

## Part II Part I. Magnetohydrodynamics.

### §5 §1 MHD Equations

I will give a quick derivation based on postulating the fluid equations (which should be familiar from undergrad. courses on fluid dynamics)

① Continuity equation (mass conservation)

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0$$

② Navier-Stokes equation (momentum conservation)

$$\rho \left( \frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} \right) = -\nabla p + \nabla \cdot \hat{\Pi} + \vec{F}$$

$\frac{d\vec{u}}{dt}$   
Lagrangian (conv.)  
derivative

pressure

viscous  
stress

all other  
forces

(per unit  
volume)

- Here  $\hat{\Pi} = \mu \left( \nabla \vec{u} + (\nabla \vec{u})^T - \frac{2}{3} \nabla \cdot \vec{u} \mathbb{1} \right)$

Newtonian viscosity (bulk viscosity = 0)

$$\nabla \cdot \hat{\Pi} = \mu \left( \nabla^2 \vec{u} + \frac{1}{3} \nabla \cdot \vec{u} \right) = \rho \nu \left( \nabla^2 \vec{u} + \frac{1}{3} \nabla \cdot \vec{u} \right)$$

kinematic viscosity

= To get pressure, we need another equation.

There are several schemes available for this

• Adiabatic (no heat flow) :  $\frac{d}{dt} \frac{p}{\rho^\gamma} = 0$  ,  $\gamma = \frac{5}{3}$  (ideal gas)  
+ transport/dissipation terms

• Isothermal :  $p = \frac{T}{m} \rho = \frac{\rho v_{th}^2}{2}$  ,  $T = \text{const}$

- Incompressible:  $u \ll c_s = \sqrt{\frac{\gamma p}{\rho}}$  sound speed

$\rho = \text{const}$ , so from ①,  $\boxed{\nabla \cdot \vec{u} = 0}$

Take divergence of ②:  $\checkmark$

$$\nabla^2 p = \nabla \cdot \vec{F} - \rho \nabla \cdot (\vec{u} \cdot \nabla \vec{u}) \quad \text{equation for pressure}$$

— Now deal with the forces.

We have a conducting fluid, so

- Lorentz force  $\vec{F}_L = \frac{1}{c} \vec{j} \times \vec{B}$   
 $\uparrow$  current density

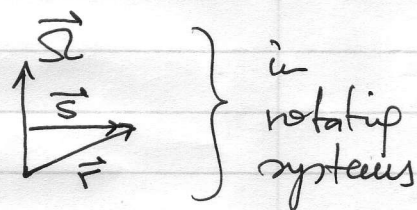
- Electrostatic force  $\vec{F}_E = \rho_e \vec{E}$   
 $\uparrow$  charge density

Other forces that often occurs are

- Gravity  $\vec{F}_g = \rho \vec{g}$

- Coriolis  $\vec{F}_c = -2\rho \vec{\Omega} \times \vec{u}$

- Centrifugal  $\vec{F}_{cf} = \rho \Omega^2 \vec{s}$



- etc.

We won't be interested in these. So:

$$\textcircled{1} \left\{ \begin{array}{l} \frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0 \end{array} \right.$$

$$\textcircled{2} \left\{ \begin{array}{l} \rho \frac{d\vec{u}}{dt} = -\nabla p + \rho \gamma (\nabla^2 \vec{u} + \frac{1}{3} \nabla \cdot \vec{u}) + \frac{1}{c} \vec{j} \times \vec{B} + \rho_e \vec{E} + \text{etc.} \\ ++ \text{ equation for } p \end{array} \right.$$

Now we need to determine  $\vec{E}$  and  $\vec{B}$ .

Maxwell's equations:

③  $\nabla \cdot \vec{E} = 4\pi \rho_e$  Gauss

④  $\nabla \cdot \vec{B} = 0$  solenoidality of  $\vec{B}$

⑤  $\frac{\partial \vec{B}}{\partial t} = -c \nabla \times \vec{E}$  Faraday's law (evolves  $\vec{B}$ )

⑥  $\nabla \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t}$  Ampère/Maxwell equ.

Non-relativistic limit

time scales  $\tau$

spatial scales  $l$

velocities  $u \sim l/\tau \ll c$

⑤  $\Rightarrow E \sim \frac{1}{c} \frac{l}{\tau} B \sim \frac{u}{c} B$

in ⑥:  $\frac{|\frac{1}{c} \frac{\partial \vec{E}}{\partial t}|}{|\nabla \times \vec{B}|} \sim \frac{\frac{1}{c} \frac{1}{\tau} \frac{u}{c} B}{\frac{1}{e} B} \sim \frac{u^2}{c^2} \ll 1$

remove Maxwell's contribution

(means we filter out EM waves)

~~So we've got a simple equation for  $\vec{J}$ :~~ So we've got a simple equation for  $\vec{J}$ :

$\vec{J} = \frac{c}{4\pi} \nabla \times \vec{B} \Rightarrow \vec{J}_L = \frac{(\nabla \times \vec{B}) \times \vec{B}}{4\pi}$   
Ampère's law

in ②:  $\frac{|\rho_e \vec{E}|}{|\frac{1}{c} \vec{J} \times \vec{B}|} \sim \frac{\rho_e E}{\frac{1}{e} B^2} \sim \frac{\frac{1}{e} E^2}{\frac{1}{e} B^2} \sim \frac{u^2}{c^2} \ll 1$

from ③  
 $\rho_e = \frac{\nabla \cdot \vec{E}}{4\pi}$

So electrostatic force is negligible.

Magnetic forces : from (6)

$$\vec{F}_L = \frac{1}{c} \vec{J} \times \vec{B} = \frac{(\nabla \times \vec{B}) \times \vec{B}}{4\pi} = \frac{\vec{B} \cdot \nabla \vec{B}}{4\pi} - \nabla \frac{B^2}{8\pi} =$$

magnetic tension                  magnetic pressure

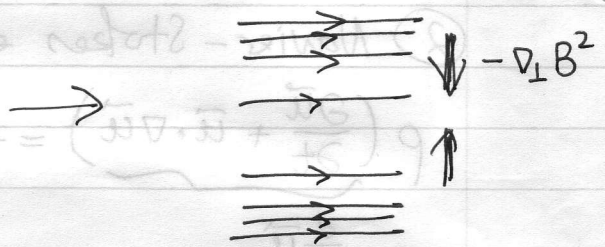
$\hat{b} = \frac{\vec{B}}{B}$

$\vec{B} \hat{b} \cdot \nabla \vec{B} \hat{b} = B^2 \hat{b} \cdot \nabla \hat{b} + \hat{b} \nabla_{\parallel} \frac{B^2}{2}$

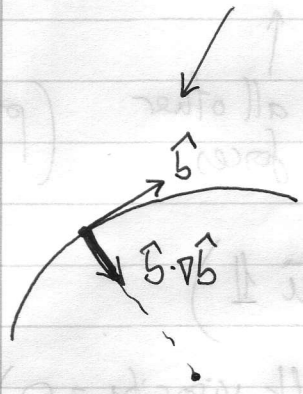
$\nabla B$ : all forces  $\perp \vec{B}$

$$= \frac{B^2}{4\pi} \hat{b} \cdot \nabla \hat{b} - \nabla_{\perp} \frac{B^2}{8\pi}$$

curvature force                  magnetic pressure



field wants to be uniform



$\hat{b} \cdot \nabla \hat{b} \perp \hat{b}$

field lines want to straighten

Magnetic field is hard to compress and to bend.

It remains to determine  $\vec{B}$  : from (5), for which we need  $\vec{E}$  in terms of known quantities.

Ohrm's law: electric field of a fluid element

$$(7) \quad \vec{E}' = \vec{E} + \frac{1}{c} \vec{u} \times \vec{B} = \frac{1}{\sigma} \vec{j} \leftarrow \text{conductivity}$$

~~magnetic resistivity~~

Substitute this into (5) and use Ampere's law:

$$\frac{\partial \vec{B}}{\partial t} = -c \nabla \times \left[ \frac{1}{\sigma} \frac{c}{4\pi} \nabla \times \vec{B} - \frac{1}{c} \vec{u} \times \vec{B} \right] =$$
$$= \nabla \times (\vec{u} \times \vec{B}) - \frac{c^2}{4\pi \sigma} \nabla \times (\nabla \times \vec{B}) \quad \leftarrow \text{we } \nabla \cdot \vec{B} = 0$$

$$(8) \quad \boxed{\frac{\partial \vec{B}}{\partial t} = \nabla \times (\vec{u} \times \vec{B}) + \eta \nabla^2 \vec{B}} \quad \text{induction equation (Hertz)}$$

$\eta$  magnetic diffusivity

Together with (1) and (2), we have a closed system.

NB:  $\nabla \cdot \vec{B} = 0$  must be satisfied initially, then it will always be satisfied (check!)