

§ 15. Degenerate Fermi Gas.

Recall that our general scheme for dealing with a quantum gas was (see pp 126-127):

1) Calculate μ from

$$N = \sum_i \bar{n}_i = \frac{(2S+1)Vm^{3/2}}{\sqrt{2}\pi^2\hbar^3} \int_0^\infty \frac{dE \sqrt{E}}{e^{\beta(E-\mu)} + 1} \quad (1)$$

Using the above,

2) Calculate mean energy:

$$U = \sum_i \varepsilon_i \bar{n}_i = \frac{(2S+1)Vm^{3/2}}{\sqrt{2}\pi^2\hbar^3} \int_0^\infty \frac{dE E^{3/2}}{e^{\beta(E-\mu)} + 1} \quad (2)$$

↑
keep just +,
for fermions

which is, completely generally, also

$$U = -\frac{3}{2}\Phi = \frac{3}{2}PV$$

so, knowing U , we immediately know the equation of state, $P = \frac{2}{3} \cdot \frac{U}{V}$ (pressure \leftrightarrow energy density)

and can also compute heat capacity:

$$C_V = \left(\frac{\partial U}{\partial T} \right)_V$$

Finally, if we want entropy, either for its own sake or, say, to calculate C_p , we have $\Phi = U - TS - \mu N$,

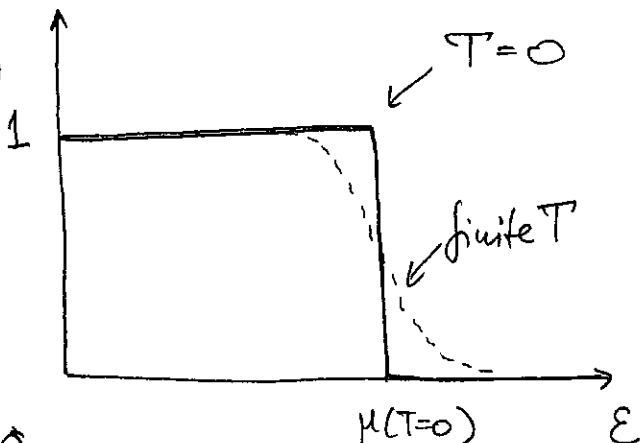
$$\text{So } S = \frac{U - \Phi - \mu N}{T} = \frac{5}{3} \frac{U - \mu N}{T} \left[\text{or } - \left(\frac{\partial \Phi}{\partial T} \right)_{V, \mu} \right]$$

Thus, full construction of thermodynamics hinges on our ability to calculate the integrals in (1) and (2).

So let us now implement this program for a degenerate Fermi gas. We will assume that $\beta \rightarrow \infty$, so

$$\frac{1}{e^{\beta(\epsilon - \mu)} + 1}$$

this is just $n_{\mathbf{k}}$ (occ. # of state with a given momentum & spin) see p.124



This distribution replaces Maxwell's in the treatment of degenerate gas

$$\mu(T=0)$$

this is called Fermi energy:

$$\boxed{\epsilon_F = \mu(T=0)}$$

So, at $T=0$, particles just stack up to some fixed energy

At finite T , the step function will be smoothed out, with the width of the smoothing region $\sim k_B T$.

15.1 Energy and Equ of State

The part of our programme pertaining to the calculation of μ (or ϵ_F as we now call it at $T=0$), energy and equ of state is now extremely easy:

$$1) \int_0^{\infty} \frac{d\epsilon \sqrt{\epsilon}}{e^{\beta(\epsilon - \mu)} + 1} \approx \int_0^{\epsilon_F} d\epsilon \sqrt{\epsilon} = \frac{2}{3} \epsilon_F^{3/2} \stackrel{\text{from (1)}}{=} \frac{\sqrt{2} \pi^2 \hbar^3}{2S+1} \left(\frac{N}{V}\right) \stackrel{=N}{=} \frac{1}{m^{3/2}}$$

$$\boxed{\epsilon_F = \frac{\hbar^2}{2m} \left(\frac{6\pi^2 n}{2S+1}\right)^{2/3} = \frac{\hbar^2 k_F^2}{2m}}$$

$$k_F = \left(\frac{6\pi^2 n}{2S+1}\right)^{1/3} \quad \text{— the max wave number up to which the particles stack up.}$$

So this

- tells us what μ is (for use in (2))
- tells us the max energy particles can have (i.e. which levels are occupied)

- tells us how good the $T=0$ approximation is :

indeed, clearly, this picture is fine as long as the smearing of the step function is narrower than the step function itself, i.e.

$$k_B T \ll \epsilon_F \Rightarrow T \ll T_F = \frac{\epsilon_F}{k_B}$$

↑ this is exactly the same (up to numerical factors) as the degeneration temperature (see p. 133) $\sim 10^4$ K for electrons in metals.

Note that "low" temperature might not actually be so low if the particle density is high (and mass small)



Eg., for electrons in white dwarves (HW), ~~energies~~

Fermi energy can be as large as \sim MeV $\Leftrightarrow T_F \sim 10^{10}$ K

and we must in fact consider relativistic electron gas

(similar calculations as ~~above~~ above, but ~~redo~~ redo everything with the relativistic formulae for $\epsilon(\vec{k})$ - see p. 125 onwards;

~~is~~ in fact it is a good exercise to go through the calculations on pp 125 - 135 for $\epsilon = \hbar kc$)

2) Now calculate energy from (2):

$$U = \frac{(2S+1)Vm^{3/2}}{\sqrt{\pi}\pi^2\hbar^3} \int_0^{\epsilon_F} d\epsilon \epsilon^{3/2} = \frac{3}{5} N\epsilon_F,$$

$N \cdot \frac{1}{\frac{2}{3}\epsilon_F^{3/2}}$

$\frac{2}{5}\epsilon_F^{5/2}$

whence, immediately, the equation of state:

$$p = \frac{2}{3} \frac{U}{V} = \frac{2}{5} n\epsilon_F = \frac{\hbar^2}{5m} \left(\frac{6\pi^2}{2S+1}\right)^{2/3} n^{5/3}$$

independent of temperature. It can be said that the gas behaves as a "pure mechanism" - i.e. entropy is not involved. Indeed (from p134),

Ex. $T S = \frac{5}{3} U - \mu N = 0$, as it should be at $T=0$.

Note. That the equation of state we have derived is actually just (keeping N const)

$$pV^{5/3} = \text{const}$$

the adiabatic law, which, as we know from pp/27-29, holds completely generally - but it is only the equation of state when $T=0$.

From this we can obtain some experimentally verifiable quantities, e.g. "the bulk modulus"

for
els
in metals

$$B = -V \frac{\partial p}{\partial V} = \frac{2}{3} n\epsilon_F.$$

15.2 Heat Capacity

We are not done, however, because we can't calculate heat capacity yet. Indeed, ~~we~~ we took $T \rightarrow 0$, so everything is independent of temperature, so we know

$$C_V = \left(\frac{\partial U}{\partial T} \right)_V = 0 \text{ at } T=0$$

But this is not really a great surprise.

We'd like to have a formula for $C_V(T)$ - and for that, we need to calculate finite-temperature corrections to our lowest-order ($T=0$) approximation of the integrals in (1) and (2).

This requires a little bit of maths:

We are interested in integrals of the form

$$I = \int_0^\infty \frac{d\varepsilon f(\varepsilon)}{e^{\beta(\varepsilon-\mu)} + 1} \quad \begin{aligned} f(\varepsilon) &= \sqrt{\varepsilon} \text{ in (1)} \\ &= \varepsilon^{3/2} \text{ in (2)} \end{aligned}$$

Change variables: $\beta(\varepsilon-\mu) = x$, i.e. $\varepsilon = \mu + k_B T x$.

Then

$$I = k_B T \int_{-\mu/k_B T}^\infty \frac{dx f(\mu + k_B T x)}{e^x + 1} =$$

$\xrightarrow{\text{here we have changed } x \rightarrow -x}$

$$= k_B T \int_0^{\mu/k_B T} \frac{dx f(\mu - k_B T x)}{e^{-x} + 1} + k_B T \int_0^\infty \frac{dx f(\mu + k_B T x)}{e^x + 1}$$

but $\frac{1}{e^{-x} + 1} = 1 - \frac{1}{e^x + 1}$, so we have

$\mu/k_B T$ - (29)

$$I = k_B T \int_0^{\mu/k_B T} dx f(\mu - k_B T x) - k_B T \int_0^{\mu/k_B T} \frac{dx f(\mu - k_B T x)}{e^x + 1}$$

invert to ϵ

$$\int_0^{\mu} d\epsilon f(\epsilon)$$

$$+ k_B T \int_0^{\infty} \frac{dx f(\mu + k_B T x)}{e^x + 1}$$

extend to ∞ at the price of an exponentially small error (because $\mu/k_B T \gg 1$)

$$\approx \int_0^{\mu} d\epsilon f(\epsilon) + k_B T \int_0^{\infty} \frac{dx}{e^x + 1} [f(\mu + k_B T x) - f(\mu - k_B T x)]$$

$$= \int_0^{\mu} d\epsilon f(\epsilon) + 2(k_B T)^2 f'(\mu) \int_0^{\infty} \frac{dx x}{e^x + 1} + \dots$$

(if necessary, can keep higher-order terms)

So, for $f(\epsilon) = \sqrt{\epsilon}$, we have, from (1),

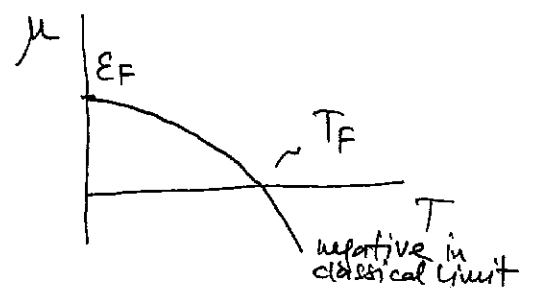
$$N = \frac{(2S+1) V m^{3/2}}{\sqrt{2\pi} \pi^2 \hbar^3} \left[\frac{2}{3} \mu^{3/2} + 2(k_B T)^2 \frac{\pi^2}{12} \frac{1}{2\sqrt{\mu}} + \dots \right]$$

pure number = $\frac{\pi^2}{12}$
~~...~~
 (see, e.g. LL §58 on how to compute this)

$$\mu^{3/2} = \epsilon_F^{3/2} - \frac{\pi^2}{8} \frac{1}{\sqrt{\mu}} (k_B T)^2 + \dots$$

$$\mu \approx \epsilon_F \left[1 - \frac{\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

So this is the corrected expression for $\mu(T)$



Now set $f(\epsilon) = \epsilon^{3/2}$ and calculate energy from (2):

$$U = \frac{N}{\frac{2}{3} \epsilon_F^{3/2}} \left[\frac{2}{5} \mu^{5/2} + 2 (k_B T)^2 \frac{\pi^2}{12} \frac{3}{2} \sqrt{\mu} + \dots \right]$$

sub. our expression for μ

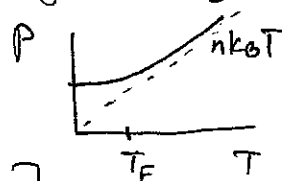
$$\approx \frac{N}{\frac{2}{3} \epsilon_F^{3/2}} \left[\frac{2}{5} \epsilon_F^{5/2} - \frac{2}{5} \epsilon_F^{5/2} \cdot \frac{5}{2} \frac{\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \right. \\ \left. + 2 \cdot \frac{\pi^2}{12} \cdot \frac{3}{2} (k_B T)^2 \epsilon_F^{1/2} + \dots \right]$$

$$= N \left[\frac{3}{5} \epsilon_F - \epsilon_F \frac{\pi^2}{8} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \epsilon_F \frac{3\pi^2}{8} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

$$= N \epsilon_F \left[\frac{3}{5} + \frac{\pi^2}{4} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

Thus, $U = \frac{3}{5} N \epsilon_F \left[1 + \frac{5\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$

corrected
expression
for $U(T)$

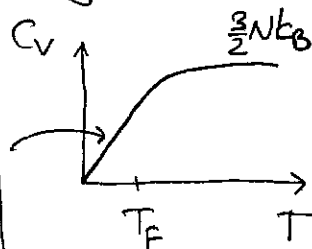


Equation of state:

$$P = \frac{2}{3} \frac{U}{V} = \frac{2}{5} n \epsilon_F \left[1 + \frac{5\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

Heat capacity:

$C_V = \left(\frac{\partial U}{\partial T} \right)_V = N k_B \frac{\pi^2}{2} \frac{k_B T}{\epsilon_F} + \dots$



Thus, the heat capacity $C_V \propto T$ at low T — this is the dominant part of the heat capacity of a metal at low T because the heat capacity due to lattice vibrations is $\propto T^3$ at low T .