

Lecture 2

24.01.05

Org. details

- office hours: after lectures or by appt.
- attendance list
- no lecture on 2.02.05
- first examples class 9.02.05 11:30, venue TRA

NOTES

§ 2 Kinetic Derivation of MHD Equations and the Limits of MHD Description

NB: summarise eqns from §1

Step.

(I)

Start from particles of species  $s (= i, e)$

position:  $\vec{r}_i(t)$ ,  $\dot{\vec{r}}_i = \vec{v}_i(t)$

velocity:  $\vec{v}_i(t)$ ,  $\dot{\vec{v}}_i = \frac{q_s}{m_s} [\vec{E}(t, \vec{r}_i) + \frac{1}{c} \vec{v}_i(t) \times \vec{B}(t, \vec{r}_i)]$

This information can be assembled into the Klimontovich distribution  $f_s$  (exact)

$$F_s(\vec{r}, \vec{v}, t) = \frac{1}{V} \sum_{i=1}^{N_s} \delta(\vec{r} - \vec{r}_i(t)) \delta(\vec{v} - \vec{v}_i(t))$$

vol. of the system

$$F_s(\vec{r}, \vec{v}, t) d^3\vec{r} d^3\vec{v} = \frac{\# \text{ particles } s \text{ with } \vec{r}_i, \vec{v}_i \text{ in } [\vec{r}, \vec{r} + d\vec{r}] \cap [\vec{v}, \vec{v} + d\vec{v}]}{\text{system volume}}$$

$$\int F_s d^3\vec{r} d^3\vec{v} = \text{mean particle density}$$

Ref.:

Klimontovich

Klimontovich equ (exact):

$$\textcircled{1} \frac{\partial F_s}{\partial t} + \vec{v} \cdot \nabla F_s + \frac{q_s}{m_s} (\vec{E} + \frac{1}{c} \vec{v} \times \vec{B}) \cdot \frac{\partial F_s}{\partial \vec{v}} = 0$$

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The fields are found from Maxwell's eqns:

$$\nabla \cdot \vec{E} = 4\pi \rho_e = 4\pi \sum_s q_s \int d^3\vec{v} F_s(\vec{r}, \vec{v}, t)$$

$$\nabla \cdot \vec{B} = 0$$

$$\frac{\partial \vec{B}}{\partial t} + c \nabla \times \vec{E} = 0$$

$$-\frac{1}{c} \frac{\partial \vec{E}}{\partial t} + \nabla \times \vec{B} = \frac{4\pi}{c} \vec{j} = \frac{4\pi}{c} \sum_s q_s \int d^3\vec{v} \vec{v} F_s(\vec{r}, \vec{v}, t)$$

Everything here has variation on very small scales in the configuration space: interparticle distances.

Step (II): Coarse-grain everything in the configuration space. Formally, this means

$$f_s(\vec{r}, \vec{v}, t) \equiv \langle F_s \rangle = \int d^3\vec{r}' d^3\vec{v}' \underbrace{G(\vec{r}', \vec{v}', t)}_{\text{smoothing fn. of some width.}} F_s(\vec{r}-\vec{r}', \vec{v}-\vec{v}', t)$$

Similarly, coarse-grain the fields

$$\vec{E} = \langle \vec{E} \rangle + \delta \vec{E}, \quad \vec{B} = \langle \vec{B} \rangle + \delta \vec{B}$$

Since Maxwell's eqns are linear, coarse-grained fields satisfy Maxwell's eqns with coarse-grained distr. fn. used to calculate coarse-grained charge and current distributions.

Similarly, fluctuating fields ... with  $\delta F$

~~Fluctuating fields ... with  $\delta F$~~

Example of

g

From Klimontovich eqn (1), get

$$\frac{\partial f_s}{\partial t} + \vec{v} \cdot \nabla f_s + \frac{q_s}{m_s} \left( \langle \vec{E} \rangle + \frac{1}{c} \vec{v} \times \langle \vec{B} \rangle \right) \cdot \frac{\partial f_s}{\partial \vec{v}} =$$

$$= - \underbrace{\left\langle \frac{q_s}{m_s} (\delta \vec{E} + \frac{1}{c} \vec{v} \times \delta \vec{B}) \cdot \frac{\partial \delta F_s}{\partial \vec{v}} \right\rangle}_{\text{collisions}} = \sum_s' \hat{C}_{ss'} [f_s, f_{s'}]$$

In principle, there is a closure problem: we don't know how to determine  $\delta F_s$  in terms of  $f_s$ .

A closure scheme (BBGKY) leads to an approximation

Get

$$\textcircled{2} \quad \frac{\partial f_s}{\partial t} + \vec{v} \cdot \nabla f_s + \frac{q_s}{m_s} \left( \vec{E} + \frac{1}{c} \vec{v} \times \vec{B} \right) \cdot \frac{\partial f_s}{\partial \vec{v}} = \sum_s' \hat{C}_{ss'} [f_s, f_{s'}]$$

omit (...)

collision operator  
(Landau  
Lenard-Balescu)

- Vlasov-Landau equation, kinetic eqn.  
(Boltzmann equation for plasmas)

++ Maxwell's equations with  $\rho_e$  and  $\vec{j}$  based on  $f_s$ .

Step (III): Now take velocity-space moments of (2).  
(so description in  $\vec{r}$  space quantities)

Density  $n_s(\vec{r}, t) = \int d^3\vec{v} f_s(\vec{r}, \vec{v}, t)$  satisfies (exact)

$$\textcircled{3} \quad \frac{\partial n_s}{\partial t} + \nabla \cdot (n_s \vec{u}) = 0, \text{ where}$$

mean velocity  $n_s \vec{u}_s(\vec{r}, t) = \int d^3\vec{v} \vec{v} f_s(\vec{r}, \vec{v}, t)$  satisfies

$$\bullet \quad m_s n_s \left( \frac{\partial \vec{u}_s}{\partial t} + \vec{u}_s \cdot \nabla \vec{u}_s \right) = - \nabla \cdot \hat{P}_s + q_s n_s \left( \vec{E} + \frac{\vec{u}_s \times \vec{B}}{c} \right) + \vec{R}_s$$

inter because collisions between same species  $\rightarrow$  interspecies collisions (friction)

The fields are found from Maxwell's eqns:

$$\nabla \cdot \vec{E} = \frac{1}{\epsilon_0} \rho = \frac{1}{\epsilon_0} \sum_s q_s n_s = \frac{1}{\epsilon_0} \rho$$

$$\nabla \cdot \vec{B} = 0$$

$$\frac{\partial \vec{B}}{\partial t} + \nabla \times \vec{E} = 0$$

$$\frac{1}{c} \frac{\partial \vec{E}}{\partial t} + \nabla \times \vec{B} = \frac{1}{c} \vec{j} = \frac{1}{c} \sum_s q_s n_s \vec{v}_s = \frac{1}{c} \vec{j}$$

Close is valid provided we assume that

$$\frac{1}{h \lambda_D^3} \ll 1$$

$$\text{where } \frac{1}{\lambda_D^2} = \sum_s \frac{4\pi q_s^2 n_s}{T_s} = \sum_s \frac{\omega_{ps}^2}{v_{th,s}^2}$$

Debye length

$$\langle \vec{E}(\vec{r}, t) \rangle = \langle \vec{E}(\vec{r}, t) \rangle = \langle \vec{E}(\vec{r}, t) \rangle$$

of some width

Similarly, coarse-grain the fields

$$\vec{E} = \langle \vec{E} \rangle + \delta \vec{E}, \quad \vec{B} = \langle \vec{B} \rangle + \delta \vec{B}$$

Since Maxwell's eqns are linear, coarse-grained fields satisfy Maxwell's eqns with coarse-grained charge and current distributions.

Similarly, fluctuate fields ... with  $\delta \vec{E}$

~~fluctuate fields~~

... where  $\hat{P}_s = m_s \int d^3\vec{v} (\vec{v} - \vec{u}_s) (\vec{v} - \vec{u}_s) f_s$

- pressure tensor for species  $s$ .

~~Again a closure problem~~

Define  $p_s = \frac{1}{3} \text{Tr} \hat{P}_s$  and  $\hat{P}_s = p_s \mathbb{1} + \hat{\Pi}_s$   
 scalar pressure ↑ viscous stress

So, we get

$$(4) \quad m_s n_s \left( \frac{\partial \vec{u}_s}{\partial t} + \vec{u}_s \cdot \nabla \vec{u}_s \right) = -\nabla p_s + \nabla \cdot \hat{\Pi}_s + q_s n_s (\vec{E} + \frac{\vec{u}_s \times \vec{B}}{c}) + \vec{R}_s$$

Again closure problem: must know  $p_s, \hat{\Pi}_s$ .

Take the second moment of (2), get

~~$$\frac{\partial p_s}{\partial t} + \vec{u}_s \cdot \nabla p_s + \frac{5}{3} p_s \nabla \cdot \vec{u}_s - \frac{2}{3} \hat{\Pi}_s : \nabla \vec{u}_s + \frac{2}{3} \nabla \cdot \vec{q}_s = \frac{2}{3} Q_s$$~~

~~$$\frac{\partial p_s}{\partial t} + \vec{u}_s \cdot \nabla p_s + \frac{5}{3} p_s \nabla \cdot \vec{u}_s - \frac{2}{3} \hat{\Pi}_s : \nabla \vec{u}_s + \frac{2}{3} \nabla \cdot \vec{q}_s = \frac{2}{3} Q_s$$~~ Ohmic heating in here

$$(5) \quad \frac{\partial p_s}{\partial t} + \vec{u}_s \cdot \nabla p_s + \frac{5}{3} p_s \nabla \cdot \vec{u}_s - \frac{2}{3} \hat{\Pi}_s : \nabla \vec{u}_s + \frac{2}{3} \nabla \cdot \vec{q}_s = \frac{2}{3} Q_s$$

interspecies heat exchange

where heat flux  $\vec{q}_s = \int d^3\vec{v} \frac{m_s |\vec{v} - \vec{u}_s|^2}{2} (\vec{v} - \vec{u}_s) f_s$

Note: temperature  $T_s = \frac{p_s}{n_s} = \frac{2}{3} \int d^3\vec{v} \frac{m_s |\vec{v} - \vec{u}_s|^2}{2} f_s$

This can go on forever:

$\vec{q}_s$  in terms of 4th moments etc.

Measures must be taken to stop this madness!

STEP (IV) : Assume

~~XXXXXXXXXX~~  $\nabla \sim \frac{1}{l}$ ,  $(l \gg \lambda_{mfp})$  - mean free path

Then  $\frac{\partial}{\partial t} \sim \frac{1}{\tau}$ ,  $\frac{\tau}{\tau_c} \sim \frac{l}{u} \frac{v_{th}}{\lambda_{mfp}} \sim \frac{v_{th}}{u} \frac{l}{\lambda_{mfp}} \gg 1$ ,  $v_{th} = \sqrt{\frac{T}{m}}$   
coll. time

In this <sup>(collisional)</sup> approximation, plasma is in local thermodynamic equilibrium and the distribution  $f_s$  is,

to lowest order in  $\frac{\lambda_{mfp}}{l}$ , a Maxwellian:

$$f_s(\vec{r}, \vec{v}, t) = \left[ \frac{n_s}{2\pi T_s(\vec{r}, t)} \right]^{3/2} e^{-\frac{m_s [\vec{v} - \vec{u}_s(t, \vec{r})]^2}{2T_s(\vec{r}, t)}}$$

This can be shown to solve  $\sum_{s'} C_{ss'} [f_s, f_{s'}] = 0$ .

Then to lowest order in  $\frac{\lambda_{mfp}}{l}$ , we have

- isotropic pressure,  $\hat{P}_s = \mathbb{1} p_s$ ,  $\hat{\Pi}_s = 0$

- no heat flow  $\vec{q}_s = 0$

- no momentum or heat exchange between species

$$\vec{u}_e = \vec{u}_i, T_e = T_i, \vec{R}_s = 0, Q_s = 0$$

~~Approximation~~ To get collisional effects, go to next order and calculate everything.

Refs:

Braginskii  
Helander &

Sigmar

The result is eqns (3), (4), (5) with

$\hat{\Pi}_s, \vec{q}_s, \vec{R}_s, Q_s$  = some expressions in terms of  $n_s, \vec{u}_s, p_s$  and <sup>transport</sup> coefficients that depend on the collision frequencies/mean free path.

→ for lecture 3

~~Revision~~ Revision of last time:

Klimontovich distr. function,

$$\textcircled{1} \quad F_s = \frac{1}{V} \sum_{i=1}^{N_s} \delta(\vec{r} - \vec{r}_i(t)) \delta(\vec{v} - \vec{v}_i(t)) \quad \text{exact}$$

+ Maxwell's equations for field varying on interparticle distances

Assume  $\frac{1}{n\lambda_D^3} \ll 1$ , coarse grain, get

$f_s(\vec{r}, \vec{v}, t)$  and  $\vec{E}, \vec{B}$  varying on macroscopic distances.

$$\textcircled{2} \quad \frac{\partial f_s}{\partial t} + \vec{v} \cdot \frac{\partial f_s}{\partial \vec{r}} + \frac{q_s}{m_s} (\vec{E} + \frac{1}{c} \vec{v} \times \vec{B}) \cdot \frac{\partial f_s}{\partial \vec{v}} = \sum_{s'} C_{ss'} [f_s, f_{s'}]$$

coll. operator

Take moments:

$$\textcircled{3} \quad \frac{\partial \vec{n}_s}{\partial t} + \nabla \cdot (n_s \vec{u}_s) = 0$$

$$\textcircled{4} \quad m_s n_s \left( \frac{\partial \vec{u}_s}{\partial t} + \vec{u}_s \cdot \nabla \vec{u}_s \right) = -\nabla p_s + \nabla \cdot \hat{\Pi}_s + q_s n_s \left( \vec{E} + \frac{\vec{u}_s \times \vec{B}}{c} \right) + \vec{R}_s$$

pressure      visc stress      Lorentz      interspecies friction

++ pressure equation ( $p_s = n_s T$ )

Assume  $\frac{\lambda_{\text{mp}}}{\ell} \ll 1$  (collisional)  $\Rightarrow$  local thermodynamic equilibrium to lowest order ( $f_s$  Maxwellian), isotropic pressure,  $\hat{\Pi}_s = 0$ ,

no heat flux, no exchange of heat/mom. between species ( $\vec{u}_i = \vec{u}_e, T_i = T_e$ )

To next order  $\mathcal{O}\left(\frac{\lambda_{\text{mp}}}{\ell}\right)$ , calculate viscous stress, interspecies friction, heat flux, heat exchange in terms of  $n_s, \vec{u}_s, p_s$  and transport coefficients. Ref: Braginskii

See revision of last time p. 10 en regard.

STEP (V): Introduce MHD variables:

electrons do not carry mass { mass density  $\rho = n_i m_i (+ n_e m_e)$   
 fluid velocity  $\vec{u} = \vec{u}_i$  (NB:  $\rho \vec{u} \approx n_i m_i \vec{u}_i + n_e m_e \vec{u}_e$ )

They do carry charge { current density  $\vec{J} = en(\vec{u}_i - \vec{u}_e)$   
 charge density  $\rho_e = e(n_i - n_e)$

pressure  $p = p_i + p_e$  etc.

(3)  $\Rightarrow \frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0$  continuity eqn (6)

NB:  $\vec{R}_e + \vec{R}_i = 0$   
 (4)<sub>i</sub> + (4)<sub>e</sub>  $\Rightarrow \rho \frac{d\vec{u}}{dt} = -\nabla p + \nabla \cdot \hat{\Pi} + \rho_e \vec{E} + \frac{1}{c} \vec{J} \times \vec{B}$  (7)  
 hold on. small when  $u \ll c$

$\frac{v}{c} \ll 1$  In the non-relat. approx, we showed in lecture 1 that the field eqns are

$\vec{J} = \frac{c}{4\pi} \nabla \times \vec{B}$  Ampere's law

$\frac{\partial \vec{B}}{\partial t} = -c \nabla \times \vec{E}$  Faraday's law (8)

We had to bring in Ohm's law in order to find  $\vec{E}$  in terms of  $\vec{B}$  and  $\vec{J}$  and close (9.8).

In this kinetic derivation, we can derive Ohm's law from the electron momentum equation.

Target:  $\vec{E} + \frac{1}{c} \vec{u} \times \vec{B} = -\frac{1}{\sigma} \vec{J}$





In (4)<sub>e</sub>, use  $\vec{u}_e = \vec{u}_i - \frac{\vec{J}}{en} = \vec{u} - \frac{\vec{J}}{en}$  and get

$$\vec{E} + \frac{\vec{u} \times \vec{B}}{c} = \frac{\vec{J} \times \vec{B}}{cen} - \frac{\nabla p_e}{en} + \frac{\nabla \cdot \hat{\Pi}_e}{en} + \frac{\vec{R}_e}{en} - \frac{m_e}{e} \left( \frac{\partial \vec{u}_e}{\partial t} + \vec{u}_e \cdot \nabla \vec{u}_e \right)$$

Hall      thermoelectric      resistive      electron inertia

$$\beta = \frac{P}{B^2/8\pi}$$

$$\frac{1}{\beta} \frac{p_i}{e}$$

$$\frac{p_i}{e}$$

$$\frac{\lambda_{mf}}{p_i} \left( \frac{p_i}{e} \right)^2$$

$$\frac{1}{\beta} \frac{p_i}{e} \sqrt{\frac{m_e}{m_i} \frac{p_i}{\lambda_{mf}}}$$

$$\sqrt{\frac{m_e}{m_i}} \frac{p_i}{e}$$

$$\frac{p_e}{\lambda_{mf}} = \frac{v_{ei}}{\Omega_e}$$

Electron-ion friction term: technically speaking, need a kinetic calculation, but roughly get

$$\vec{R}_e = m_e n_e v_{ei} (\vec{u}_i - \vec{u}_e) = \frac{m_e}{e} v_{ei} \vec{J}$$

↑  
e-i coll. freq.

[From kinetics, we learn  $v_{ei} \sim 10 \frac{n_e}{T_e^{3/2}} \frac{1}{\text{sec}}$ ]

Assume a subsidiary ordering:

1)  $\nabla \sim \frac{1}{l}$ ,  $l \gg \rho_i = \frac{v_{thi}}{\Omega_i}$  gyro/cyclotron/Larmor radius

where  $\Omega_i = \frac{eB}{m_i c}$  ion gyro/cyclotron freq.

NB:  $\frac{p_e}{p_i} = \frac{v_{the}}{v_{thi}} \frac{\Omega_i}{\Omega_e} = \sqrt{\frac{T_e}{T_i}} \sqrt{\frac{m_e}{m_i}} \ll 1$

↑ in collisional plasma

Then  $\tau \Omega_i \sim \frac{l}{u} \Omega_i \sim \frac{l}{v_{thi}} \frac{v_{thi}}{u} \Omega_i \sim \frac{l}{\rho_i} \frac{v_{thi}}{u} \gg 1$

2) Also assume  $u \sim v_{thi}$  (same order!),  $\beta \lesssim 1$

3) Also assume  $p_e \gg \lambda_{mf}$  (or  $v_{ei} \gg \Omega_e$ ) [ $\Rightarrow \rho_i \gg \lambda_{mf}$ ]

[NB:  $l \gg \lambda_{mf} \Rightarrow T_e \sim T_i$  - assumed before]



With these assumptions, all terms except resistive drop out and we get

$$\vec{E} + \frac{\vec{u} \times \vec{B}}{c} = \frac{m_e \nu_{ei}}{e^2 n} \vec{j} \equiv \frac{1}{\sigma} \vec{j}$$

- our old Ohm's law.

Note that we also need  $\rho_i \gg \lambda_{mfp}$  to get the old isotropic expression for the viscous stress:

$$\hat{\Pi} = \rho \nu \left[ \nabla \vec{u} + (\nabla \vec{u})^T - \frac{2}{3} \nabla \cdot \vec{u} \mathbb{1} \right]$$

where kin. viscosity  $\nu \sim \nu_{thi} \lambda_{mfp}$

### Summary of assumptions:

$l \gg \rho_i \gg \rho_e \gg \lambda_{mfp}$	collisional particles unmagnetised.
$u \sim v_{thi} \ll c$	non-relativistic

### Extensions of MHD:

- if particles are magnetised,  $\rho_i \leq \lambda_{mfp}$
  - viscous stress becomes anisotropic
  - Ohm's law much more complicated
- } Braginskii theory...
- ~~if  $\frac{1}{\beta} \rho_i \gg l \gg \lambda_{mfp} \gg \rho_e$  ( $\beta$  small or  $\rho_i$  big)~~
- if  $\frac{1}{\beta} \rho_i \gg l \gg \lambda_{mfp} \gg \rho_e$  ( $\beta$  small or  $\rho_i$  big)
  - Hall term important  $\Rightarrow$  Electron MHD
  - if neutrals are present, 3-fluid theory  
(for low ionisations, e-i fluid drifts wrt neutrals  $\Rightarrow$  ambipolar diffusion)

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$$\vec{E} + \vec{v} \times \vec{B} = \frac{1}{\epsilon_0} \nabla \rho - \frac{1}{c^2} \frac{d\vec{J}}{dt}$$

$$\vec{E} + \vec{v} \times \vec{B} = \frac{1}{\epsilon_0} \nabla \rho - \frac{1}{c^2} \frac{d\vec{J}}{dt}$$

$$\frac{1}{\epsilon_0} \nabla \rho - \frac{1}{c^2} \frac{d\vec{J}}{dt} = \frac{1}{\epsilon_0} \nabla \rho - \frac{1}{c^2} \frac{d\vec{J}}{dt}$$

Electron-ion factor term: technically speaking need a kinetic calculation, but roughly fit

Note that we also need  $l \gg \lambda_D = \frac{v_{th}}{\omega_p} = \sqrt{\frac{T}{4\pi e^2 n}}$  same for  $e_i$

But  $\frac{\rho_i^2}{\lambda_D^2} \sim \frac{\omega_{pi}^2}{\Omega_i^2} \sim \frac{4\pi e^2 n m_i^2 c^2}{m_i e^2 B^2} \sim \frac{4\pi \rho_i c^2}{B^2}$

$$\sim \frac{\rho_i v_{th}^2}{B^2} \frac{c^2}{v_{th}^2} \sim \beta \frac{c^2}{v_{th}^2} \gg 1 \text{ if } \beta \text{ not too small}$$

$$\frac{\rho_e^2}{\lambda_D^2} \sim \frac{4\pi e^2 n m_e^2 c^2}{m_i e^2 B^2} \sim \frac{m_e n c^2}{B^2} \sim \frac{m_e}{m_i} \beta \frac{c^2}{v_{th}^2} \gg 1$$

$$1 \gg \frac{v_{th}^2}{c^2} \ll \frac{m_i}{m_e}$$

$$1 \ll \frac{v_{th}^2}{c^2} \gg \frac{m_i}{m_e}$$

Also assume  $v \sim v_{th}$  (same order),  $\rho \ll 1$

Also assume  $\rho \gg \lambda_{De}$  (or  $\lambda_{Di} \gg \lambda_{De}$ )

[No:  $l \gg \lambda_{De} \rightarrow T \sim T_i$  - assumed before]