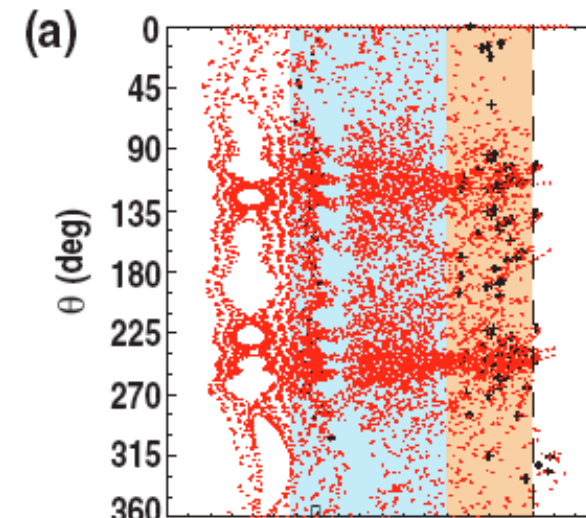
T. H. Stix, Nucl. Fusion 183,353 (1978)

A. B. Rechester and M. N. Rosenbluth,

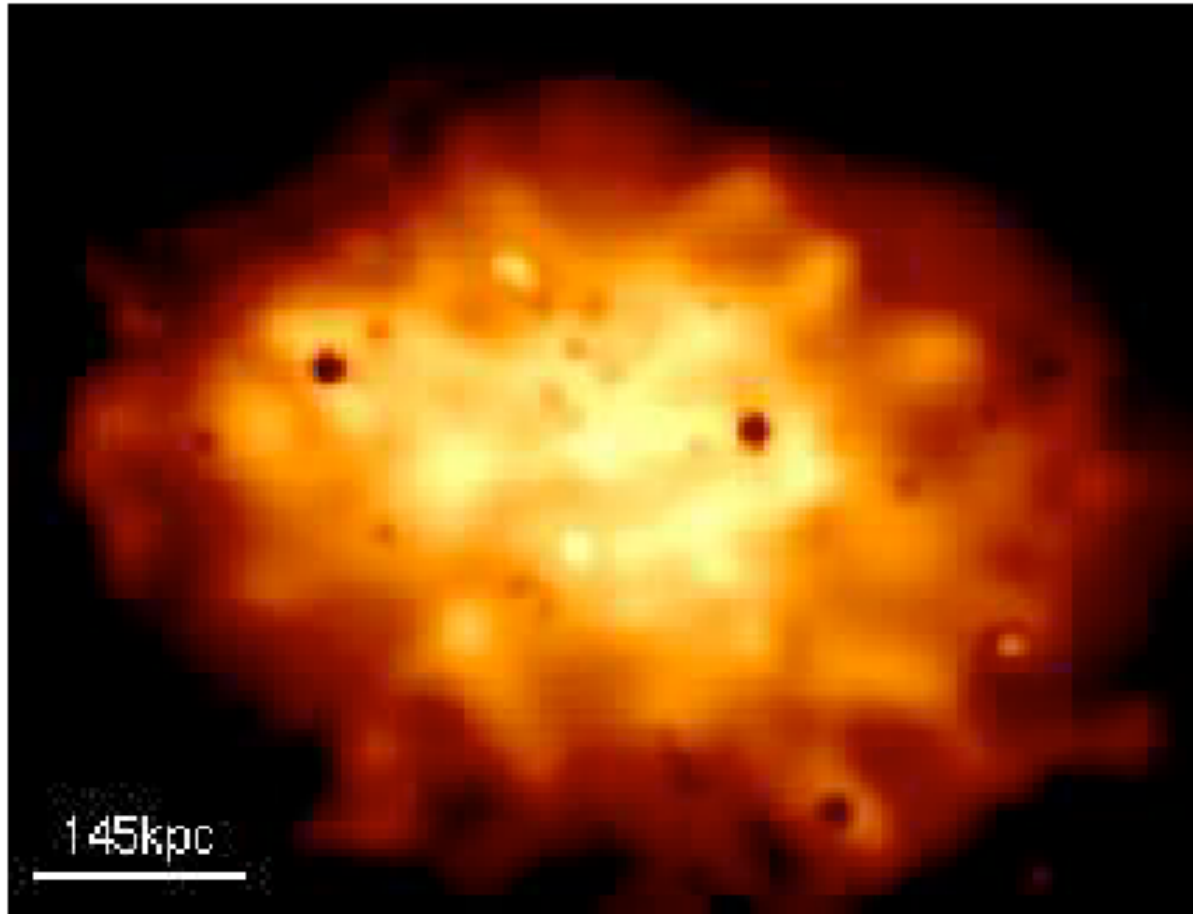
Phys. Rev. Lett. **3**, 38 (1978)

$$\chi_r \sim v_{\parallel} D_B \quad \text{Collisionless}$$

$$\sim \frac{\chi_{\parallel}}{L_c} D_B \quad \text{Collisional}$$



# Galaxy Clusters -- Confined by stochastic field?



*Outside is 10KeV  
Inside is <1KeV  
Thermal equilibration  
Shorter than lifetime,  
Magnetic confinement  
With stochastic fields?*

**The Coma Cluster: pressure map**  
[Schuecker *et al.* 2004, *A&A* **426**, 387]

# Electrostatic or Electromagnetic Turbulence?

Current tokamaks and ITER will have  $\beta \sim \epsilon/q^2 \sim a/q^2R$  and therefore turbulence will have strong magnetic component. We expect:

$$\frac{\delta B}{B} \sim \frac{\Delta}{qR}, \quad k_{\parallel} \sim \frac{1}{qR}$$

Field line crosses a whole Eddy going once around the tokamak.

$$D_B = \frac{\Delta^2}{qR} \quad ??$$

Field line diffusion if field is incoherent each time around.

# Electrostatic or Electromagnetic Turbulence?

Estimating the thermal diffusivities we get:

$$\chi_i \sim v_{thi} D_B$$

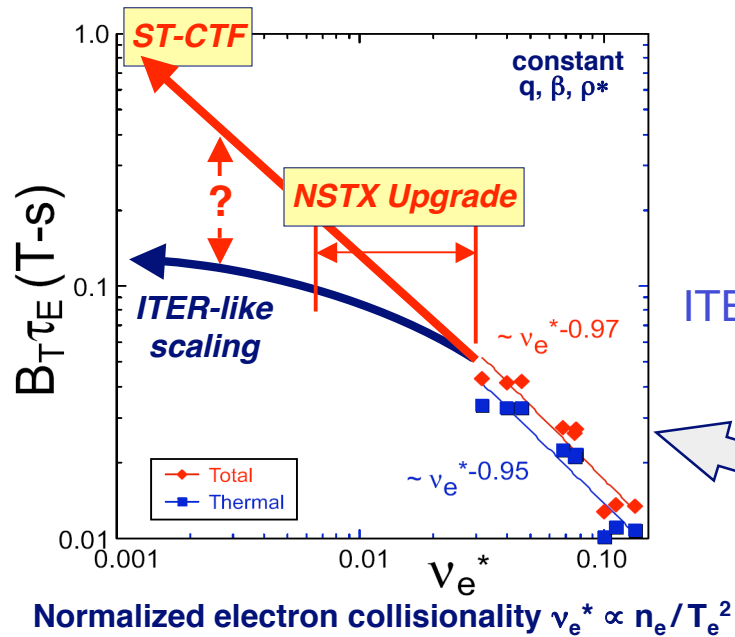
Ion thermal diffusivity from both cross field and along field line motion.

$$\chi_e \sim v_{the} D_B \sim 60 \chi_i$$

Electron thermal diffusivity from motion along stochastic field line.



# Spherical Tokamaks



Stan Kaye et. al NF 2007.

- Future ST's are projected to operate at 10-100 times lower normalized collisionality  $\nu^*$

ITER  $B\tau_E$  (e-static g-Bohm)  $\propto \rho_*^{-3} \beta^0 \nu_*^{-0.14} q^{-1.7}$   
 Petty et al., PoP, Vol. 11 (2004)

- Conventional tokamaks observe weak inverse dependence of confinement on  $\nu^*$
- NSTX observes much stronger scaling vs.  $\nu^*$ 
  - Does favorable scaling extend to lower  $\nu^*$  ?
  - What modes dominate e-transport in ST ?

## Self Consistent Transport.

The motion of the particles affects the field -- test particle transport typically ignores the correlation of motion and field fluctuations.

Electrons provide constraints on the fluctuating field.

This area is much less researched although there is some -- *e.g. Itoh and Itoh.*

## Frozen in Magnetic Field.

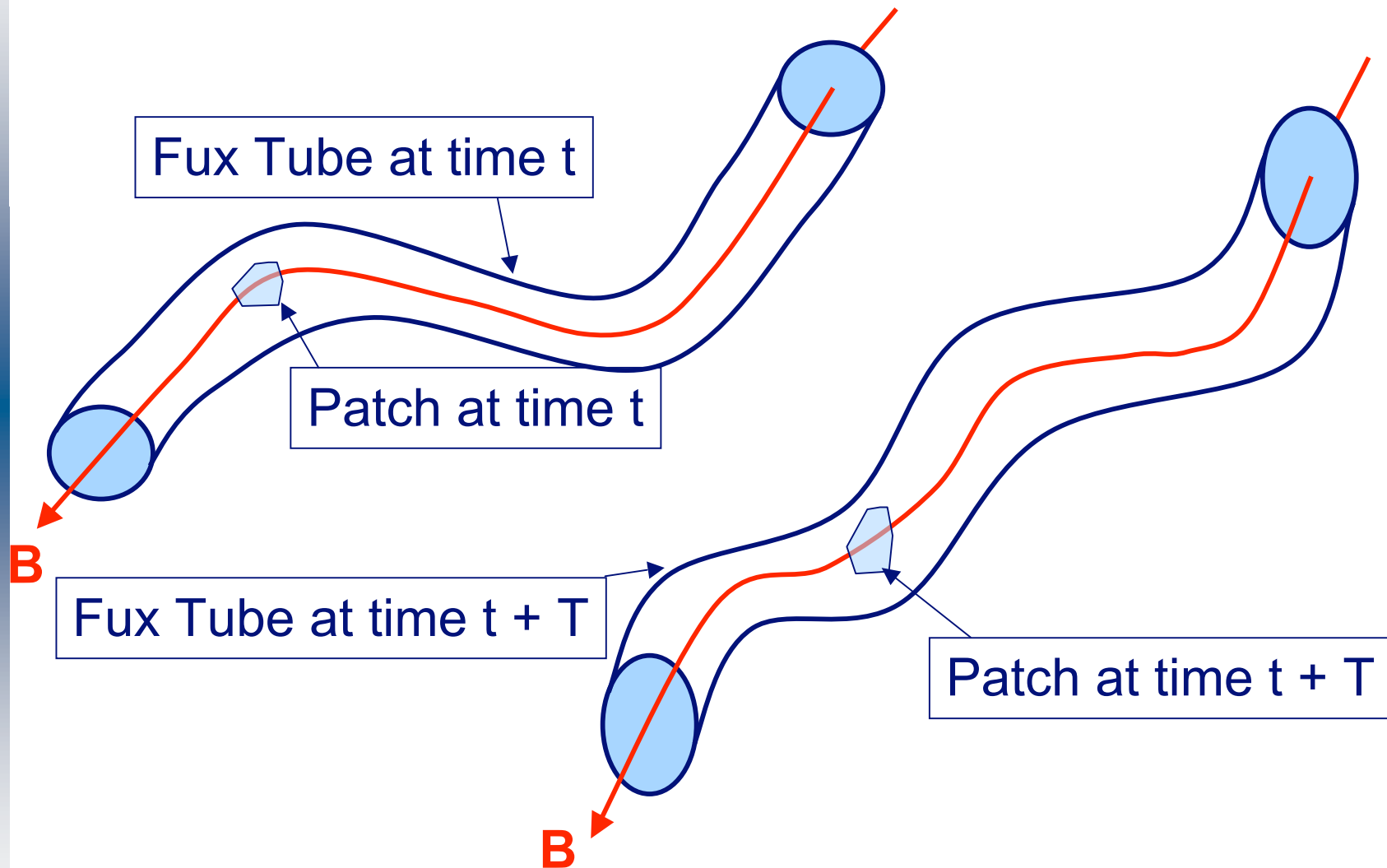
Alfven 1945, Kelvin, Cauchy Ideal Ohm's law:  $\mathbf{E} + \mathbf{v} \times \mathbf{B} = 0$ .

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E} = \nabla \times (\mathbf{v} \times \mathbf{B})$$

**Flux Freezing.** If  $\mathbf{E} + \mathbf{v} \times \mathbf{B} = 0$  the magnetic flux through a loop that moves with the plasma flow,  $\mathbf{v}$ , is constant in time.

**Frozen Field Lines.** If  $\mathbf{E} + \mathbf{v} \times \mathbf{B} = 0$  the magnetic field lines change as though they are simply convected with velocity,  $\mathbf{v}$ . Thus we say that the field lines are frozen to the plasma

Velocity of Magnetic Field lines = Velocity of Plasma =  $v$ .



# When are Field Lines Frozen?

Newcomb 1958

Suppose we can write

$$\mathbf{E} = \mathbf{E}_{\perp} + E_{\parallel} \mathbf{b}, \quad E_{\parallel} = \mathbf{b} \cdot \nabla \zeta$$

$$\tilde{\mathbf{v}} = \frac{1}{B^2} [(\mathbf{E} - \nabla \zeta) \times \mathbf{B}],$$



$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\tilde{\mathbf{v}} \times \mathbf{B}).$$

velocity

This only makes sense if  $\zeta$  is finite and single valued.

## Frozen Field Lines with Pressure and Hall Terms

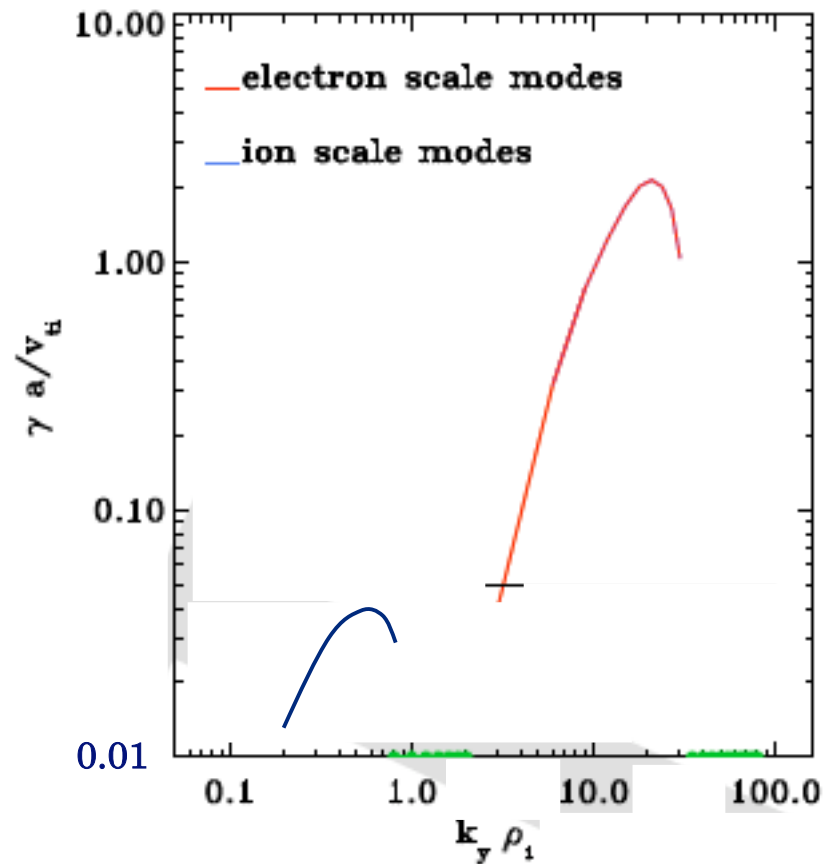
$$\mathbf{E} + \mathbf{v}_e \times \mathbf{B} = -\frac{\nabla p_e}{en_e}, \quad \mathbf{B} \cdot \nabla T_e = 0$$

$$\tilde{\mathbf{v}} = \mathbf{v}_e + \frac{\ln n_e}{eB^2} (\nabla T_e \times \mathbf{B})$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\tilde{\mathbf{v}} \times \mathbf{B}).$$

Field lines are frozen but not to either species.

## Ion Scale -- Electron Scale Turbulence?



Short wavelength  
dominates growth  
rate.

Long wavelength  
dominates  
turbulent transport

$$\chi \sim \frac{\Delta^2}{\tau} \sim \frac{\gamma}{k^2}$$

# Ion Scale Turbulence in Tokamaks

Most of the tokamak turbulence seems to satisfy:

$$\omega \sim \omega^* \sim k_{\parallel} v_{thi} \sim \frac{v_{thi}}{qR} \ll k_{\parallel} v_{the} \left( \sim \frac{v_{the}}{qR} \right)$$

Electron response to highest order is due to rapid motion along the field. The guiding center motion holds.

$$v_{\parallel} = \sqrt{\frac{2}{m_e} (\mathcal{E} - \mu B)}$$



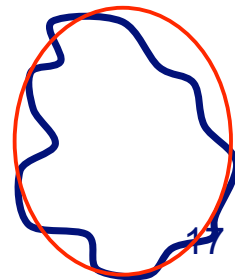
# Condition on $E_{\parallel}$

Collisionless electron drift kinetic equation:

perturbed electron distribution function,  $\delta f_e$ , satisfies:

$$v_{\parallel} \mathbf{b} \cdot \nabla (\delta f_e) = -v_{\parallel} \frac{e E_{\parallel} F_M}{T_0} - v_{\parallel} \mathbf{b} \cdot \left\{ \frac{\nabla n_0}{n_0} + \frac{\nabla T_0}{T_0} \left( \frac{\mathcal{E}}{T_0} - \frac{3}{2} \right) \right\} F_M$$

where  $\mathbf{b}$  is the unit vector along the full field (perturbed plus unperturbed),  $E_{\parallel} = \mathbf{b} \cdot \mathbf{E}$  and parallel derivative is at fixed  $\mathcal{E}$  and  $\mu$ .



# Condition on $\mathbf{E}_{\parallel}$

Let

$$\delta f_e = \left\{ \frac{\delta n}{n_0} + \frac{\delta T}{T_0} \left( \frac{\mathcal{E}}{T_0} - \frac{3}{2} \right) \right\} F_M + \delta f_H$$

where the homogeneous solution satisfies:

$$v_{\parallel} \mathbf{b} \cdot \nabla (\delta f_H) = 0$$

Then:

$$v_{\parallel} \mathbf{b} \cdot \nabla \left( \left\{ \frac{\delta n}{n_0} + \frac{\delta T}{T_0} \left( \frac{\mathcal{E}}{T_0} - \frac{3}{2} \right) \right\} \right) = -v_{\parallel} \frac{eE_{\parallel}}{T_0} - v_{\parallel} \mathbf{b} \cdot \left\{ \frac{\nabla n_0}{n_0} + \frac{\nabla T_0}{T_0} \left( \frac{\mathcal{E}}{T_0} - \frac{3}{2} \right) \right\}$$

Any dependence on  $\varepsilon$  and  $\mu$  can be absorbed into  $\delta f_H$

# Condition on $\mathbf{E}_{\parallel}$

$$\rightarrow \mathbf{b} \cdot \nabla(\delta T + T_0) = 0$$

$$\mathbf{b} \cdot \nabla \left\{ (\delta T + T_0) \left( \frac{\delta n}{n_0} + \ln n_0 \right) \right\} = -e E_{\parallel}$$

and clearly

$$E_{\parallel} = \mathbf{b} \cdot \nabla \zeta$$

and

$$\tilde{\mathbf{v}} = \frac{1}{B^2} [(\mathbf{E} - \nabla \zeta) \times \mathbf{B}],$$

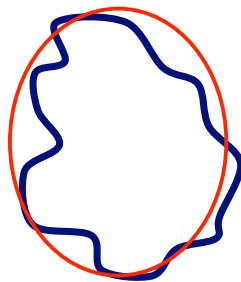
Field lines are frozen to the  
“funny velocity”

# Moving Surfaces

If surfaces aren't broken then they are frozen to  $\tilde{\mathbf{v}}$

$$\frac{\partial \tilde{\psi}}{\partial t} + \tilde{\mathbf{v}} \cdot \nabla \tilde{\psi} = 0 \quad \text{Equation for moving flux surfaces}$$

$\delta f_H$  can be found by averaging along moving field line -- flux surface average for passing particles and bounce average for trapped particles. E.g. in the collisionless case the passing particles have:



Integral over constant  $\tilde{\psi}$  surface.

$$\text{with } \bar{\zeta} = \frac{\oint ds \frac{B\zeta}{v_{\parallel} |\nabla \tilde{\psi}|}}{\oint ds \frac{B}{v_{\parallel} |\nabla \tilde{\psi}|}}$$

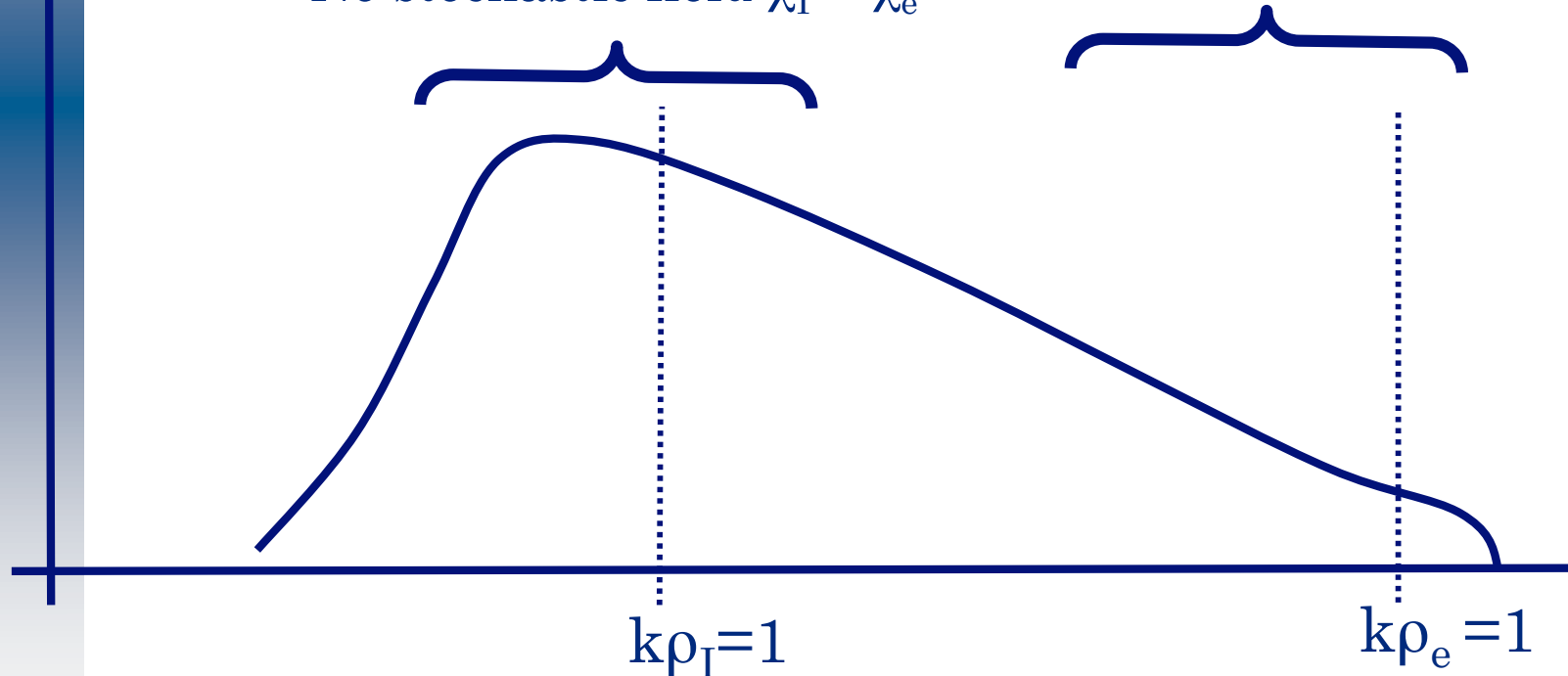
$$\delta f_e = \frac{e(\zeta - \bar{\zeta})}{T_e} F_0$$

# No Stochasticity?

$E_k^2$

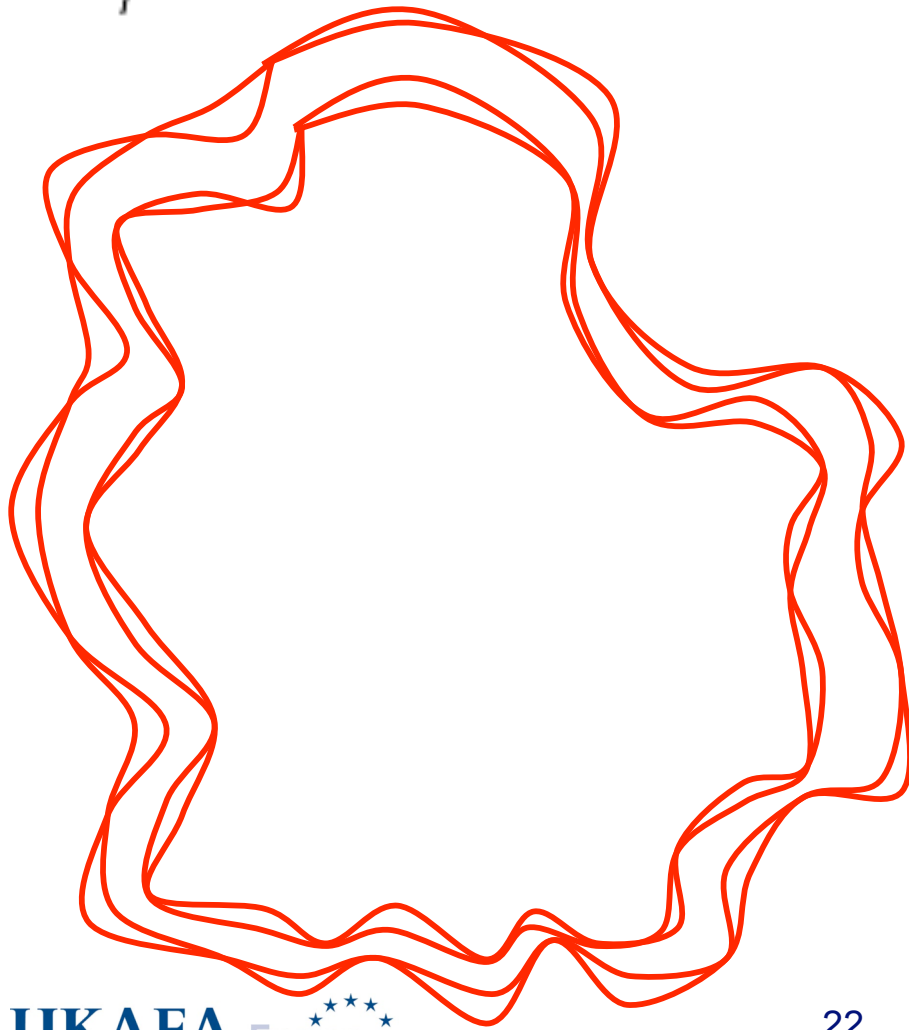
Ion scales with frozen  
Fields ITG, TEM etc.  
Transport scales --  
No stochastic field  $\chi_I \sim \chi_e$

Electron scales  
Stochastic fields?  
Micro-tearing layers.  
reconnection scales



# Embedded Micro-tearing

$\tilde{\psi}$  surfaces



Rational surfaces distorted by electromagnetic ITG/TEM can still tear and make electron scale islands. How would we know they are there?

Electron heat transport from micro-tearing:

$$\chi_e \sim \rho_e^2 \frac{v_{the}}{qR} \sim \chi_i \sqrt{m_e/m_i}$$

# Conclusions

- In most Tokamak Turbulence the field lines are frozen to a flow.
- Electrons can still be transported since they don't move with the lines.
- Small scales could be affecting the large scales. We might imagine small scales stochasticity is always present.